#### **Chapter 6 - Radiative Corrections**

There are two kinds of radiative corrections.

<mark>Virtual corrections:</mark> Feynman diagrams have loops.

Real corrections:

Feynman diagrams have additional final state particles, i.e, photons.

We will see that *these two kinds of corrections cannot be separated.* At first sight this seems very surprising. But on further thought, it's not so surprising. **Example**. The lowest order corrections to electron scattering. Tree diagram (LO of perturbation theory)



Virtual corrections (NLO  $\rightarrow$  loops)

**Real radiation** 



We'll see that the real and virtual effects must be combined.

#### Divergences

There are two kinds of divergences (actually three[†] but we'll only consider two)

# <mark>UV divergences</mark>

The integral over an internal loop momentum qµ may be divergent (i.e., infinite) due to the contribution from large q<sup>µ</sup>. For example,

 $\int \frac{d\tau_g}{(q^2+m^2)^2} \sim \int \frac{dq}{q}$ 

"Miraculously", the UV divergences cancel out after <u>RENORMALIZATION</u>.

[**†** the third kind are *collinear divergences*]

#### IR divergences

(i) The integral over an internal loop momentum qμ may be divergent (i.e., infinite) due to the contribution from small q<sup>μ</sup>. For example,

$$\int \frac{d^4g}{g^2 \left(q^2 + m^2\right)^2}$$

(ii) Also, the integral over final states in a real radiative process, e.g., Bremsstrahlung, may be divergent due to the contribution from small q<sup>µ</sup>.

"Miraculously", the IR divergences cancel out after COMBINING REAL AND VIRTUAL CORRECTIONS.

## SECTION 6.1. Bremsstrahlung of soft photons.

What is a <u>soft photon?</u> A soft photon is a low-energy photon. now, how low is low? that depends on the process. a photon that is emitted in a scattering process with characteristic momentum/energy scale Q is considered to be a soft photon if  $\omega \ll Q$ .

#### CLASSICAL CALCULATION

For details, see P.&Sch.

The electron trajectory is

 $p_{/}$ 

$$j^\mu(x)=e{\int}d au\,{dy^\mu( au)\over d au}\,\,\delta^{(4)}ig(x-y( au)ig).$$

$$y^\mu( au) = egin{cases} (p^\mu/m) au & ext{for } au < 0; \ (p'^\mu/m) au & ext{for } au > 0. \end{cases}$$

Use *classical* electrodynamics to calculate the vector potential

$$\widetilde{A}^{\mu}(k) = -rac{1}{k^2}\,\widetilde{\jmath}^{\mu}(k).$$

$$A^{\mu}(x) = \int \frac{d^4k}{(2\pi)^4} e^{-ik \cdot x} \frac{-ie}{k^2} \Big( \frac{p'^{\mu}}{k \cdot p' + i\epsilon} - \frac{p^{\mu}}{k \cdot p - i\epsilon} \Big).$$

And use the vector potential to calculate the radiation fields (**E** and **B**).

$$oldsymbol{\mathcal{E}}(\mathbf{k}) = -i\mathbf{k}\mathcal{A}^0(\mathbf{k}) + ik^0\mathcal{A}(\mathbf{k});$$
  
 $oldsymbol{\mathcal{B}}(\mathbf{k}) = i\mathbf{k}\times\mathcal{A}(\mathbf{k}) = \hat{k}\timesoldsymbol{\mathcal{E}}(\mathbf{k}).$ 

Energy = 
$$\frac{1}{2} \int d^3x \left( |\mathbf{E}(x)|^2 + |\mathbf{B}(x)|^2 \right).$$

Using the explicit form of  $\mathcal{A}(\mathbf{k})$  (6.7), we finally arrive at an expression for the energy radiated\*:

Energy = 
$$\int \frac{d^3k}{(2\pi)^3} \sum_{\lambda=1,2} \left. \frac{e^2}{2} \right| \boldsymbol{\epsilon}_{\lambda}(\mathbf{k}) \cdot \left( \frac{\mathbf{p}'}{k \cdot p'} - \frac{\mathbf{p}}{k \cdot p} \right) \right|^2.$$
(6.12)

