Particle Dynamics in Special Relativity

In classical dynamics (Newton's
laws of motion) there is a
universal time coordinate t;
then
x(t) = particle position
u(t) = dx/dt = velocity

p = m u (momentum)
and dp/dt = F (Newton's second law)

The equations are written in terms of *3-vectors*; the time is universal.

In relativistic dynamics, we must follow the particle in the 4dimensional spacetime, $x^{\mu}(\tau) = 4$ -position, a function of proper time τ .

> Proper time (from the French "propre") is the time coordinate in a frame of reference (F") co-moving with the particle. $d\tau = dt$ "

Then we have $\eta^{\mu} = dx^{\mu}/d\tau = 4$ -velocity $p^{\mu} = m \eta^{\mu} = 4$ -momentum $dp^{\mu}/d\tau = K^{\mu}$ = the "Minkowski force", a 4-vector We must express the dynamical equations in terms of 4vectors, because that's the only way to ensure that the laws of physics (equations) are the same in all frames of reference.

The metric tensor $g_{\mu\nu} = diag(1,-1,-1,-1)$ $g_{00} = 1 ; g_{11} = g_{22} = g_{33} = -1$ $g_{\mu\nu} = 0 \text{ if } \mu \text{ not equal to } \nu$ **Proper time** (from the French "propre")

- τ is a scalar
- In an arbitrary frame of reference (F) $c (d\tau) = \sqrt{c^2(dt)^2 - (dx)^2 - (dy)^2 - (dz)^2}$
- or, $c^{2} (d\tau)^{2} = c^{2} (dt)^{2} - (dx)^{2} - (dy)^{2} - (dz)^{2}$ or, $c^{2} (d\tau)^{2} = g_{\mu\nu} dx^{\mu} dx^{\nu}$

(* the Einstein summation convention)

• $d\tau$ is a scalar! $d\tau' = d\tau$

I.e., $d\tau$ has the same value in all inertial frames. Transformation: $d\tau' = d\tau$

Theorem : $d\tau$ is a scalar.

Proof: Consider an arbitrary inertial frame *F*. We define proper time,

 $c^{2} (d\tau)^{2} = c^{2} (dt)^{2} - (dx)^{2} - (dy)^{2} - (dz)^{2}$

Now consider the *co-moving* frame *F"*.

The displacement 4-vector in this frame is

dt'', with dx''=0, dy''=0, dz''=0;

Apply the Lorentz transformation from **F**" to **F**:

(W.L.O.G., assume the relative velocity is v i)

$$c(dt) = \gamma \{ c(dt'') + (v/c) dx'' \} = \gamma c(dt'');$$

$$dx = \gamma \{ dx'' + v dt'' \} = \gamma v dt'';$$

$$dy = dy^{\prime\prime} = 0$$

$$dz = dz'' = 0$$

Thus

 $c^{2} (d\tau)^{2} = \gamma^{2} \{ c^{2} - v^{2} \} (dt'')^{2}$ But the right-hand side is just $c^{2} (dt'')^{2}$. So $d\tau = dt''$ for any inertial frame **F**. Q.E.D.

Scalars and Vectors

- $d\tau$ is a scalar (i.e., invariant)
- x^{μ} is a vector (by definition)
- anything that transforms in the same way as x^μ is a vector (by definition)
- therefore dx^{μ} is a vector; it's just $x_2^{\mu} x_1^{\mu}$
- $\eta^{\mu} = dx^{\mu}/d\tau$ is a vector
- $p^{\mu} = m \eta^{\mu}$ is a vector (m is the rest mass, invariant)
- $dp^{\mu}/d\tau$ is a vector
- So if we have a dynamical equation $dp^{\mu}/d\tau = K^{\mu}$, then K^{μ} must be a vector (the Minkowski force)

Energy and momentum of a particle

Exercise. Express these 4-vectors in terms of the particle 3-velocity **u**: η^{μ} and p^{μ} Define particle velocity **u** as it would be measured by an observer in frame *F*: **u** = d**x** /dt (a 3-vector)

