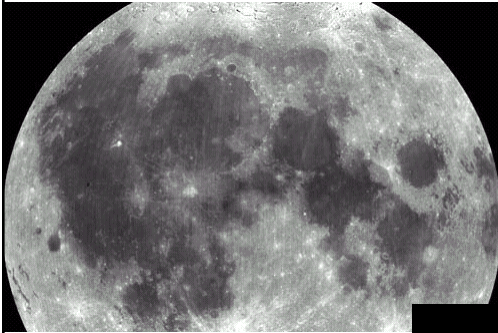


## The Moon and Mercury



The Moon

	Moon	Mercury	Earth
Mass	0.01	0.06	1
Diameter	0.27	0.38	1
Density	0.60	0.98	1

Mercury



## The Lunar Surface



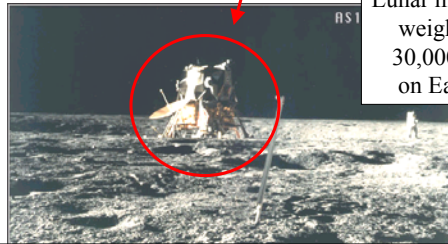
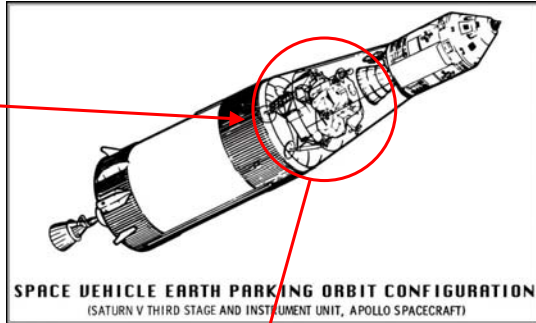
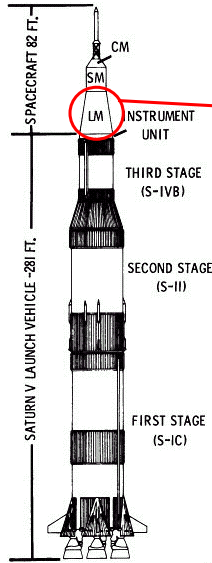
- Heavily Cratered Highlands
- Smooth Maria

# Apollo --- Our Visit to the Moon

6 manned landings on Moon, 1969 - 1972

6,371,000 lbs  
total weight  
at liftoff

Space shuttle  
to same scale



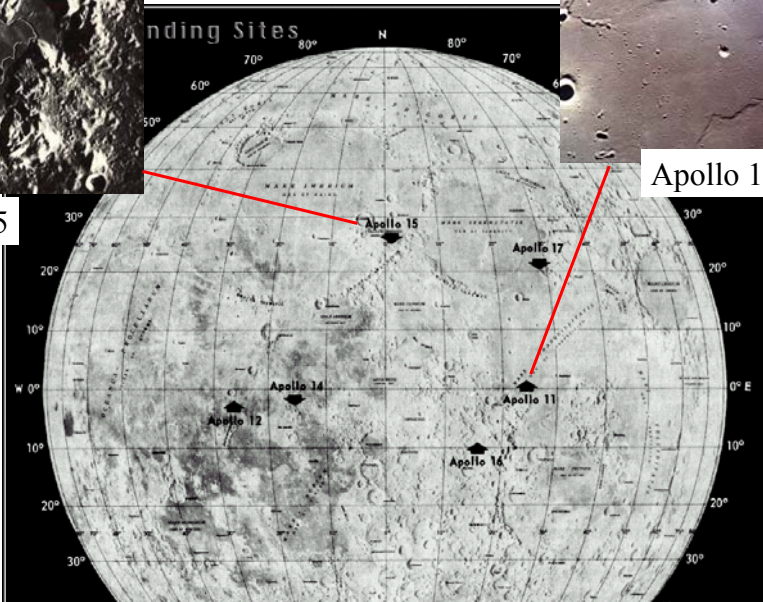
Lunar module  
weighed  
30,000 lbs  
on Earth.

## Apollo Landing Sites

6 landings



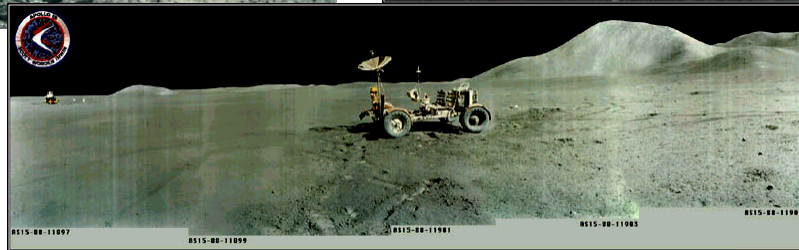
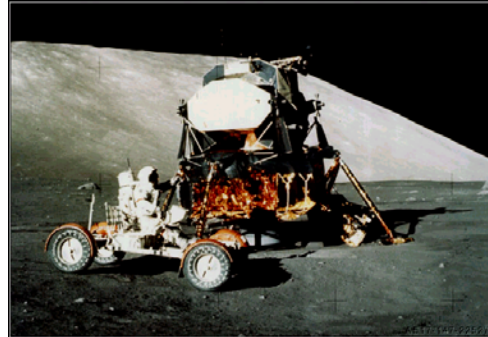
Apollo 15



Apollo 11



# Prospecting on the Moon

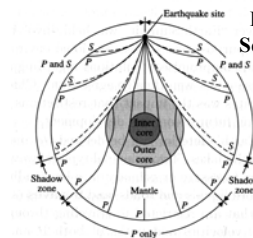


# Apollo's Accomplishments

- PR
  - beat Soviets to Moon
- Science
  - Brought back 400 kg of Lunar rocks
    - Composition
    - Age
  - Left behind seismographs and other instruments.
    - “Moonquakes” show internal structure of Moon
  - No life on Moon.



