

## Measuring Motion, Doppler Effect—23 Oct

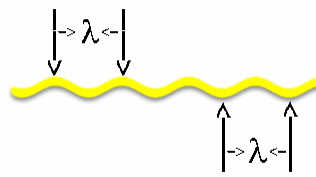
- Where are the elements in the baby created?
- Measuring motion

## Measuring speed without seeing motion

- You are driving 80mph. Just over the crest of a hill, you see a cop car in the distance. In an instant, the cop's computer writes you a ticket.
- Astronomers can measure the speed of a star in orbit around the Milky Way without seeing it move very far. (The orbit takes 200Myr.)
- Q: How can cops & astronomers figure out speed without seeing the object move?
  - A. Measure the wavelength of light from object
  - B. Measure the intensity of light from the object

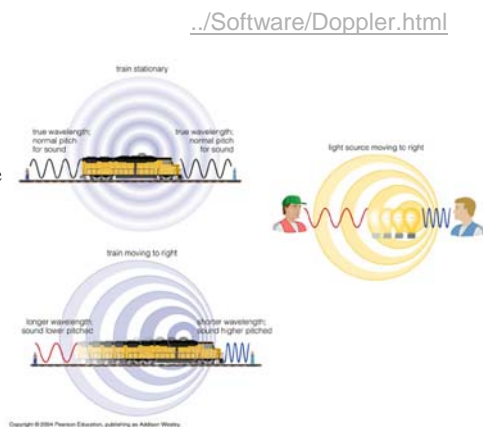
## Wavelength, Frequency

- Wavelength  $\lambda$  = distance between successive crests.
  - m meter
  - nm nanometer ( $10^{-9}\text{m}$ )
  - Å angstrom ( $10^{-10}\text{m}$ )
- Wave moves at speed of light  $c$ .
- Frequency is rate at which crests pass.
  - $f = c/\lambda$
  - Cycles/second; Hertz



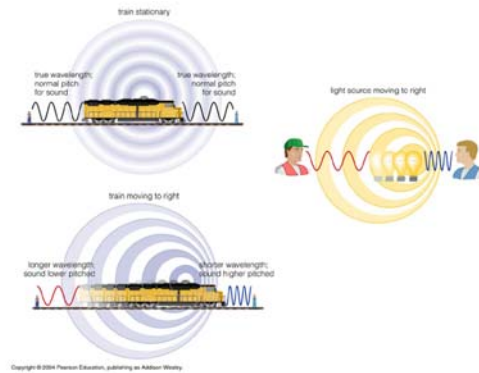
## Measuring Motion: Doppler effect

- How do you measure the velocity of a star?
- Velocity = (change in position)/time
  - Measuring how much star moves is not possible, since we cannot go to the star.
- Velocity is encoded in the light that the stars emits.
- Waves emitted from a star moving towards us are bunched together.
  - Star moves between emitting one wave crest and another. Therefore wavelength is shorter.



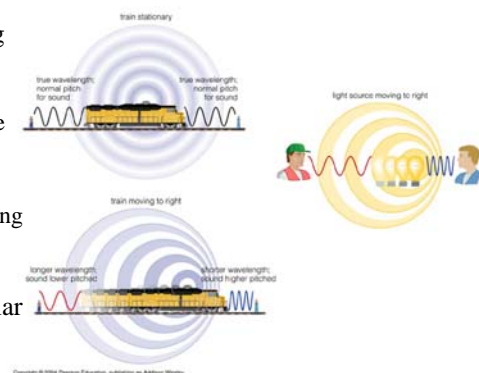
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- $\lambda_{\text{observed}} / \lambda_{\text{rest}} = 1 + v/c$ 
  - $v$  is speed, positive if star is moving toward us.
  - $c$  is speed of light.
- $\Delta\lambda = \lambda_{\text{observed}} - \lambda_{\text{rest}}$  is called the shift in wavelength.



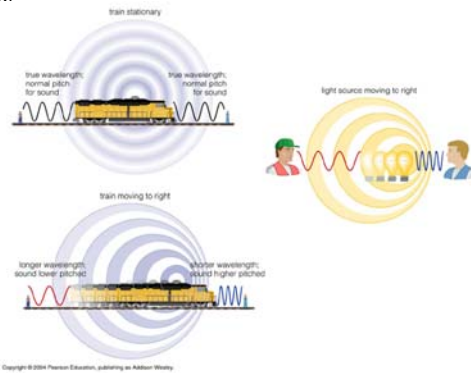
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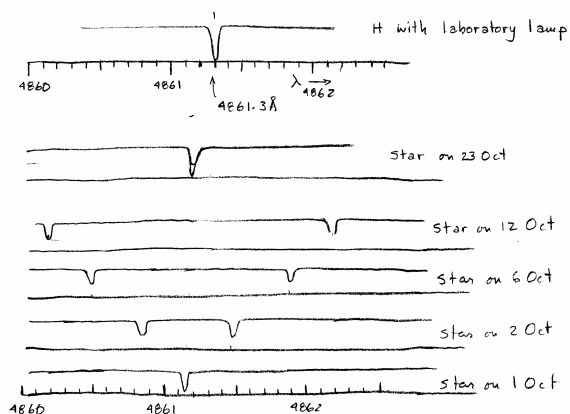
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- Terminology
  - $v/c = (\lambda_{\text{observed}} - \lambda_{\text{rest}}) / \lambda_{\text{rest}}$
  - is called a redshift if positive (star is moving away)
  - is called a blueshift if negative (star is moving toward)



## Pickering's discovery

- We are interpreting E. C. Pickering's spectra of Mizar (a star in the Big Dipper) in 1889.
  - Spectra showing the H $\beta$  line of hydrogen, which is the blue-green line that we saw with the hydrogen tubes.
  - These are *absorption* spectra: The amount of light is high except at wavelengths where hydrogen absorbs.
- Describe the changes in the spectra.

