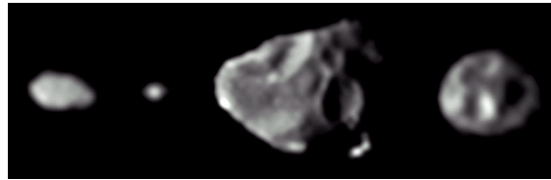


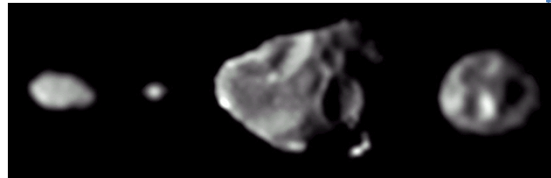
Rings vs. moons—28 Jan



Metis Adrastea Amalthea Thebe

- Why are there no big moons near planets? Why are rings close to planets?
- Differences between Jupiter's inner and outer moons
- Tides
- Roche's limit: If a moon is near its planet, the tidal force of the planet is greater than the gravitational force of the planet: the moon cannot hold together.

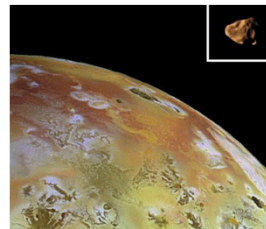
The Innermost Moons of Jupiter



Metis Adrastea Amalthea Thebe

- Q5: What holds me together?
 - a. Gravity
 - b. Atomic bonds between the atoms
- Q6: What holds Io & Metis together?
(Think about the shapes of Io & Metis.)
 - a. Gravity for both
 - b. Bonds for both
 - c. Gravity for Io; bonds for Metis
 - d. Gravity for Metis; bonds for Io.

Amalthea
& Io



Tidal force

- The force of the sun on a rock or a drop of water on Earth is

$$F = -GM_s m / R^2$$

- Consider rocks along the line between the earth and sun. $R = R_{es} + r$.

$$F = -GM_s m (R_{es} + r)^{-2}$$

- Important math trick: If $x \ll 1$, $(1 + x)^n \approx 1 + nx$
- Taylor series: $f(x) = f(x_0) + (x - x_0) f'(x_0)$

$$F = -GM_s m R_{es}^{-2} + 2GM_s m R_{es}^{-3} r$$

- First term is force at center of earth.
- Second term is tidal force.

Tidal force

- Consider a rock along the line between the earth and sun.
 $R = R_e + r$.

$$F = -GM_s m R_{es}^{-2} \left(1 - \frac{2r}{R_{es}}\right)$$

- First term is force at center of earth.
- Second term is tidal force.
- 1. To think about the tides, consider the force on a drop of water. What does the first term do to the drop of water? Answer: The first term accelerates the drop to keep it orbiting the sun.
- 2. Why are there two high tides every day? Answer: At noon ($r < 0$) the tidal force pulls the drop away from Earth. At midnight, the tidal force ($r > 0$) also pulls the drop away from Earth.
- Write the answers on paper. Turn your papers in.

1. If a planet is twice as far from the sun, the ratio of the tidal force to the force on the planet is ____ .
 - A. $\frac{1}{2}$
 - B. $\frac{1}{4}$
 - C. 2
 - D. 4
2. If a planet has twice the radius, the ratio of the tidal force to the force on the planet is ____ .

How big is the tidal force?

- Consider rocks along the line between the earth and sun. $R = R_e + r$.

$$F_{\text{Tidal}}/m = 2(GM_s R_{es}^{-2})\left(\frac{r}{R_{es}}\right)$$

- The term $GM_s R_{es}^{-2}$ is the acceleration of the earth in its orbit. It must be $\frac{v^2}{R_{es}} = 4\pi^2 R_{es} P^{-2}$.

$$\frac{F_{\text{Tidal}}}{m} = 8\pi^2 P^{-2} r.$$

- At the equator, $\frac{F_{\text{Tidal}}}{m} = 8 \cdot \pi^2 \frac{6300\text{km}}{(\pi \times 10^7\text{s})^2} = 5 \times 10^{1+6-14} = 5 \times 10^{-7} \text{m/s}^2$

Roche limit. Approximation #1

- A moon orbits Jupiter. The moon is a fluid having no strength. Consider a mass m on a line joining Jupiter and the moon, where the tidal force is greatest.
- The tidal force on mass m is

$$F_{\text{Tidal}}/m = 2(GM_J R_{jm}^{-2}) \left(\frac{r_m}{R_{jm}} \right)$$

- The gravitational force is

$$\frac{F_{\text{moon}}}{m} = GM_{\text{moon}} r_m^{-2}$$

- If the moon holds itself together, then

$$\begin{aligned} F_{\text{Tidal}} &\leq F_{\text{moon}} \\ 2M_J R_{jm}^{-3} &\leq M_{\text{moon}} r_m^{-3} \end{aligned}$$

Roche limit. Approximation #1

- A moon orbits Jupiter. The moon is a fluid having no strength. Consider a mass m on a line joining Jupiter and the moon, where the tidal force is greatest.
- Express in term of mass density ρ .

$$2\rho_j r_j^3 R_{jm}^{-3} \leq \rho_{\text{moon}}$$

- For the moon to hold itself together, it must be outside the Roche limit:

$$R_{jm} \geq 2^{1/3} (\rho_j / \rho_{\text{moon}})^{1/3} r_j$$

- For the case where Jupiter and its moon have the same density,

$$R_{jm} \geq 1.26 r_j$$

- The moon must be outside 1.26 of the radius of the planet.