

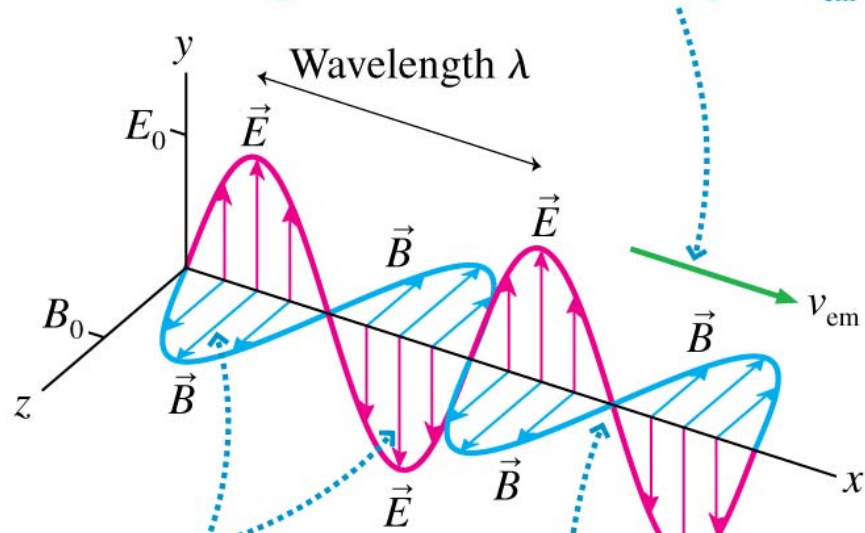
# Electromagnetic Waves

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All electromagnetic waves travel in a vacuum with the same speed, a speed that we now call the *speed of light*.

$$v_{\text{em}} = \frac{1}{\sqrt{\epsilon_0 \mu_0}} = 3.00 \times 10^8 \text{ m/s} = c$$

1. A sinusoidal wave with frequency  $f$  and wavelength  $\lambda$  travels with wave speed  $v_{em}$ .



2.  $\vec{E}$  and  $\vec{B}$  are perpendicular to each other and to the direction of travel. The fields have amplitudes  $E_0$  and  $B_0$ .

3.  $\vec{E}$  and  $\vec{B}$  are in phase. That is, they have matching crests, troughs, and zeros.

# Properties of Electromagnetic Waves

Any electromagnetic wave must satisfy four basic conditions:

1. The fields  $\mathbf{E}$  and  $\mathbf{B}$  are perpendicular to the direction of propagation  $\mathbf{v}_{\text{em}}$ . Thus an electromagnetic wave is a transverse wave.
2.  $\mathbf{E}$  and  $\mathbf{B}$  are perpendicular to each other in a manner such that  $\mathbf{E} \times \mathbf{B}$  is in the direction of  $\mathbf{v}_{\text{em}}$ .
3. The wave travels in vacuum at speed  $v_{\text{em}} = c$
4.  $\mathbf{E} = c\mathbf{B}$  at any point on the wave.

# Properties of Electromagnetic Waves

The energy flow of an electromagnetic wave is described by the **Poynting vector** defined as

$$\vec{S} \equiv \frac{1}{\mu_0} \vec{E} \times \vec{B}$$

The magnitude of the Poynting vector is

$$S = \frac{EB}{\mu_0} = \frac{E^2}{c\mu_0}$$

The intensity of an electromagnetic wave whose electric field amplitude is  $E_0$  is

$$I = \frac{P}{A} = S_{\text{avg}} = \frac{1}{2c\mu_0} E_0^2 = \frac{c\epsilon_0}{2} E_0^2$$

# Radiation Pressure

It's interesting to consider the force of an electromagnetic wave exerted on an object per unit area, which is called the **radiation pressure**  $p_{\text{rad}}$ . The radiation pressure on an object that absorbs all the light is

$$p_{\text{rad}} = \frac{F}{A} = \frac{P/A}{c} = \frac{I}{c}$$

$$\Delta p = \frac{\text{energy absorbed}}{c} \quad (E = pc)$$

$$F = \frac{\Delta p}{\Delta t} = \frac{(\text{energy absorbed}) / \Delta t}{c} = \frac{P}{c}$$

where P is the power (joules per second) of the light.

where  $I$  is the intensity of the light wave. The subscript on  $p_{\text{rad}}$  is important in this context to distinguish the radiation pressure from the momentum  $p$ .

## Example Solar sailing

A low-cost way of sending spacecraft to other planets would be to use the radiation pressure on a solar sail. The intensity of the sun's electromagnetic radiation at distances near the earth's orbit is about  $1300 \text{ W/m}^2$ . What size sail would be needed to accelerate a 10,000 kg spacecraft toward Mars at  $0.010 \text{ m/s}^2$ ?

**SOLVE** The force that will create a  $0.010 \text{ m/s}^2$  acceleration is  $F = ma = 100 \text{ N}$ . We can use Equation 35.39 to find the sail

area that, by absorbing light, will receive a 100 N force from the sun:

$$A = \frac{cF}{I} = \frac{(3.00 \times 10^8 \text{ m/s})(100 \text{ N})}{1300 \text{ W/m}^2} = 2.3 \times 10^7 \text{ m}^2$$

Intermediate/Advanced Concepts

# Wave equations in a medium

The induced polarization in Maxwell's Equations yields another term in the wave equation:

$$\frac{\partial^2 E}{\partial z^2} - \mu\epsilon \frac{\partial^2 E}{\partial t^2} = 0 \quad \frac{\partial^2 E}{\partial z^2} - \frac{1}{v^2} \frac{\partial^2 E}{\partial t^2} = 0$$

This is the **Inhomogeneous Wave Equation**.

The polarization is the driving term for a new solution to this equation.

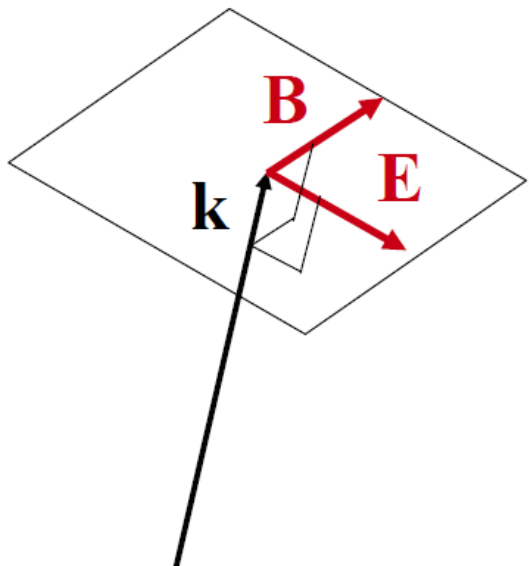
$$\frac{\partial^2 E}{\partial z^2} - \mu_0\epsilon_0 \frac{\partial^2 E}{\partial t^2} = 0 \quad \frac{\partial^2 E}{\partial z^2} - \frac{1}{c^2} \frac{\partial^2 E}{\partial t^2} = 0$$

**Homogeneous (Vacuum) Wave Equation**

$$\begin{aligned} \mathbf{E}(z, t) &= \text{Re}\{\mathbf{E}_0 e^{i(kz - \omega t)}\} \\ &= \frac{1}{2}\{\mathbf{E}_0 e^{i(kz - \omega t)} + \mathbf{E}_0^* e^{-i(kz - \omega t)}\} \\ &= |\mathbf{E}_0| \cos(kz - \omega t) \end{aligned} \quad n^2 = \frac{c^2}{v^2} = \frac{\mu\epsilon}{\mu_0\epsilon_0} \quad \frac{c}{v} = n$$



# Propagation of EM Waves



$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t} \quad \text{where} \quad \mathbf{E} = \hat{\mathbf{x}} E_0 e^{i(\mathbf{k} \cdot \mathbf{r} - \omega t)}$$

$$\Rightarrow \nabla \times \equiv i\mathbf{k} \times \quad \text{and} \quad \frac{\partial}{\partial t} \equiv -i\omega$$

$$\Rightarrow \mathbf{B} = \frac{1}{\omega} \mathbf{k} \times \mathbf{E}$$

Vectors  $\mathbf{k}$ ,  $\mathbf{E}$ ,  $\mathbf{B}$  form a right-handed triad.

Note: free space or isotropic media only

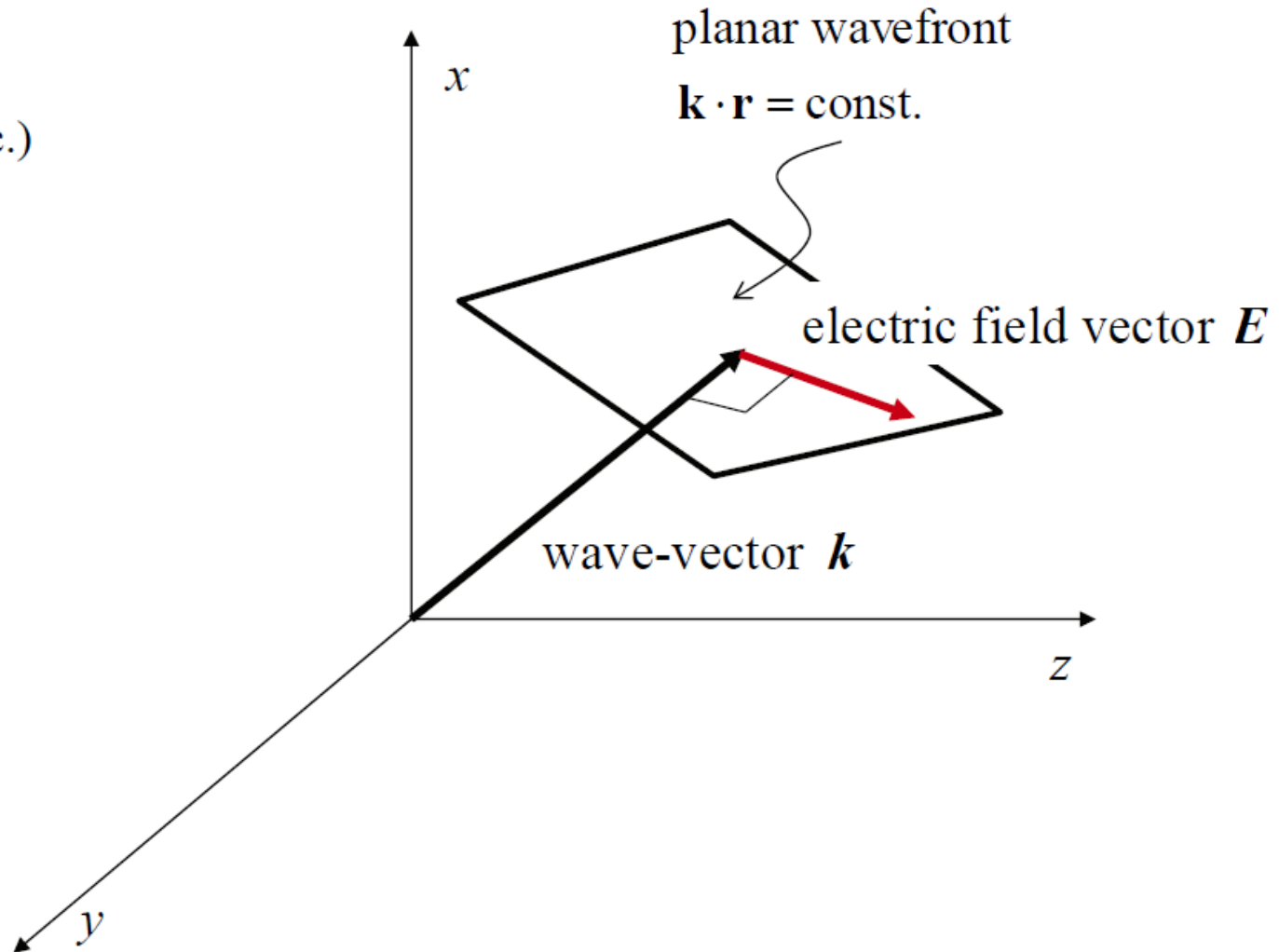
# Polarization and Propagation

In isotropic media  
(e.g. free space,  
amorphous glass, etc.)

$$\mathbf{k} \cdot \mathbf{E} = 0$$

i.e.  $\mathbf{k} \perp \mathbf{E}$

More generally,  
 $\mathbf{k} \cdot \mathbf{D} = 0$   
(reminder: in  
anisotropic media,  
e.g. crystals, one  
could have  
 $\mathbf{E}$  not parallel to  $\mathbf{D}$ )

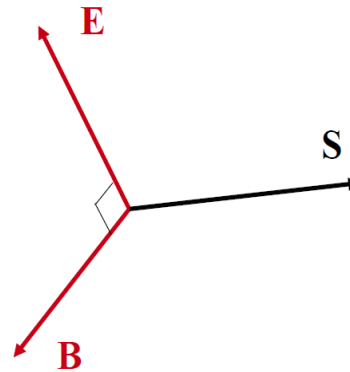


# Energy and Intensity

Summary (free space or isotropic media)

$$\mathbf{S} = \frac{1}{\mu_0} \mathbf{E} \times \mathbf{B}; \quad \|\mathbf{S}\| = c\epsilon_0 \|\mathbf{E}\|^2 \quad \text{Poynting vector}$$

$$\langle\langle \mathbf{S} \rangle\rangle = \frac{1}{T} \int_t^{t+T} \|\mathbf{S}\| dt \quad \text{Irradiance (or intensity)}$$



$$\mathbf{S} = \mathbf{E} \times \mathbf{H}$$

$$\mathbf{S} = \frac{1}{\mu_0} \mathbf{E} \times \mathbf{B} = c^2 \epsilon_0 \mathbf{E} \times \mathbf{B}$$

so in free space

$$\mathbf{S} \parallel \mathbf{k}$$

$\mathbf{S}$  has units of  $\text{W}/\text{m}^2$

so it represents energy flux (energy per unit time & unit area)

$$\langle \sin^2(kx - \omega t) \rangle$$

$$= \langle \cos^2(kx - \omega t) \rangle = \frac{1}{2}$$

- **Poynting vector** describes flows of E-M power
- Power flow is directed along this vector (usually parallel to  $\mathbf{k}$ )
- Intensity is average energy transfer (i.e. the time averaged Poynting vector:  $I = \langle \mathbf{S} \rangle = P/A$ , where  $P$  is the power (energy transferred per second) of a wave that impinges on area  $A$ .)