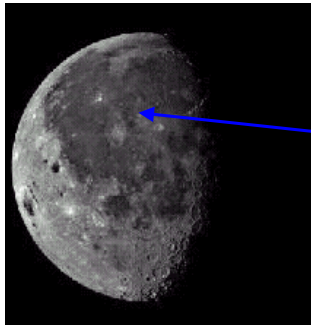
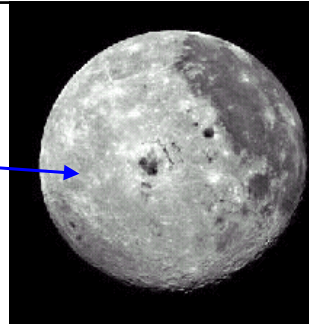
*A one kilogram (2.2 lb) Apollo 16 breccia rock formed from micrometeorite impact. Shiny, black, impact-generated glass was splashed on the side.*



**Example:  
Seismology of  
Earth's  
interior**

## The Lunar Surface

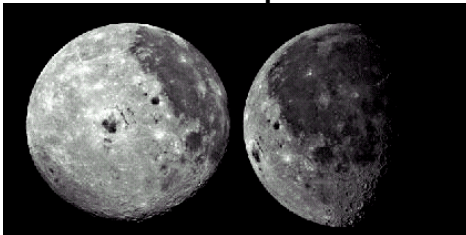
- Highlands
  - low density rocky slag formed on top of molten Moon
  - heavily cratered



### Maria

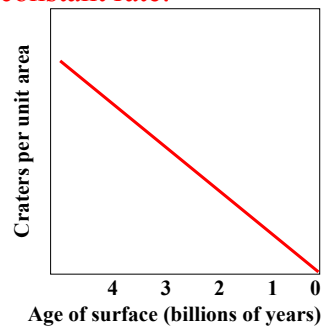
- 17% of Lunar surface
- Huge lava flows from volcanoes
- Basalts - similar to volcanic rock on Earth
  - but no water
  - fewer volatile (= easily melted) elements.
  - Much less iron.

## Impact Craters as Clocks



- Constant rain of meteors continuously makes craters
- Geologic activity
  - lava flows
  - covers over craters
- So number of craters per unit area proportional to time span since surface was last covered.

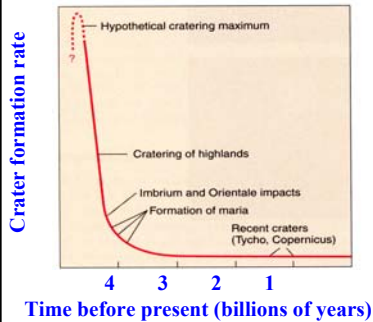
If craters form at constant rate:



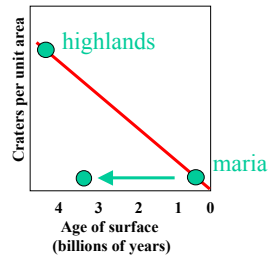
## Crater counts + Radioactive Dating of rock samples → History

10x more craters per unit area on highlands than on maria.

→ Highlands 10x older?



If craters form at constant rate:

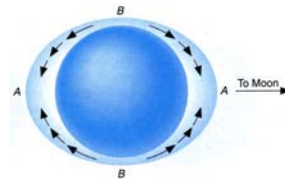
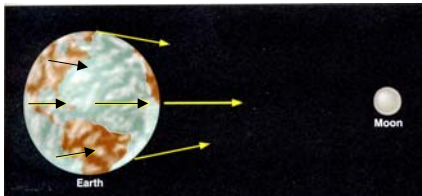


**BUT:**

Highlands

- Rock samples → 4.1 - 4.4 billion yrs old
- Maria
  - Rock samples → 3.3 - 3.8 billion yrs old

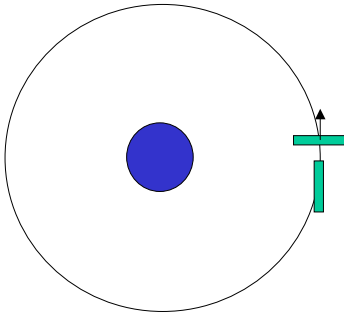
## Tides [3.6]



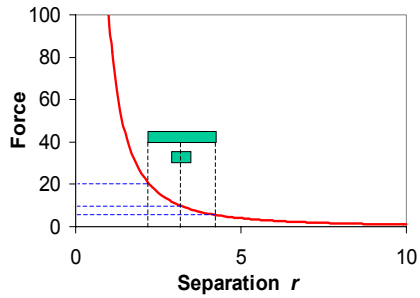
- Why do we always see same side of Moon?
  - Moon rotates in same period that it evolves around the Earth. *Why?*
  - Tidal locking.
- Moon is receding from Earth (see pg. 73)
  - Earth rotates faster than tides do.
  - Friction on tides slows Earth's rotation.
  - Conservation of angular momentum =  $mvr$  causes moon to move outwards.

