

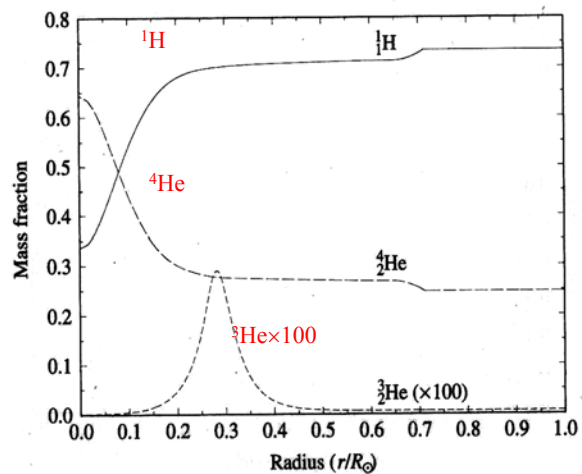
Life of the Sun—16 Oct

- Sun will use up the hydrogen in the center in 5Byr
- Center of sun must shrink to get hotter to balance gravity
- Sun will become a red giant. Surface expands.
- Sun will become a planetary nebula
- Sun will become a white dwarf



Composition of the sun

- In center, hydrogen is half used up.



A Balancing Act

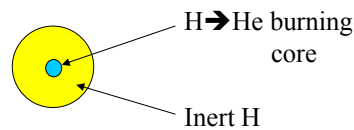
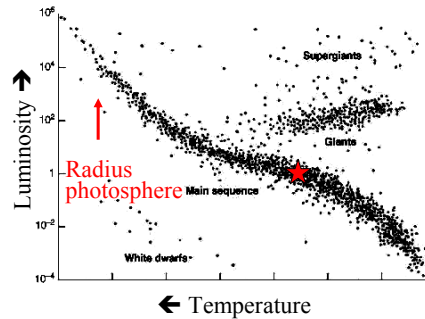
- All astronomical objects do a balancing act.
 - Gravity pulls inward.
 - Something else pushes outward or gravity causes acceleration to change the motion.
- 1. The Earth does a balancing act. What prevents the Earth from collapsing?
 - A. Gas pressure
 - B. The strength of the materials
 - C. Atoms change their directions of motion.
- 1. What prevents the Earth's atmosphere from being dense at my feet but sparse at my head?
 - A. Gas pressure
 - B. The strength of the materials
 - C. Atoms change their directions of motion.

A Balancing Act: Gravity vs. Gas Pressure

- Force of gravity balances gas pressure in the sun.
 - Force of gravity GM^2/R^2
 - Force of gas $PV=nkT$
 - k is Boltzmann's constant. $k = [R, \text{ not radius}] / (\text{number in a mole})$
 - Details (m is mass of gas particle)
 $P = (nm)kT/m/V = M kT/(mR^3)$
 $F = \text{area } P = R^2 M kT/(mR^3) = M kT/(mR)$
 - In balance
 $GMm/R = kT$
- 1. We are watching the birth of the sun. The not-yet sun is a gas cloud slowly shrinking. It is getting
 - A. warmer
 - B. cooler

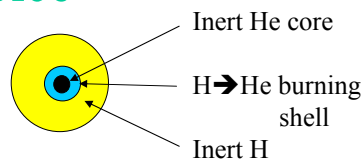
Sun as a main-sequence star

- $H \rightarrow He$ in the core
- $T=15MK$
- Fuel will last another 5 Byr.



