

Name (please print): \_\_\_\_\_

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PHY321

Practice Final Exam

Spring 2008

There are four questions. Total points = 50. Show all of your work!

1. [15 points] A damped oscillator consists of a spring (with a spring constant  $k = 6.00 \text{ N/m}$ ), a block of mass  $m = 1.50 \text{ kg}$ , and a damping force given by  $-b\dot{x}$  (where  $b = 0.60 \text{ kg/s}$ ). Suppose that the block is initially pulled so that the spring is extended a distance of  $x = 0.12 \text{ m}$  and then released from rest.

(a) [3 points] What is the equation of motion for the block?

$$\ddot{x} + 2\beta\dot{x} + \omega_0^2 x = 0$$

$$\ddot{x} + 2(0.2)\dot{x} + 2^2 x = 0$$

$$\Rightarrow \boxed{\ddot{x} + 0.4\dot{x} + 4x = 0}$$

$$\omega_0 = \sqrt{\frac{k}{m}} = \sqrt{\frac{6.0}{1.5}} = 2.0 \text{ s}^{-1}$$

$$\beta = \frac{b}{2m} = \frac{0.6}{2 \cdot 1.5} = 0.2 \text{ s}^{-1}$$

$\omega_0 > \beta \Rightarrow$  underdamped

(b) [3 points] What is the solution to the equation of motion i.e. what is  $x(t)$ ?

$$x = A e^{-\beta t} \cos \omega_1 t \quad \text{with } \omega_1 = \sqrt{\omega_0^2 - \beta^2} = 1.99 \text{ s}^{-1}$$

$$\text{At } t=0 \quad x = A = 0.12 \text{ m}$$

$$\Rightarrow \boxed{x(t) = 0.12 e^{-0.2t} \cos(1.99t)}$$

(c) [3 points] Calculate the time required for the amplitude of the oscillations to fall to one-fifth of its initial value.

$$A/5 = A e^{-\beta t} \Rightarrow \frac{1}{5} = e^{-0.2t}$$

$$\therefore \ln \frac{1}{5} = -1.609 = -0.2t \quad \therefore t = \frac{1.609}{0.2}$$

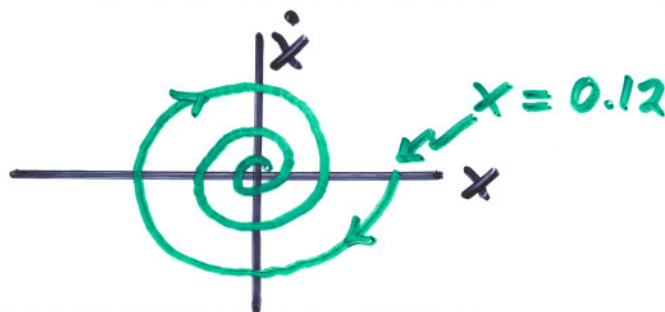
$$\Rightarrow \boxed{t = 8.05 \text{ sec}}$$

(d) [3 points] How many oscillations are made by the block in this time?

$$T = \frac{2\pi}{\omega_1} = \frac{2\pi}{1.99} = 3.16 \text{ sec}$$

$$\# \text{ oscillations in } 8.05 \text{ sec} = \frac{8.05}{3.16} = \boxed{2.55}$$

(e) [3 points] Sketch (roughly) the phase path ( $\dot{x}$  vs.  $x$ ).



$$V = \frac{4}{3} \pi r^3 \Rightarrow dV = 4\pi r^2 dr$$

2. [10 points] A sphere of radius  $R = 15.0$  m has a density that varies as  $\rho(r) = \rho_0 r/R$  and  $\rho_0 = 3.0 \text{E}3 \text{ kg/m}^3$ .

$$dM = \rho dV = 4\pi \rho r^2 dr$$

(a) [4 points] What is the total mass of the sphere?

$$\Rightarrow M = \int_0^R 4\pi \rho_0 \frac{r}{R} r^2 dr = 4\pi \frac{\rho_0}{R} \int_0^R r^3 dr$$

$$\therefore M = \frac{4\pi \rho_0}{R} \cdot \frac{1}{4} r^4 \Big|_0^R = \frac{\pi \rho_0}{R} \cdot R^4 = \pi \rho_0 R^3 = \boxed{3.18 \text{E}7 \text{ kg}}$$

(b) [6 points] What is the gravitational force on an object of mass  $m = 2.0$  kg situated at a radius of  $r = 10.0$  m, i.e.  $5.0$  m beneath the outer surface of the sphere? (You can use Gauss' Law or the Shell Theorem.)