# Tidal locking

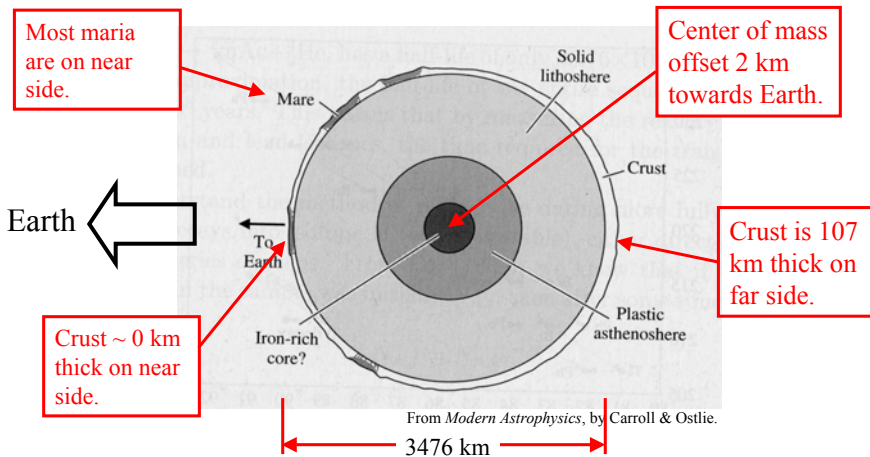
- Moon in it's orbit:



$$F_{\text{gravity}} = \frac{Gm_1m_2}{r^2}$$



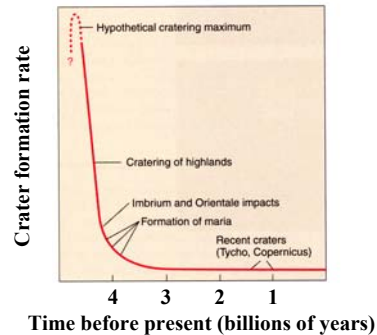
# The Interior of the Moon



- Geologically inactive.
- No magnetic field → no molten core.

# The Early History of the Moon

- The Moon and Earth are genetically related .
  - formed from different proportions of a common reservoir of materials. **Moon's composition is like Earth's mantle.**
- Magma Ocean
  - Early in its history, the Moon was melted to great depths to form a "magma ocean."
  - The lunar highlands
    - early, low density rocks
    - floated to surface of magma ocean  
4.1 - 4.4 billion yrs ago.
- Followed by a series of huge asteroid impacts
  - created basins (the maria)
  - Filled by lava flows  
3.2 - 3.9 billion years ago.



# The Origin of the Moon

## Four theories

At least three wrong theories

- **The Fission Theory** --- Moon ejected from Earth.
  - But Moon's orbit is not in plane of Earth's rotation.
  - Why aren't Moon and Earth's composition *exactly* identical?
- **The Sister Theory** --- Moon and Earth formed side by side.
  - But why aren't compositions not *exactly* identical?
- **Capture Theory** --- Moon came in from other part of Solar System
  - No way for Moon to lose energy in order to go into orbit.
  - Why are compositions so similar?
- **The Current Theory** --- the Giant Impact Theory



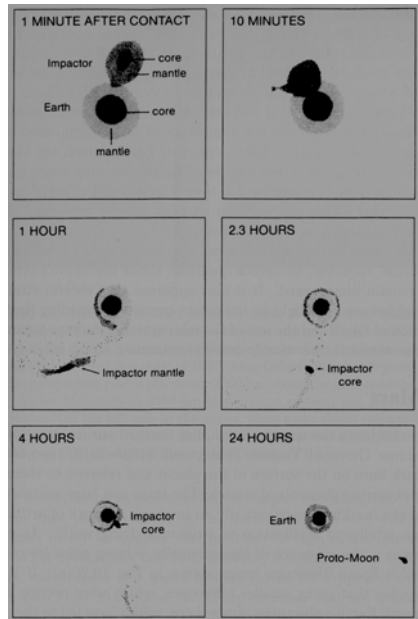


# The Giant Impact Theory

## The problem:

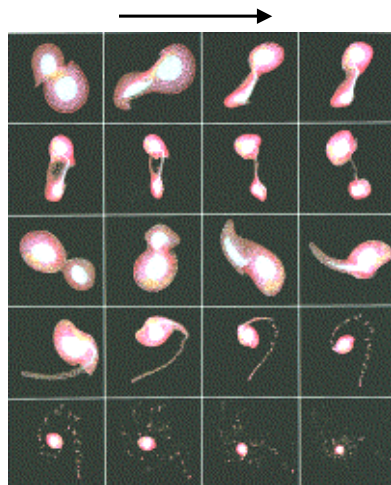
- The Moon is mostly made out of mantle material.
- Mantles are formed through *differentiation* of original mix of material.
- How do you get a mantle without having an iron core?

movie: [\(A. Cameron's simulations\)](#)



From *Modern Astrophysics*, by Carroll & Ostlie.

## Computer simulation of formation of the Moon



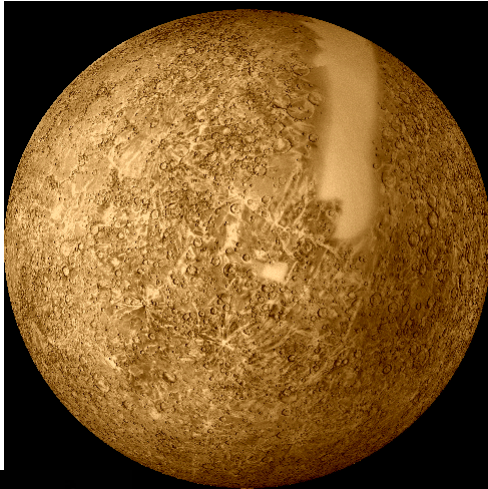
Red = mantle material  
 Orange = heated mantle  
 Blue = core material  
 Green = heated core

movie: [//www.xtec.es/recursos/astronom/moon/camerone.htm](http://www.xtec.es/recursos/astronom/moon/camerone.htm) [\(A. Cameron's simulations\)](#)

paintings: [//www.psi.edu/projects/moon/moon.html](http://www.psi.edu/projects/moon/moon.html) [paintings](#)



## Mariner 10 Mosaic Image of Mercury



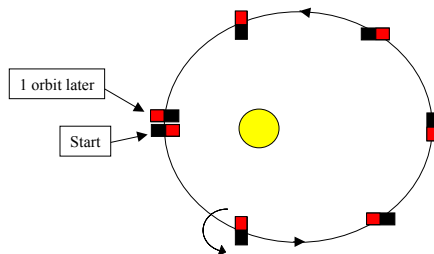
“High resolution”  
ground-based image.

