

## When the Sun Dies—8 Oct

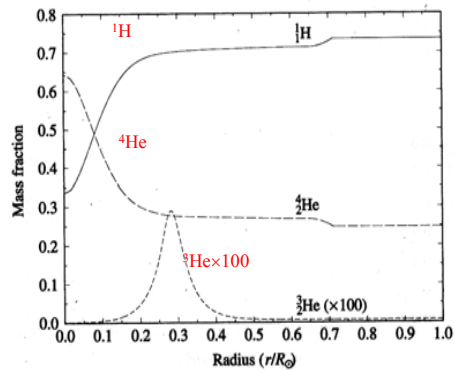
- Sun will use up the hydrogen in the center in 5Byr
- Center of sun must shrink to get hotter to balance gravity
  - Sun will become a red giant. Surface expands.
- Sun will become a planetary nebula
- Sun will become a white dwarf



- New schedule
  - <http://www.pa.msu.edu/courses/AST207/>
- Office hours 10/13-10/17
  - Jack Baldwin, 3270 BPS, MWF 12:00-13:00
- Observing (weather permitting)
  - Fri & Sat, 9:00-11:00pm
  - MSU Observatory, Forest & College Rd

## Composition of the sun

- In center, hydrogen is half used up.



## A Balancing Act

- All astronomical objects do a balancing act.
  1. What prevents the Earth's atmosphere from being dense at my feet but sparse at my head?
    - A. Gas pressure
    - B. The strength of the materials
    - C. Atoms change their directions of motion.
- 1. The Earth does a balancing act. What prevents the Earth from collapsing?
  - A. Gas pressure
  - B. The strength of the materials
  - C. Atoms change their directions of motion.

## A Balancing Act: Gravity vs. Gas Pressure

- What prevents the Earth's atmosphere from being dense at my feet but sparse at my head?
  - Gas pressure
  - The strength of the materials
  - Atoms change their directions of motion.
- Force of gravity balances gas pressure.
  - Force of gravity  $GM^2/R^2$
  - Force of gas  $PV=nkT$ 
    - $k$  is Boltzmann's constant.  $k=R/(\text{number in a mole})$
    - Details ( $m$  is mass of gas particle)
 
$$P = (nm)kT/m/V = M kT/(mR^3)$$

$$F = \text{area } P = R^2 M kT/(mR^3) = M kT/(mR)$$
  - Balance:  $GMm/R=kT$
- We are watching the birth of the sun. The not-yet sun is a gas cloud slowly shrinking. It is getting
  - warmer
  - cooler

## The sun's choice

- How does the sun produce energy at the present time?
  - Fuse hydrogen to produce helium
  - Fuse He to produce carbon
  - Fuse carbon with helium
  - Fuse neon
  - Fuse oxygen

Reaction	Min. Temp.
$4 \text{ } ^1\text{H} \rightarrow \text{}^4\text{He}$	10 MK
$3 \text{ } ^4\text{He} \rightarrow \text{}^{12}\text{C}$	200 MK
$\text{}^{12}\text{C} + \text{}^4\text{He} \rightarrow \text{}^{16}\text{O}, \text{Ne}, \text{Na}, \text{Mg}$	800 MK
$\text{Ne} \rightarrow \text{O}, \text{Mg}$	1500MK
$\text{O} \rightarrow \text{Mg}, \text{S}$	2000MK
$\text{Si} \rightarrow \text{Fe peak}$	3000MK

## The sun's choice

- Why does fusion of helium require a higher temperature?
  - Helium is heavier
  - Helium has 2 protons
  - Helium has two neutrons

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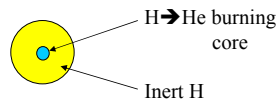
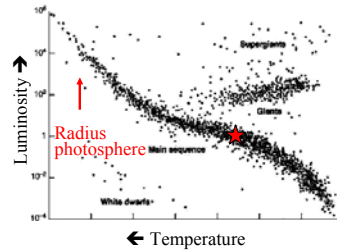
## The sun's choice

- Why does fusion of helium require a higher temperature?
  - Helium is heavier
  - Helium has 2 protons
  - Helium has two neutrons
  - With more charge, it takes higher speeds to bring two He nuclei close enough to fuse.
    - Carbon has 6 protons.

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## Sun as a main-sequence star

- H → He in the core
- T=15MK
- Fuel will last another 5 Byr.



## The sun's choice

- Sun does a balancing act.
    - $RT = k / (GMm)$
  - Sun must produce energy to replenish the energy radiated away.
  - If H → He shuts off, source of energy to maintain pressure shuts off, and gravity wins.
1. What does the sun do to adjust for gravity's victory?

