## White dwarfs & supernovae — Oct 19





Cygnus Loop Supernova 20,000 yr ago

- What causes pressure in white dwarfs?
- Where were the elements in the baby made?
  - Carbon was made and expelled by giants
  - Iron was made in massive stars and expelled by supernovae
  - Heavier elements were made in supernovae & in giants, where there are free neutrons. Nuclei capture neutrons.

- Observing
  - Wed & Thurs, Oct 21 & 22, 8:00-10:00pm.
    - Elevator cannot go up after 10:00pm.
  - Attend only if stars are visible. See angel at 6:00pm, if weather is ambiguous.
  - Quiz. You will be asked to locate a star using the Abrams Planetarium star chart. Quiz counts as one clicker assignment.
  - Go to the south end of the building (toward Wilson Rd.) & take the elevator up to the penthouse.

- Test 2 on Wed.
  - Through Adams' discover of a white dwarf (12 Oct)
  - Energy production in sun not included (12 Oct).
  - One 8 <sup>1</sup>/<sub>2</sub> x 11" cheat sheet.
- Open house nights at the observatory
  - Friday and Saturday, Oct. 23 and 24, 9-11pm, weather permitting.
  - Observatory:
    - Go south on Farm Lane to the end
    - Turn right. Observatory is 100 yards to the west.













## Guest star of 1054

- Records of Sung Dynasty
  - In the first year of the period Chihho, ..., a guest star appeared severa degrees SE of Thien-kuan. After more than a year it gradually becam invisible.-p578.
- Gas expelled in 1054AD, still glowing
- Other SN
  - 1572 Tycho
  - 1604 Kepler







## What is a supernova? Why sun becomes a white dwarf, not a supernova

• In future double-shell burning sun, hot enough to burn

 $3^{4}\text{He}\rightarrow^{12}\text{C}$ 

- When He exhausted, gravity wins, and core contracts.
- Temperature rises.
- Electrons are so tight that they become degenerate.
- New source of pressure to resist gravity.
- Temperature not hot enough to burn carbon.

Reaction	Min. Temp.
4 <sup>1</sup> H <b>→</b> <sup>4</sup> He	10 <sup>7</sup> ° K
3 <sup>4</sup> He → <sup>12</sup> C	2x10 <sup>8</sup>
$^{12}\text{C} + {}^{4}\text{He} \rightarrow {}^{16}\text{O}$ , Ne, Na, Mg	8x10 <sup>8</sup>
Ne 🗲 O, Mg	1.5x10 <sup>9</sup>
O ➔ Mg, S	2x10 <sup>9</sup>
Si <b>→</b> Fe peak	3x10 <sup>9</sup>



## What is a supernova? Why massive star becomes a supernova

- Hot enough to burn  ${}^{4}\text{He} + {}^{12}\text{C} \rightarrow {}^{16}\text{O}$ , etc
- When C exhausted, gravity wins, and core contracts.
- Temperature rises.
- Temperature hot enough to burn neon.
  - $^{20}$ Ne +  $^{4}$ He  $\rightarrow$   $^{24}$ Mg
- Disaster with iron
  - Burning releases energy
  - Fusing iron takes up energy
- Gravity finally wins.

Reaction	Min. Temp.
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Si →Fe peak	3x10 <sup>9</sup>
	hydrogen fusion helium fusion
carbon fusion oxygen fusi	
	neon fusion magnesium fusior silicon fusion
inert iron core	