- Semi-major axis = 0.39 au.  
= 58 million km
- But high eccentricity (0.2)
  - Distance from Sun varies between 46 and 58 million km.
- 88 day period of revolution about Sun.
- Only 1 visit: Mariner 10 (1974).
- Messenger orbiter planned
  - 2004 launch, 2009 arrival.

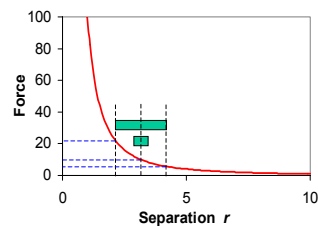
## Mercury's Rotation

- Tidal locking
- ...but Mercury rotates in 2/3 of its period of revolution around Sun. *Why?*
  - Orbit is elliptical ( $e = 0.2$ )
  - strongest tidal effect when closest to Sun
  - Another way to put it.... Mercury rotates 1.5 times per Mercury year, so elongation always points towards Sun at closest approach.

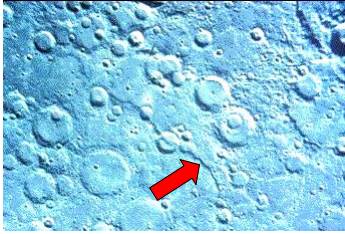
[Kepler1 simulation](#)



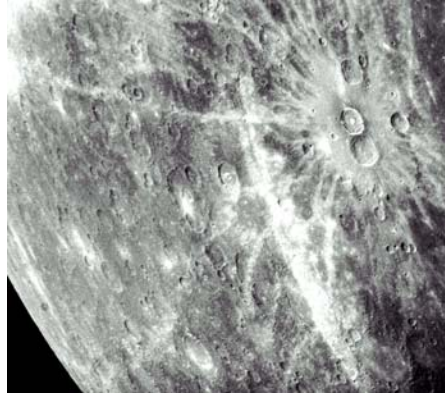
$$F_{\text{gravity}} = \frac{Gm_1m_2}{r^2}$$



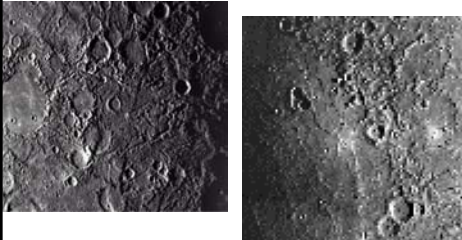
## Features on Mercury



Discovery Rupes  
(1 km-high cliff)



Rayed impact craters



Lava flows

## Structure of Mercury

- No seismological data... just educated guesses.
- Rocky mantle/crust
  - Perhaps 700 km thick.
- Metallic core
  - 75% of planet's radius
  - Weak magnetic field
    - ==> molten material in core.
- Faults & Ridges → Mercury shrank
  - by ~ 1-2 km as it cooled.
  - Crater counts
  - → this happened after first 0.5 billion yrs.

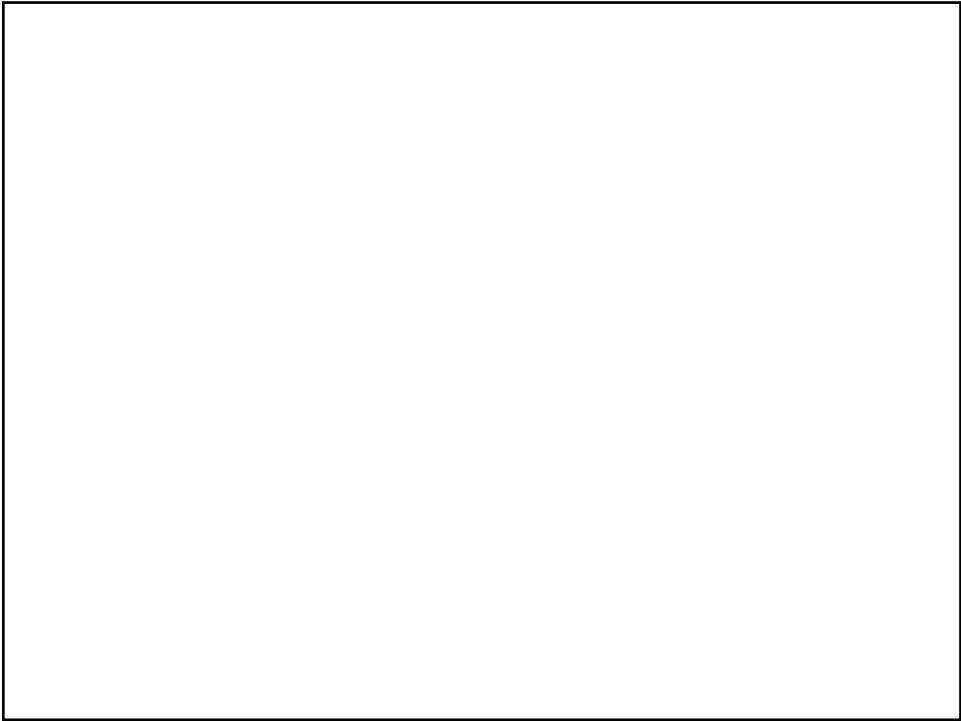


450 km-long  
ridge (rupes)