Changes:

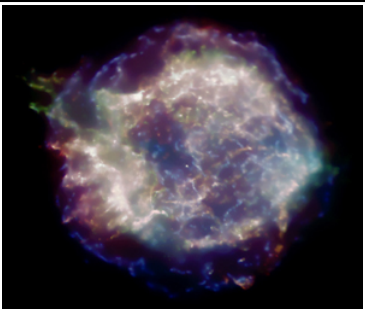




## Neutron capture

- In a supernova, there are free neutrons made by destroying nuclei.
  - Nucleus captures neutrons and turns into a heavier nucleus.  
Inside a nucleus,  
$$\text{nucleus} + \text{n} \rightarrow \text{heavier nucleus}$$
  - Nucleus may decay into a more stable one.  
$$\text{n} \rightarrow \text{p} + \text{e}^- + \nu$$
  - Nucleus may capture more neutrons.
  - Eventually unstable nuclei decay into stable ones. Some heavy as uranium.
1. If  $^{197}\text{Au}$  captures a neutron, it becomes \_\_\_\_\_. (Au has 79p. Hg has 80p. Pt has 78p.)
    - A.  $^{197}\text{Hg}$
    - B.  $^{198}\text{Au}$
    - C.  $^{198}\text{Hg}$
    - D.  $^{198}\text{Pt}$
  2. If a neutron in  $^{198}\text{Au}$  decays, it becomes \_\_\_\_\_. (Au has 79p. Hg has 80p. Pt has 78p.)
    - A.  $^{198}\text{Hg}$
    - B.  $^{198}\text{Au}$
    - C.  $^{198}\text{Pt}$
- The net effect is to turn gold into mercury.

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- Calculation of nuclear reactions in a supernova.
  - Start with iron and add neutrons
  - Look at gold
    - 79 protons,  $197 - 79 = 118$  neutrons

## Questions on the Supernova Movie

1. What is the only element at the start? How many neutrons does it have? • “R process movie” at [www.jinaweb.org/html/gallery3.html](http://www.jinaweb.org/html/gallery3.html)
2. At what time did some gold form? Gold has 79 protons. Is this gold stable?
3. At the end of the calculation, how many protons does the nucleus with the most protons have?
4. What is the time at the end of the calculation?
5. Are the end products stable?

## Where were the elements in the baby made?

- Lighter elements (He, O, C, Ne, Mg, etc) are made by fusion with a release of energy
  - $4\text{H} \rightarrow \text{He} + \text{energy}$
  - $3\text{He} \rightarrow \text{C} + \text{energy}$
- Elements heavier than iron are made in supernovae and in giant

Periodic Table  
1990 Dr. Michael Ehaber

1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18
1 H 1.008	2 He 4.0026	3 Li 6.941	4 Be 9.0122	5 B 10.811	6 C 12.011	7 N 14.007	8 O 15.999	9 F 18.998	10 Ne 20.180	11 Na 22.990	12 Mg 24.305	13 Al 26.982	14 Si 28.086	15 P 30.974	16 S 32.06	17 Cl 35.453	18 Ar 39.948
19 K 39.098	20 Ca 40.078	21 Sc 44.956	22 Ti 47.88	23 V 50.942	24 Cr 51.996	25 Mn 54.938	26 Fe 55.845	27 Co 58.933	28 Ni 58.693	29 Cu 63.546	30 Zn 65.38	31 Ga 69.723	32 Ge 72.64	33 As 74.922	34 Se 78.96	35 Br 79.904	36 Kr 83.80
37 Rb 85.468	38 Sr 87.62	39 Y 88.906	40 Zr 91.224	41 Nb 92.906	42 Mo 95.94	43 Tc 98.906	44 Ru 101.07	45 Rh 102.91	46 Pd 106.42	47 Ag 107.868	48 Cd 112.411	49 In 114.818	50 Sn 118.710	51 Sb 121.757	52 Te 127.6	53 I 126.905	54 Xe 131.29
55 Cs 132.905	56 Ba 137.327	57 La 138.905	58 Ce 140.12	59 Pr 140.908	60 Nd 144.24	61 Pm 144.913	62 Sm 150.36	63 Eu 151.964	64 Gd 157.25	65 Tb 158.925	66 Dy 162.50	67 Ho 164.930	68 Er 167.259	69 Tm 168.930	70 Yb 173.054	71 Lu 174.967	
87 Fr 223.018	88 Ra 226.025	89 Ac 227.028	90 Th 232.038	91 Pa 231.036	92 U 238.029	93 Np 237.048	94 Pu 244.041	95 Am 243.061	96 Cm 247.070	97 Bk 247.070	98 Cf 251.080	99 Es 252.083	100 Fm 257.103	101 Md 258.103	102 No 259.108	103 Lr 260.105	

Lanthanides: La, Ce, Pr, Nd, Pm, Sm, Eu, Gd, Tb, Dy, Ho, Er, Tm, Yb, Lu

Actinides: Ac, Th, Pa, U, Np, Pu, Am, Cm, Bk, Cf, Es, Fm, Md, No, Lr