Shell Theorem  $\Rightarrow$

$$M' = \pi \frac{\rho_0}{R} \cdot 10^4 = 6.28 \text{E}6 \text{ kg}$$

Acts as if all at center

$$\Rightarrow F = \frac{GmM'}{10^2} = \boxed{8.38 \text{E}-6 \text{ N}}$$

$$\text{G's Law} \Rightarrow 4\pi 10^2 \vec{g} = -4\pi G M' \vec{e}_r$$

$$\therefore \vec{g} = -\frac{G M'}{10^2} \vec{e}_r$$

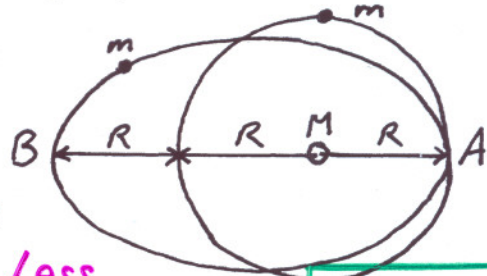
$$\Rightarrow F = mg = -\frac{GmM'}{10^2}$$

3. [12 points] Spaceships of equal mass  $m$  are in orbits around a planet of mass  $M$  (with  $M \gg m$ ). One spaceship is in a circular orbit (of radius  $R$ ) and one in an elliptical orbit. The orbits intersect at point A. (Note: you have been given  $G$ ,  $m$ ,  $M$  and  $R$ . Please express your answers in terms of these variables.)

In circular orbit

$$U = -\frac{GmM}{R} \quad KE = +\frac{1}{2} \frac{GmM}{R}$$

$$\Rightarrow E_{\text{circ}} = -\frac{1}{2} \frac{GmM}{R}$$



(a) [2 points] Which spaceship has the larger total energy?

$$E_{\text{ellipse}} = -\frac{GmM}{2a} = -\frac{GmM}{3R} \Rightarrow \text{Less negative so } E_{\text{ellipse}} > E_{\text{circ}}$$

(b) [2 points] What is the velocity of the spaceship in the circular orbit?

$$\frac{1}{2} m v^2 = \frac{1}{2} \frac{GmM}{R} \Rightarrow \boxed{v = \sqrt{\frac{GM}{R}}}$$

(c) [2 points] What is the velocity of the spaceship in the elliptical orbit at point A?

$$E_{\text{ellipse}} = K_A + U_A \Rightarrow K_A = \frac{1}{2} m v_A^2 = -\frac{GmM}{3R} + \frac{GmM}{R} = \frac{2}{3} \frac{GmM}{R}$$

(d) [2 points] What is the velocity of the spaceship in the elliptical orbit at point B?

$$E_{\text{ellipse}} = K_B + U_B \Rightarrow K_B = \frac{1}{2} m v_B^2 = -\frac{GmM}{3R} + \frac{GmM}{2R} \Rightarrow \boxed{v_A = \sqrt{\frac{4GM}{3R}}}$$

(e) [4 points] What is the eccentricity of the ellipse?

$$\left. \begin{aligned} r_{\min} &= R = \frac{a}{1+E} \\ r_{\max} &= 2R = \frac{a}{1-E} \end{aligned} \right\} \begin{aligned} r_{\max} &= \frac{2R}{1-E} = \frac{1+E}{1-E} \\ r_{\min} &= R = \frac{1+E}{1-E} \end{aligned}$$

$$\Rightarrow 2(1-E) = 1+E \quad \therefore 3E = 1 \quad \therefore \boxed{E = \frac{1}{3}}$$

