
PHY492: Nuclear & Particle Physics

Lecture 21

Neutrinos and neutrino oscillations

HW hints

Beyond the Standard Model

Tier II Assignments
in E-mail this afternoon

Propagation of neutrino mass states

- Remember, Schrodinger wave equation solutions ?
- Remember, time dependent Schrodinger wave equation solutions ?

$$e^{-i(\omega t - kx)}; \quad \omega = \frac{E}{\hbar}; \quad k = \frac{p}{\hbar}$$

$$\psi(x, t) = \psi(x, 0) e^{-\frac{i}{\hbar}(Et - px)}; \quad \text{let } x = L, \quad t = L / c$$

$$\psi(L) = \psi(0) e^{-\frac{i}{\hbar c}(E - pc)L}$$

$$\psi(L) = \psi(0) e^{-\frac{i}{\hbar c} \frac{m^2 c^4}{2E} L}$$

$$pc = E \sqrt{1 - \frac{m^2 c^4}{E^2}} \approx E - \frac{m^2 c^4}{2E}, \quad E \gg mc^2$$

$$E - pc \approx \frac{m^2 c^4}{2E}$$

- Phase factor $e^{-\frac{i}{\hbar c} \frac{m^2 c^4}{2E} L}$ depends on distance L from production, particle energy, and mass **squared !!**

Three ν flavors and three ν masses

- As is the case for quarks, the mixing matrix is 3×3
- Very different from quarks, off-diagonal mixing angles are LARGE.

Missing one crucial angle, $\sin\theta_{13}$, and a CP violating phase δ

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = M \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$|M| \approx \begin{pmatrix} 0.83 & 0.55 & s_{13}e^{i\delta} \\ 0.39 & 0.59 & .7 \\ 0.39 & 0.59 & .7 \end{pmatrix}; \quad s_{13} < 0.25 \\ \delta \sim ?$$

- Result for muon neutrinos oscillating to electron neutrinos is the same as in the two neutrino case.

appearing

electron neutrinos

$$\left| \langle \nu_e | \nu(L) \rangle \right|^2 = \sin^2 2\theta_{13} \sin^2 \left(1.27 \Delta m_{23}^2 L / E \right)$$

$$\Delta m_{23}^2 \sim 3 \times 10^{-3} \text{ eV}^2; \quad \text{pick } E \sim 2 \text{ GeV}$$

$$\lambda = \frac{2\pi E}{1.27 \Delta m_{23}^2} = \frac{4\pi}{4 \times 10^{-3}} \text{ km} \approx 3000 \text{ km}$$

First maximum at $\lambda / 4 \approx 750 \text{ km}$

Fermilab neutrino beam
points to Sudan MN which
is 750 km away

Mass hierarchy of three mass states

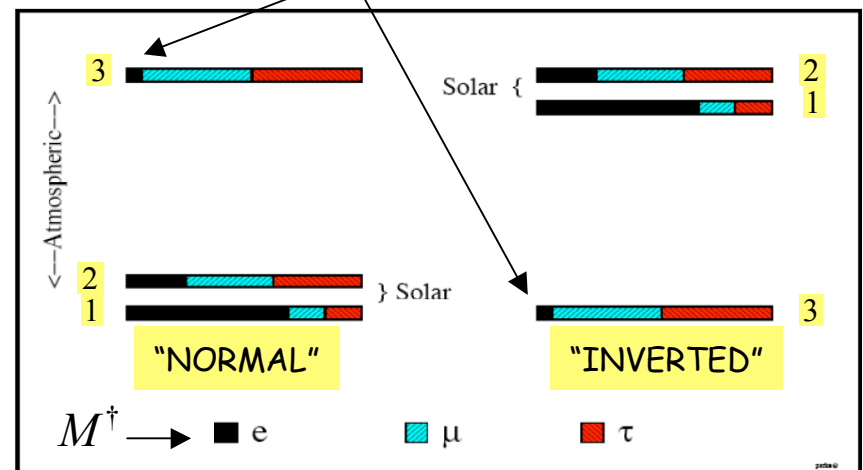
- Mass splitting $\Delta m_{ij}^2 = |m_j^2 - m_i^2|$

