Thin Lenses \rightarrow Thick Lenses



Lens maker's formula

$$\frac{1}{f} = (n-1) \left[\frac{1}{R_1} - \frac{1}{R_2} + \frac{(n-1)d}{nR_1R_2} \right].$$

"Thin" lens \rightarrow d is negligible

$$\frac{1}{f} \approx (n-1) \left[\frac{1}{R_1} - \frac{1}{R_2} \right].$$

Paraxial approximation $\sin(\theta) \approx \tan(\theta) \approx \theta$ $\cos(\theta) \approx 1$

See Hecht Ch. 5 and review the following Equations. Refer to lecture given on 10/01 for derivation of the following equations

$$\frac{1}{f} = \frac{1}{s_0} + \frac{1}{s_i}$$
$$x_0 x_i = f^2$$
$$M_T \equiv \frac{y_i}{y_0} = -\frac{s_i}{s_0}$$
$$M_L \equiv \frac{dx_i}{dx_0} = -\frac{f^2}{x_0^2}$$

"Sign" convention is of paramount importance! (See Hecht Table 5.1, Fig. 5.12, Table 5.2)

Recall : Real and Virtual Images



See also Hecht Table 5.3

Numerical Aperture



 θ : half-angle subtended by the imaging system from an *axial* object

Numerical Aperture (NA) = $n \sin \theta$

Speed (f/#)=1/2(NA)pronounced f-number, e.g. f/8 means (f/#)=8.

Aperture stop

the physical element which limits the angle of acceptance of the imaging system

We will learn that

the spatial resolution limit due to diffraction $\approx 1.22 \times f \lambda / D = 0.61 \times \lambda / NA$ [Rayleigh Criterion].

When Paraxial Approximation Fails: Ray Tracing + Diffraction



ray scatter diagram (\Leftrightarrow defocus)

Databases of common lenses and elements
Simulate aberrations and ray scatter diagrams for various points along the field of the system (PSF, point spread function)

•Standard optical designs (e.g. achromatic doublet)

•Permit optimization of design parameters (e.g. curvature of a particular surface or distance between two surfaces) vs designated functional requirements (e.g. field curvature and astigmatism coefficients)

•Also account for diffraction by calculating the at different points along the field modulation transfer function (MTF) [Fourier Optics]

Chromatic

- is due to the fact that the refractive index of lenses, etc. varies with wavelength; therefore, focal lengths, imaging conditions, etc. are wavelength-dependent
- Geometrical (monochromatic)
 - are due to the deviation of non-paraxial rays from the approximations we have used so far to derive focal lengths, imaging conditions, etc.; therefore, rays going through imaging systems typically do not focus perfectly but instead scatter around the "paraxial" (or "Gaussian") focus

Refractive index n is dispersive!

 $n(\omega)$

Deteriorate the image: •Spherical aberration •Coma •Astigmatism

Deform the image: •Field curvature •Distortion

Departures from the idealized conditions of Gaussian Optics (e.g. paraxial regimes).

Hecht 6.3.2

$$\frac{1}{f} \approx (n-1) \left[\frac{1}{R_1} - \frac{1}{R_2} \right].$$



Chromatic Aberration



Solutions:

- 1. Combine lenses (achromatic doublets)
- 2. Use mirrors



Figure 1.21 Longitudinal chromatic aberration





Melles Griot "Fundamental Optics"

Spherical Aberration



Figure 1.15 Spherical aberration of a plano-convex lens







Solution I: Aspheric Mirrors or Lenses

Hubble Telescope

It was probably the most precisely figured mirror ever made, with variations from the prescribed curve of only 10 nanometers, it was too flat at the edges by about 2.2 microns. Source: wikipedia





Lens Shape

 $q = \frac{\left(R_1 + R_2\right)}{\left(R_2 - R_1\right)}$

Solution II: Chose a proper shape of a singlet lens for a given image-object distance. $\frac{1}{f} \approx (n-1) \left[\frac{1}{R_1} - \frac{1}{R_2} \right].$

For a given desired focal length, there is freedom to choose one of the radii for a singlet. The spherical aberration and coma depend on the particular choice, so these aberrations can be minimized by the designed form.



Figure 1.23 Aberrations of positive singlets at infinite conjugate ratio as a function of shape

Lens Selection Guide



http://www.newport.com/Lens-Selection-Guide/140908/1033/catalog.aspx#

Astigmatism



Figure 1.16 Astigmatism represented by sectional views

Coma and Deformation



Figure 1.18 Positive transverse coma



Figure 1.19 Field curvature









Figure 1.20 Pincushion and barrel distortion