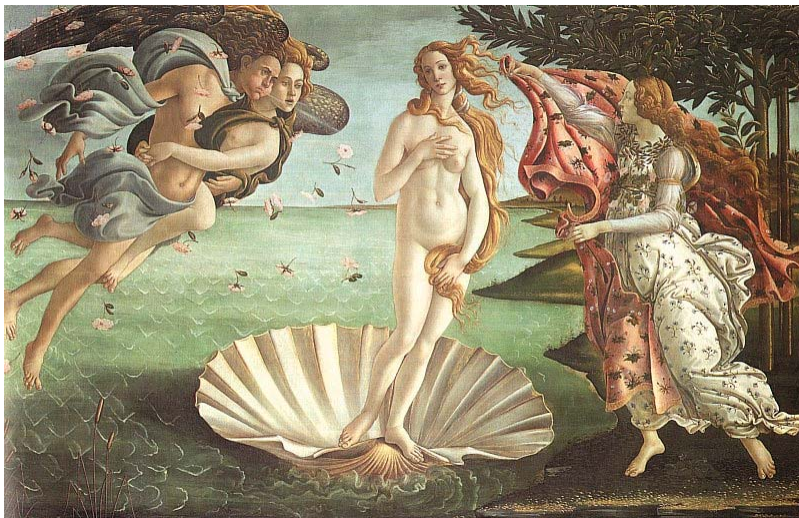


200 km-wide  
view showing  
Santa Maria  
Rupes.



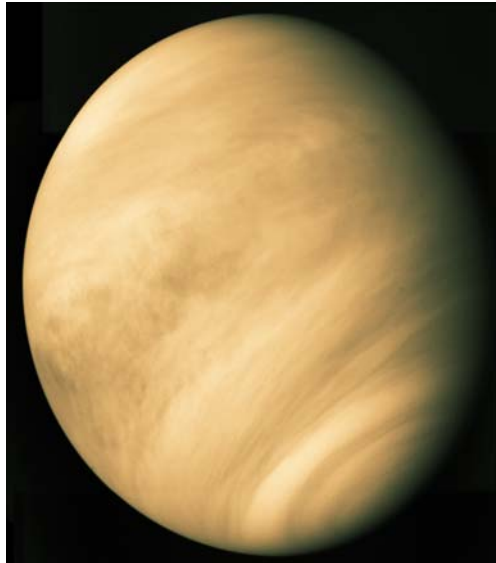


Venus  
(according to Botticelli)



# Venus (according to Galileo)

	Venus	Earth
Diameter	0.95	1
Mass	0.81	1
Semi-major axis	0.72	1
Surface temperature	480° C	20° C
Surface Air Pressure	92	1
Atmosphere	CO <sub>2</sub>	N <sub>2</sub>



Venera 7 (1970)  
Venera 10,11 (1975)  
Venera 11,12 (1978)  
Venera 13,14 (1981)

## Venera Landers (USSR)

Venera 13, 14 soil samples: basalts



ВЕНЕРА-14 ОБРАБОТКА ИППИ АН СССР И ЦДКС



ВЕНЕРА-14 ОБРАБОТКА ИППИ АН СССР И ЦДКС

The view from Venera 14

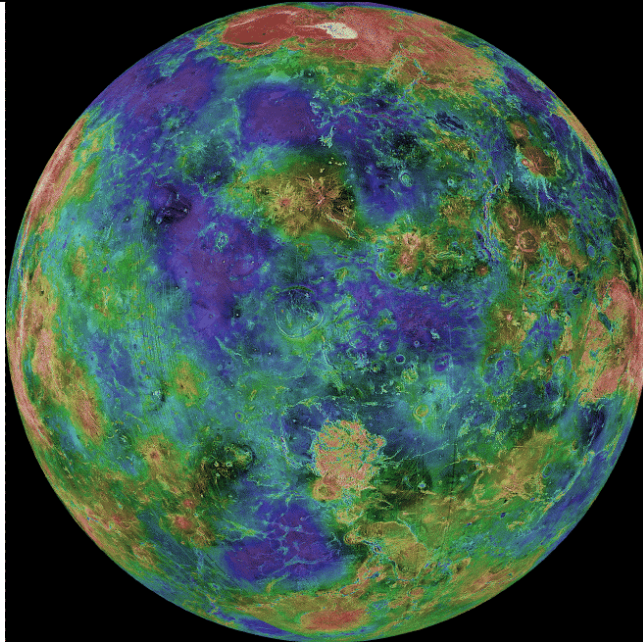
## Radar Map of Venus

Made by Magellan orbiter in 1991-93.



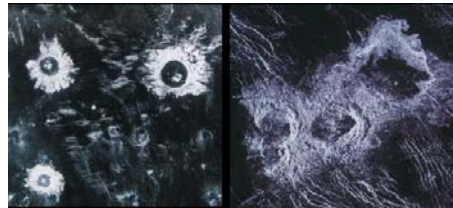
Blue = lower  
Brown/red = higher.