Solar: $\Delta m_{12}^2 \sim 5 \times 10^{-5} \text{ eV}^2$ **Small**

Atmospheric: $\Delta m_{23}^2 \sim 2.5 \times 10^{-3} \text{ eV}^2$ **Big**

e component not to scale

- Mass hierarchy **unknown**
 - "NORMAL" as in quark masses
 - "INVERTED" as in P-states?
 - Can't do CP violation on earth without hierarchy resolution



- MSW effect
 - In a long baseline experiment ν_e and $\bar{\nu}_e$ passing through the earth have oscillation probability affected with opposite signs.
 - Good for determination of mass hierarchy, but complicates CP violation analysis. Unavoidable, even in Japan.

HW hints

14.1 Top quark lifetime

$$\delta E \delta t \sim \hbar; \quad \tau \sim \frac{\hbar}{\Gamma}$$

Top quark decay


$$t \rightarrow W + b$$

Top produced \sim at rest

Total energy is top mass

W momentum = b-quark momentum

Parton-parton collision

$$p_a = x_a \frac{\sqrt{s}}{2c} \quad p_b = -x_b \frac{\sqrt{s}}{2c}$$


s = center of mass energy squared of pp collision

\hat{s} = center of mass energy of parton(a) parton(b) collision

$$= x_a x_b s$$

14.3

$$\Lambda = (uds); \quad \Sigma^+ = (uus)$$

$$K^0 = (\bar{s}d); \quad K^+ = (\bar{s}u)$$

$$\pi^+ = (u\bar{d}); \quad \pi^+ = \frac{1}{\sqrt{2}}(u\bar{u} + d\bar{d})$$

HW hints

14.5 Q=-1/3 Quark mixing matrix

Strong Interaction state

$$K^0 = (d, \bar{s})$$

$$\begin{pmatrix} d \\ s \end{pmatrix} = \begin{pmatrix} \cos \theta_c & -\sin \theta_c \\ \sin \theta_c & \cos \theta_c \end{pmatrix} \begin{pmatrix} d' \\ s' \end{pmatrix}$$

Weak interaction acts on mixed states: d', \bar{s}'

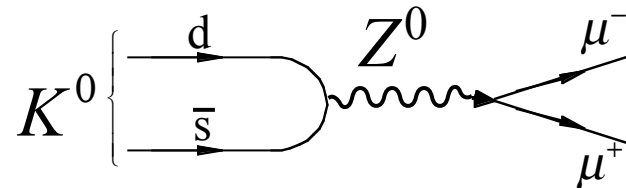
Z boson couples only to the same weak flavor quarks $d' + \bar{d}'$
 $s' + \bar{s}'$

$\langle d', \bar{s}' | Z \rangle$ and $\langle s', \bar{d}' | Z \rangle$ are both ZERO

$$\langle d', \bar{d}' | Z \rangle = \langle s', \bar{s}' | Z \rangle \neq 0$$

Calculate

$$\langle d, \bar{s} | Z \rangle$$



HW hint: Deep Inelastic Scattering (DIS)

14.7 a) 4-vector dot product (lab frame)

$$Q = [(\vec{k}' - \vec{k}), \nu]; \quad P = [0, m_p c^2]$$
$$Q \cdot P = ?$$

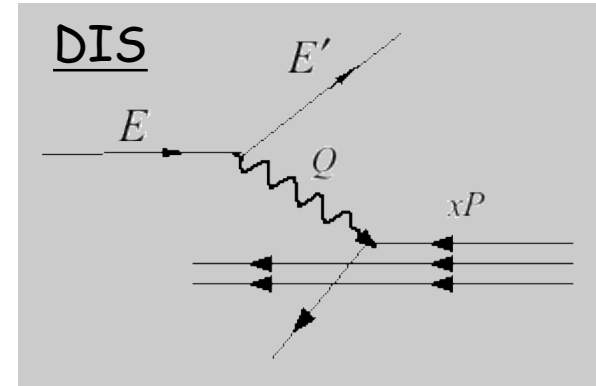
b) parton absorbs Q but remains massless

$$(x\vec{P} + \vec{Q})^2 = 0$$

c) from 14.6

W = mass of hadronic system

$$W^2 = m_p^2 + \frac{2m_p c^2 \nu}{c^4} - \frac{Q^2}{c^4}$$



Looking for the answers

Standard model is continually being tested for evidence of something that doesn't fit, that might shed light on one or more of these questions:

- Why is the boson symmetry broken; W/Z boson masses large, but the photon mass zero?
- Why do quarks have such wildly different masses?
- Why is the # of quark and lepton and generations equal to 3?
- Why do quarks and neutrinos have the observed mixing angles θ ? Quarks have small mixing θ 's, but neutrinos have large mixing θ 's.
- Why are neutrinos masses so small compared to anything?
- Why is gravity such a puny force?
- Why doesn't the proton decay, or does it?
- Why does matter dominate over antimatter? Why CP violation?
- What is dark matter?

Dynamical symmetry breaking, the Higgs Mechanism

- Weinberg, Glashow & Salam

- Basic theory has 4 *massless* bosons: triplet W_1, W_2, W_3 , and singlet B
- Dynamical Symmetry Breaking via the Higgs Mechanism
- W_1, W_2 become W^\pm , but W_3 , and B mix with "Weinberg angle" = θ_W

$$\begin{aligned} \gamma &= B^0 \cos \theta_W + W_3 \sin \theta_W \\ Z^0 &= -B^0 \sin \theta_W + W_3 \cos \theta_W \end{aligned}$$

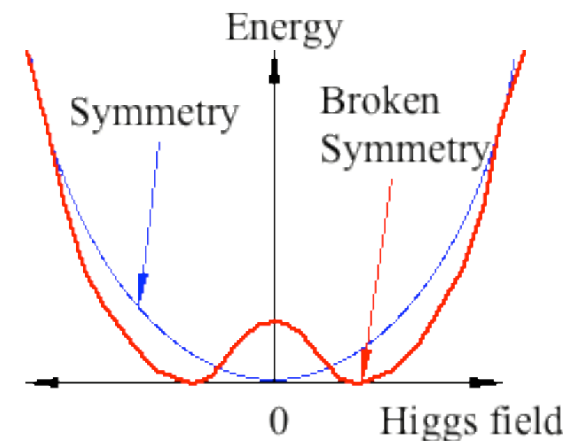
- W and Z masses are related

$$m_W / m_Z = \cos \theta_W$$

- θ_W measured in neutrino scattering

$$\sin^2 \theta_W = 0.233$$

- Predict: $m_W / m_Z = 0.88$ $m_W / m_Z = 80.4 / 91.2 = 0.88$



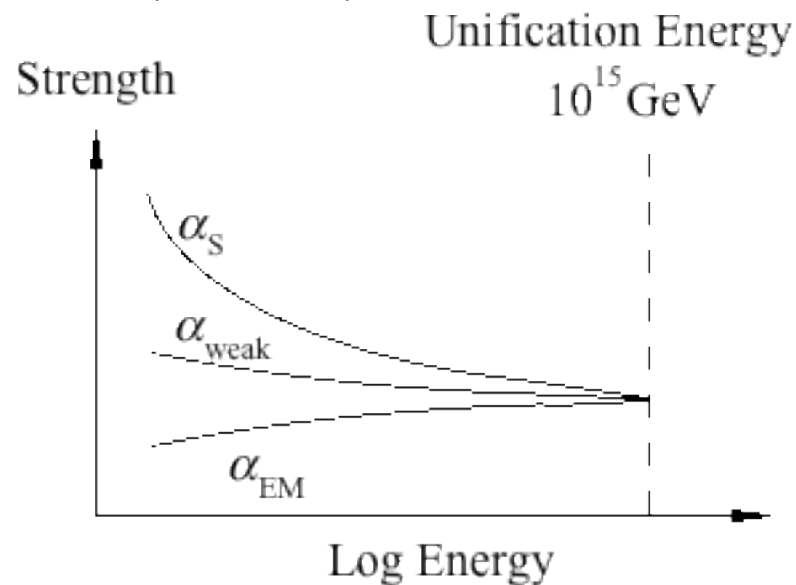
- Simplest dynamical symmetry breaking would have 1 Higgs scalar (spin 0) particle

$$H^0$$

$$200 \text{ GeV} > m_H > 115 \text{ GeV}$$

Grand Unified Theory (GUT)