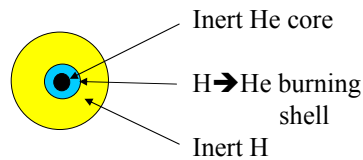
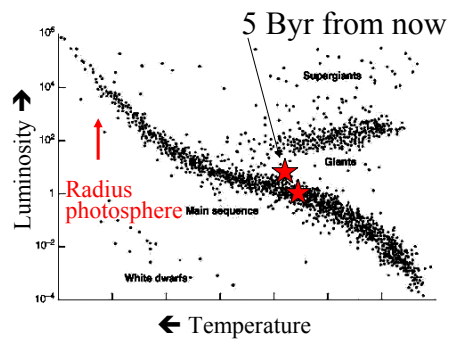
The sun's choice

- Sun does a balancing act.
 - $RT=k/(GMm)$
- Sun must produce energy to replenish the energy radiated away.
- If $H \rightarrow He$ shuts off, source of energy to maintain pressure shuts off, and gravity wins.
- What does the sun do to adjust for gravity's victory?
- Core of the sun shrinks.
 - Core of sun gets hotter
 - $H \rightarrow He$ in the a shell surrounding inert core
 - Balance restored.



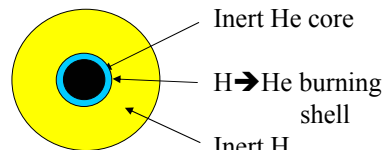
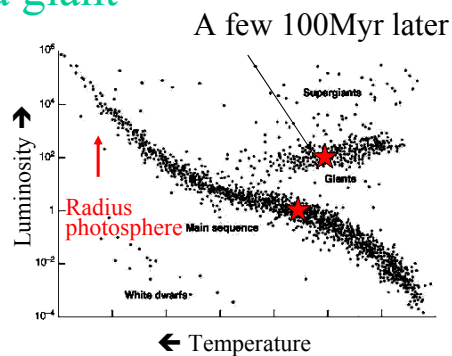
Sun as a subgiant

- H is gone in the core
- The never-ending battle between gravity and pressure. How does the sun adjust?
 - Without burning fuel to keep temperature up, pressure ($PV=nRT$) would fall and gravity would win.
 - Core shrinks, gets hotter
 - $H \rightarrow He$ in the a shell surrounding inert core
 - Balance restored.



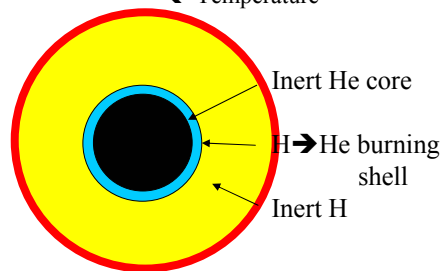
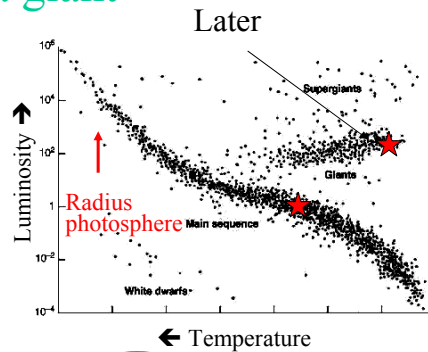
Sun as a giant

- H is gone in the core
- The never-ending battle between gravity and pressure. How does the sun adjust?
 - Without burning fuel to keep temperature up, pressure ($PV=nRT$) would fall and gravity would win.
 - Core shrinks, gets hotter
 - $H \rightarrow He$ in the a shell surrounding inert core
 - Balance restored.
- Inert He core expands



Sun as a giant

- H is gone in the core
- The never-ending battle between gravity and pressure. How does the sun adjust?
 - Without burning fuel to keep temperature up, pressure ($PV=nRT$) would fall and gravity would win.
 - Core shrinks, gets hotter
 - $H \rightarrow He$ in the a shell surrounding inert core
 - Balance restored.
- Inert He core expands



