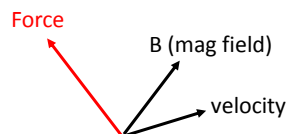


## Comet tails—18 Feb

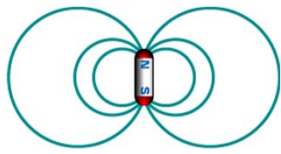
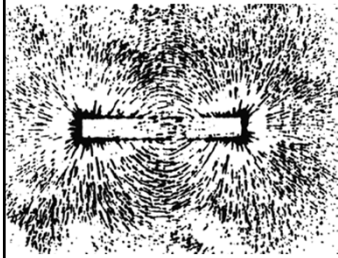
- Mike Velbel, *The geology of comets: An early look at results from the Stardust-NExT mission's encounter with comet 9P/Tempel 1*  
– Today, 4:00 PM, 204 Nat Sci Building.
- Finish comet tails
- Pluto & relation to objects beyond Neptune
- Next week: Read about derivation of Kepler's Laws (Textbook §4.4 & Carroll & Ostlie §2.1-2.3) Needed for understanding Kuiper Belt.

## Motion in a magnetic field

- Force of a particle with charge  $e$  moving with velocity  $\vec{v}$  in a magnetic field  $\vec{B}$
  - $\vec{F} = e\vec{v} \times \vec{B}$
1. Assume  $\vec{v}$  and  $\vec{B}$  are parallel. What is the path of the particle?
  2. In a magnetic field, does the speed of the particle change?
    - A. Yes
    - B. No
  3. Assume  $\vec{v}$  and  $\vec{B}$  are perpendicular. What is the path of the particle?



## Magnetic Field Lines



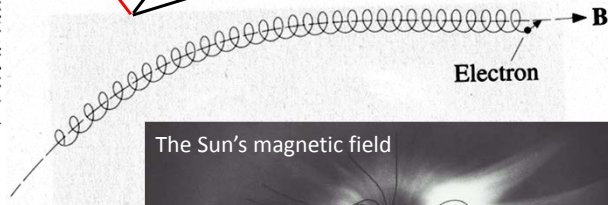
Magnetic field  
lines of force

Force

B (mag field)

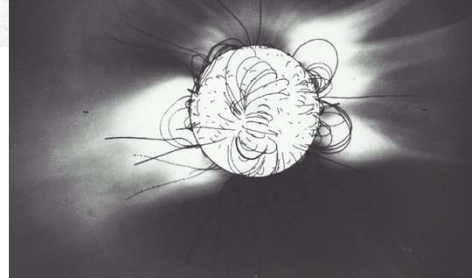
velocity

Motion of a charged  
particle around a  
magnetic field line

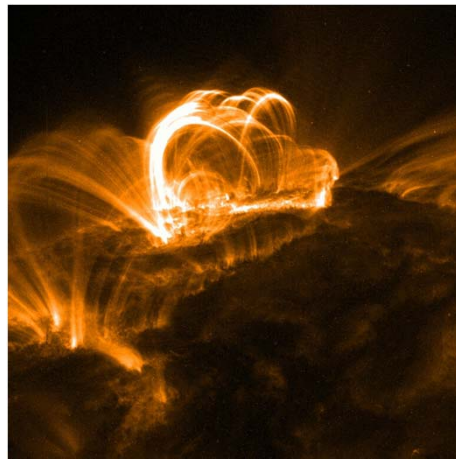


Electron

The Sun's magnetic field



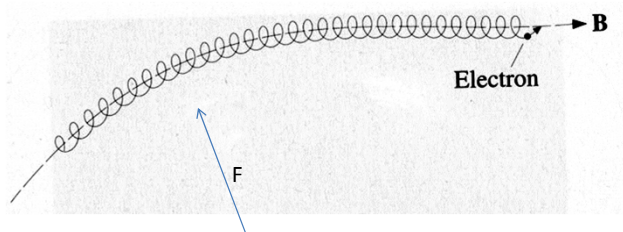
- Solar flares produce ionized gas. Why is the ionized gas in arcs?



NASA/LMSAL TRACE Satellite

## Magnetic field is “frozen in”

- If I push on the electrons, the magnetic field lines move.
  - The electrons can move parallel to magnetic field.
  - They cannot move perpendicular to it. They circle a magnetic field line with the Larmor radius. Typical Larmor radius in the solar wind is 300km.  $r = mv/(qB)$ .
- Magnetic field is “frozen in” ionized matter.

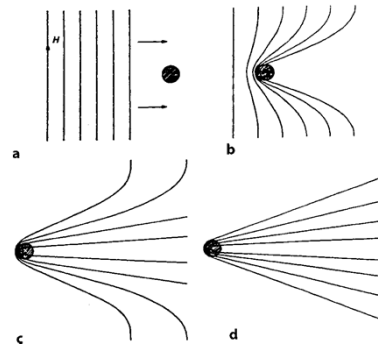


## Solar wind & comets' ion tail

- Biermann (early 1950s): The sun's ultraviolet radiation ionized gas from the comet. Suppose there is a stream of charged particles from the sun. The collisions with the stream of particles accelerate the ions away from the sun.
- Chapman (1950s): temperature of sun's corona is very high  $\Rightarrow$  conductivity is high.
- Parker (1950s): Since pressure is nearly constant & gravity falls as  $1/R^2$ , particles can escape the sun's gravity as a wind.

