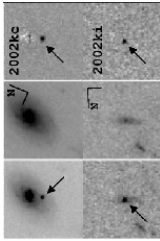


## Weighing the Universe—30 Nov

- Weighing the universe means to find mass density or density parameter
  - $\Omega = PE/KE = 8\pi G/3 \text{ mass density} / H^2$
- Why?
  - What is the universe made of? Is there mass that we cannot see?
  - What is the fate of the universe? Will it expand forever or fall back on itself?
    - Expand forever if  $\Omega \leq 1$
- How?
  - Mass in a large sphere surrounding us pulls on a galaxy on the surface
  - Measure how much the galaxy slows.
  - Use supernovae
- What we will find: Galaxies speed up!
  - “Dark energy” is repulsive whereas matter and radiation are attractive.



Distant supernovae

Riess et al, 2004, ApJ 607, 665.

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## What is the Universe Made of?

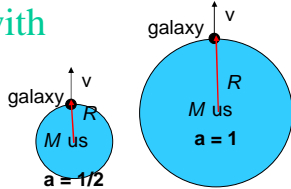
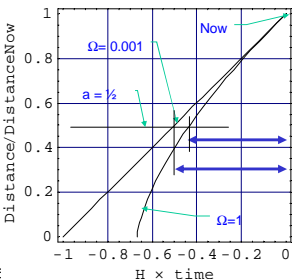
- Spherical sample of universe.  $R$ =moon's orbit. Sample has
  - 3 oz of ordinary matter
  - 1 lb of dark matter
  - 3 lb of dark energy
- Ordinary matter—protons, neutrons, electrons
  - Stars, gas, dust, planets, us
  - $\Omega_{\text{matter}} = 4\%$
- Dark matter—not detected except through gravity
  - $\Omega_{\text{dark matter}} = 23\%$
- Light
  - Mass density is small now. Dominant before universe was 1 Million years old
- Cosmological constant or dark energy
  - Repulsive
  - $\Omega_{\text{cosmological constant}} = 73\%$
- $\Omega_{\text{matter}} + \Omega_{\text{dark matter}} + \Omega_{\text{cosmological constant}} = 1$



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## Timing the motion with supernovae

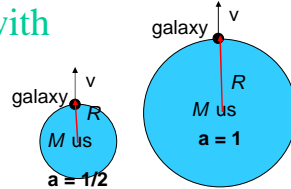
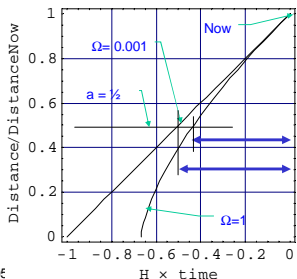
- Method for astronomical weighing:
  - Define a motion
    - Universe expands by a factor.
  - Time the motion
    - If the motion takes longer, the mass is less.
- Look at Type I supernova in a distant galaxy.
- Type I supernovae have same luminosity
- For which model is the density higher?
- For the universe with  $\Omega=1$ , how much time does it take for the universe to expand by a factor of 5? Same question with  $\Omega=0.001$ . Express time in units of  $1/H=14\text{Byr}$ .
- Identify these events on the graph: (1) SN at  $z=4$  or  $a=1/5$  emits some light in higher density universe. (2) Light gets to  $a=1/2$ . (3) We see that light.

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## Timing the motion with supernovae

- Method for astronomical weighing:
  - Define a motion
    - Universe expands by a factor.
  - Time the motion
    - If the motion takes longer, the mass is less.
- Distance to SN is greater if time for light to travel is longer.
- Suppose you discovered a supernova at  $z=2$  and found it to be faint. Then is the density parameter  $\Omega$  bigger or smaller. What on the graph gave you the answer?

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