$$\langle\langle \mathbf{S} \rangle\rangle = I \equiv \langle \mathbf{E}(t) \times \mathbf{H}(t) \rangle = \frac{c\epsilon_0}{2} E^2 = \frac{c\epsilon_0}{2} (E_x^2 + E_y^2)$$

$$c\epsilon_0 \approx 2.654 \times 10^{-3} \text{ A/V}$$

**example**  $E = 1 \text{ V/m}$

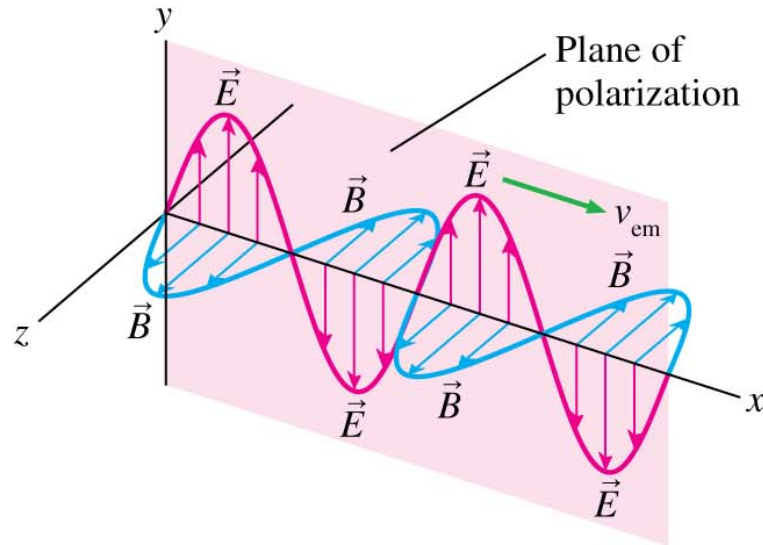
$$I = ? \text{ W/m}^2$$

$$h\omega[\text{eV}] = \frac{1239.85}{\lambda[\text{nm}]}$$

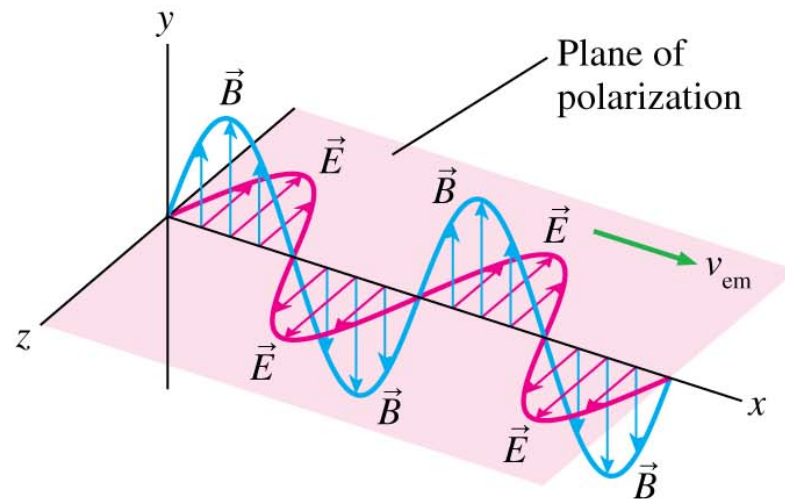
$$h = 1.05457266 \times 10^{-34} \text{ Js}$$

# Polarization & Plane of Polarization

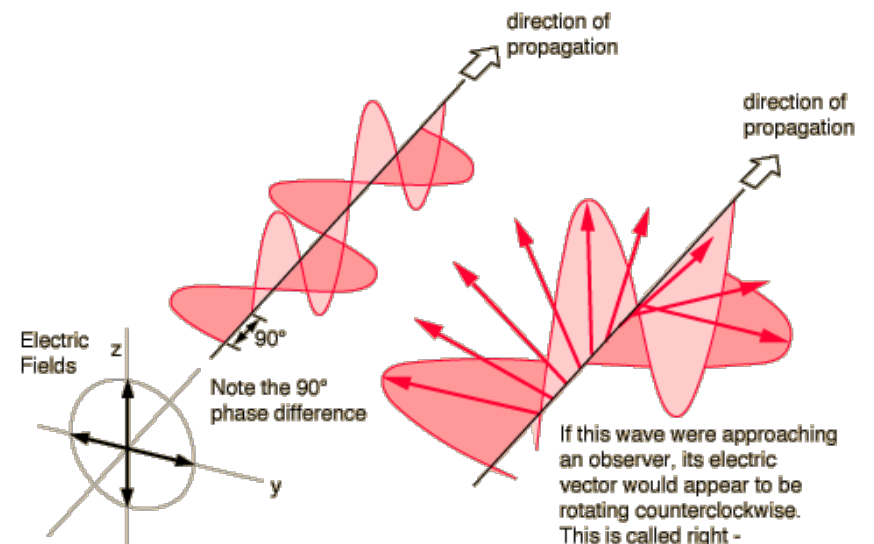
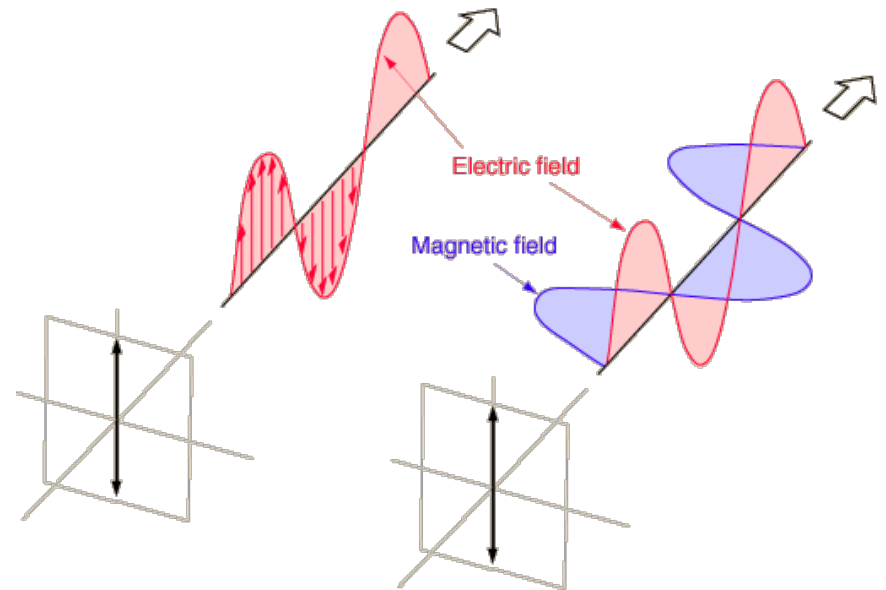
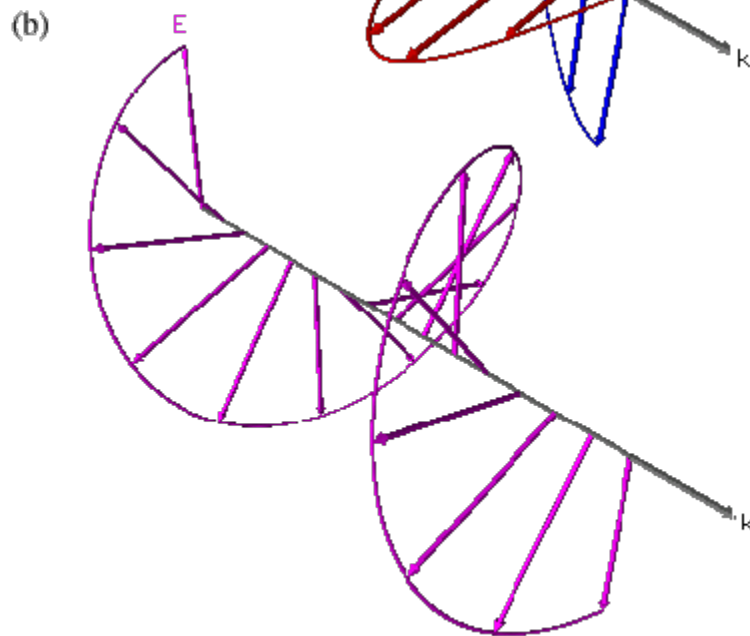
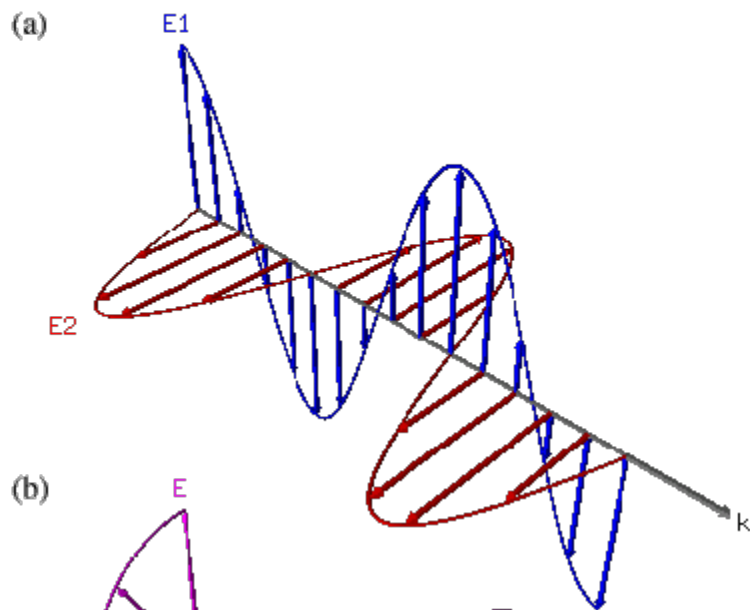
(a) Vertical polarization