## Solar wind & comet's ion tail

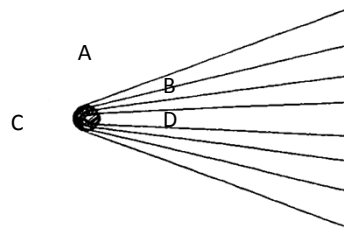
- Magnetic fields amplifies effect of solar wind.
- The sun's ultraviolet radiation ionized gas from the comet. Suppose there is a solar wind, a stream of charged particles from the sun.
  - The solar wind carries the magnetic field with it. (a)
  - The magnetic field wraps around the ionized gas from the comet. (b)
  - The ionized gas can move only into the region without magnetic fields.



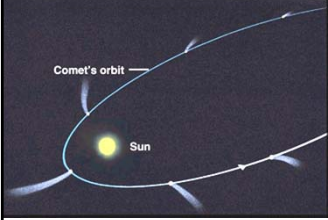
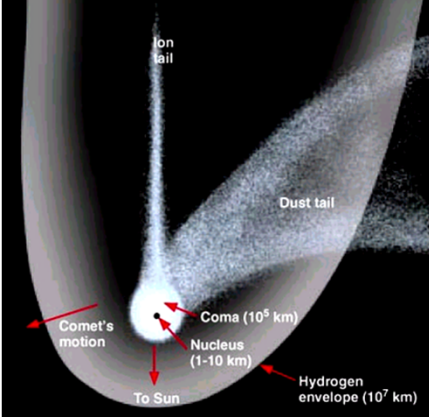
Alfven, H. (1957)  
On the theory of comet tails,  
Tellus, 9, 92

## Comet

1. Where are ions that the comet emits?
  - E. More than one location is correct.
2. Where are hydrogen atoms? Assume they come off the comet in all directions. (Hydrogen atoms do not absorb the sun's radiation.)



## Tails

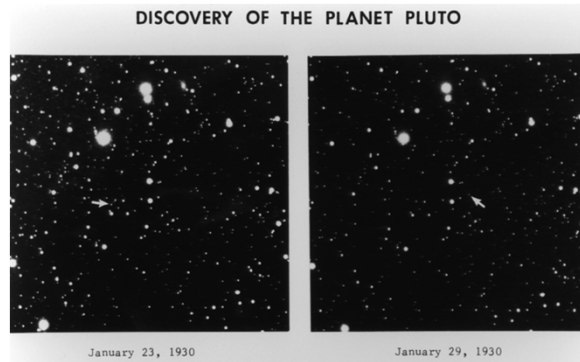
- **dust tail**
  - up to 10 million km long
  - smoke-sized dust particles
  - driven off nucleus by escaping gases
  - pushed outwards by Sun's radiation
  - competing force of Sun's gravity → curve in tail.
- **ion tail**
  - Up to 100's of millions km long
  - Ions, pushed out by charged particles from Sun (solar wind).

## Pluto

- **Outline**
  - The ninth planet
  - Pluto and Neptune are in a 2:3 resonance
  - Pluto's orbit is tilted  $17^\circ$  relative to the ecliptic.
  - Pluto has same orbits as the Kuiper Belt objects.
  - Other objects in the Kuiper Belt are comets and large bodies similar to Pluto.

## Discovery of Pluto

- Tombaugh discovered Pluto in a dedicated search in 1930s.