Triple-alpha process

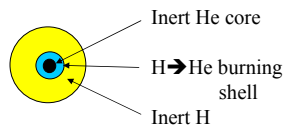
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## The sun's choice

- Sun does a balancing act.
  - $RT = k / (GMm)$
- Sun must produce energy to replenish the energy radiated away.
- If H → He shuts off, source of energy to maintain pressure shuts off, and gravity wins.
  - Sun shrinks.
  - Core of sun gets hotter
  - H → He in the a shell surrounding inert core
  - Balance restored.

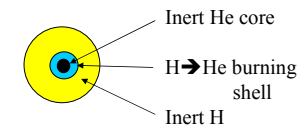
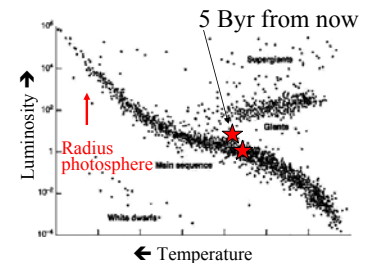
Triple-alpha process

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## Sun as a subgiant

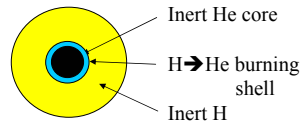
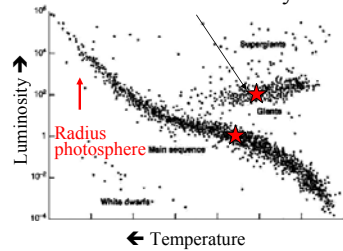
- H is gone in the core
- The never-ending battle between gravity and pressure. How does the sun adjust?
  - Without burning fuel to keep temperature up, pressure ( $PV = nRT$ ) would fall and gravity would win.
  - Core shrinks, gets hotter
  - H → He in the a shell surrounding inert core
  - Balance restored.



## Sun as a giant

A few 100Myr later

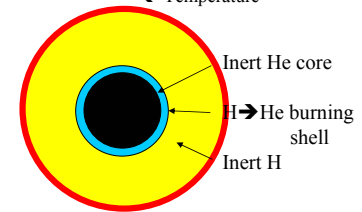
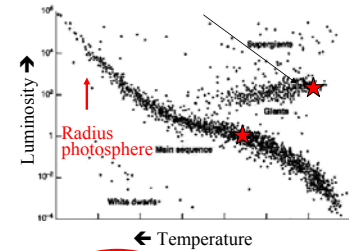
- H is gone in the core
- The never-ending battle between gravity and pressure. How does the sun adjust?
  - Without burning fuel to keep temperature up, pressure ( $PV=nRT$ ) would fall and gravity would win.
  - Core shrinks, gets hotter
  - $H \rightarrow He$  in the a shell surrounding inert core
  - Balance restored.
- Inert He core expands



## Sun as a giant

Later

- H is gone in the core
- The never-ending battle between gravity and pressure. How does the sun adjust?
  - Without burning fuel to keep temperature up, pressure ( $PV=nRT$ ) would fall and gravity would win.
  - Core shrinks, gets hotter
  - $H \rightarrow He$  in the a shell surrounding inert core
  - Balance restored.
- Inert He core expands



## The sun's choice

- Sun does a balancing act.
  - $RT=k/(GMm)$
- Sun must produce energy to replenish the energy radiated away.
- Without burning fuel to keep temperature up, pressure ( $PV=nRT$ ) would fall and gravity would win.
  - Core shrinks, gets hotter  $T=200MK$

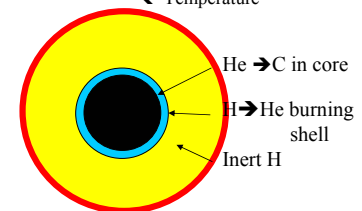
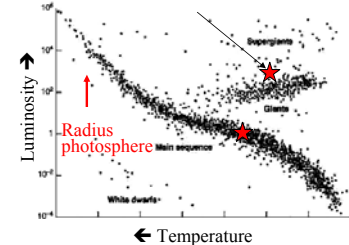
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$Ne \rightarrow O, Mg$	1500MK
$O \rightarrow Mg, S$	2000MK
$Si \rightarrow Fe\ peak$	3000MK

Triple-alpha process

## Sun Burns Helium

Later

- H is gone in the core & shell is exhausted
- The never-ending battle between gravity and pressure. How does the sun adjust?
  - Without burning fuel to keep temperature up, pressure ( $PV=nRT$ ) would fall and gravity would win.
  - Core shrinks, gets hotter  $T=200MK$
  - $3He \rightarrow C$  in the core (triple alpha process)
  - Balance restored.



## Other fusion reactions?



- Sun has one more trick after He is exhausted in core.
  - Burn He in a shell
- Sun is not massive enough to shrink further and get hotter
  - Core is supported by pressure of degenerate electrons.
  - Temperature does not rise to burn anything else.
- End of the road: planetary nebula & white dwarf core

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triple-alpha process