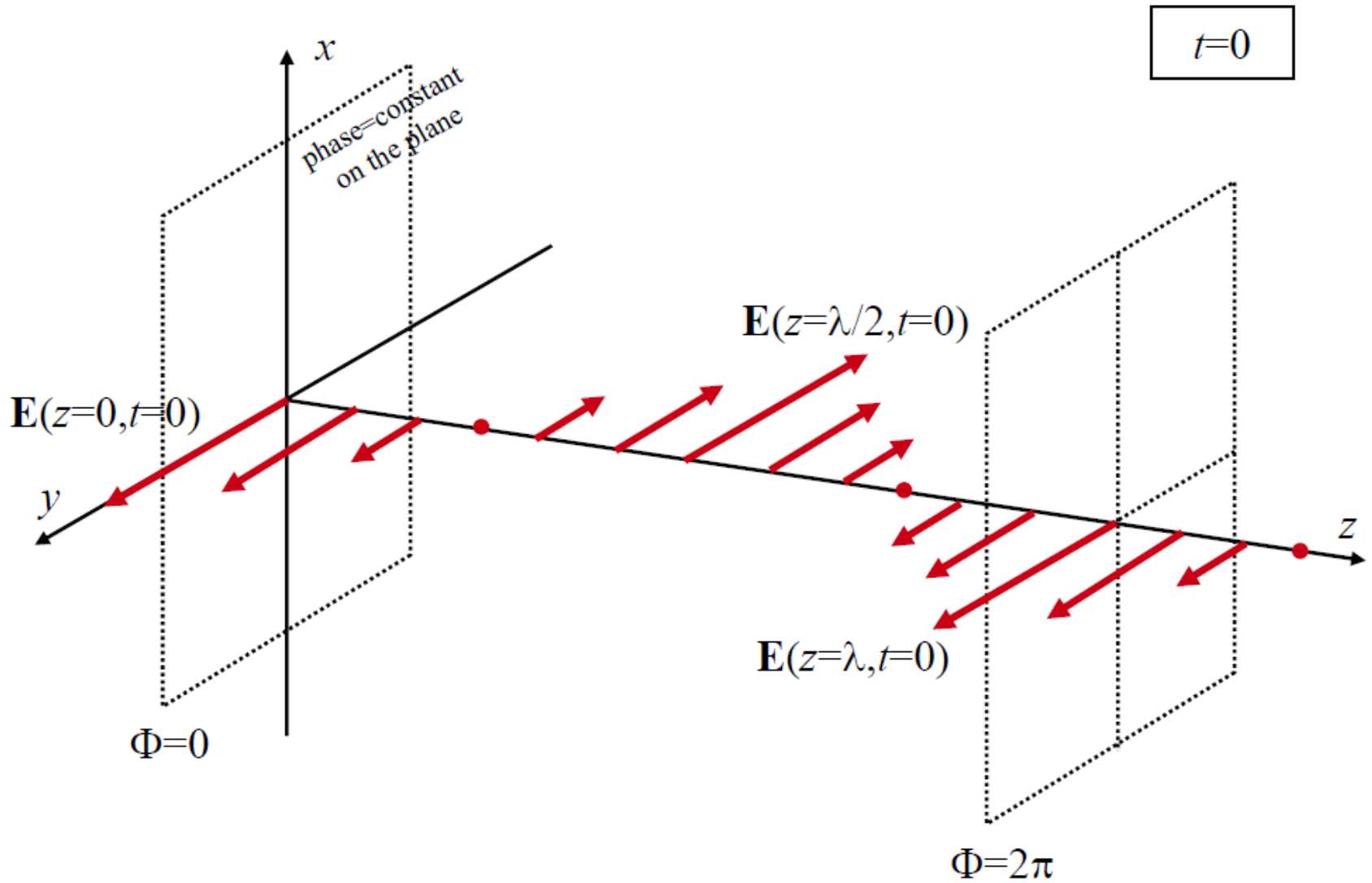
(b) Horizontal polarization



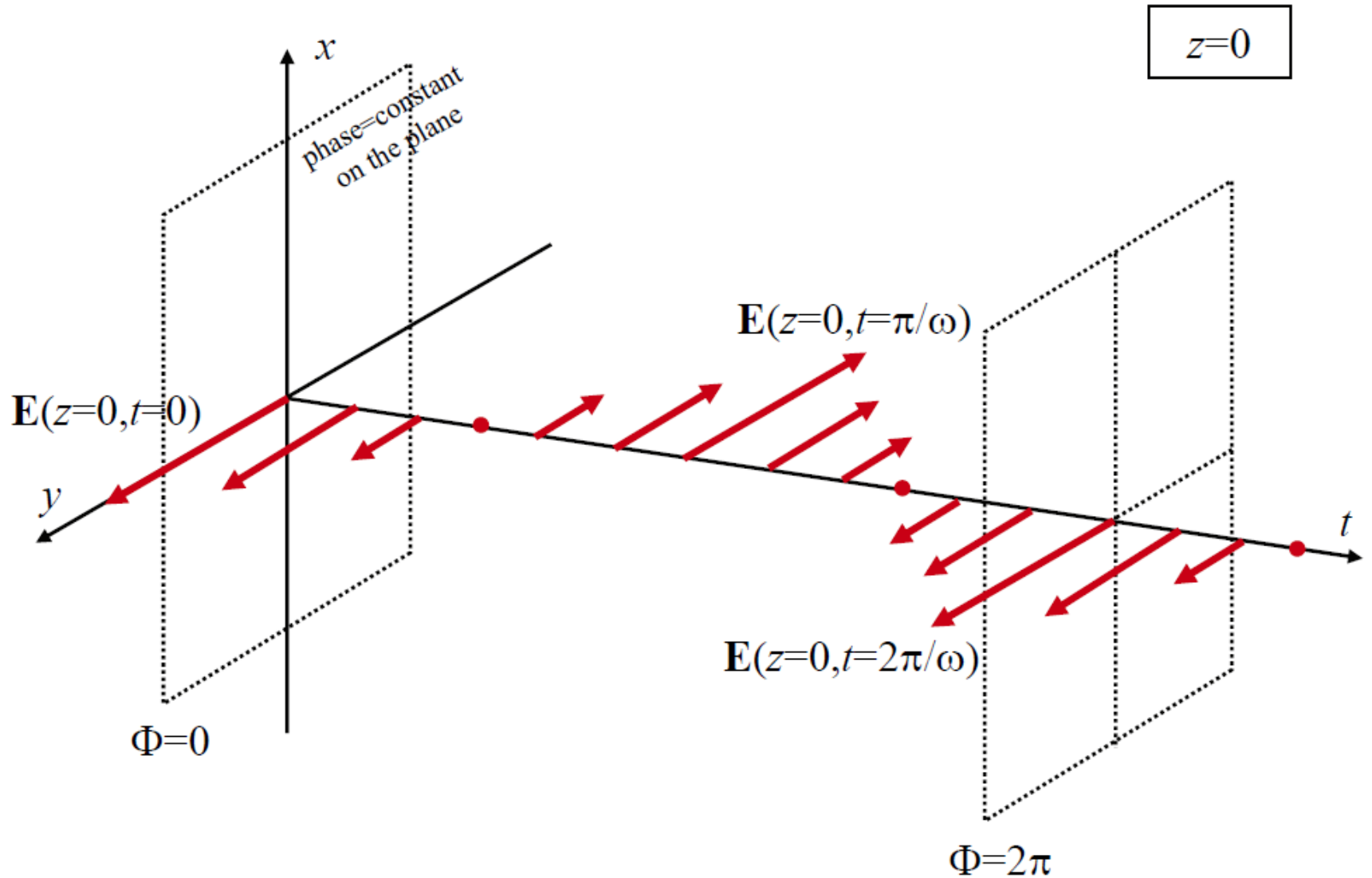
# Linear versus Circular Polarization



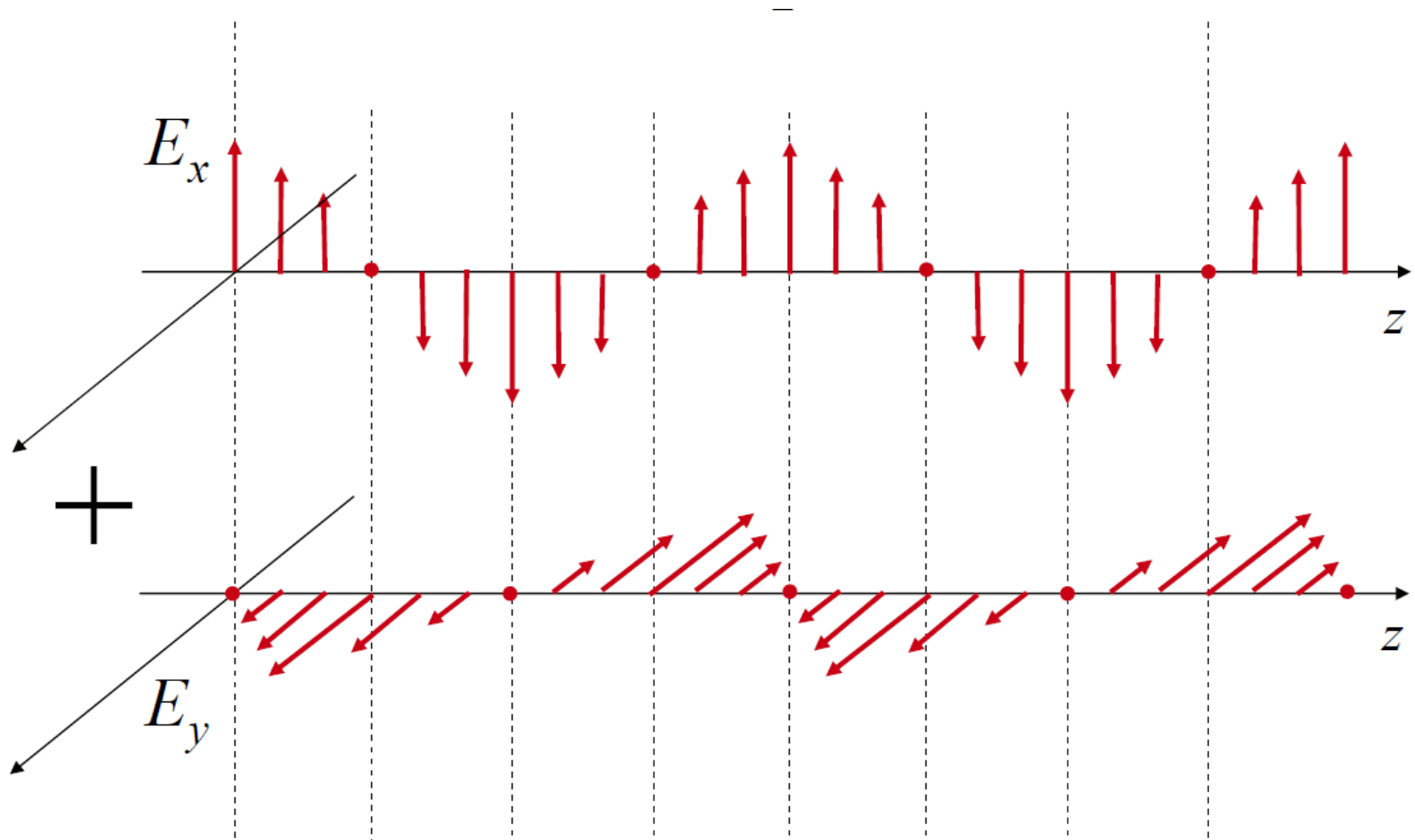
# Linear polarization (frozen time)



# Linear polarization (fixed space)

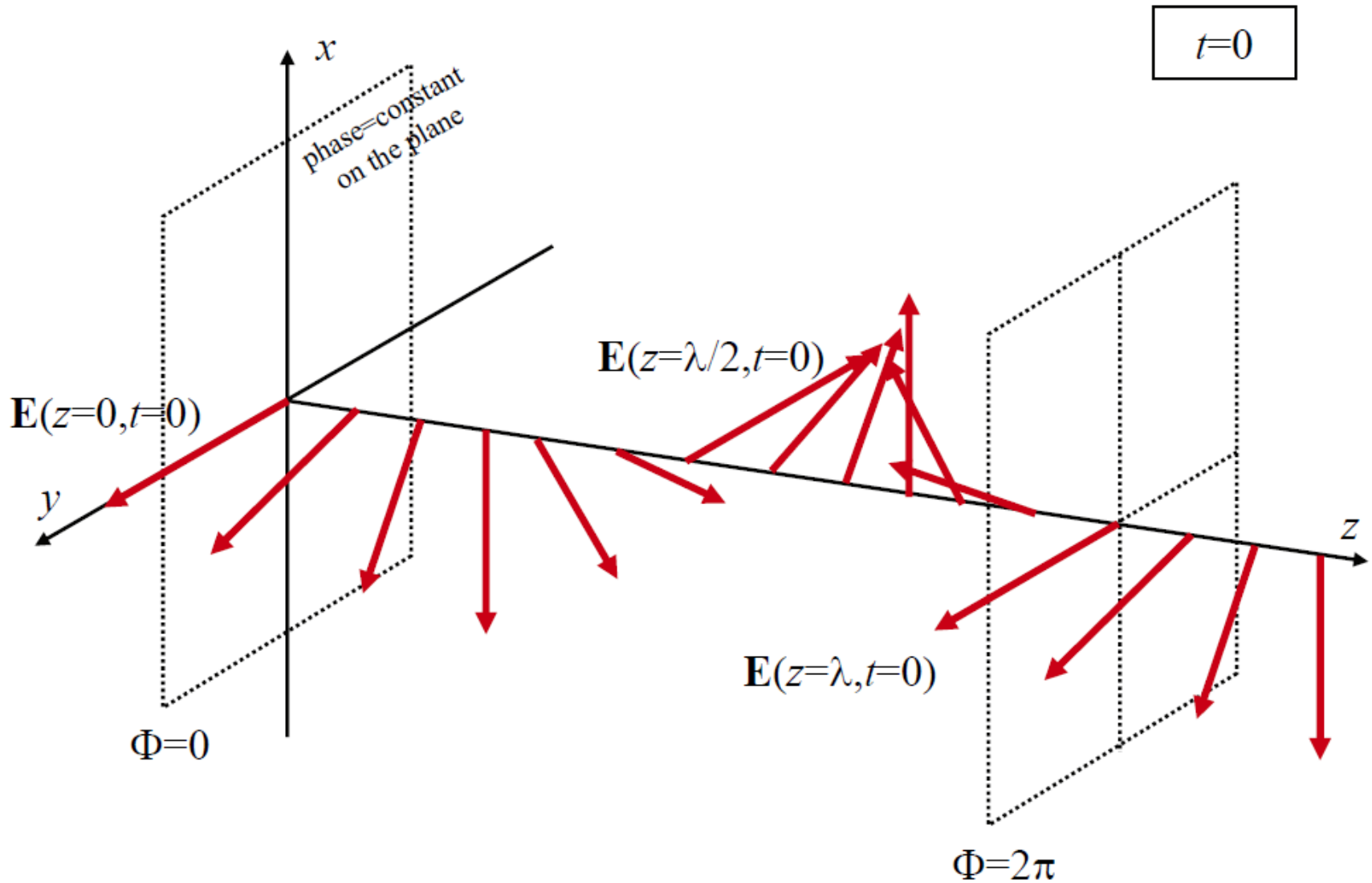


# Circular polarization (linear components)

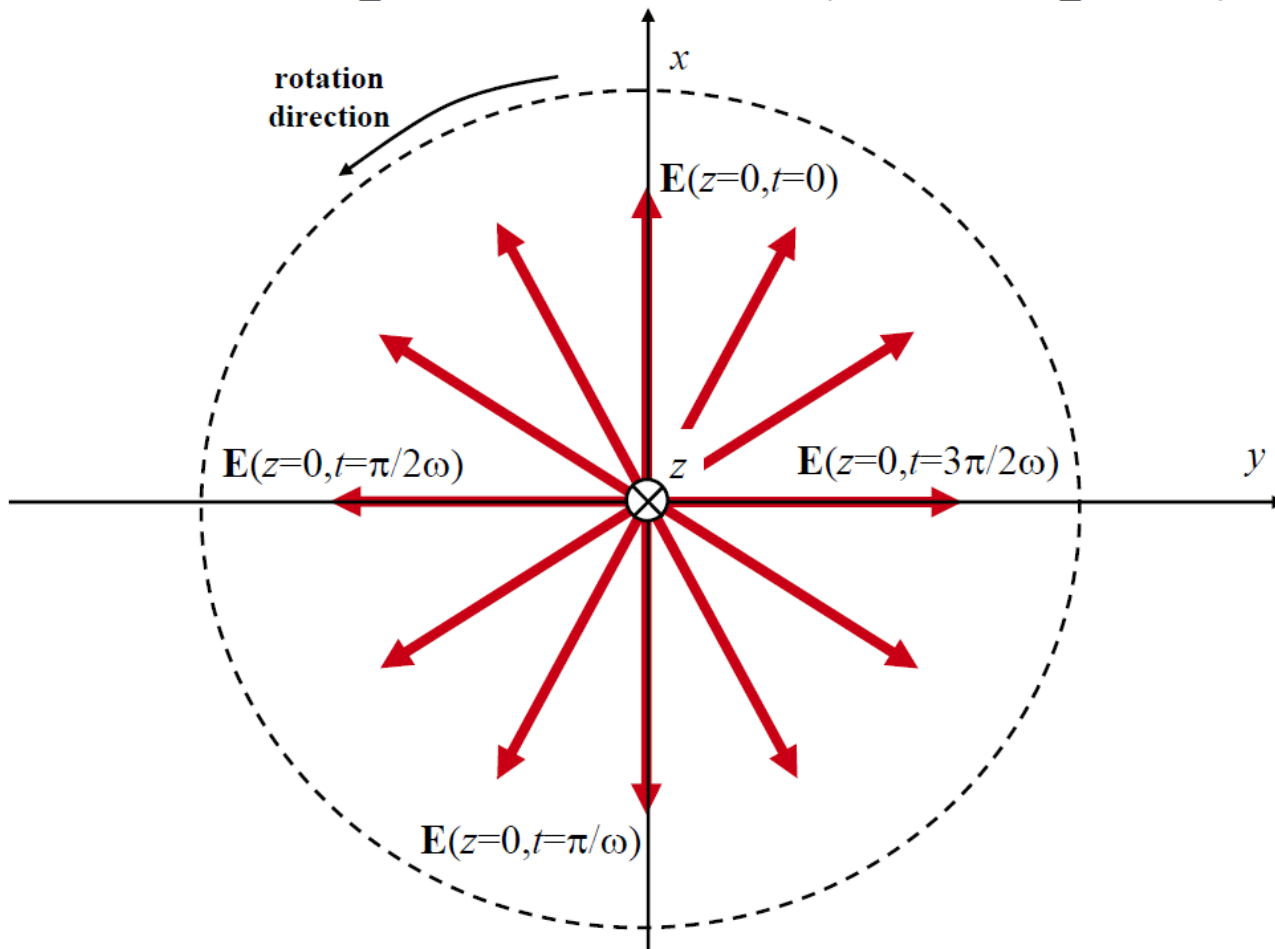




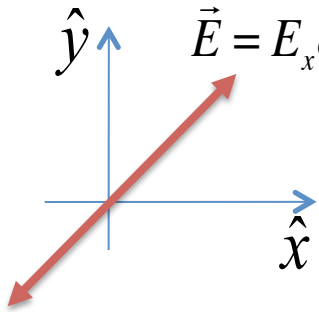
# Circular polarization (frozen time)