[Rotating Venus](#)



## The surface of Venus

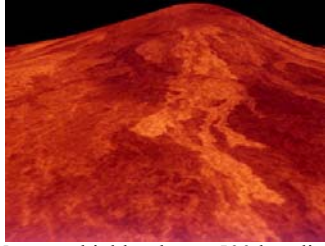
- Impact craters
  - ➔ age dating of surface
    - only 15% as many craters as lunar maria.
- ➔ Oldest terrain only 800 million yrs old
  - compare to 3.8 billion yrs on Earth
- Constant resurfacing by volcanic action.
  - but appears to have ceased ~ 500 million yrs ago



[Fig 9.5]  
Magellan Radar  
Imaging.

# Volcanic Activity on Venus

Radar Imaging: 100 m resolution



Sif Mons, a shield volcano 500 km diameter x 3 km high.

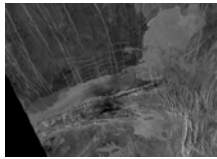


(a) Terrestrial volcano

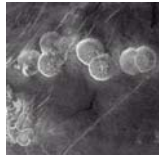


(b) Lunar impact crater

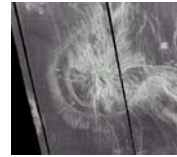
[Fig 8.13]



Lava flow



"Pancake" volcanoes, due to very thick lava.

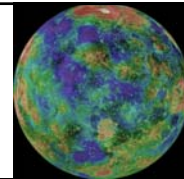


Corona: a collapsed dome over a magma chamber.

## Interior Structure

- Similar to Earth
- Iron core 3000 km in radius
- Molten mantle
- Crust

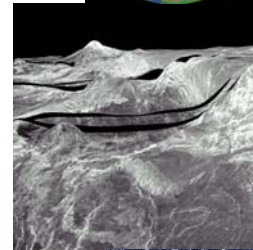
Magellan Radar Images



## Tectonics

- No plates as on earth
- But much shearing, compression and stretching of crust by convection currents in mantle.
- Has pushed up "continents"
  - Aphrodite and Ishtar
- Rift valleys and cracks

Lakshmi Planum  
Hilly area on Ishtar



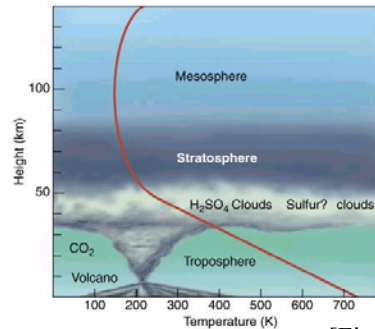
[Fig 9.9]  
Ridges & cracks



## The Atmosphere of Venus

- Surface Pressure = 92 x Earth's
- Surface Temperature = 482° C
  - melting point of lead: 327°
- Sulfuric acid cloud layer at 30-60 km

	Venus	Earth
CO <sub>2</sub>	96%	0.03%
N <sub>2</sub>	3.5	78.1
Ar	0.006	0.93
O <sub>2</sub>	0.003	21.0



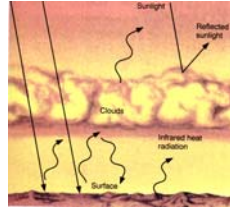
[Fig 9.11]

## Some Surface Temperatures in °F

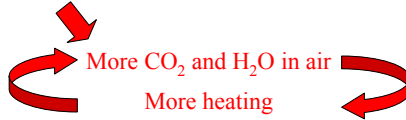
- Mercury (Mariner 10) 800F
- Venus (Mariner 2; Venera landers) 900F
- Hell (Revelations 21:8) 832F
  - “But the fearful and unbelieving shall have their part in the lake which burneth with fire and brimstone”
  - boiling point of brimstone (sulfur); 832F

## The Runaway Greenhouse Effect

- Suppose Venus started out with Earth-like conditions
  - CO<sub>2</sub> originally dissolved in oceans or combined with rocks
- Modest extra heating



CO<sub>2</sub> blocks escape of infrared radiation



- H<sub>2</sub>O then breaks down into hydrogen and oxygen.
  - *Equipartition of energy*...
    - basketball - tennis ball collision: tennis ball flies off at high velocity
      - many such collisions → all bodies have same kinetic energy =  $\frac{mv^2}{2}$
- light bodies (hydrogen atoms) move much faster and escape.

**Atmosphere evolves... no going back.**

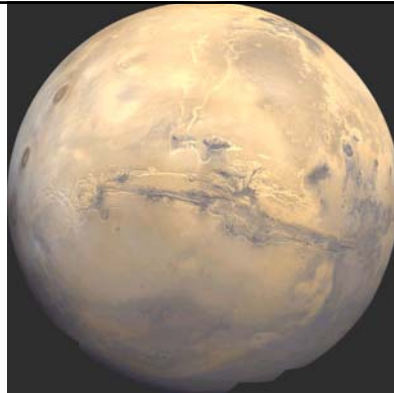
## The Odd Rotation of Venus

- Slow rotation
  - 1 Venus day = 243 Earth days.
    - (not same as orbital period = 225 days)
- Retrograde (backwards) rotation
  - Conjectured to be due to massive impact during planet's formation.