Conservation of mom. in x direction:  $m_1 u_1 + 0 = m_1 v_1 \cos \psi + m_2 v_2 \cos \xi$   
 $\Rightarrow m_2 v_2 \cos \xi = 8 \cdot 15 - 8 \cdot 8 \cdot \cos 60^\circ = 88.0$

4. [13 points] A particle of mass  $m_1 = 8.0$  kg traveling at an initial velocity  $u_1 = 15.0$  m/s in the positive x direction collides with a stationary particle of mass  $m_2 = 12.0$  kg. After the collision, the first particle is traveling with speed  $v_1 = 8.0$  m/s at an angle  $\psi = 60.0^\circ$  with respect to the x direction.

conservation of mom. in y direction:  $0 = m_1 v_1 \sin \psi - m_2 v_2 \sin \xi$

(a) [4 points] What are the final speed and direction of the second particle? (i.e. find  $v_2$  and  $\xi$ .)  
 Warning: do not assume that the collision is elastic; it is not.)

$$\Rightarrow m_2 v_2 \sin \xi = 8 \cdot 8 \cdot \sin 60^\circ = 55.4$$

Dividing  $\Rightarrow \frac{\sin \xi}{\cos \xi} = \tan \xi = \frac{55.4}{88.0} = 0.630 \Rightarrow \xi = 32.2^\circ$

substituting  $\Rightarrow v_2 = \frac{88.0}{12 \cdot \cos 32.2} = 8.67 \text{ m/s}$

(b) [2 points] How much kinetic energy is lost in the collision?

$$K_i = \frac{1}{2} \cdot 8 \cdot 15^2 = 900 \text{ J}$$

$$K_f = \frac{1}{2} \cdot 8 \cdot 8^2 + \frac{1}{2} \cdot 12 \cdot 8.67^2 = 256 + 451 = 707$$

$$\Rightarrow \Delta K_{\text{lost}} = 900 - 707 = 193 \text{ J}$$

(c) [2 points] What is the velocity of the center of mass (CM) frame? Draw a picture of the situation in the CM frame before the collision, showing clearly the initial velocities  $u'_1$  and  $u'_2$ .

$$V_{\text{cm}} = \frac{m_1 u_1}{m_1 + m_2} = \frac{8 \cdot 15}{8 + 12} = 6 \text{ m/s}$$

$u'_1 = u_1 - V = 9 \text{ m/s}$        $u'_2 = 0 - 6 \text{ m/s} = -6 \text{ m/s}$

(d) [5 points] Draw a picture of the situation in the CM frame after the collision. What are the final speeds  $v'_1$  and  $v'_2$  of the two particles in the CM frame? What is the scattering angle  $\theta$  in the CM frame?

$$v'_{1y} = v_{1y} = v_1 \sin \psi = 8 \cdot \sin 60 = 6.93$$

$$v'_{1x} = v_{1x} - V = 8 \cdot \cos 60 - 6 = 4 - 6 = -2$$

$$\Rightarrow v'_1 = \sqrt{6.93^2 + (-2)^2} = 7.21 \text{ m/s}$$

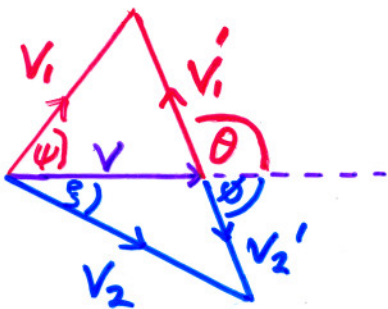
$$\tan \theta = \frac{6.93}{-2} \Rightarrow \theta = 106^\circ$$

$$v'_{2y} = v_{2y} = v_2 \sin \xi = 8.67 \sin 32.2 = 4.62 \text{ m/s}$$

$$v'_{2x} = v_{2x} - V = 8.67 \cos 32.2 - 6 = 1.34 \text{ m/s}$$

$$\Rightarrow v'_2 = \sqrt{4.62^2 + 1.34^2} = 4.81 \text{ m/s}$$

$$\tan \phi = \frac{4.62}{1.34} = 74^\circ \Rightarrow \theta + \phi = 106 + 74 = 180^\circ$$



