## Definitions, results, etc.

$$r = R(t) \boldsymbol{\varpi}$$

$$H = \frac{1}{R} \frac{dR}{dt}$$

## \*Densities:

Matter: 
$$\rho_m = \rho_{o,m} R^{-3}$$

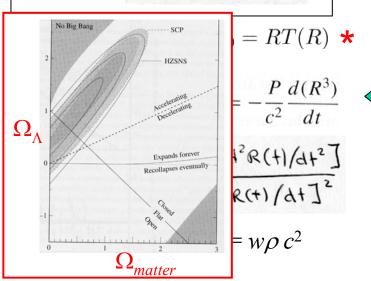
Radiation: 
$$\rho_r = \rho_{o,r} R^{-4}$$

Dark energy: 
$$\rho_{\rm A} = \rho_{o,\Lambda} R^0$$

$$\rho_{\rm c}(t) = \frac{3H^2(t)}{8\pi G}$$

$$\Omega(t) = \frac{\rho(t)}{\rho_c(t)}$$

$$\Omega \equiv \Omega_m + \Omega_{\rm rel} + \Omega_{\Lambda}$$



## **Physics**

## Per unit mass:

$$\left( \left( \frac{1}{R} \frac{dR}{dt} \right)^2 - \frac{8}{3} \pi G \rho \right) R^2 = -kc^2$$

$$\rho = \frac{u}{c^2} \star$$

$$(ds)^2 = (c dt)^2 - R^2(t) \left[ \left( \frac{d\varpi}{\sqrt{1 - k\varpi^2}} \right)^2 + (\varpi d\theta)^2 + (\varpi \sin\theta d\phi)^2 \right]$$

$$\left[ \left( \frac{1}{R} \frac{dR}{dt} \right)^2 - \frac{8}{3} \pi G \rho - \frac{1}{3} \Lambda c^2 \right] R^2 = -kc^2$$

Cosmological Constant (a.k.a. *Dark Energy*) Curvature  $k = \frac{1}{R_o^2} \times \begin{bmatrix} 0 \\ -1 \end{bmatrix}$ 

 $k = \frac{1}{\mathcal{R}_o^2} \times 0$ 

$$dU = -PdV \star$$

$$\frac{d^{2}R}{dt^{2}} = \left\{ -\frac{4}{3}\pi G \left[ \rho_{m} + \rho_{\text{rel}} + \frac{3(P_{m} + P_{\text{rel}})}{c^{2}} \right] + \frac{1}{3}\Lambda c^{2} \right\} R$$

★ = you should be able to write these down from memory.