$$\begin{split} & \operatorname{Energy} = \frac{e^2}{(2\pi)^2} \int \! dk \, \mathcal{I}(\mathbf{v}, \mathbf{v}'), \\ & \mathcal{I}(\mathbf{v}, \mathbf{v}') = \int \! \frac{d\Omega_{\hat{k}}}{4\pi} \left( \frac{2(1\!-\!\mathbf{v}\cdot\mathbf{v}')}{(1\!-\!\hat{k}\cdot\mathbf{v})(1\!-\!\hat{k}\cdot\mathbf{v}')} - \frac{m^2/E^2}{(1\!-\!\hat{k}\cdot\mathbf{v}')^2} - \frac{m^2/E^2}{(1\!-\!\hat{k}\cdot\mathbf{v})^2} \right). \\ & \mathcal{I}(\mathbf{v}, \mathbf{v}') \approx \log \Bigl( \frac{1-\mathbf{v}'\cdot\mathbf{v}}{1-|\mathbf{v}|} \Bigr) + \log\Bigl( \frac{1-\mathbf{v}'\cdot\mathbf{v}}{1-|\mathbf{v}'|} \Bigr) = \log\Bigl( \frac{(E^2-\mathbf{p}\cdot\mathbf{p}')^2}{E^2(E-|\mathbf{p}|)^2} \Bigr) \\ & \approx 2\log\Bigl( \frac{p\cdot p'}{(E^2-|\mathbf{p}|^2)/2} \Bigr) = 2\log\Bigl( \frac{-q^2}{m^2} \Bigr), \\ & \text{where } q^2 = (p'-p)^2. \end{split}$$

## Result of the classical calculation

In conclusion, we have found that the radiated energy at low frequencies is given by

Energy = 
$$\frac{\alpha}{\pi} \int_{0}^{k_{\max}} dk \ \mathcal{I}(\mathbf{v}, \mathbf{v}') \xrightarrow[E\ggm]{} \frac{2\alpha}{\pi} \int_{0}^{k_{\max}} dk \ \log\left(\frac{-q^2}{m^2}\right).$$
 (6.18)

If this energy is made up of photons, each photon contributes energy k. We would then expect

Number of photons = 
$$\frac{\alpha}{\pi} \int_{0}^{k_{\text{max}}} dk \frac{1}{k} \mathcal{I}(\mathbf{v}, \mathbf{v}').$$
 SC (6.19)

We hope that a quantum-mechanical calculation will confirm this result.

Of course there are no photons in classical electromagnetism. We need a QCD calculation to find the number of photons.

#### Q.E.D. CALCULATION

For single photon emission, there are two Feynman diagrams in lowest order.

Now we'll make lots of approximations, which are valid for soft photons; i.e. the asymptotic behavior as  $\omega \rightarrow 0$ . (Not strictly 0! But small compared to the momentum transfer that occurs in the scattering process.)

#### Quantum Computation

Consider now the quantum-mechanical process in which one photon is radiated during the scattering of an electron:



Let  $\mathcal{M}_0$  denote the part of the amplitude that comes from the electron's interaction with the external field. Then the amplitude for the whole process is

$$i\mathcal{M} = -ie\bar{u}(p') \left( \mathcal{M}_0(p', p-k) \frac{i(\not\!\!p-k\!\!\!/+m)}{(p-k)^2 - m^2} \gamma^{\mu} \epsilon^*_{\mu}(k) + \gamma^{\mu} \epsilon^*_{\mu}(k) \frac{i(\not\!\!p'+k\!\!\!/+m)}{(p'+k)^2 - m^2} \mathcal{M}_0(p'+k,p) \right) u(p).$$
(6.20)

$$\mathcal{M}_0(p', p-k) \approx \mathcal{M}_0(p'+k, p) \approx \mathcal{M}_0(p', p),$$

and we can ignore k in the numerators of the propagators. The numerators can be further simplified with some Dirac algebra. In the first term we have

$$(\not p + m)\gamma^{\mu}\epsilon^{*}_{\mu} u(p) = \left[2p^{\mu}\epsilon^{*}_{\mu} + \gamma^{\mu}\epsilon^{*}_{\mu}(-\not p + m)\right]u(p)$$
$$= 2p^{\mu}\epsilon^{*}_{\mu} u(p).$$

Similarly, in the second term,

$$\bar{u}(p')\,\gamma^{\mu}\epsilon^*_{\mu}(\not\!\!p'+m) = \bar{u}(p')\,2p'^{\mu}\epsilon^*_{\mu}.$$

The denominators of the propagators also simplify:

$$(p-k)^2 - m^2 = -2p \cdot k;$$
  $(p'+k)^2 - m^2 = 2p' \cdot k.$ 

So in the soft-photon approximation, the amplitude becomes

$$i\mathcal{M} = \bar{u}(p') \left[ \mathcal{M}_0(p', p) \right] u(p) \cdot \left[ e \left( \frac{p' \cdot \epsilon^*}{p' \cdot k} - \frac{p \cdot \epsilon^*}{p \cdot k} \right) \right].$$
(6.22)

This is just the amplitude for elastic scattering (without bremsstrahlung), times a factor (in brackets) for the emission of the photon.

$$d\sigma(p \to p' + \gamma) = d\sigma(p \to p') \cdot \int \frac{d^3k}{(2\pi)^3} \frac{1}{2k} \sum_{\lambda=1,2} e^2 \left| \frac{p' \cdot \epsilon^{(\lambda)}}{p' \cdot k} - \frac{p \cdot \epsilon^{(\lambda)}}{p \cdot k} \right|^2.$$
(6.23)

$$d(\mathrm{prob}) = rac{d^3k}{(2\pi)^3} \sum_{\lambda} rac{e^2}{2k} \bigg| \, oldsymbol{\epsilon}_{\lambda} \cdot \Big( rac{\mathbf{p}'}{p' \cdot k} - rac{\mathbf{p}}{p \cdot k} \Big) \bigg|^2.$$