PHY 321 Formulae and Constants

$$\omega_0^2 = \frac{k}{m} \quad \beta = \frac{b}{2m}$$

Oscillations: SHM  $\ddot{x} + \omega_0^2 x = 0$   $x(t) = A \sin(\omega_0 t - \delta)$

Damped SHM  $\ddot{x} + 2\beta\dot{x} + \omega_0^2 x = 0$  The general solution is:-  
 $x(t) = e^{-\beta t} [A_1 \exp(\sqrt{(\beta^2 - \omega_0^2)t}) + A_2 \exp(-\sqrt{(\beta^2 - \omega_0^2)t})]$  depending on the relative value of  $\omega_0^2$  and  $\beta^2$ .

$$T = \frac{2\pi}{\omega}$$

for underdamped  $x(t) = e^{-\beta t} [A \cos \omega_1 t + B \sin \omega_1 t]$   $\omega_1^2 = \omega_0^2 - \beta^2$

Gravity:  $\vec{F} = -G \frac{Mm}{r^2} \hat{e}_r$   $U(r) = -G \frac{Mm}{r}$   $\vec{g} = \frac{\vec{F}}{m}$   $\Phi = \frac{U}{m}$   $\vec{g} = -\nabla\Phi$   
 $V = \frac{4}{3}\pi r^3 \Rightarrow dV = 4\pi r^2 dr$

Gauss' Law:  $\nabla \cdot \vec{g} = -4\pi G\rho$   
 $\int \vec{g} \cdot d\vec{a} = -4\pi GM$

Know the Shell Theorem!

Calculus in spherical coordinates:  $\nabla\psi = \hat{e}_r \frac{\partial\psi}{\partial r} + \hat{e}_\theta \frac{1}{r} \frac{\partial\psi}{\partial\theta} + \hat{e}_\phi \frac{1}{r \sin\theta} \frac{\partial\psi}{\partial\phi}$

$$\nabla \cdot \vec{A} = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 A_r) + \frac{1}{r \sin\theta} \frac{\partial}{\partial\theta} (A_\theta \sin\theta) + \frac{1}{r \sin\theta} \frac{\partial A_\phi}{\partial\phi}$$

Central force motion: reduced mass  $\mu = \frac{Mm}{M+m}$

$$\vec{L} = \vec{r} \times \vec{p} \quad l = |\vec{L}| = \mu r^2 \dot{\theta} \quad K = \frac{1}{2} \mu (\dot{r}^2 + r^2 \dot{\theta}^2) = \frac{1}{2} \mu \dot{r}^2 + \frac{l^2}{2\mu r^2}$$

$$E = K + U = \frac{1}{2} \mu v^2 + U(r) = \frac{1}{2} \mu \dot{r}^2 + \frac{l^2}{2\mu r^2} + U(r) = \frac{1}{2} \mu \dot{r}^2 + V(r)$$

Orbits in gravitational field (Kepler's problem):  $U(r) = -G \frac{Mm}{r} = \frac{-k}{r}$

$$\frac{\alpha}{r} = 1 + \epsilon \cos\theta \quad \alpha = \frac{l^2}{\mu k} \quad \epsilon = \sqrt{1 + \frac{2El^2}{\mu k^2}} \quad a = \frac{k}{2|E|} = \frac{\alpha}{1 - \epsilon^2}$$

$\epsilon = 0$ : circle  $0 < \epsilon < 1$ : ellipse  $\epsilon = 1$ : parabola  $\epsilon > 1$ : hyperbola

Elliptical orbits:  $\langle K \rangle = \frac{-1}{2} \langle U \rangle$   $E = \frac{-k}{2a}$  Kepler's 3rd Law:  $\tau^2 = 4\pi^2 \frac{\mu}{k} a^3$

Astronomical Data:  $G = 6.67 \times 10^{-11} \text{ Nm}^2/\text{kg}^2$

$$L = T - U \quad \frac{\partial L}{\partial q} - \frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}} \right) = 0$$