# Mars

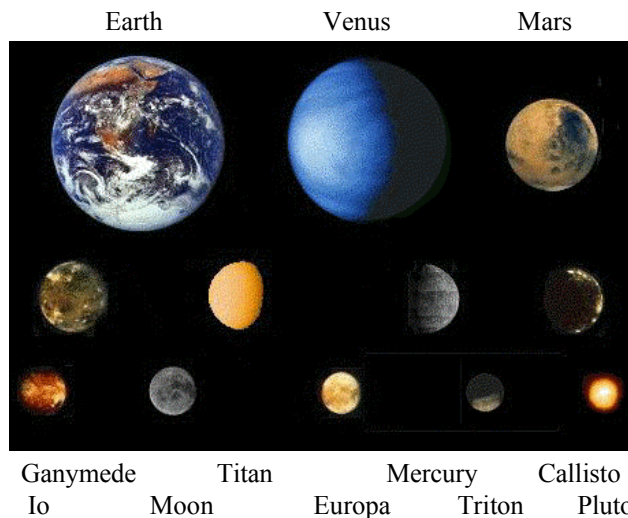
	Venus	Earth	Mars
Diameter	0.95	1	0.53
Mass	0.81	1	0.11
Semi-major axis	0.72	1	1.52
Density	0.96	1	0.71



- Some of the 16 spacecraft that have gone to Mars:
  - **Mariner 9** orbiter (1971-72)
  - **Viking 1,2** landers (1976-80)
  - **Pathfinder** lander + rover (1997)
  - **Climate Orbiter, Polar lander** (crashed, 1999).
  - **Mars Global Surveyor**: orbiting Mars since March 1999.
  - **Odyssey**: orbiting Mars since October 2001.

[Rotating Mars](#)

## Some planets and moons shown in correct relative sizes

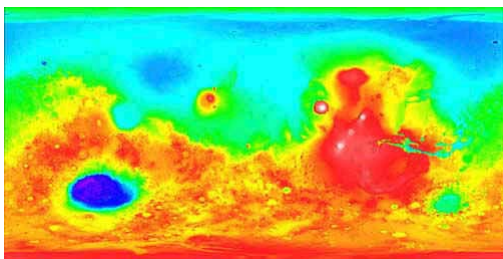


## Geology

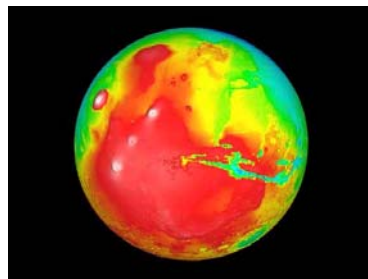
- Density suggests mostly silicates, but small metal core
- No detectable magnetic field
- Continental highlands
  - cover ~ 50% of planet.
- Low-lying lava plains
  - average of 4 km lower than continents.
  - Same age as lunar maria - 3-4 billion yrs old.

## Topographic Map

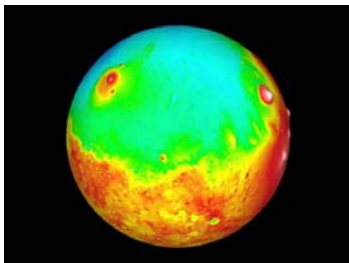
From Mars Global Surveyor orbiter



Red = high areas.  
Blue = low areas.

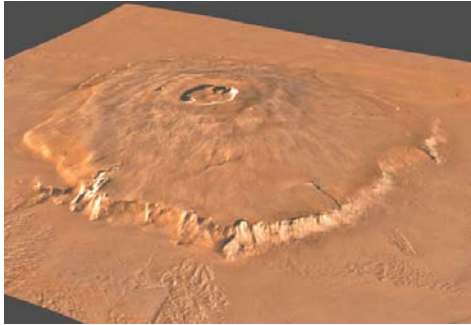


Tharsis bulge



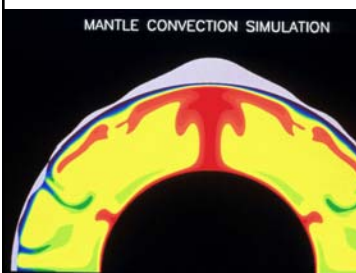
- uplifted continent 10 km high.
- has 4 huge volcanoes, 15 km high.

## Olympus Mons



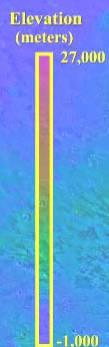
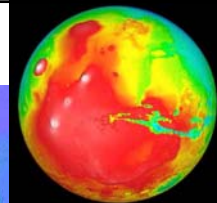
- 500 km diameter
  - would cover MI lower peninsula
- 25 km above surrounding plains
- largest mountain in Solar System.
  - 100 x volume of Mauna Loa
- < 100 million yrs old (impact crater counts)
  - so Mars is still geologically active.

## Tharsis Montes



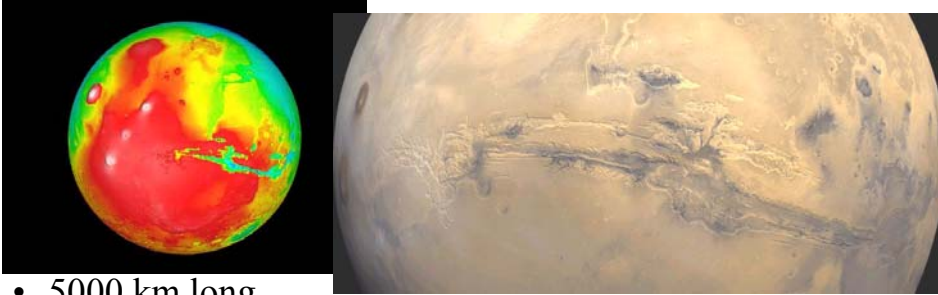
- Three shield volcanoes along crest of Tharsis bulge.
- Reach same height as Olympus Mons, but start out at higher elevation.

Olympus Mons



Tharsis Montes

## Valles Marineris




- 5000 km long
  - 1/4 way around Mars
  - would stretch clear across US.
- Huge tectonic crack in Tharsis bulge
  - no outlet for water
    - but some minor role of water erosion in side canyons.

