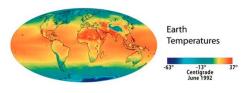
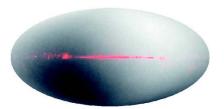
## Universe at 400,000yr—1 Dec

- Cosmic background radiation is a snapshot of U at 400,000 yr.
- What is in the snapshot?
  - History of MW & local group: motion of 300km/s
  - Fluctuations at an angular scale of 1°.
  - Growing clumps of mass  $\Rightarrow$  mass density of the universe
  - Evidence of first stars





Temperature of radiation from Big Bang, measured by WMAP satellite

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### Universe at 13.7Byr & 11.7Byr

- M101 is a snapshot of the universe.
  - v = 1050 km/s
  - D=v/H=1050/60=18Mpc=18\*3.2=56MillionLight year
  - Light took 18Myr to get to us.
- Light from Abell2218 took 2Byr to get to us. Are all of the objects in the picture of Abell 2218 from the time when the universe was 11.7Byr old?

A. Y

B. N



M104



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Abell2218

#### Universe at 4min

- At 3min
  - Neutrons got incorporated into <sup>4</sup>He.
  - Temperature was 1BK.
- 1. If you could go back in time to when the U was 4min old and take a picture, what would you see? Stars. Very hot black-body radiation.
  - A. S and V
  - B. Sonly
  - C. V only
  - D. Neither S nor V.

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#### Universe at 4min

- At 4min, U was slightly cooler than 1BK.
- There were no stars, no galaxies, no planets.
  These could not form when the temperature was so hot.

# Ionized/un-ionized gas

• Ionization is the loss of an electron.

H atom  $\rightarrow p + e^{-}$ 

 Recombination is when electron and nucleus combine.

 $p + e^{-} \rightarrow H$  atom

- lonization occurs if the temperature is hot enough.
- 1. Name one thing in this room that is/has ionized gas.
  - A. Gas in fluorescent light
  - B. Air
  - C. Air in my lungs

- Light scatters poorly off of electrons bound in an atom or molecule.
- Light scatters readily off of free electrons.

photon electron

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- If the air were ionized, light would travel 500m before it is scattered. Light scatters nearly equally in all directions.
- 1. If the air were ionized, could you, sitting in Michigan Stadium, watch the Spartans beating Michigan? Could you see stars at night?
  - A. YY
  - B. YN
  - C. NY
  - D. NN
- If the air were ionized, the air would be a fog.
- We do not have a picture of the U at 4min, because at such a high temperature, the U is ionized. Light could travel only a short distance before scattering of an electron.

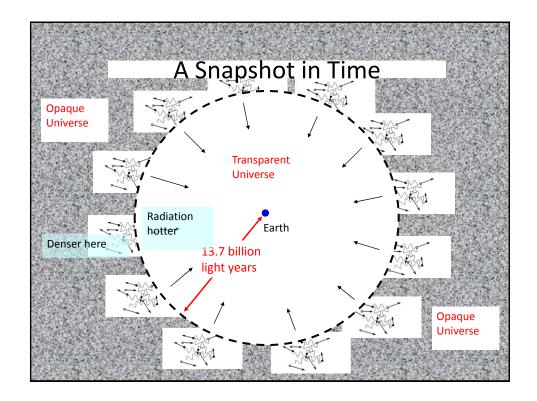
## Matter and Radiation Decoupled

- Early, hot universe
  - Hydrogen is ionized.
    - Protons and electrons are free
  - Universe is opaque.
    - Photons travel only short distances.
    - Scattered by free electrons.
- Decoupling: p + e<sup>-</sup> → H atom
  - T = 3000 K; universe 400,000 yr old.
  - Universe becomes transparent
    - Photons decouple from matter, continue in whatever direction they were moving.
- After decoupling
  - Photons travel in same direction. Preserves information about the matter at decoupling.
  - In certain directions, matter & light were slightly more dense.
    Light preserves that information.

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# Before & after decoupling

- Adam was sent back in time to when the U was 300,000yr old. Noah was sent back to when the U was 500,000yr old. A & N sent us a message that they are toasty warm.
- 1. Can we get their messages?
  - A. Y for both
  - B. Y for A, N for N
  - C. N for A, Y for N
  - D. N for both.



#### Wilkinson Microwave Anisotropy Probe (WMAP) Satellite

- Goal: Take a picture of the U when it was 400,000yr old.
  - Measure spatial variations in temperature of the CBR.
- Measure anisotropy
  - Penzias & Wilson found radiation to be isotropic.
  - Goal is to measure small deviations.
- Sensitivity is 0.000035K (a part in 100,000).
- Anything in the instrument even 0.0001K warmer is fatal.
- Symmetric design
  - Record temperature difference between left & right channels.
     Temperature difference is small.
  - Rotate entire instrument.
    - Instrumental problems do not change; radiation from the sky does change.



