

Model of Stars—5 Oct

- Hot-plate model of a star
- Thermal radiation
- Hertzsprung-Russell diagram
- Homework 4
 - Due Mon, Oct 12
 - Missouri Club on Fri.

The Hot-plate Model of a Star

- The surface of a star is made of tiles of hot plates.
- How does the energy from the hot-plate get to my hand?
 - Key observation: I can hold my hand much closer to the hot plate when it faces to the side, rather than up.



http://www.acemart.com/graphics/00000001/products/WELLh70_01.jpg

The Hot-plate Model of a Star

- How does the energy from the hot-plate get to my hand?
 - Key observation: I can hold my hand much closer to the hot plate when it faces to the side, rather than up.
- Energy moves from the hot plate to my hand by
 - movement of hot air
 - by radiation (mostly infrared light)
- 1. How does energy move from the sun to the earth?
 - A. By radiation only
 - B. By movement of hot air only
 - C. Both A & B



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The Hot-plate Model of a Star

- The surface of a star is made of tiles of hot plates.
- 1. How does energy move from the sun to the earth?
 - A. By radiation only
 - B. By movement of hot air only
 - C. Both A & B
- Energy leaves stars primarily by radiation.
 - For the sun, the radiation is mostly ultraviolet light, visible light and infrared light.



http://www.acemart.com/graphics/00000001/products/WELLh70_01.jpg

The Hot-plate Model of a Star

- The surface of a star is made of tiles of hot plates.
- We concentrate on the radiation produced by the hot plate.
- 1. What is a way to make hot plates produce more energy per second? (The same question applies to a star: What are two ways to make a star brighter or more luminous?)
 - A. Make the plates hotter.
 - B. Make the plates bigger.
 - C. Make plates hotter & bigger.
 - D. None of the above answers.
- The luminosity of a star (the energy produced every second) depends on temperature and size.



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The Hot-plate Model of a Star

- The surface of a star is made of tiles of hot plates.
- The luminosity of a star (the energy produced every second) depends on temperature and size.
- 1. What can I do to make the same hot-plate at the same setting burn my hand and not burn my hand? (Without modifying the sun, what can I do to make the sun brighter or fainter?)
 - A. Move my hand closer or farther.
 - B. It is not possible.
- The flux of a star (the energy received at the earth every second) depends on temperature, size, and distance to the star.



http://www.acekart.com/graphics/00000001/products/WELLh70_01.jpg

Thermal radiation (Blackbody Radiation)

- Any object that absorbs light also emits light.
- Do people emit light?
 - People emit light in the “thermal infrared” part of the spectrum.
 - Your eyes cannot see infrared radiation. $8000 < \lambda < 12000 \text{nm}$
 - You can see visible light
 - Blue 440nm
 - Green 550nm
 - Red 620nm
- A perfect absorber (perfectly black) emits a characteristic spectrum of light. (Called thermal or black-body radiation.)
 - Intensity depends only on
 - Temperature
 - Area
- A non-perfect absorber (grey body) with emissivity ϵ absorbs a fraction ϵ and reflects a fraction $(1-\epsilon)$.
 - Intensity is ϵ that of thermal radiation.



Picture taken with an infrared camera
ornitorinko.org:8080/.../portrait-bits.jpg

Infrared camera—Seeing with infrared eyes

- A perfect absorber (perfectly black) emits a characteristic spectrum of light. (Called thermal or black-body radiation.)
 - Intensity depends only on
 - Temperature
 - Area
 - A non-perfect absorber (grey body) with emissivity ϵ absorbs a fraction ϵ and reflects a fraction $(1-\epsilon)$.
 - Intensity is ϵ that of thermal radiation.
 - Thermal infrared
 - Wavelength is 8,000-12,000 nm
 - An object with a temperature of 300K emits most of its light in the thermal infrared.
 - Does infrared light show the same thing as visible light?
- Q Which is the hottest part of the man's face?
 - A. His hair.
 - B. His forehead.
 - C. His eyeglasses.
 - D. His moustache.



