

## Rings & Moons—31 Jan

- Homework 3 is due Mon, 7 Feb.
- Roche's limit: If a moon is near its planet, inside the Roche limit, the tidal force of the planet is greater than the gravitational force of the planet: the moon cannot hold together.
- Roche limit for nonrotating moon:
 
$$R_{jm} \geq 1.26 (\rho_j / \rho_{moon})^{1/3} r_j$$
- Better calculation of Roche limit. Moon rotates.
- Even better calculation: Moon is oblate.
- Compare rings and moons of the Jovian planets.

## Roche limit. Approximation #1

- A moon orbits Jupiter. The moon is a fluid having no strength. Consider a mass  $m$  on a line joining Jupiter and the moon, where the tidal force is greatest.
- The tidal force on mass  $m$  is

$$F_{\text{Tidal}}/m = 2(GM_J R_{jm}^{-2}) \left( \frac{r_m}{R_{jm}} \right)$$

- The gravitational force is

$$\frac{F_{\text{moon}}}{m} = GM_{\text{moon}} r_m^{-2}$$

- If the moon holds itself together, then

$$\begin{aligned} F_{\text{Tidal}} &\leq F_{\text{moon}} \\ 2M_J R_{jm}^{-3} &\leq M_{\text{moon}} r_m^{-3} \end{aligned}$$

## Roche limit. Approximation #2

- A moon is likely to rotate with the same period as its orbit, which is the case with Earth's moon.
1. If the moon rotates, the Roche limit be \_\_\_  $1.26r_j$ .
    - A. Greater than
    - B. Less than
    - C. Same as
  2. Write Newton's 2<sup>nd</sup> Law ( $F=ma$ ) for a rotating moon that is barely held together by its own gravity. You may use symbols, rather than finding expressions for the symbols.

## Roche limit. Approximation #2

- Better assumption: The moon rotates with the same period as its orbit. Then

$$F_{moon} - F_{tidal} = m \frac{v^2}{r_{moon}} = m 4\pi^2 r_{moon} / P^2$$

- Substitute Kepler's 3<sup>rd</sup> Law

$$P^2 = 4\pi^2 R_{jm}^3 / (GM_j)$$

- RHS becomes

$$GM_j r_{moon} / R_{jm}^3$$

– Half of the tidal force.

- Roche's limit is then

$$R_{jm} \geq 3^{1/3} (\rho_j / \rho_{moon})^{1/3} r_j$$

$$3^{1/3} = 1.44$$

## Roche's limit. Approximation #3

- Near breakup, the moon is oblate.
- 1. If the moon is oblate, the force of the moon on mass  $m$  is  $\frac{4\pi}{3} G m \rho_{moon} r_m$  and Roche's limit is \_\_\_\_\_.
  - A. Less than. Greater
  - B. Greater than. Greater
  - C. Less than. Less.
  - D. Greater than. Less
- The force of the moon is less than that of a spherical moon. Result:



$$R_{jm} \geq 2.44 (\rho_j / \rho_{moon})^{1/3} r_j$$

## Inner moons of Jupiter

- For moons having the same density as Jupiter, the Roche limit for Jupiter's moons is  $2.44r_j = 169\text{Mm}$
- R/Roche accounts for the density of Jupiter and the moon.
- 1. Metis is held together by its own gravity. Amalthea is held together by its own gravity.
  - A. T for Metis. F for A
  - B. F for Metis. T for A
  - C. T for both
  - D. F for both

Moon	R[Mm]	R/Roche
Metis	128	0.94
Adrastea	129	0.95
Amalthea	181	0.88
Thebe	222	1.3
Io	422	



