Review Sessions

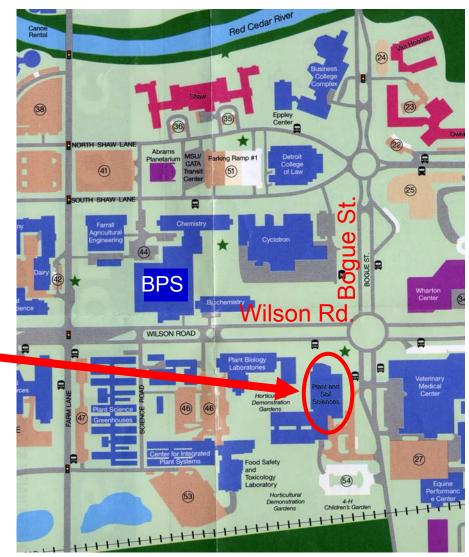
Two choices:

- Monday: 7PM
- or Tuesday: 6PM
- Both in

Plant & Soil Science (PSS) room A101 (corner of Wilson & Bogue)

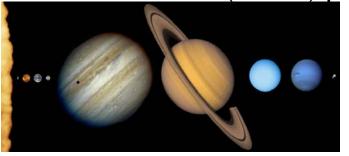
Print & bring Study Guide

• on course web site (reachable through Angel)



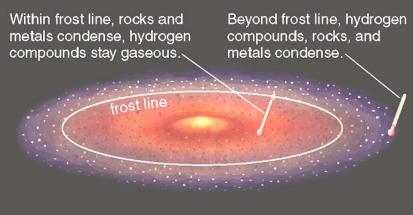
Overview of Solar System

- The solar system is a disk
 - Rotation of sun, orbits of planets all in same direction.
 - Most planets rotate in this same sense. (Venus, Uranus, Pluto are exceptions).
 - Angular momentum of pre-solar gas cloud.
- Terrestrial vs. Jovian (Giant) planets



- High vs. low density
 - Rocks vs. mostly gas
- Composition
 - heavy elements vs. primarily H/He
- Difference due to distance from Sun.

Object	% Total Mass
Sun	99.8
Jupiter	0.1
Comets	0.05
All other planets	0.04
Satellites & rings	0.00005
Asteroids	0.000002
Cosmic dust	0.0000001



Within the solar nebula, 98% of the material is hydrogen and helium gas that doesn't condense anywhere.

During planet formation in Solar Nebula: Presence of ice

- → more material for core
- → could gravitationally attract large masses of hydrogen & helium gas.

Terrestrial Planets

- Earth
 - Differentiated:
 - Iron/nickel core
 - Mantle of lighter rock
 - Thin crust on top
 - Plate Tectonics
 - Evolution of atmosphere
 - Thick $CO_2 \rightarrow life \rightarrow N_2, O_2$
 - Current global warming
 - Greenhouse effect
 - Man-made CO₂

- (Moon)
 - Impact craters as clocks
 - Old highlands (4.1-4.4 billion yrs)
 - Heavily cratered
 - Maria (3.3- 3.8 billion yrs)
 - Fewer craters
 - Rocks from each brought back by Apollo astronauts.
 - Age dating
 - Chemical composition
 - Tidally locked to Earth
 - Formation of Moon
 - Giant Impact is current favorite theory... collision between Earth & Mars-sized object.
- Mercury
 - Closest to Sun, eccentric orbit.
 - Airless, heavily cratered.
 - Hot, but (slightly) colder than Hell.
 - Very dense mostly iron-nickel core.
 - Geologically dead (probably)
 - But rupes → shrinkage at early time.
 - Rotates in 2/3 of its orbital period
 - Tidal locking with a twist.

Terrestrial Planets (continued)

Venus

- Differentiated like Earth
 - But no tectonic plates.
- Surface mostly studied by radar
 - Large volcanoes
 - "Continents" pushed up by tectonic flows in mantle.
 - Recent lava flows, constant resurfacing.
 - Crater density → very young surface
 - only 800 million yrs old.
- Thick CO₂ atmosphere
 - Result of runaway greenhouse effect.
 - Keeps surface very hot (900F).
 - Lead, brimstone (sulfer) are molten.
- Retrograde rotation
 - Probably due to giant impact.

Mars

- 50% smaller diameter than Earth
- 1.5 times further from Sun.
- Gigantic volcanoes.
- 50% highland "continents"
 - Tharsis bulge.
 - Cracked open to form Valles Marineris.
- 50% low-lying lava plains.
- Atmosphere

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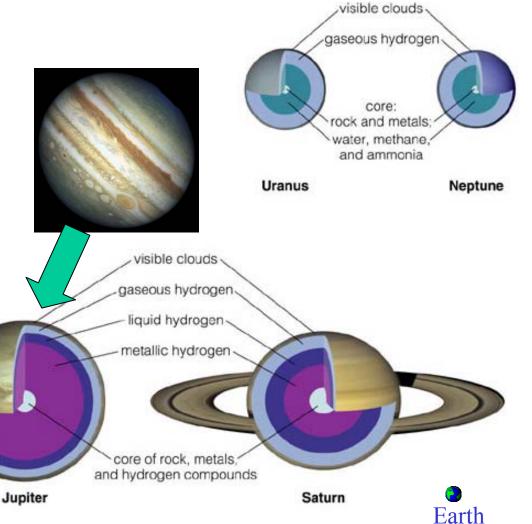
- CO₂, like Venus, but very thin.
- Liquid water currently impossible.
- Climate change
 - Loss of atmosphere
 - Low escape velocity
 - Solar wind
 - Could not retain heat
 - Water froze out
 - even less heat retained
 - 2 Rovers are finding evidence of past water.
 - Life?