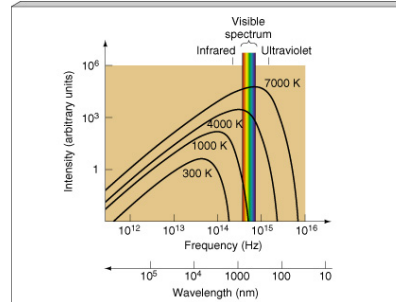
Model of a Star: Thermal Radiation

- Wien's Law

$$\lambda_{\text{peak}} T = 2.9 \text{ mm K}$$

- Brighter for hotter objects
- Wavelength changes with temperature

- For the sun, $T=5700\text{K}$.
 - λ_{peak}
 $= 2.9 \text{ mm} / 5700 \text{ K} = .0005 \text{ mm} = 500 \text{ nm}$
- For a person, $T=273+37=310\text{K}$.
 - $\lambda_{\text{peak}} = 2.9 \text{ mm} / 310 \text{ K} = .01 \text{ mm}$
(infrared)

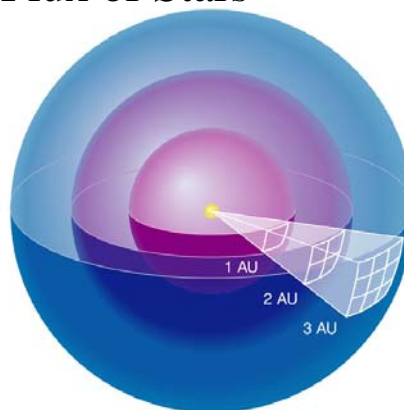


- Stefan-Boltzmann Law

- Energy emitted per second depends on $\text{Area} \times T^4$.

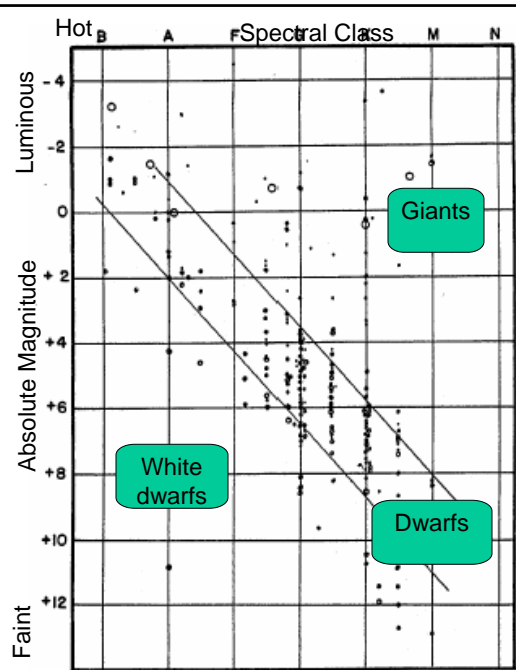
Luminosity & Flux of Stars

- Luminosity = amount of energy per second (Watt) produced by the star
 - $L=R^2T^4$
- Flux = energy per second received by a detector on earth (Watt/m^2)
 - $F=L/D^2$
- At greater distances from star, light is spread over larger area. Flux is lower.



Hertzsprung-Russell diagram

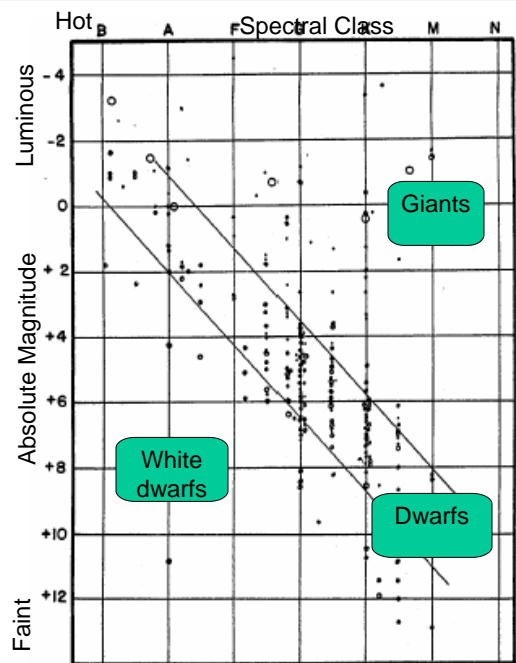
- HR Diagram is plot of temperature & luminosity
 - Spectral Class is related to temperature
 - OBeAFineGirlKissMe.
 - Hottest stars on left
 - Most luminous stars are on the top.
- S-B Law: $L=AT^4$.
- 1. A star is moved 10 times farther away. In the HRD, it moves
 - up
 - down
 - left
 - right
 - not at all