The sun's choice

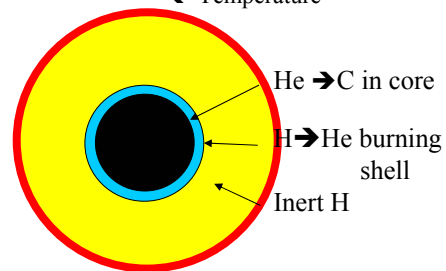
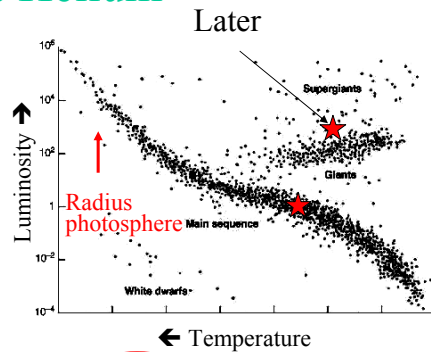
- Sun does a balancing act.
 - $RT=k/(GMm)$
- Sun must produce energy to replenish the energy radiated away.
- Without burning fuel to keep temperature up, pressure ($PV=nRT$) would fall and gravity would win.
 - Core shrinks, gets hotter
 $T=200MK$

| Reaction | Min. temp. |
|--|------------|
| $4\ ^1H \rightarrow\ ^4He$ | 10 MK |
| $3\ ^4He \rightarrow\ ^{12}C$ | 200 MK |
| $^{12}C +\ ^4He \rightarrow\ ^{16}O, Ne, Na, Mg$ | 800 MK |
| $Ne \rightarrow O, Mg$ | 1500MK |
| $O \rightarrow Mg, S$ | 2000MK |
| $Si \rightarrow Fe\ peak$ | 3000MK |

Triple-alpha process

Sun Burns Helium

- H is gone in the core & shell is exhausted
- The never-ending battle between gravity and pressure. How does the sun adjust?
 - Without burning fuel to keep temperature up, pressure ($PV=nRT$) would fall and gravity would win.
 - Core shrinks, gets hotter $T=200\text{MK}$
 - $3\text{He}\rightarrow\text{C}$ in the core (triple alpha process)
 - Balance restored.



The sun's choice

1. Why does fusion of helium require a higher temperature?
 - A. Helium is heavier
 - B. Helium has 2 protons
 - C. Helium has two neutrons
- With more charge, it takes higher speeds to bring two He nuclei close enough to fuse.
 - Carbon has 6 protons.

| Reaction | Min. temp. |
|---|------------|
| $4\ ^1\text{H} \rightarrow\ ^4\text{He}$ | 10 MK |
| $3\ ^4\text{He} \rightarrow\ ^{12}\text{C}$ | 200 MK |
| $^{12}\text{C} +\ ^4\text{He} \rightarrow\ ^{16}\text{O},\ \text{Ne},\ \text{Na},\ \text{Mg}$ | 800 MK |
| $\text{Ne} \rightarrow\ \text{O},\ \text{Mg}$ | 1500MK |
| $\text{O} \rightarrow\ \text{Mg},\ \text{S}$ | 2000MK |
| $\text{Si} \rightarrow\ \text{Fe peak}$ | 3000MK |

Triple-alpha process

Other fusion reactions?



- Sun has one more trick after He is exhausted in core.
 - Burn He in a shell
- Sun is not massive enough to shrink further and get hotter
 - Core is supported by pressure of degenerate electrons.
 - Temperature does not rise to burn anything else.
- End of the road: planetary nebula & white dwarf core

| Reaction | Min. Temp. |
|---|------------|
| $4 \text{ } ^1\text{H} \rightarrow \text{}^4\text{He}$ | 10 MK |
| $3 \text{ } ^4\text{He} \rightarrow \text{}^{12}\text{C}$ | 200 MK |
| $\text{}^{12}\text{C} + \text{}^4\text{He} \rightarrow \text{}^{16}\text{O}, \text{Ne}, \text{Na}, \text{Mg}$ | 800 MK |
| $\text{Ne} \rightarrow \text{O}, \text{Mg}$ | 1500MK |
| $\text{O} \rightarrow \text{Mg}, \text{S}$ | 2000MK |
| $\text{Si} \rightarrow \text{Fe peak}$ | 3000MK |

Triple-alpha process