# Circular polarization (fixed space)



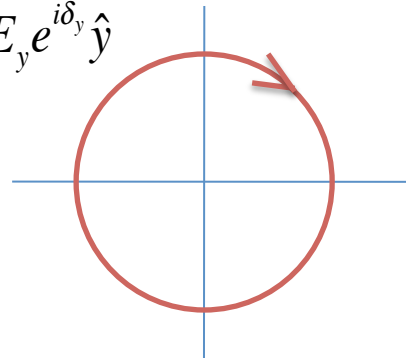
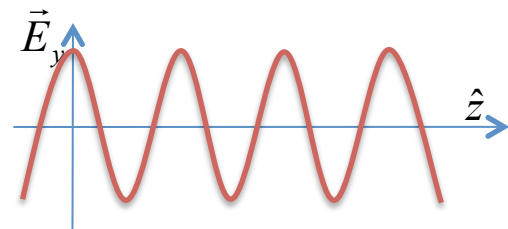
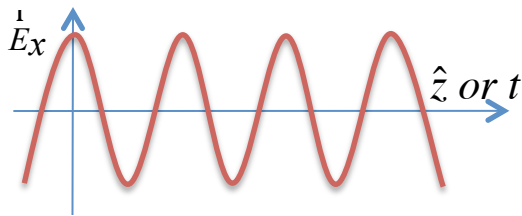
# Polarization: Summary



linear polarization  
y-direction

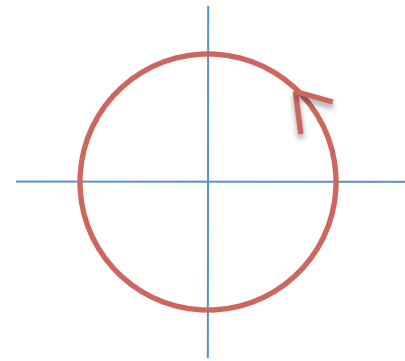
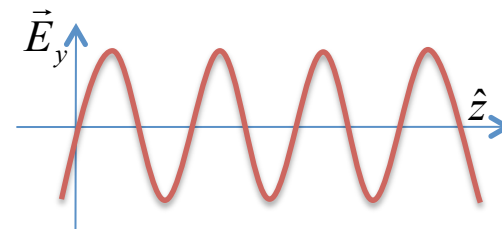
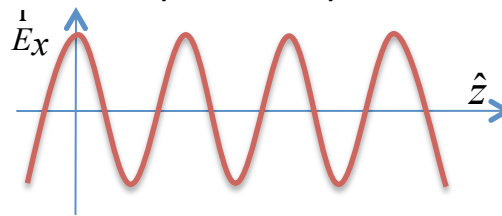
Phase difference  $\delta = \delta_x - \delta_y$

Phase difference =  $0^\circ$



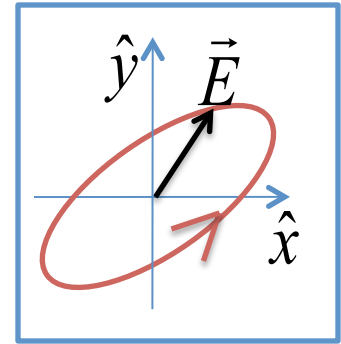
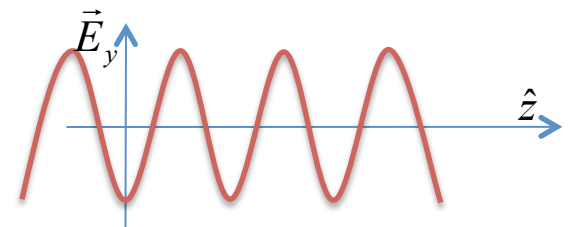
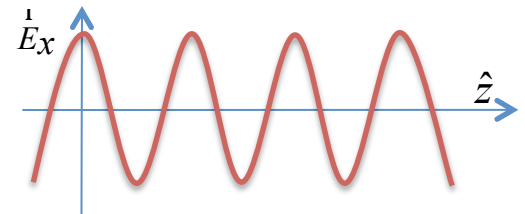
right circular  
polarization

Phase difference  $\rightarrow$   
 $90^\circ (\pi/2, \lambda/4)$



left circular  
polarization  
(+: positive helicity)

Phase difference  $\rightarrow$   
 $180^\circ (\pi, \lambda/2)$



left elliptical  
polarization

# Polarization Applets

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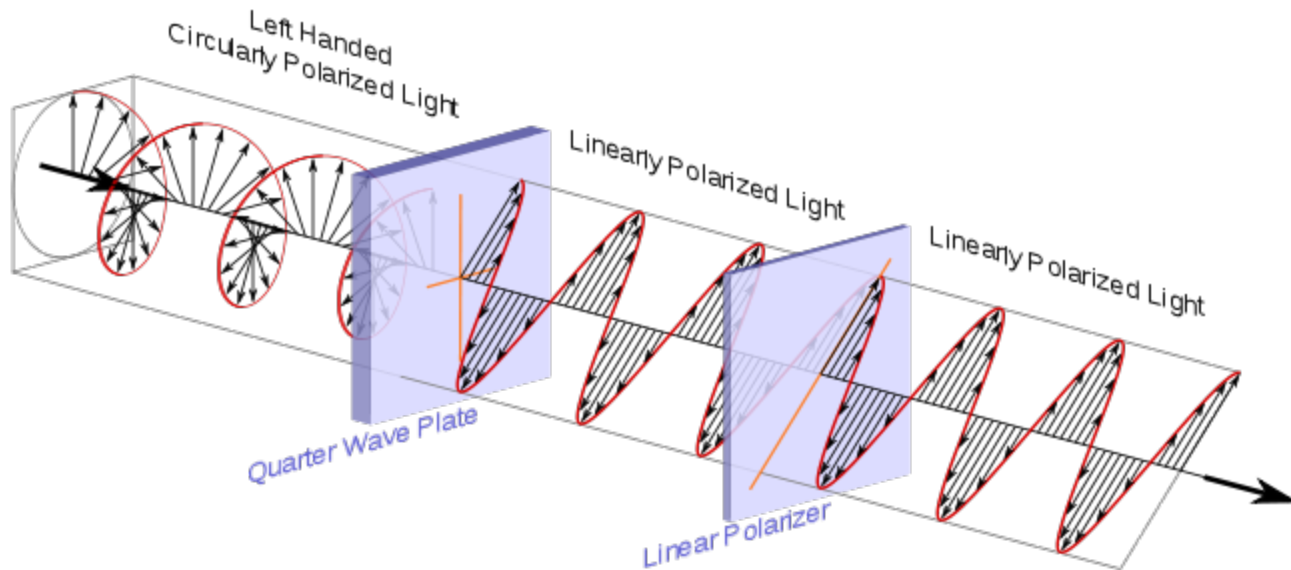
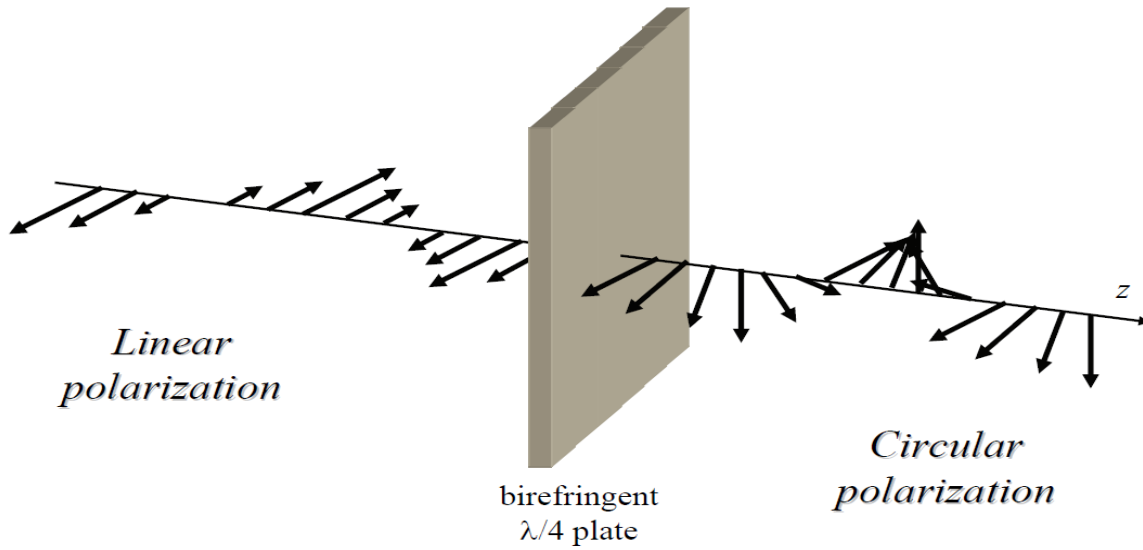
- [Polarization Exploration](#)

[http://webphysics.davidson.edu/physlet\\_resources/dav\\_optics/Examples/polarization.html](http://webphysics.davidson.edu/physlet_resources/dav_optics/Examples/polarization.html)

- 3D View of Polarized Light

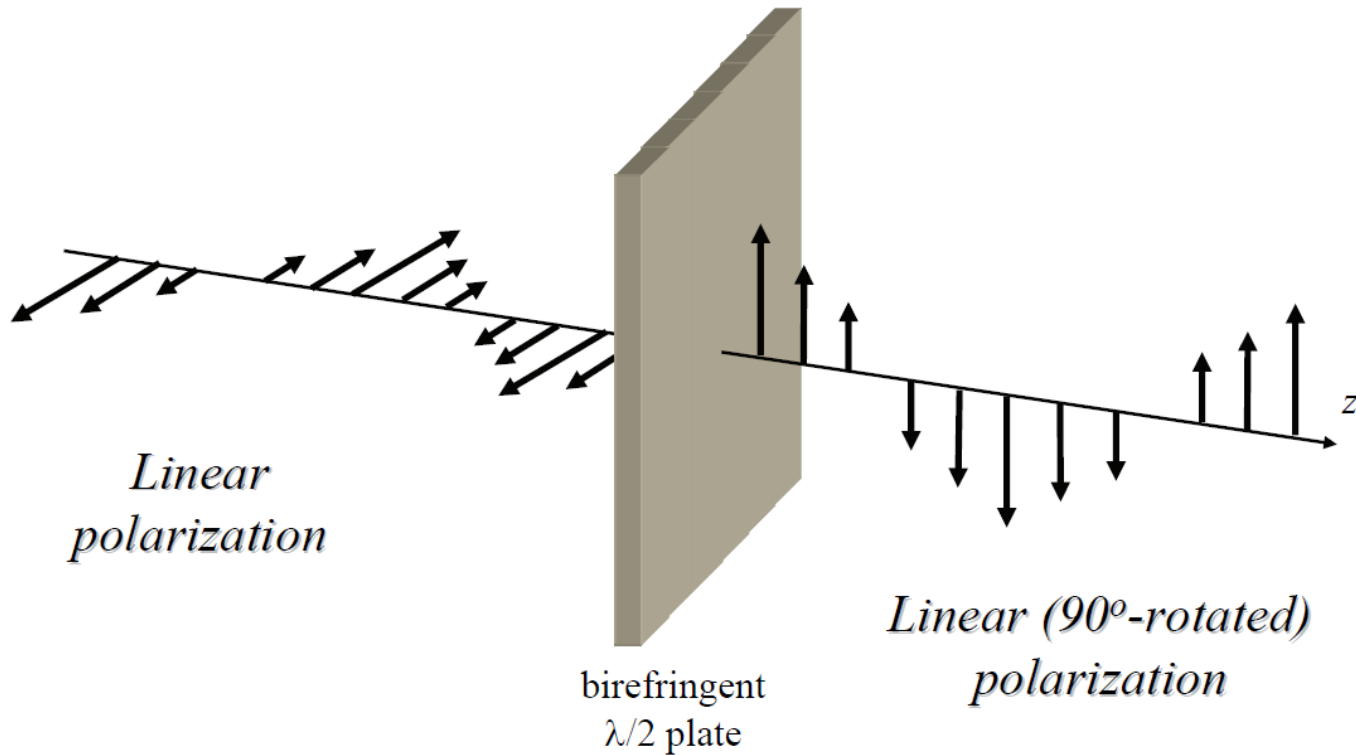
<http://fipsgold.physik.uni-kl.de/software/java/polarisation/index.html>