## The Martian Atmosphere

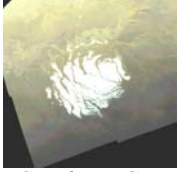
	Venus	Earth	Mars
Surface temperature	482° C	20° C	-100° C
Surface Air Pressure	92	1	0.007
CO <sub>2</sub>	96%	0.03%	95%
N <sub>2</sub>	3.5%	78%	2.7%

- Little air
- Very cold
- (almost) no liquid water.
  - At Mars' low atmospheric pressure, water should go straight from ice to vapor.
- No Greenhouse effect because there is so little atmosphere.

# Polar Ice Caps




Mars - February 1995  
HST - WFPC2



**Southern Cap**

- Always below 150° K (-279°F), so CO<sub>2</sub> frozen all year.
- Unknown mix of CO<sub>2</sub> and H<sub>2</sub>O ice.



Mars  
North Polar Cap  
HST - WFPC2

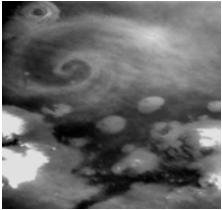
**Northern Cap**

Frozen CO<sub>2</sub> layer in winter

Only underlying H<sub>2</sub>O ice left in summer, 3 km thick


## Clouds

Cyclonic storm near North Polar ice cap  
(+ some frozen CO<sub>2</sub> on ground)



[Rotating Mars](#)

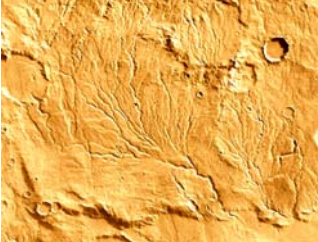
Giant dust storm



100 km

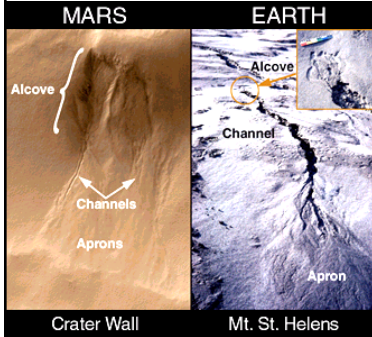
# Climate change

- Used to be *lots* of running water
  - Runoff channels.
    - From rainstorms billions of years ago.
  - Outflow channels.
    - From more recent meltings of huge volumes of permafrost.
- Loss of atmosphere
  - due to low escape velocity.
  - Low air pressure, then water freeze-out
    - ==> cannot retain heat.
  - “runaway refrigerator effect”

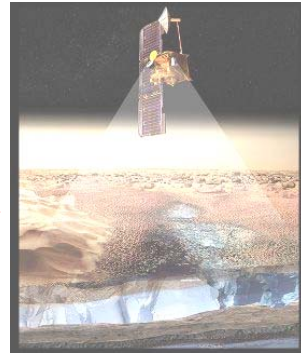


	Venus	Earth	Mars
Escape velocity	0.93	1	0.45
Surface temperature	482° C	20° C	-100° C
Surface Air Pressure	92	1	0.007

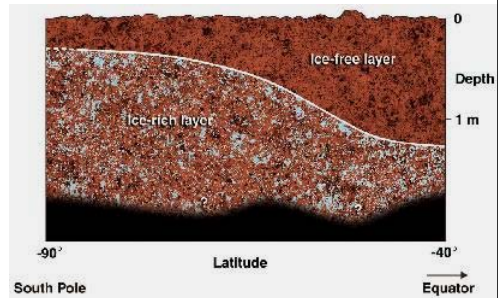
## Sub-surface Water



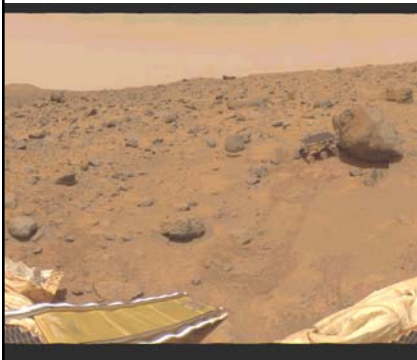
Mars Odyssey is finding lots of water just below surface.



Recently made runoff channels,  
seen by Mars Global Surveyor.  
... but these features are very rare.



## Results from the Landers



Mars Pathfinder



Soil analysis: clays & iron oxides.  
Weather stations:  $-33^{\circ}$  to  $-100^{\circ}$  C.

Search for life.

## Life on Mars?



Meteorite from Mars.

- Formed on Mars 4.5 billion yrs ago.
- Ejected from Mars by meteor impact 15 million yrs ago.
- Eventually captured by Earth (!!)
- Found in Antarctica.

### Viking landers analyzed soil samples:

- Extremely sensitive search.
- No signs of life.
- But they only measured in 2 locations.



Possible discovery of organic compounds in Martian meteorites, and even a possible (micro) fossil.

- *Unclear!* Considerable skepticism among many scientists.
- Occam's Razor: Extraordinary claims require extraordinary proof.

## Coming soon... to a planet near you:

### • Beagle 2

- Launch: June 2, 2003; Landing: Dec 24, 2003
- Robot arm:
  - Grinder, drill, "mole", microscope, spectrometer (measure chemical composition).
  - Gas analyzer to search for life
    - $C^{12}/C^{13}$  ratio?
    - Methane from bacteria?



### • MER A, B

- Two separate missions, to different parts of Mars:
  - Launch: June 8, 2003; Landing: Jan 6, 2004
  - Launch: June 25, 2003; Landing: Jan 24, 2003
- Will travel 40 meters/day, for 3 months.
- Carries cameras, spectrometers, alpha-particle detector, grinder.
- Will determine history of climate & water.

