## Universe at 400,000yr: Weighing the universe—4 Dec Test 3 has been graded. Will be returned on Mon. Final exam Covers entire course with emphasis on 20<sup>th</sup> century cosmology (Oct 28 to end of term, Hwk 7–10) One 8.5×11" cheat sheet Mon, 14<sup>th</sup>, 3:00-5:00, 1410 BPS (large classroom next door) Please fill out on-line SOCT (Student Opinion of Courses and Teaching) http://rateyourclass.msu.edu Will close when grades are submitted. Review

- Map of radiation from the Big Bang is mottled with an angular scale of 1°.
- Physical size of hotter, denser regions is c×(age of U).
- If universe has lots of matter, expansion is fast at beginning and slows down.
- Today
  - angle = physical size / distance ⇒ weighing the universe Ast 207 F2009









## Weighing Universe using fluctuations in radiation from Big Bang Principle for astronomical weighing: Define a motion Time the motion If the motion takes longer, the mass is less. To find mass density of the universe, measure the time it takes for the U to expand by a factor of 1000 by looking at angular size of fluctuations in radiation from Big Bang. 3. If the angle is small, the U has less mass. Result: Angle is large $\Rightarrow$ Age of U is less than 1/H The inference " $\Rightarrow$ Mass density is high" is incorrect. The correct statement is "⇒ Sum of mass density and dark energy is high" Gravity is more complicated than what Newton thought and we have discussed. More complete description of gravity next week. Ast 207 F2009