# Quarter wave plate

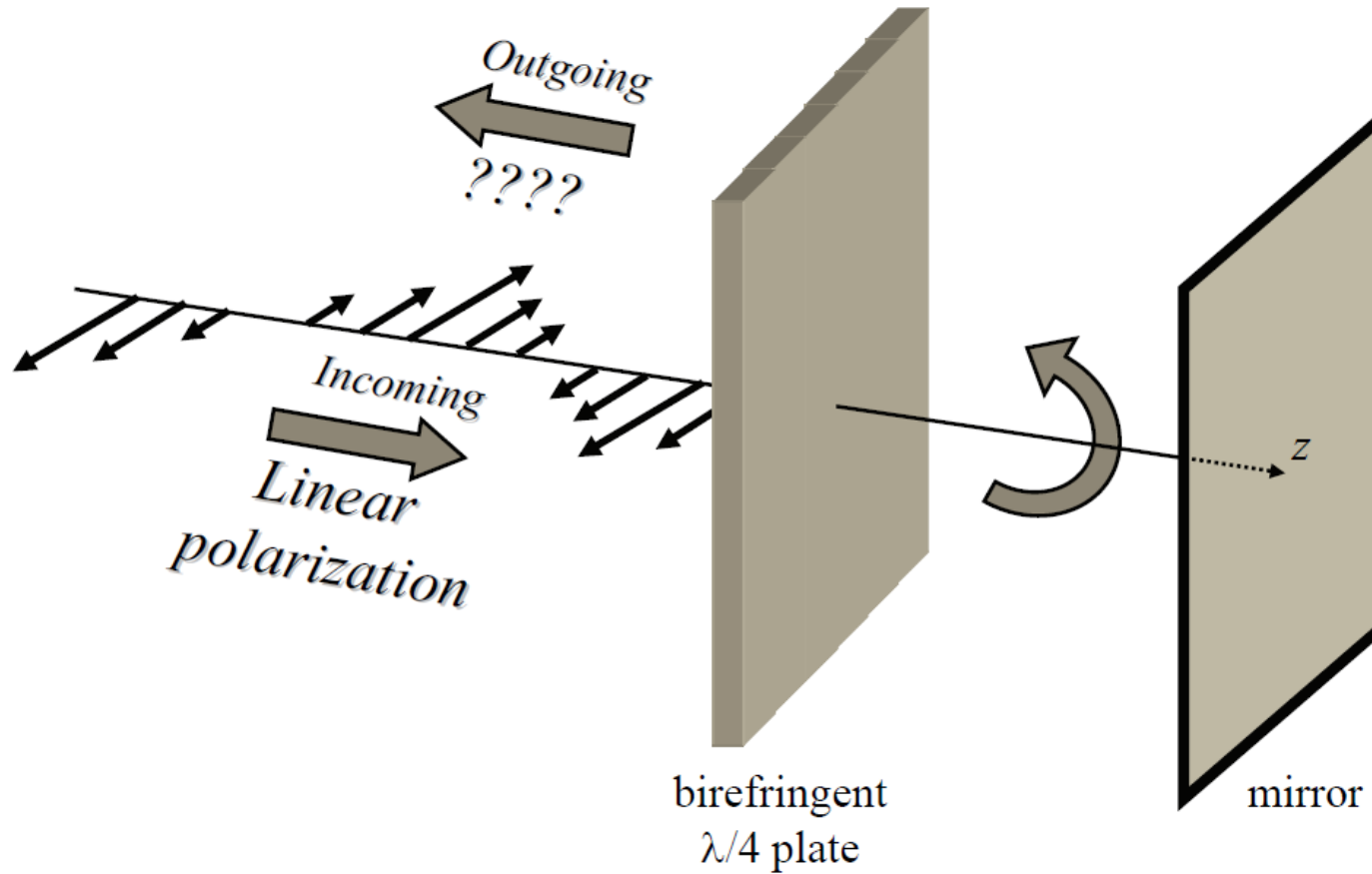


# Half wave plate

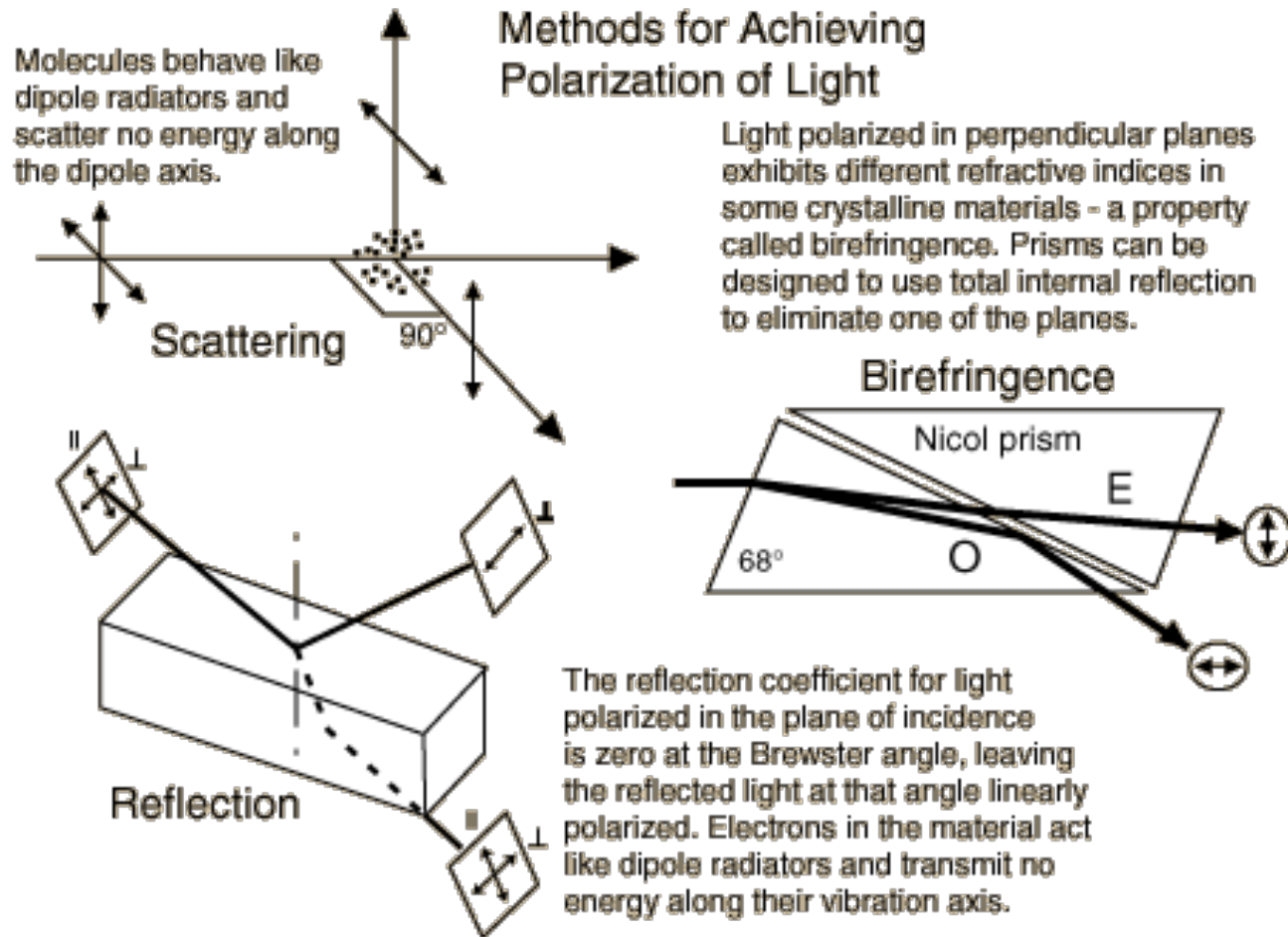
$\lambda/2$  plate



Quiz for the Lab – Bonus Credit 0.2 pts



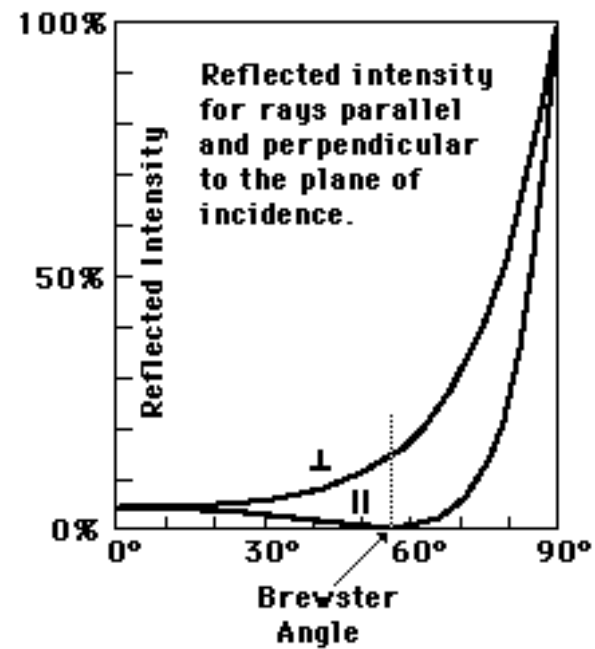
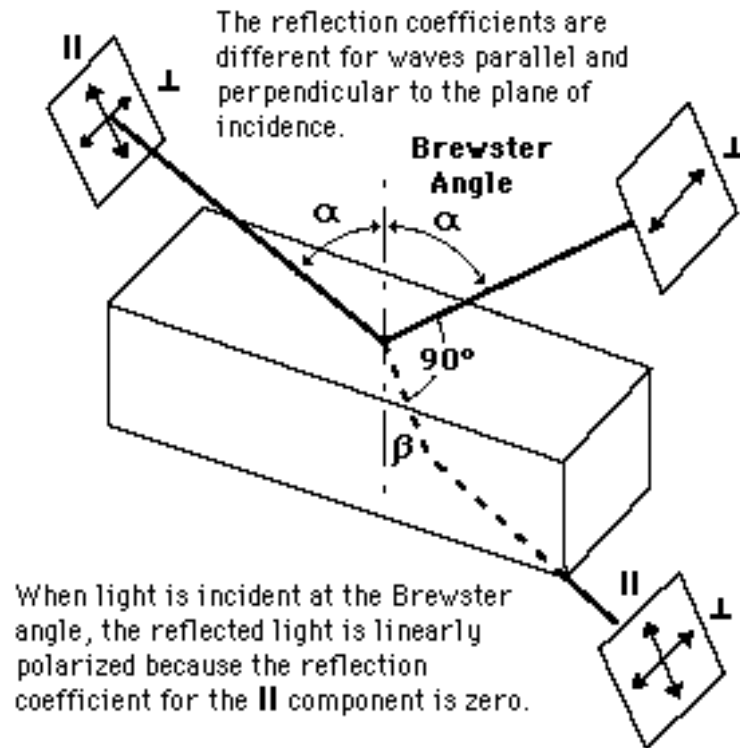
# Methods for generating polarized light



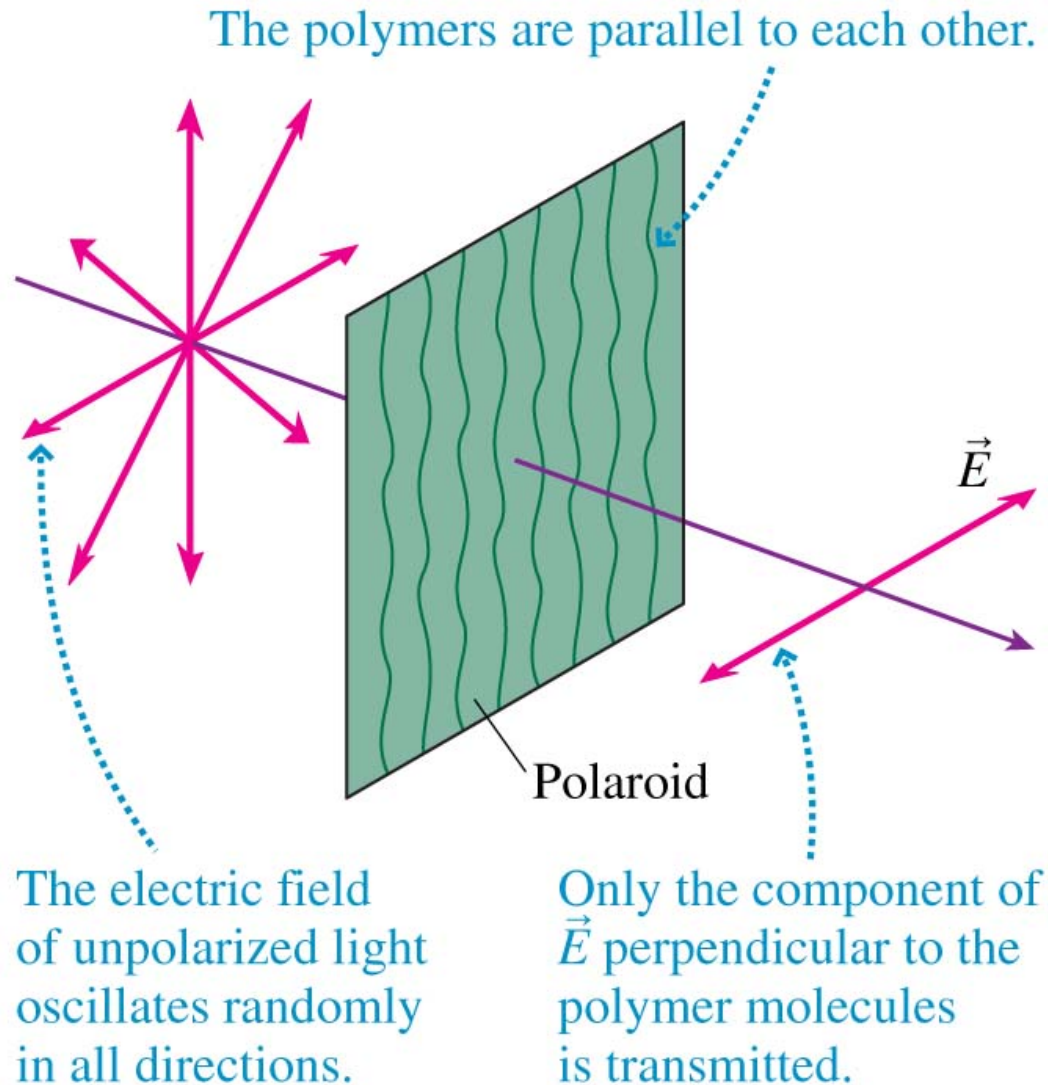


# Polarization by Reflection

<http://hyperphysics.phy-astr.gsu.edu/hbase/phyopt/polar.html>



# A Polarizing Filter



# Malus' s Law

Suppose a *polarized* light wave of intensity  $I_0$  approaches a polarizing filter.  $\vartheta$  is the angle between the incident plane of polarization and the polarizer axis. The transmitted intensity is given by Malus' s Law:

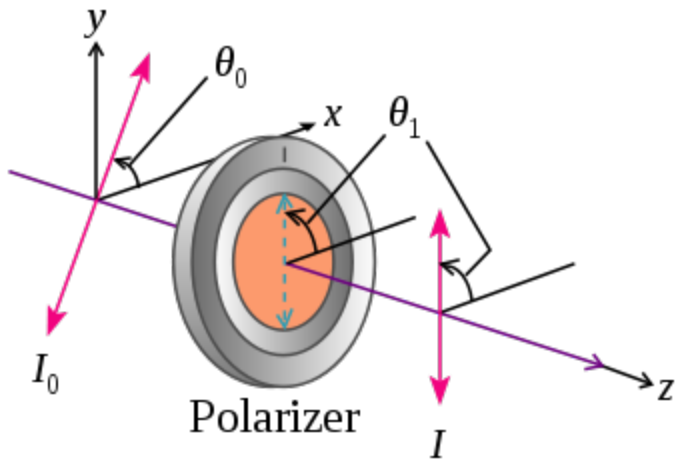
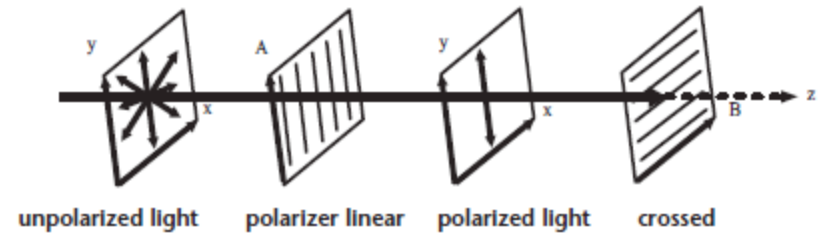
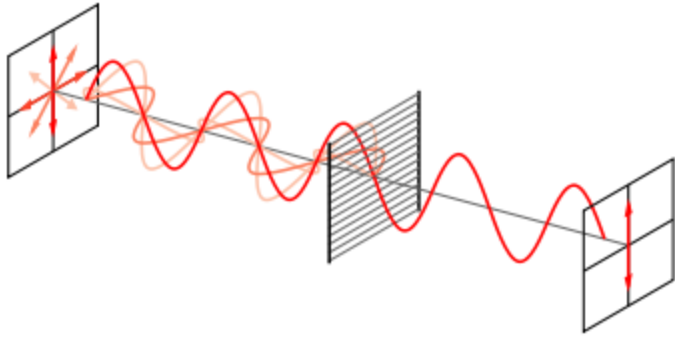
$$I_{\text{transmitted}} = I_0 \cos^2 \theta \quad (\text{incident light polarized})$$

If the light incident on a polarizing filter is *unpolarized*, the transmitted intensity is

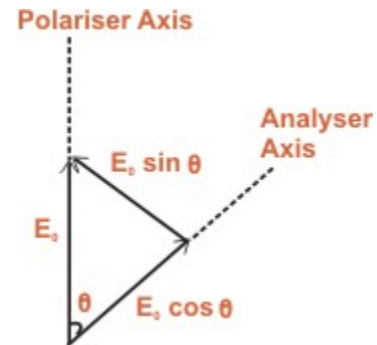
$$I_{\text{transmitted}} = \frac{1}{2} I_0 \quad (\text{incident light unpolarized})$$

In other words, a polarizing filter passes 50% of unpolarized light and blocks 50%.

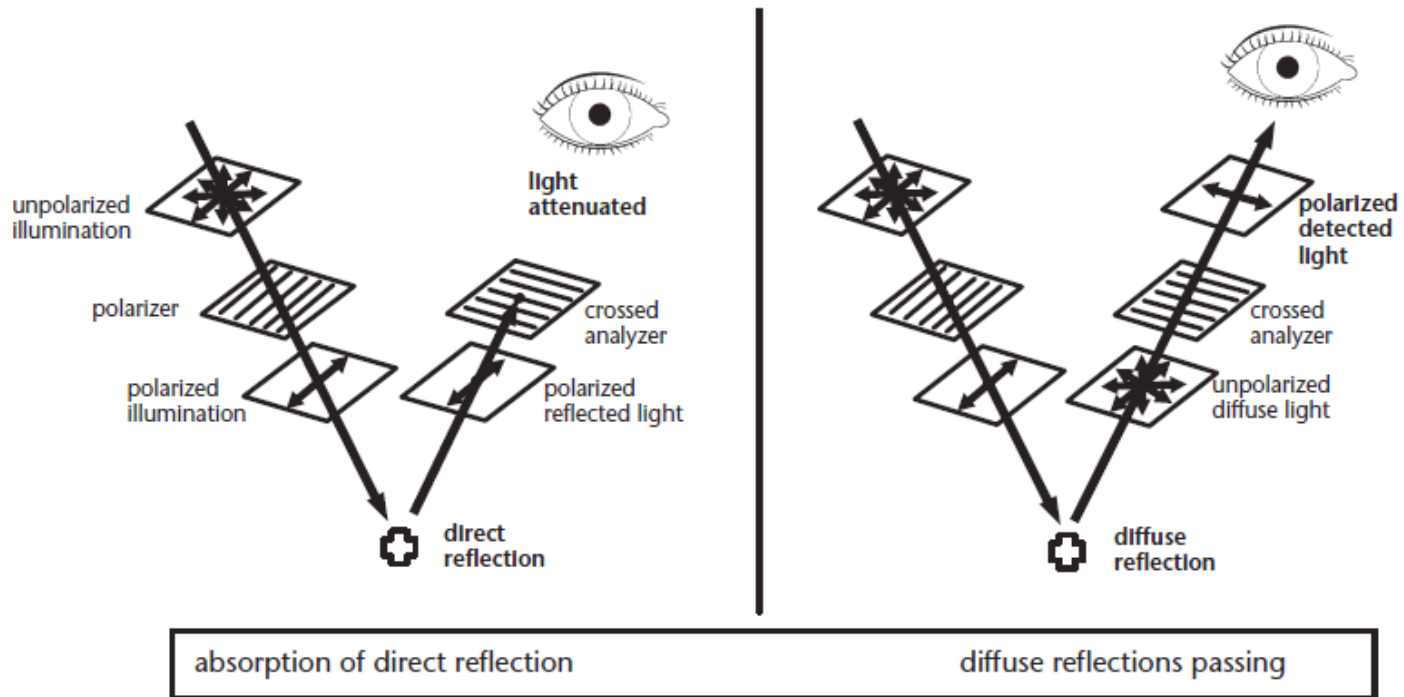
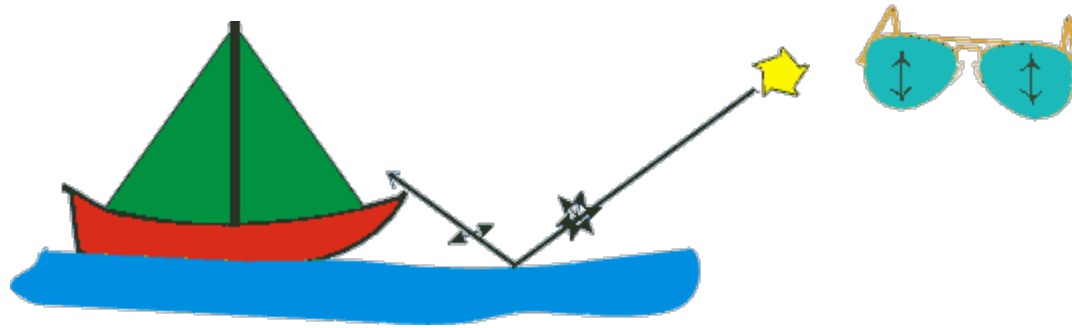
# Malus's Law



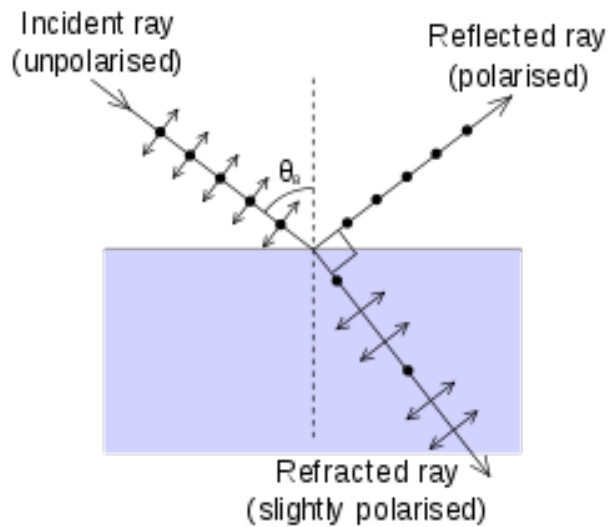
$$I = \frac{1}{2} c \epsilon_0 E_0^2 \cos^2 \theta = I_0 \cos^2 \theta,$$



# Polarized sunglasses



# Brewster Angle



$$\theta_1 + \theta_2 = 90^\circ,$$

$$n_1 \sin(\theta_1) = n_2 \sin(\theta_2),$$

$$n_1 \sin(\theta_B) = n_2 \sin(90^\circ - \theta_B) = n_2 \cos(\theta_B).$$

$$\theta_B = \arctan\left(\frac{n_2}{n_1}\right),$$

# Polarization by scattering (Rayleigh scattering/Blue Sky)

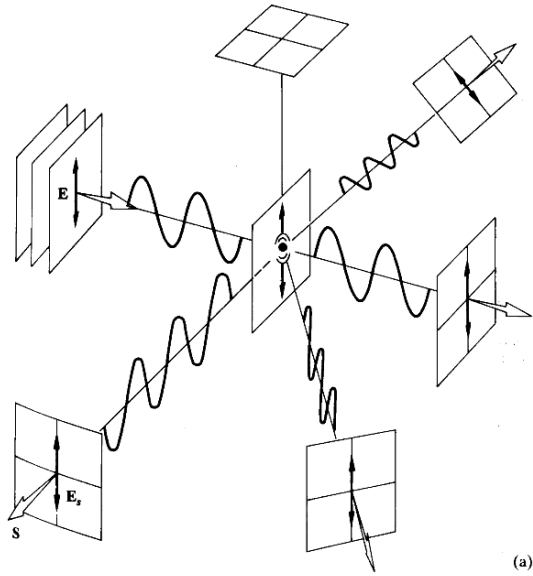


FIGURE 8.35a Scattering of polarized light by a molecule.