Astrophysics and twentieth-century astronomy to 1950, O Gingerich, ed., Cambridge, 1984

Hertzsprung-Russell diagram

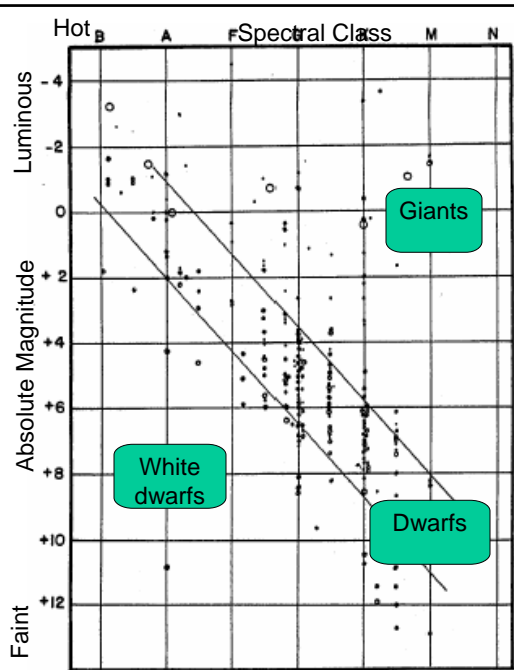
- HR Diagram is plot of temperature & luminosity
 - Spectral Class is related to temperature
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 - Hottest stars on left
 - Most luminous stars are on the top.
- S-B Law: $L=AT^4$.
- 1. A star is gets hotter and its size does not change. In the HRD, it moves
 - up & left
 - up & right
 - up-down
 - left-right
 - not at all



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Hertzsprung-Russell diagram

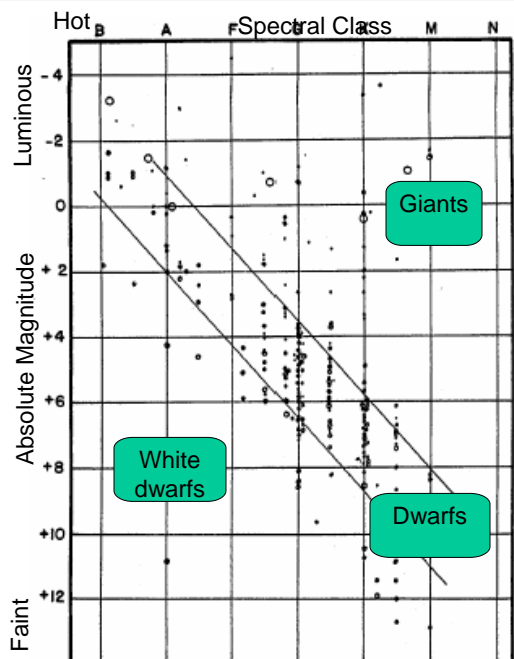
- HR Diagram is plot of temperature & luminosity
 - Spectral Class is related to temperature
 - OBeAFineGirlKissMe.
 - Hottest stars on left
 - Most luminous stars are on the top.
- S-B Law: $L=AT^4$.
- 1. Can two stars of the same spectral class have different luminosities?
 - No. No such cases exist on the H-R diagram.
 - Yes, temperatures differ
 - Yes, sizes differ
 - Yes, both size & temperatures differ.



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Hertzsprung-Russell diagram

- HR Diagram is plot of temperature & luminosity
 - Spectral Class is related to temperature
 - OBeAFineGirlKissMe.
 - Hottest stars on left
 - Most luminous stars are on the top.
- S-B Law: $L=AT^4$.
- H-R diagram reveals stars are hot-plates.
 - Dwarfs have differing temperatures and approximately the same size.
 - Giants are large.
 - White dwarfs are small.



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