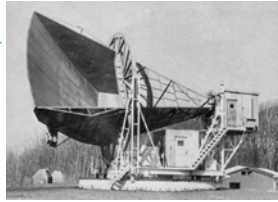


When Radiation Ruled— 24 Oct

- Homework 7 is due on Wed, not Mon.
- Picture of Penzias & Wilson's equipment in Deutsches Museum, Munich, Germany (Thanks to Joel Adelsberg)

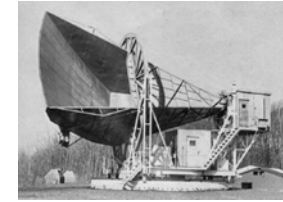


When Radiation Ruled—24 Oct

- At present, radiation from the Big Bang is weak
 - $T = 2.7 \text{ K}$
 - Has no affect on history of universe
- In past, radiation from the Big Bang was
 - Hot enough to change matter
 - Denser than matter
- Temperature and expansion

$$T / T_{\text{now}} = 1/a$$

$$a = \text{Dist} / \text{Dist}_{\text{now}}$$



Universe now



Matter: 0.1mg
 $T=0.8 \times 10^9 \text{ K}$



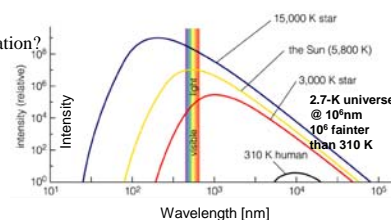
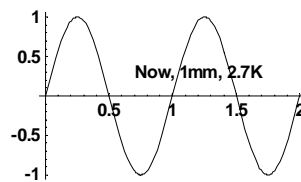
Rad: 0.6kg
 $T=0.8 \times 10^9 \text{ K}$

Universe at 3min

Expansion stretches wavelength of light

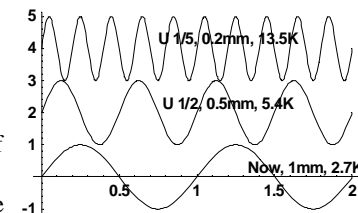
- We see black-body radiation with $T=2.7 \text{ K}$, and wavelength at the peak intensity $\lambda_{\text{max}} = 1 \text{ mm}$.

$$\lambda_{\text{max}} = 2.7 \text{ mm} \cdot \text{K} / T \text{ (Wein's Law)}$$
 - Principle: Wavelength of radiation stretches by the same factor as the universe expands.
1. When the U was half the present size, what was the wavelength at the peak intensity?
 - A. 0.5 mm
 - B. 1 mm
 - C. 2 mm.
 2. What was the temperature of the radiation?
 - A. 1.3 K
 - B. 2.7 K
 - C. 5.4 K



Expansion stretches wavelength of light

- Wavelength of radiation stretches by the same factor as the universe expands.
1. When the U was half the present size, what was the wavelength at the peak intensity? 0.5 mm
 2. What was the temperature of the radiation? 5.4 K
- Key idea: When the universe was smaller (when the distance between us and some object was smaller), the temperature was hotter.



Book-burning Universe

- Key idea: When the universe was smaller (when the distance between us and some object was smaller), the temperature was hotter. There is no obvious limit to the temperature.
- At one time, the universe was too hot to have paper. (Paper burns.)
 - Occurs at $451\text{ F} = 500\text{ K}$.
 - (In reality, there was no carbon and no paper at that time.)
- Define the expansion parameter a to be
 - $a = \text{distance between two objects} / \text{present distance}$
- 2. Hoag's object is 300 Mpc from the Milky Way. How far was it when the U was just hot enough to burn paper?
 - A. 30 Mpc, $a=1/10$, $T=2.7 * 10 = 27\text{K}$
 - B. 10 Mpc, $a=1/30$
 - C. 3 Mpc, $a= 1/100$
 - D. 1 Mpc, $a= 1/300$



Book-burning Universe

- Key idea: When the universe was smaller (when the distance between us and some object was smaller), the temperature was hotter. There is no obvious limit to the temperature.
- At one time, the universe was too hot to have paper.
- 1. What other familiar things were not possible at one time? What other reactions might have occurred when the universe was smaller & hotter.



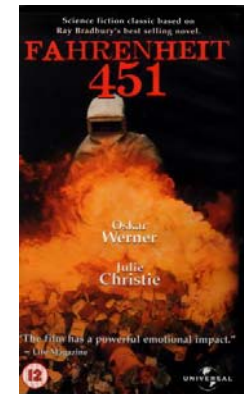
Book-burning Universe

- Key idea: When the universe was smaller (when the distance between us and some object was smaller), the temperature was hotter. There is no obvious limit to the temperature.
- At one time, the universe was too hot to have paper.
- 1. What other familiar things were not possible at one time? What other reactions might have occurred when the universe was smaller & hotter.
 - U was too hot to have stars.
 - U was too hot to have molecules.
 - U was so hot that atoms were ionized.
 - U was too hot to have nuclei other than hydrogen.



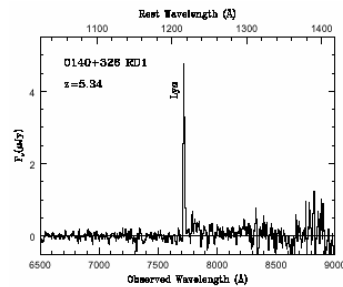
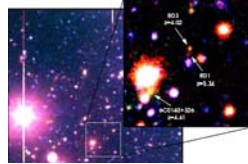
Book-burning Universe

- Key idea: When the universe was smaller (when the distance between us and some object was smaller), the temperature was hotter. There is no obvious limit to the temperature.
- What other reactions might have occurred when the universe was smaller & hotter?
- Events in the universe's life
- First stars formed
 - When U cooled enough, gravity was able to overcome pressure.
- Recombination: U changed from opaque to transparent
 - Ionization & recombination
 - Free $p + e \rightarrow$ hydrogen atom
- Production of the first nuclei other than H
 - Nuclear reaction
 - Free protons + neutrons \rightarrow helium nucleus



Expansion parameter & redshift

- Expansion parameter
 - a = distance between two galaxies / present distance
 - The expansion parameter a changes from 0 at the Big Bang to 1 at the present.
- Universe expand the same as wavelength of light.
- 1. When the light that we see left galaxy 0140+326 RD1, its wavelength was 1215 Å (121.5nm). When the light that we see now left the galaxy, the expansion parameter of the universe was
 - less than 1
 - 1
 - greater than 1.



Expansion parameter & redshift

- Expansion parameter
 - a = distance between two galaxies / present distance
 - The expansion parameter a changes from 0 at the Big Bang to 1 at the present.
- Universe expand the same as wavelength of light.
- 1. When the light that we see left Galaxy 0140+326 RD1, its wavelength was 1215 Å (121.5nm). By what factor has the universe expanded since the time the light left that galaxy?
 - We see its wavelength to be 7710Å. U has expanded by a factor of 6.35 since the time the light left that galaxy.
 - $a = 1215/7710 = 1/6.35$.
- Redshift z
 - $z = 1/a - 1$
 - $z = \lambda_{\text{received}} / \lambda_{\text{emitted}} - 1$
 - Speed = $z \times \text{speed of light}$ (for small z)
- For galaxy 0140+326 RD1,
 - $z = 7710/1215 - 1 = 6.34 - 1 = 5.34$