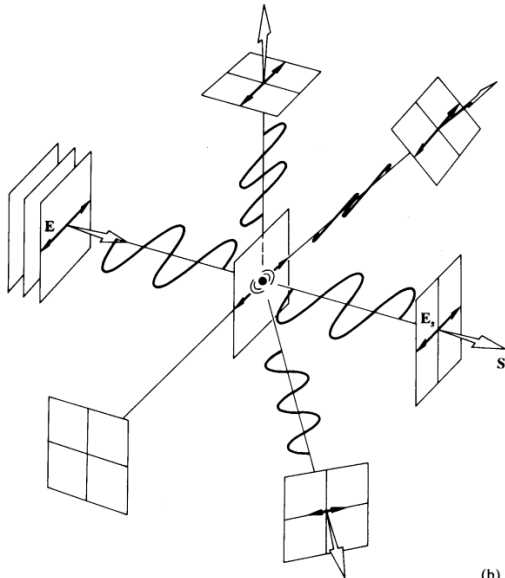


FIGURE 8.35b

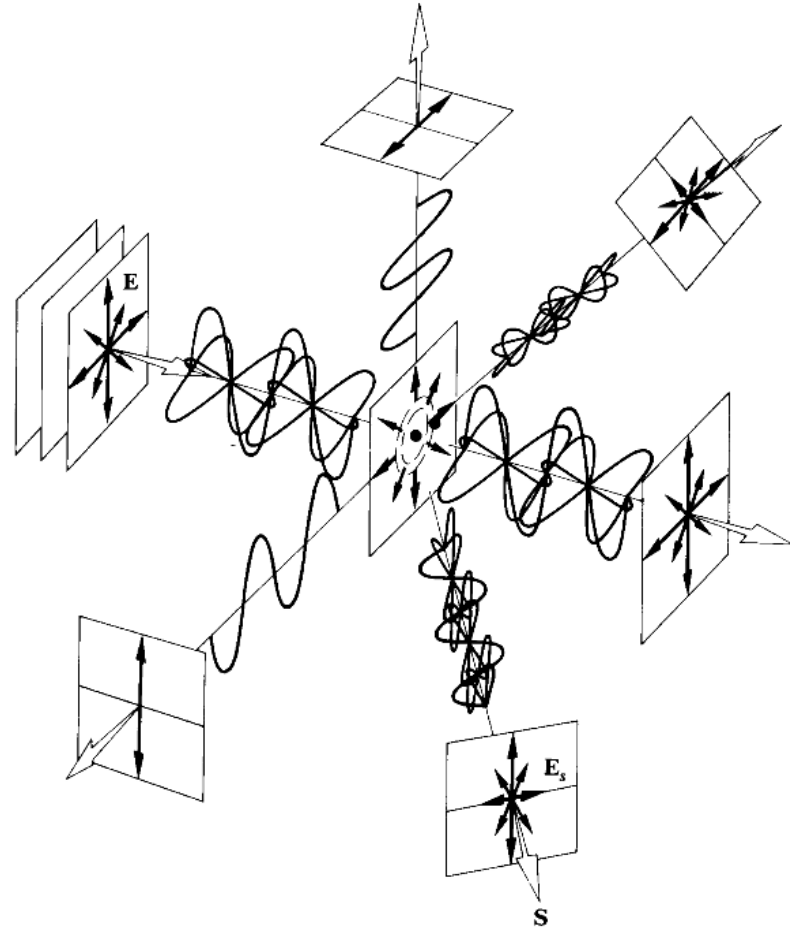
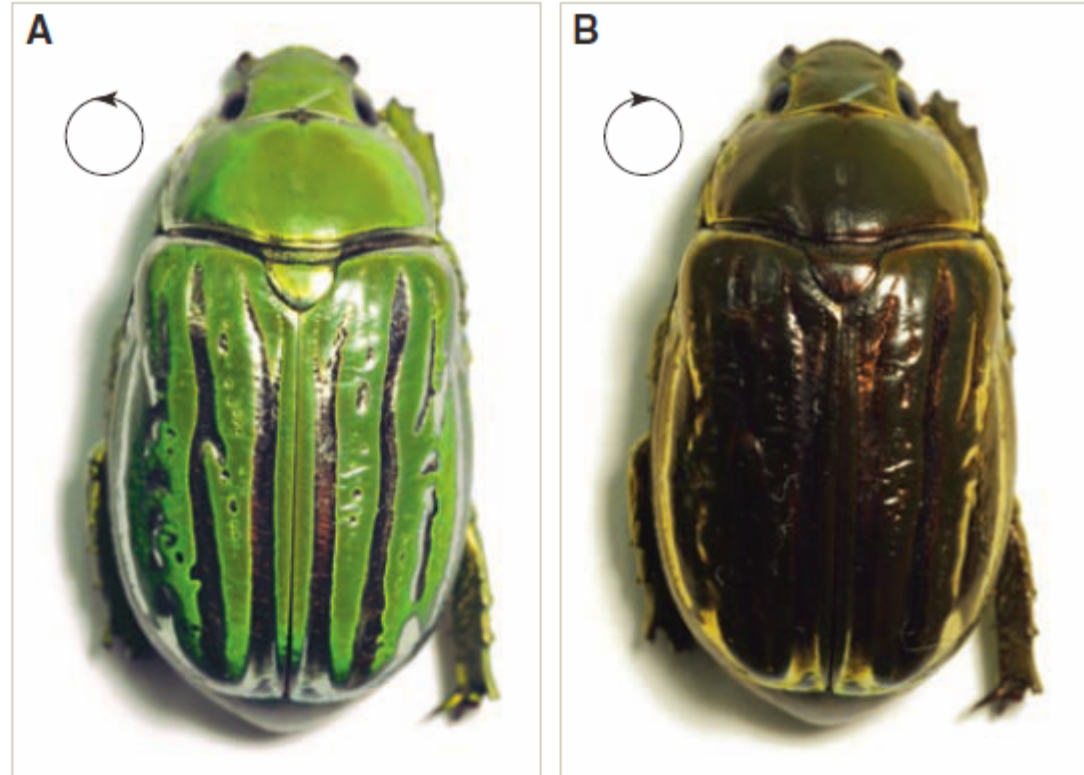


FIGURE 8.36 Scattering of unpolarized light by a molecule.

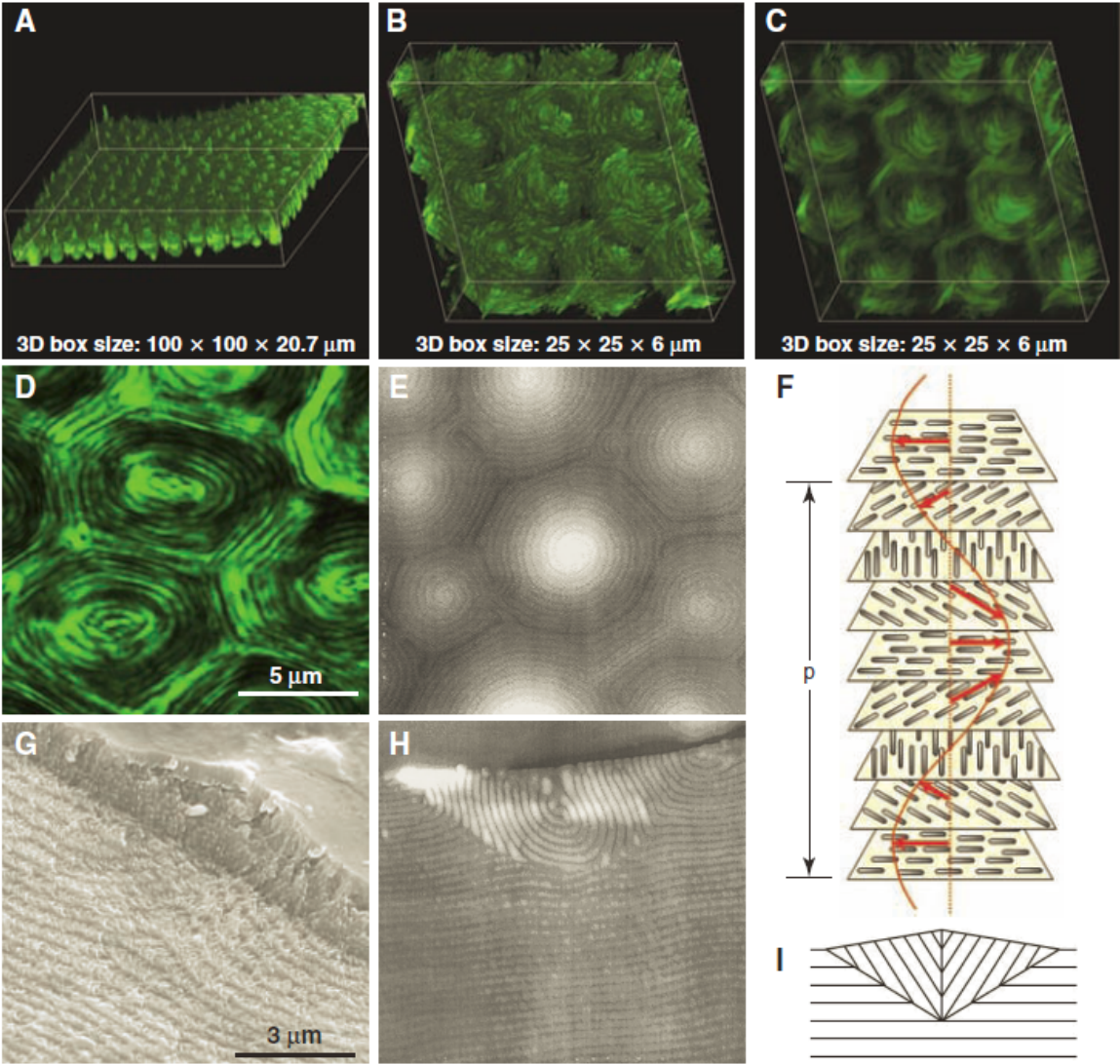
# Circularly polarized light in nature

**Fig. 1.** Photographs of the beetle *C. gloriosa*. **(A)** The bright green color, with silver stripes as seen in unpolarized light or with a left circular polarizer. **(B)** The green color is mostly lost when seen with a right circular polarizer.

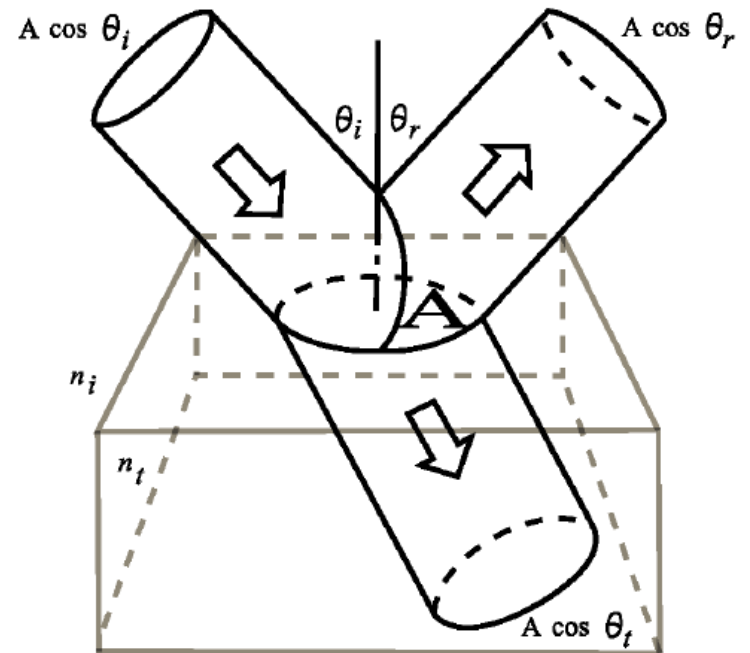
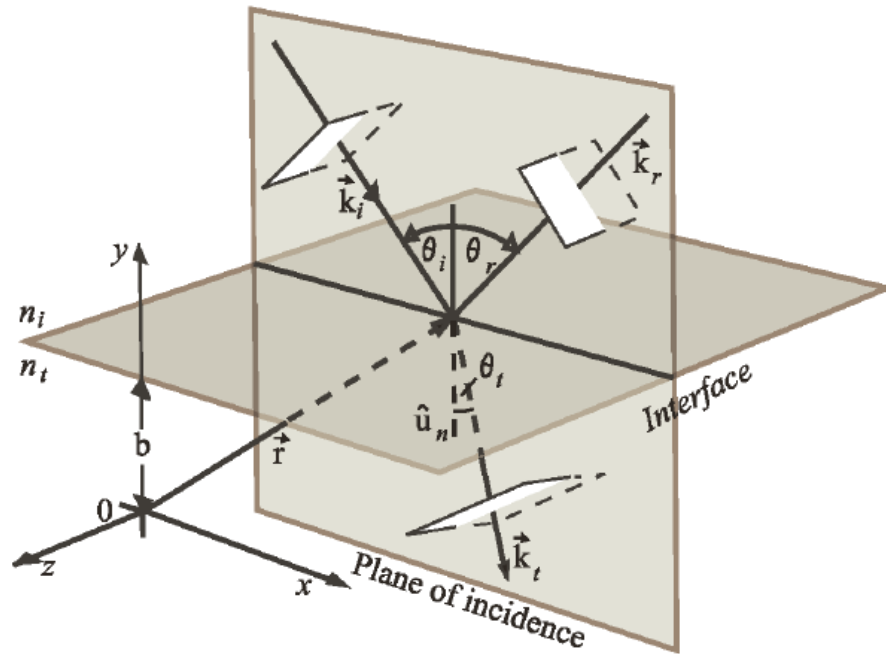




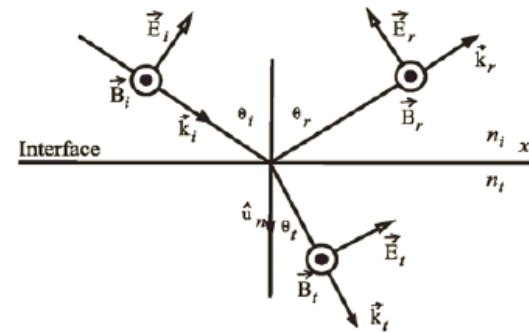
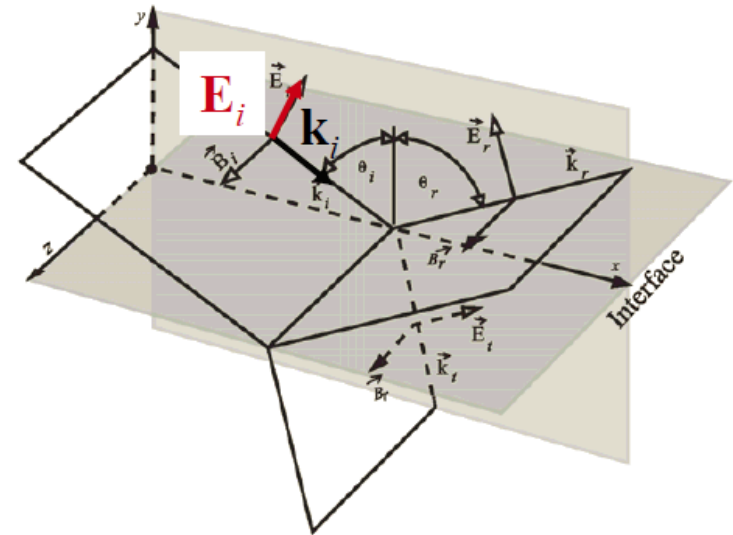
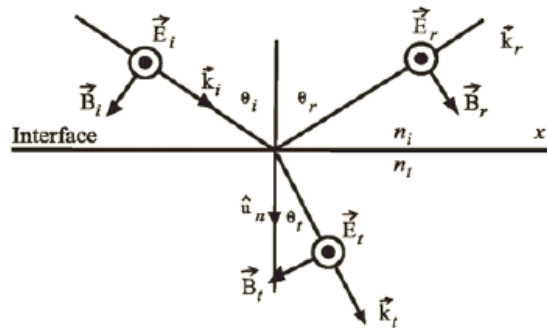
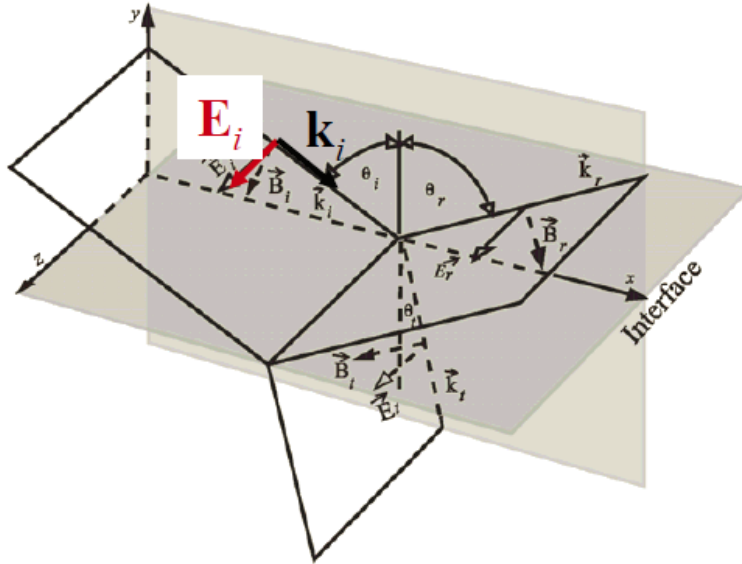
# Morphology and microstructure of cellular pattern of *C. gloriosa*