Now recall the 4-vector position: $\boldsymbol{X}^{\boldsymbol{\mu}}$

dτ

$$c^{2} (d\tau)^{2} = c^{2}(dt)^{2} - (dx)^{2} - (dy)^{2} - (dz)^{2}$$

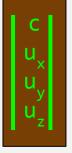
$$c^{2} (d\tau)^{2} = c^{2}(dt)^{2} - u_{x}^{2}(dt)^{2} - u_{y}^{2}(dt)^{2} - u_{z}^{2}(dt)^{2}$$

$$c^{2} (d\tau)^{2} = (c^{2} - u^{2})(dt)^{2}$$

$$d\tau = dt (1 - u^{2}/c^{2})^{1/2} \text{ (the time dilation formula)}$$

$$\begin{split} \mathbf{\eta}^{\underline{\mu}} &= d\mathbf{x}^{\underline{\mu}}/d\tau \\ \eta^{\mu} &= \{\eta^{0}, \, \eta\} = \{c \, dt/d\tau, \, d\mathbf{x}/d\tau\} \\ &= \{c \, /(1-u^{2}/c^{2})^{\frac{1}{2}}, \, \mathbf{u} \, /(1-u^{2}/c^{2})^{\frac{1}{2}}\} \end{split}$$

 η^{μ} = 1/(1-u^2/c^2)^½



$$\underline{p}^{\underline{\mu}} = \underline{m} \ \underline{\eta}^{\underline{\mu}}$$

$$p^{\mu} = \{p^{0}, p\}$$

 $p^{0} = mc / (1-u^{2}/c^{2})^{\frac{1}{2}} = particle energy / c$ $p = mu / (1-u^{2}/c^{2})^{\frac{1}{2}} = particle momentum$

4- morahim momentum In classical physics, Ip = mu = m dx dt In relativity we use 4-monation $p^{\mathcal{M}} = m \frac{dx^{\mathcal{A}}}{d\tau} = \begin{pmatrix} p^{*} \\ p_{x} \\ p_{y} \end{pmatrix}$ To = mu and p = mc To and p = mc to particle particle monochum energy/c $E = cp^{\circ} = \frac{mc^{2}}{\sqrt{1-u^{2}c^{2}}} particle energy$ For velocity ie = 0, E = mc2.

Pynamics - the Minkowski force de 4 = K (which comes from some de Luteraction there) · Space components An observer defines the force on a particle by $d\vec{p}/dt = \vec{F}$. (def. of \vec{F}) Non R = de de de - F If F=0 then p= anstant. . The time composent $K^{\circ} = \frac{dp^{\circ}}{d\tau} = \frac{dp^{\circ}}{d\tau} \frac{dt}{d\tau} = \frac{dp^{\circ}}{VI - u_{TC}^{2}} \frac{dp^{\circ}}{d\tau}$ and po = mc .

 $K^{0} = \frac{mc}{\sqrt{1-u^{2}/c^{2}}} \left(\frac{-1}{2} - \frac{2u}{(1-u^{2}/c^{2})} \frac{du/dt}{\sqrt{1-u^{2}/c^{2}}} - \frac{mcu}{(1-u^{2}/c^{2})^{2}} \frac{du/dt}{(2-u^{2}/c^{2})^{2}} - \frac{mcu}{(1-u^{2}/c^{2})^{2}} \frac{du/dt}{(2-u^{2}/c^{2})^{2}} + \frac{mcu}{(1-u^{2}/c^{2})^{2}} \frac{du/dt}{(2-u^{2}/c^{2})^{2}} + \frac{mcu}{(1-u^{2}/c^{2})^{2}} + \frac{mcu}{(1-u^{2}/c^{2})^{2}$ Compete : $\vec{F} = \frac{d\vec{k}}{dt} = \frac{d}{dt} \left(\frac{m\vec{u}}{N-u^{2}c^{2}} \right) = m \left\{ \frac{d\vec{u}/dt}{\sqrt{1-u^{2}c^{2}}} + \frac{\vec{u}(\vec{u}\cdot d\vec{u}/dt)}{(1-u^{2}c^{2})^{2/2}c^{2}} \right\}$ Thus U.F - m <u>U. (Lu/dF)</u> (1-u²/c²)^{8/2} Result K° = U.F. [Everyon de = K° 11-13/02 = U.F = Work/hime To be consistent with the conservation of every p_{μ}^{D} subst le E/c; $E = Cp^{\circ} = \frac{MC^{2}}{\sqrt{1-VETC^{2}}}$.