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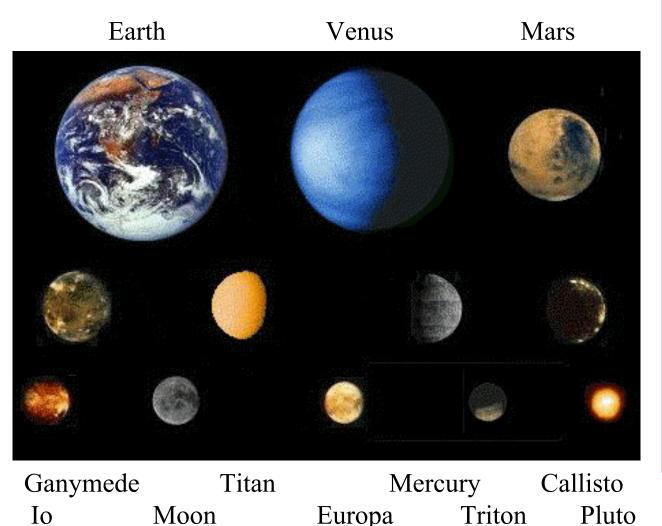
- Viking landers found no sign.
- Questionable data in meteorite.

The Giant Planets Jupiter – Saturn – Uranus - Neptune

- 14-300 x more massive than Earth.
- Massive H, He atmospheres
 - By far the most abundant elements in the solar system.
- On top of rock/ice core with 10-15 x mass of Earth.
- Lots of weather on Jupiter
 - Ammonia (NH₃) clouds.
 - Strong winds at different latitudes.
 - (differential rotation)
 - Cyclonic storms
 - Great Red Spot
 - -2 x size of Earth
 - 400 yrs so far
 - Investigated by Galileo probe.



Some planets and moons shown in correct relative sizes

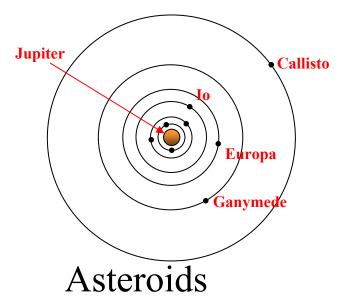


Europa

Planets: orbit around Sun Moons: orbit around planets But... asteroids, comets, dwarf planets = small bodies orbiting sun.

Moons

- Jupiter's Galilean moons, as we get closer to Jupiter:
 - Callisto ice, geologically dead.
 - Ganymede ice, but geologically active.
 - Europa rock, but covered by ice pack over liquid water.
 - Io rock, extreme volcanic activity.
- Gradient of properties due to increased tidal effects & heating from Jupiter.
- Jupiter's 59 other moons are much smaller.
- Saturn: 33 known moons
 - largest is Titan
 - N₂ atmosphere.
 - Similar to Earth's, but very cold (methane lakes).
 - Cassini/Huygens visit.



- Small rocky bodies in orbit about sun.
 - Left over from formation of Solar System.
- Most, but not all, in asteroid belt.
 - Some cross Earth's orbit

Meteorites

- Asteroids that hit Earth and don't burn up in atmosphere.
- Analyzing them \rightarrow
 - Age of solar system (4.5 billion yrs)
 - Initial chemical composition of solar system.

Rings

- All 4 giant planets have rings.
- Rings form inside Roche limit:
 - $P^2 = a^3 \rightarrow$ different parts of a moon try to move in orbits with different periods.
 - This tears bodies apart unless gravity (+ internal tensile strength) can hold them together.
 - For orbits inside Roche limit, prospective moons are torn apart.
- But rings constantly replenished by material abraded off small moons.
- Jupiter, Uranus, Neptune have very thin rings. Saturn has much larger rings.
- Shepherd satellites
 - moons sweep out divisions, contain rings through gravitational resonances.
- Rings made of ice and small bits of dust.

Comets

- Mostly ice
- Some on highly eccentric orbits
 - Spectacular tails when close to Sun.
 - Melted ice is driven off by solar radiation, solar wind.
- Most come from Oort Comet Cloud at edge of solar system.
 - Some from Kuiper Belt, just beyond Pluto.
- Pluto (& Charon)
 - No spacecraft visits, so little is known
 - Pluto probably quite similar to Triton.
 - Triton is Neptune's largest moon.
 - Retrograde orbit around Neptune.
 - 75% rock, 25% ice.
 - Very thin N_2 atmosphere.
 - Charon is half as big as Pluto.
 - Pluto now called a "Dwarf Planet"
 - Just one of several large Kuiper belt objects, all of which have:
 - Very low mass.
 - Eccentric, tilted orbits