# Reflection and Transmission @ dielectric interface



# Beyond Snell's Law: Polarization?

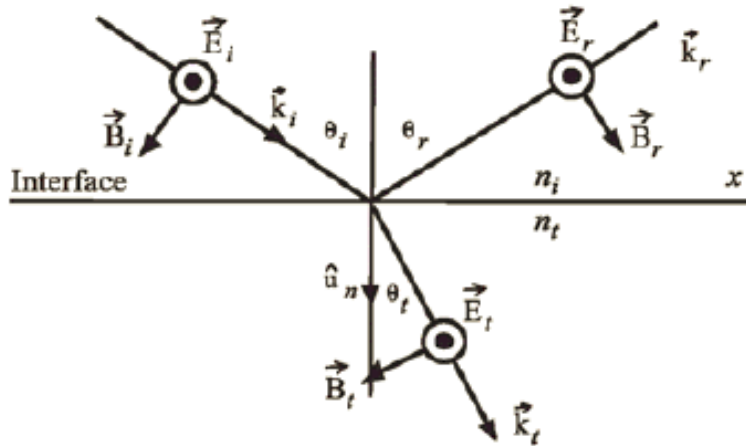


# Reflection and Transmission (Fresnel's equations)

Can be deduced from the application of boundary conditions of EM waves.

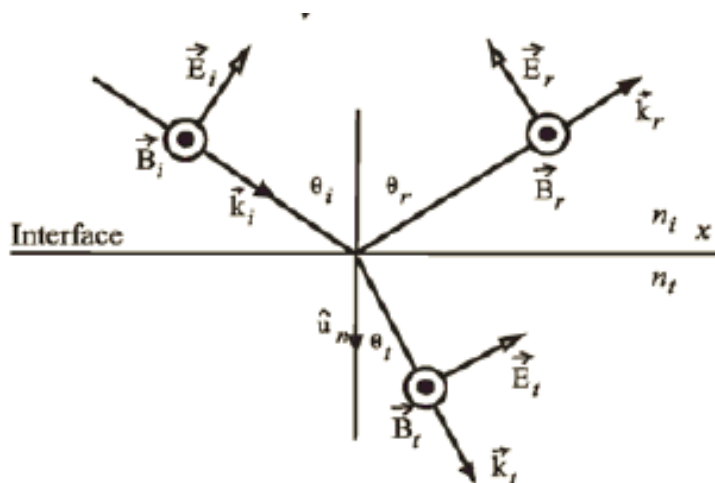
An online calculator is available at

<http://hyperphysics.phy-astr.gsu.edu/hbase/phyopt/freseq.html>



$$r_{\perp} = \left( \frac{E_{0r}}{E_{0i}} \right)_{\perp} = \frac{n_i \cos \theta_i - n_t \cos \theta_t}{n_i \cos \theta_i + n_t \cos \theta_t}$$

$$t_{\perp} = \left( \frac{E_{0t}}{E_{0i}} \right)_{\perp} = \frac{2n_i \cos \theta_i}{n_i \cos \theta_i + n_t \cos \theta_t}$$



$$r_{\parallel} = \left( \frac{E_{0r}}{E_{0i}} \right)_{\parallel} = \frac{n_t \cos \theta_i - n_i \cos \theta_t}{n_t \cos \theta_i + n_i \cos \theta_t}$$

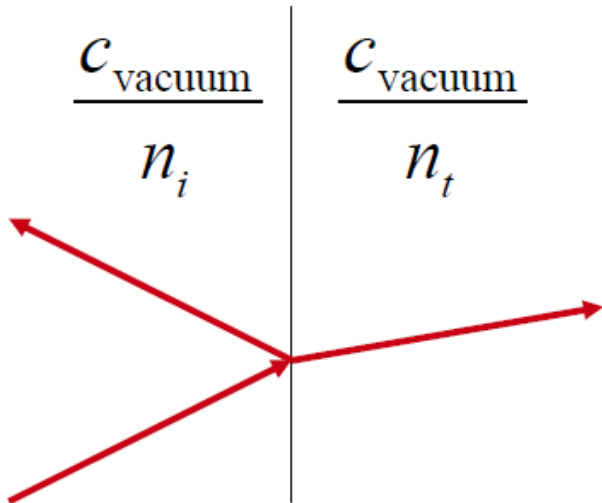
$$t_{\parallel} = \left( \frac{E_{0t}}{E_{0i}} \right)_{\parallel} = \frac{2n_i \cos \theta_t}{n_t \cos \theta_i + n_i \cos \theta_t}$$

# Reflection and Transmission of Energy @ dielectric interfaces

Recall Poynting vector definition:

$$\|\mathbf{S}\| = c \epsilon_0 \|\mathbf{E}\|^2$$

different on the two sides of the interface

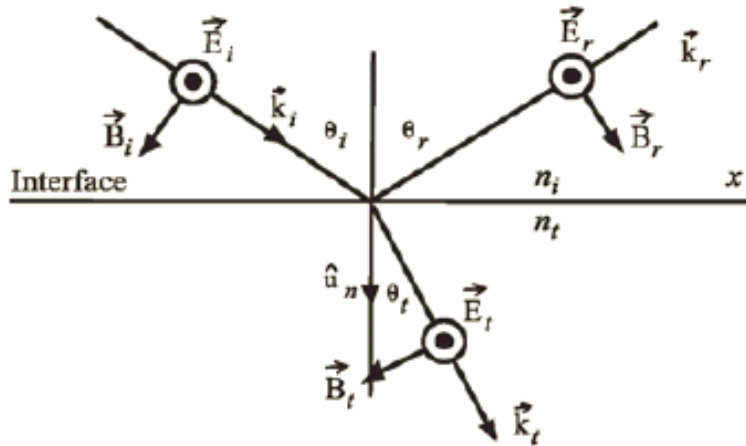


$$R = \left( \frac{E_{0r}}{E_{0i}} \right)^2 = r^2$$

$$T = \frac{n_t \cos \theta_t}{n_i \cos \theta_i} \left( \frac{E_{0t}}{E_{0i}} \right)^2 = \frac{n_t \cos \theta_t}{n_i \cos \theta_i} t^2$$

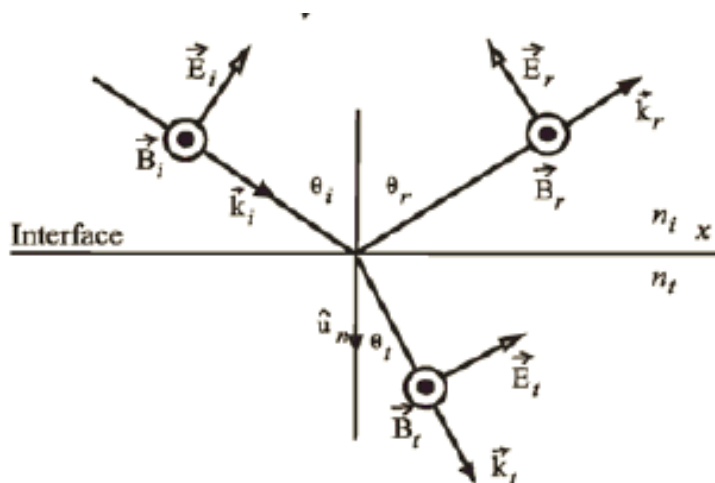
# Reflection and Transmission (Fresnel's equations)

Can be deduced from the application of boundary conditions of EM waves.



$$r_{\perp} = \left( \frac{E_{0r}}{E_{0i}} \right)_{\perp} = \frac{n_i \cos \theta_i - n_t \cos \theta_t}{n_i \cos \theta_i + n_t \cos \theta_t}$$

$$t_{\perp} = \left( \frac{E_{0t}}{E_{0i}} \right)_{\perp} = \frac{2n_i \cos \theta_i}{n_i \cos \theta_i + n_t \cos \theta_t}$$



$$r_{\parallel} = \left( \frac{E_{0r}}{E_{0i}} \right)_{\parallel} = \frac{n_t \cos \theta_i - n_i \cos \theta_t}{n_t \cos \theta_i + n_i \cos \theta_t}$$

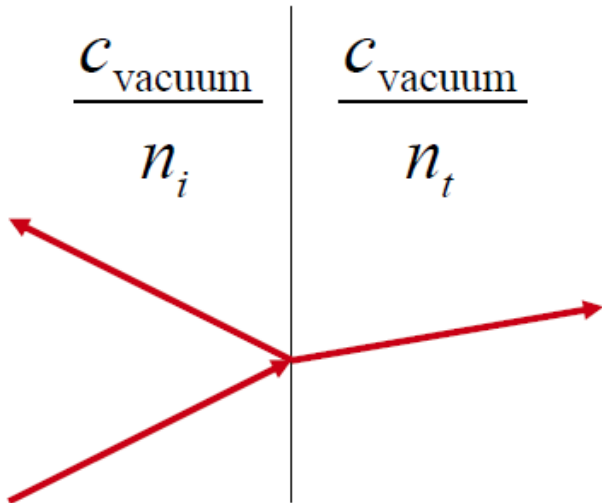
$$t_{\parallel} = \left( \frac{E_{0t}}{E_{0i}} \right)_{\parallel} = \frac{2n_i \cos \theta_t}{n_t \cos \theta_i + n_i \cos \theta_t}$$

# Reflection and Transmission of Energy @ dielectric interfaces

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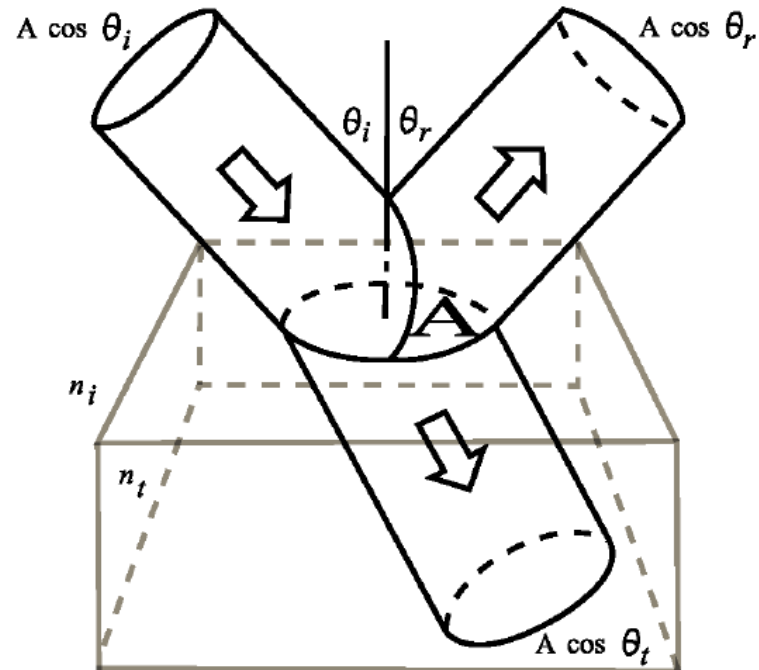


$$R = \left( \frac{E_{0r}}{E_{0i}} \right)^2 = r^2$$

$$T = \frac{n_t \cos \theta_t}{n_i \cos \theta_i} \left( \frac{E_{0t}}{E_{0i}} \right)^2 = \frac{n_t \cos \theta_t}{n_i \cos \theta_i} t^2$$

# Energy Conservation

$$R + T = 1, \text{ i.e. } r^2 + \frac{n_t \cos \theta_t}{n_i \cos \theta_i} t^2 = 1$$





# Normal Incidence

$$r_{\perp} = \left( \frac{E_{or}}{E_{oi}} \right)_{\perp} = \frac{n_i \cos \theta_i - n_t \cos \theta_t}{n_i \cos \theta_i + n_t \cos \theta_t}$$

$$t_{\perp} = \left( \frac{E_{ot}}{E_{oi}} \right)_{\perp} = \frac{2n_i \cos \theta_i}{n_i \cos \theta_i + n_t \cos \theta_t}$$

$$r_{\parallel} = \left( \frac{E_{or}}{E_{oi}} \right)_{\parallel} = \frac{n_t \cos \theta_i - n_i \cos \theta_t}{n_t \cos \theta_i + n_i \cos \theta_t}$$

$$t_{\parallel} = \left( \frac{E_{ot}}{E_{oi}} \right)_{\parallel} = \frac{2n_i \cos \theta_t}{n_t \cos \theta_i + n_i \cos \theta_t}$$

Note: independent of polarization

$$\theta_i = 0 \text{ and } \theta_t = 0$$



$$r_{\perp} = r_{\parallel} = \frac{n_t - n_i}{n_t + n_i}$$

$$t_{\perp} = t_{\parallel} = \frac{2n_i}{n_t + n_i}$$

$$R_{\perp} = R_{\parallel} = \left( \frac{n_t - n_i}{n_t + n_i} \right)^2$$

$$T_{\perp} = T_{\parallel} = \frac{4n_t n_i}{(n_t + n_i)^2}$$

# Reflectance and Transmittance @ dielectric interfaces

