

### CC—5 Dec

- Distant SN are 15% fainter than model with  $\Omega = 0$ !
  - Longer time to expand than universe having no mass at all!
  - Shorter time means expansion slowed down; longer time means expansion sped up.
- If motion takes longer than model with no mass, need “negative gravity” or “repulsive gravity.”
- “Though a good deal is too strange to be believed, nothing is too strange to have happened.” —Thomas Hardy

MLCS

$m-M$  (mag)

$\Delta(m-M)$  (mag)

compared to model with  $\Omega = 0.2$

$\Omega_0=0.24, \Omega_\Lambda=0.76$

$\Omega_0=0.20, \Omega_\Lambda=0.00$

$\Omega_0=1.00, \Omega_\Lambda=0.00$

$z$

$a=0.99$     $a=0.9$     $a=0.5$

Ast 207 F2005

### Einstein’s General Relativity

- What causes gravity?
- Newton’s answer: mass.
  - Force of gravity between what’s in the sphere and test mass  $m$
  - $F = G M m/R^2$ .
- Einstein’s answer: mass and pressure
  - Force of gravity between what’s in the sphere and test mass  $m$
  - $F = G (M + 3PV/c^2) m/R^2$

$v$

$M$     $R$

Ast 207 F2005

### Source of Gravity

- Einstein’s answer: mass and pressure
  - Force:  $F = G (M + 3PV/c^2) m/R^2$ .
- Newton’s Law of gravity
  - $F = G M m/R^2$ .
- Einstein’s Law of gravity
  - Curvature of space =  $8\pi G$  Mass-Pressure
  - $G = 8\pi G T$
  - Object feels curvature of space and changes its momentum
  - $G$  and  $T$  are tensors having 16 components
- Ordinary matter has little pressure because speed is much smaller than  $c$ .
  - $3PV/c^2 = M (v/c)^2$ .
- Radiation has positive pressure
  - $3PV/c^2 = M$ .
  - Force  $F = G 2M m/R^2$ .

$$T_{\text{matter}} = M/V \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & (v/c)^2 & 0 & 0 \\ 0 & 0 & (v/c)^2 & 0 \\ 0 & 0 & 0 & (v/c)^2 \end{pmatrix}$$

$$T_{\text{rad}} = M/V \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & \frac{1}{3} & 0 & 0 \\ 0 & 0 & \frac{1}{3} & 0 \\ 0 & 0 & 0 & \frac{1}{3} \end{pmatrix}$$

Ast 207 F2005

### Cosmological Constant

- Einstein’s answer: mass and pressure
  - Force:  $F = G (M + 3PV/c^2) m/R^2$ .
- Einstein’s Law of gravity
  - Curvature of space =  $8\pi G$  Mass-Pressure
  - $G = 8\pi G T$
  - Object feels curvature of space and changes its momentum
- Ordinary matter has little pressure because speed is much smaller than  $c$ .
  - $3PV/c^2 = M (v/c)^2$ .
- Radiation has positive pressure
  - $3PV/c^2 = M$ .
  - Force  $F = G 2M m/R^2$ .
- Einstein in 1920s: My equations of gravity allow “cosmological constant”
  - $T_{\text{cc}}$  has same mathematical properties as  $T_{\text{matter}}$  and  $T_{\text{rad}}$ .

- Write the force of gravity for the case of the cosmological constant. Watch the signs.

$$T_{\text{cc}} = M/V \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}$$

Ast 207 F2005

## Cosmological Constant

- Einstein's answer: mass and pressure  
Force:  $F = G (M + 3PV/c^2) m/R^2$ .
- Einstein's Law of gravity  
Curvature of space =  $8\pi G$  Mass-Pressure  
 $G = 8\pi G T$ 
  - Object feels curvature of space and changes its momentum
- Einstein in 1920s: My equations of gravity allow "cosmological constant"
  - $T_{cc}$  has same mathematical properties as  $T_{matter}$  and  $T_{rad}$ .
  - Pauli: "What is not forbidden is mandatory."
- 1. Write the force of gravity for the case of the cosmological constant. Watch the signs.

$$T_{matter} = M/V \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & (\rho/c^2) & 0 & 0 \\ 0 & 0 & (\rho/c^2) & 0 \\ 0 & 0 & 0 & (\rho/c^2) \end{pmatrix}$$

$$T_{rad} = M/V \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & \frac{1}{3} & 0 & 0 \\ 0 & 0 & \frac{1}{3} & 0 \\ 0 & 0 & 0 & \frac{1}{3} \end{pmatrix}$$

$$F = G (M - 3M) m/R^2$$

$$F = -G 2M m/R^2$$

- Repulsive gravity

$$T_{cc} = M/V \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}$$

Ast 207 F2005

## What is the Universe Made of?

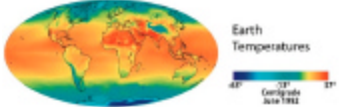
- Spherical sample of universe. R=moon's orbit. Sample has
  - 3 oz of ordinary matter
  - 1 lb of dark matter
  - 3 lb of dark energy
- Ordinary matter—protons, neutrons, electrons
  - Stars, gas, dust, planets, us
  - $\Omega_{matter} = 4\%$
- Dark matter—not detected except through gravity
  - $\Omega_{dark\ matter} = 23\%$
- Light
  - Mass density is small now. Dominant before universe was 1 Million years old
- Cosmological constant or dark energy
  - Repulsive
  - $\Omega_{cosmological\ constant} = 73\%$

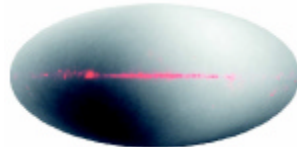
$\Omega_{matter} + \Omega_{dark\ matter} + \Omega_{cosmological\ constant} = 1$



## Weighing the Universe with the Wilkinson Microwave Anisotropy Probe (WMAP)

- Radiation from the Big Bang separated from matter when universe became neutral (not ionized) at  $a=0.001$ .
- CBR is a snapshot of U at 300,000 yr.
- Motion: universe expands from  $a=0.001$  to 1
- Timing: Observe the temperature ripples, which have a length (age of Universe)\*(speed of light)





Temperature of radiation from BB  
Ast 207 F2005 from WMAP