- At very high energies everything would be highly symmetric
- All masses and couplings would be the same
- At low energies the symmetry is broken

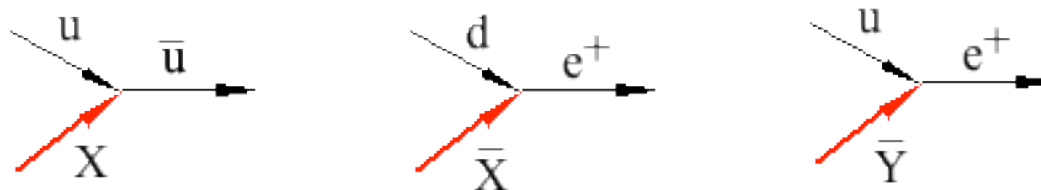


- EW group is $SU(2) \times U(1)$: 4 gauge bosons (γ , W^+ , W^- , Z)
- QCD group is $SU(3)_{\text{color}}$: 8 gauge bosons (gluons)
- Simplest GUT is $SU(5)$: ---> **24 gauge bosons** (12 more bosons)

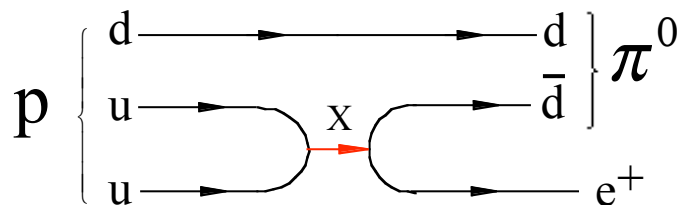
SU(5) has 12 more gauge particles

- Three with charge $= -1/3$, in three "colors" (Y_R, Y_G, Y_B)
- Three with charge $= -4/3$, in three "colors" (X_R, X_G, X_B)
- Six anti-particles of these
- These gauge particles can change quarks into leptons or change quarks into antiquarks

Examples



- Can lead to proton decay (lifetime is $< 6 \times 10^{30}$ years) **N.G.**



Measured lifetime lower limit
 $> 10^{31} - 10^{32}$ years

- Some success. **But no proton decay, no 3 generations, has monopoles**

Supersymmetry (SUSY)

- Every fermion has a supersymmetric boson partner and vice-versa
 - electron (e) \rightarrow selectron (\tilde{e})
 - quark (q) \rightarrow squark (\tilde{q})
 - photon (γ) \rightarrow photino ($\tilde{\gamma}$)
 - gluon (g) \rightarrow gluino (\tilde{g})
 - W, Z \rightarrow wino (\tilde{W}), zino (\tilde{Z})
 - Higgs (H) \rightarrow Higgsino (\tilde{H})
 - Higgs* (H^\pm) \rightarrow Higgsino* (\tilde{H}^\pm)
- R-parity $R=+1$ (particles) $R=-1$ (supersymmetric particles)
- R parity product conserved
 - supersymmetric particles produced in pairs ($A + B \rightarrow \tilde{A} + \tilde{B}$)
 $(1 \times 1 = -1 \times -1)$
 - Heavy supersymmetric particles decays ($\tilde{A} \rightarrow \tilde{B} + A$)
 $(-1 = -1 \times +1)$
 - Lightest supersymmetric particle cannot decay (might be **stop** quark)
 - Good **Dark Matter** candidate

Other qualities of Supersymmetry (SUSY)

- SUSY partners can cancel infinities that arise in calculations
- Unifies fermions and bosons
- Involves angular momentum inherently
 - rotational symmetries --- space translations
 - geometry of spacetime --- **GRAVITY**
- One major problem -- NOT a shred of evidence for it.
- Mass limit is greater than about 100 GeV
- Still looking at Fermilab Tevatron collider (CDF and D-Zero)
- Will look for it at LHC

Theories of everything