Total probability 
$$\approx \frac{\alpha}{\pi} \int_{0}^{|\mathbf{q}|} dk \frac{1}{k} \mathcal{I}(\mathbf{v}, \mathbf{v}').$$

## An infrared divergence

We can artificially make the integral in (6.25) well-defined by pretending that the photon has a very small mass  $\mu$ . This mass would then provide a lower cutoff for the integral, allowing us to write the result of this section as

$$d\sigma(p \to p' + \gamma(k)) = d\sigma(p \to p') \cdot \frac{\alpha}{2\pi} \log\left(\frac{-q^2}{\mu^2}\right) \mathcal{I}(\mathbf{v}, \mathbf{v}')$$

$$\approx d\sigma(p \to p') \cdot \frac{\alpha}{\pi} \log\left(\frac{-q^2}{\mu^2}\right) \log\left(\frac{-q^2}{m^2}\right).$$
(6.26)

# Result of the quantum calculation

Does this agree with the semiclassical result (SC) derived earlier?

The two results are sort of similar mathematically; but they are actually very different physically.

In Section 6.5 we'll see that the semiclassical result is correct, when we add all the higher order contributions to soft photon bremsstrahlung.

## Section 6.2. The electron vertex function



$$i\mathcal{M} = ie^{2} \left( \bar{u}(p') \Gamma^{\mu}(p', p) u(p) \right) \frac{1}{q^{2}} \left( \bar{u}(k') \gamma_{\mu} u(k) \right).$$

Eventually we'll calculate the lowest order vertex correction. But first let's understand some general features of the correction, valid to all orders of perturbation theory.

$$i\mathcal{M}(2\pi)\delta(p^{0\prime}-p^{0}) = -ie\bar{u}(p')\gamma^{\mu}u(p)\cdot\widetilde{A}^{\mathrm{cl}}_{\mu}(p'-p),$$

$$i\mathcal{M}(2\pi)\delta(p^{0\prime}-p^0)=-iear{u}(p^{\prime})\,\Gamma^{\mu}(p^{\prime},p)\,u(p)\cdot\widetilde{A}^{\mathrm{cl}}_{\mu}(p^{\prime}-p).$$

$$\Gamma^{\mu}(p',p) = \gamma^{\mu}F_1(q^2) + rac{i\sigma^{\mu
u}q_{
u}}{2m}F_2(q^2),$$

**Electric and Magnetic Form Factors** 

limit  $\mathbf{q} \to 0$  in the spinor matrix element. Only the form factor  $F_1$  contributes. Using the nonrelativistic limit of the spinors,

$$\bar{u}(p')\gamma^0 u(p) = u^{\dagger}(p')u(p) \approx 2m\xi'^{\dagger}\xi,$$

the amplitude for electron scattering from an electric field takes the form

$$i\mathcal{M} = -ieF_1(0)\tilde{\phi}(\mathbf{q}) \cdot 2m\xi^{\dagger}\xi.$$
(6.34)

This is the Born approximation for scattering from a potential

$$V(\mathbf{x}) = eF_1(0)\phi(\mathbf{x}).$$

Thus  $F_1(0)$  is the electric charge of the electron, in units of e. Since  $F_1(0) = 1$  already in the leading order of perturbation theory, radiative corrections to  $F_1(q^2)$  should vanish at  $q^2 = 0$ .

Again we can interpret  $\mathcal{M}$  as the Born approximation to the scattering of the electron from a potential well. The potential is just that of a magnetic moment interaction,

$$V(\mathbf{x}) = -\langle \boldsymbol{\mu} 
angle \cdot \mathbf{B}(\mathbf{x}),$$

where

$$\langle \boldsymbol{\mu} \rangle = \frac{e}{m} \big[ F_1(0) + F_2(0) \big] \xi^{\prime \dagger} \frac{\boldsymbol{\sigma}}{2} \xi.$$

This expression for the magnetic moment of the electron can be rewritten in the standard form

$$oldsymbol{\mu} = g\Big(rac{e}{2m}\Big) \mathbf{S},$$

where S is the electron spin. The coefficient g, called the Landé g-factor, is

$$g = 2[F_1(0) + F_2(0)] = 2 + 2F_2(0).$$
(6.37)

Since the leading order of perturbation theory gives no  $F_2$  term, QED predicts  $g = 2 + \mathcal{O}(\alpha)$ . The leading term is the standard prediction of the Dirac equation. In higher orders, however, we will find a nonzero  $F_2$  and thus a small difference between the electron's magnetic moment and the Dirac value. We will compute the order- $\alpha$  contribution to this anomalous magnetic moment in the next section.

Section 6.3. Evaluate the lowest order vertex correction.