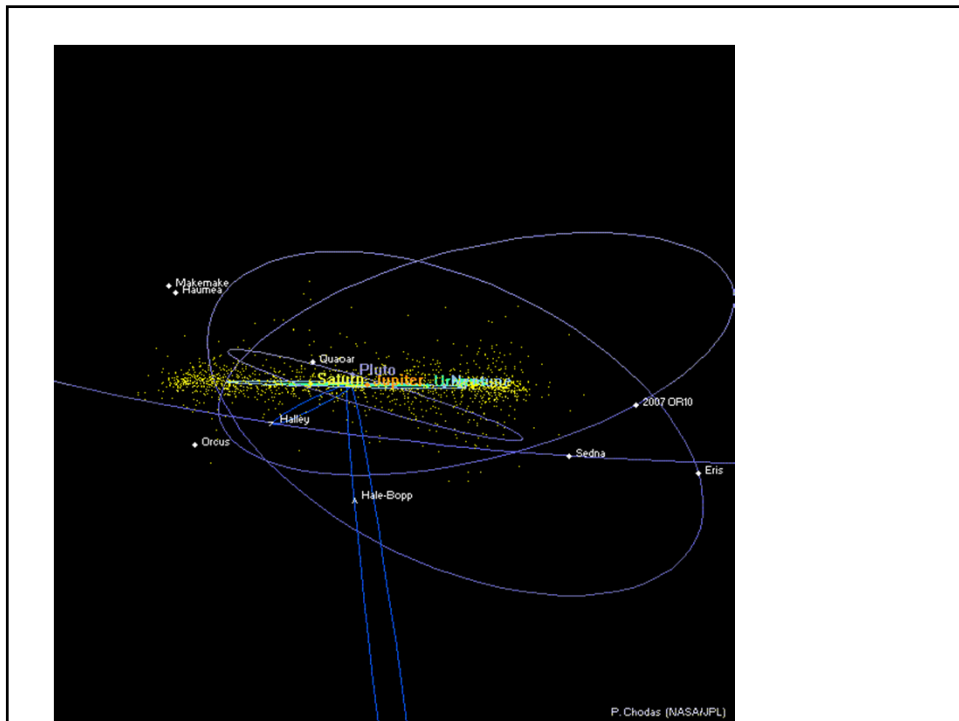
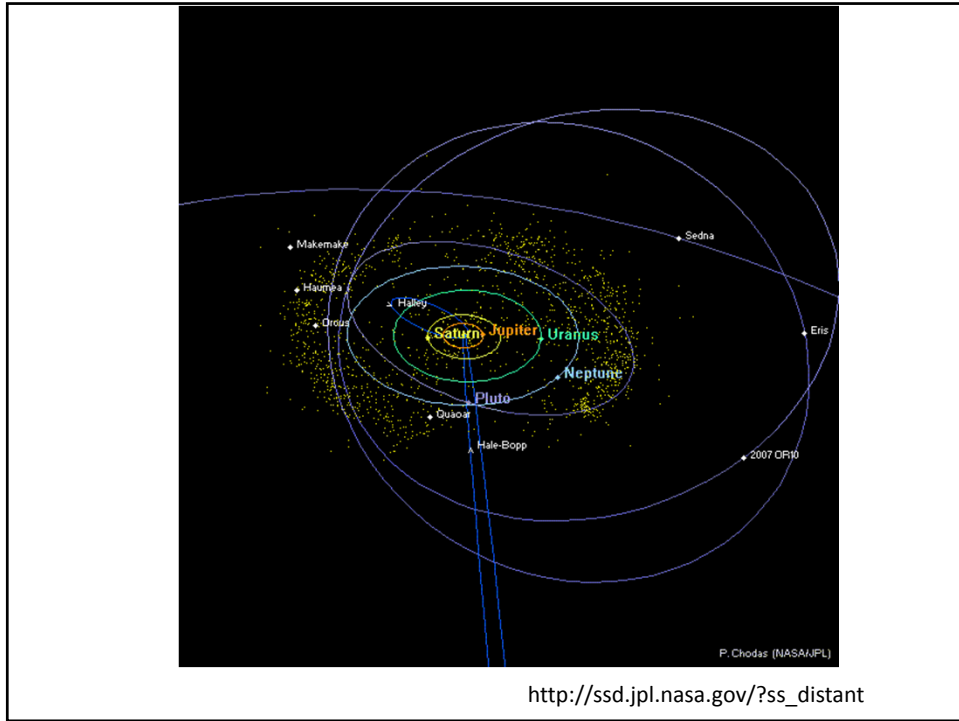


Lowell Observatory Archives

## Orbit of Pluto

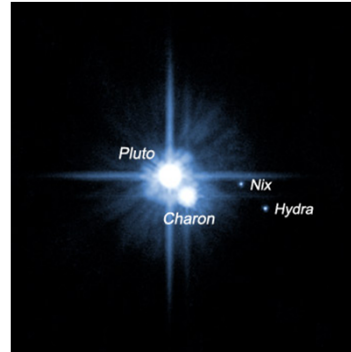
- Semi major axis
- High eccentricity
- Tilted
- $P_{\text{Pluto}}/P_{\text{Neptune}} = 248/165 = 3/2$

Planet	ecc	A [AU]	P[yr]
Earth	0.0167	1	1
Mars	0.0934	1.52	1.88
Jupiter	0.0483	5.20	11.9
Saturn	0.0560	9.54	29.4
Uranus	0.0461	19.2	84.0
Neptune	0.0097	30.1	165
Pluto	0.249	39.5	248



## Pluto has 3 moons

- Charon discovered in 1978.
  - Used to determine the mass of Pluto
- Pluto
  - Diameter 1200km ( $0.18R_{\text{Earth}}$ )
  - Density  $2.0\text{gm/cm}^3$
  - Mass  $1.30 \times 10^{22}\text{kg} = 0.0024M_{\text{Earth}}$
- Pluto & Charon both in synchronous rotation
  - always show same faces to each other
- Pluto's inclination =  $118^\circ$  (i.e. tipped on its side)
  - but Charon orbits in Pluto's equatorial plane.



[NASA](#), [ESA](#), H. Weaver (JHU/APL),  
A. Stern (SwRI),  
HST Pluto Companion Search Team