

## LAGRANGIAN FIELD THEORY AND CANONICAL QUANTIZATION (CHAPTER 2)

In the history of science, the first field theory was electromagnetism. (Maxwell) ●

There are 2 vector fields,  $\mathbf{E}$  and  $\mathbf{B}$ .

In spacetime we have a field tensor.

$$F^{\mu\nu} = \begin{pmatrix} 0 & -E_x & -E_y & -E_z \\ E_x & 0 & B_z & -B_y \\ E_y & -B_z & 0 & B_x \\ E_z & B_y & -B_x & 0 \end{pmatrix}$$

$$\text{Or, } F^{\mu\nu} = \partial^\mu A^\nu - \partial^\nu A^\mu$$

$$\text{where } \partial^\mu = g^{\mu\nu} \partial_{x^\nu}$$

■ The *classical field theory* describes electromagnetic waves with  $\omega = ck$ .

■ The *quantum field theory* describes photons. (Chapter 1)

■ We can derive the theory from a Lagrangian, and then quantize it. But there are some subtleties, due to gauge invariance! (Chapter 5)

Electromagnetism isn't very interesting without sources, i.e., *charges*.

■ We'll add the electron field in PHY 955. That's **Q**uantum **E**lectro**D**ynamics. (Ch. 7)

Recall the example of the Schroedinger equation

Classical field theory:  $\psi(\mathbf{x},t)$  is a complex function.

$$A = \int_{t_1}^{t_2} dt \int d^3x \left\{ \frac{-i\hbar}{2} \left( \frac{\partial \psi^*}{\partial t} \psi - \psi^* \frac{\partial \psi}{\partial t} \right) - \frac{\hbar^2}{2m} \nabla \psi^* \cdot \nabla \psi - V \psi^* \psi \right\}$$

$$i\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \nabla^2 \psi + V \psi$$

Quantum:  $\psi(\mathbf{x},t)$  is a non-hermitian operator.

$$[\psi(x), \psi^\dagger(x')] = \delta^3(x-x')$$

$$[\psi(x), \psi(x')] = 0$$

Now another example: (SECTION 2.2 - 2.3)

A REAL SCALAR FIELD  $\phi = \phi(\mathbf{x},t)$

This example is relativistically covariant.

$$\mathcal{L} = \frac{1}{2} \left( \frac{\partial \phi}{\partial t} \right)^2 - \frac{c^2}{2} (\nabla \phi)^2 - \frac{1}{2} \left( \frac{mc^2}{\hbar} \right)^2 \phi^2$$

Derive the field equation from Hamilton's principle,

$$A = \int \left\{ \frac{1}{2} \dot{\phi}^2 - \frac{c^2}{2} (\nabla \phi)^2 - \frac{1}{2} \left( \frac{mc^2}{\hbar} \right)^2 \phi^2 \right\} d^3x dt$$

$$\delta A = \int \left\{ \dot{\phi} (\delta \dot{\phi}) - c^2 \nabla \phi \cdot \nabla (\delta \phi) - \left( \frac{mc^2}{\hbar} \right)^2 \phi \delta \phi \right\} d^3x dt$$

$$= \int \delta \phi \left\{ -\ddot{\phi} + c^2 \nabla^2 \phi - \left( \frac{mc^2}{\hbar} \right)^2 \phi \right\} d^3x dt$$

= 0 for any variation  $\delta \phi$ . Therefore

$$\ddot{\phi} - c^2 \nabla^2 \phi + \left( \frac{mc^2}{\hbar} \right)^2 \phi = 0$$

the Klein Gordon equation

We can solve the Klein-Gordon equation, in plane waves,

$$\phi(\vec{x}, t) = C e^{i(\vec{k} \cdot \vec{x} - \omega t)}$$

where

$$-\omega^2 + c^2 k^2 + \left(\frac{mc^2}{\hbar}\right)^2 = 0$$

$$\omega = \pm \sqrt{c^2 k^2 + m^2 c^4 / \hbar^2}$$

I.e.,

$$\hbar \omega = \pm \sqrt{c^2 \hbar^2 k^2 + m^2 c^4}$$

Note that this is the energy ( $\hbar\omega$ ) and momentum ( $\hbar\mathbf{k}$ ) relation of special relativity.

**(What are the negative energy solutions?)**

The general solution (Hermitian) is

$$\phi(\vec{x}, t) = \sum_{\vec{k}} N \left\{ e^{i(\vec{k} \cdot \vec{x} - \omega t)} a_{\vec{k}} + e^{-i(\vec{k} \cdot \vec{x} - \omega t)} a_{\vec{k}}^{\dagger} \right\}$$

### Quantization

We can anticipate

$$[a_{\vec{k}}, a_{\vec{k}'}^{\dagger}] = \delta_{\vec{K}}(\vec{k}, \vec{k}')$$

$$[a_{\vec{k}}, a_{\vec{k}'}] = 0$$

Derive this from Dirac's canonical quantization. Recall,

$$[q, p] = i \hbar \quad \text{where} \quad p = \partial L / \partial q'$$

$$\Pi(\vec{x}) = \frac{\delta L}{\delta \dot{\phi}(\vec{x})} = \frac{\partial \mathcal{L}}{\partial \dot{\phi}(\vec{x})} = \dot{\phi}(\vec{x})$$

The E.T.C.R. should be

$$[\phi(\vec{x}, t), \Pi(\vec{x}', t)] = i\hbar \delta^3(\vec{x} - \vec{x}')$$

$$\phi(\vec{x}, t) = \sum_{\vec{k}} N \{ a_{\vec{k}} e^{i(\vec{k}\cdot\vec{x} - \omega t)} + a_{\vec{k}}^\dagger e^{-i(\vec{k}\cdot\vec{x} - \omega t)} \}$$

$$\Pi(\vec{x}, t) = \sum_{\vec{k}} N(-i\omega) \{ a_{\vec{k}} e^{i(\vec{k}\cdot\vec{x} - \omega t)} - a_{\vec{k}}^\dagger e^{-i(\vec{k}\cdot\vec{x} - \omega t)} \}$$

$$\begin{aligned} [\phi(\vec{x}), \Pi(\vec{x}')] &= \sum_{\vec{k}} \sum_{\vec{k}'} NN'(-i\omega') \\ & [a_{\vec{k}} e^{i\vec{k}\cdot\vec{x}} + a_{\vec{k}}^\dagger e^{-i\vec{k}\cdot\vec{x}}, a_{\vec{k}'} e^{i\vec{k}'\cdot\vec{x}'} - a_{\vec{k}'}^\dagger e^{-i\vec{k}'\cdot\vec{x}'}] \\ &= \sum_{\vec{k}} \sum_{\vec{k}'} NN'(-i\omega') \{ -\delta(\vec{k}, \vec{k}') e^{i\vec{k}\cdot(\vec{x}-\vec{x}')} \\ & \quad - \delta(\vec{k}, -\vec{k}') e^{-i\vec{k}\cdot(\vec{x}-\vec{x}')} \} \\ &= \sum_{\vec{k}} N^2(i\omega) \{ e^{i\vec{k}\cdot(\vec{x}-\vec{x}')} + e^{-i\vec{k}\cdot(\vec{x}-\vec{x}')} \} \\ &= \text{should be } i\hbar \delta^3(\vec{x}-\vec{x}') = \frac{i\hbar}{(2\pi)^3} \int d^3k e^{i\vec{k}\cdot(\vec{x}-\vec{x}')} \end{aligned}$$

$$\text{Infinite volume limit } \sum_{\vec{k}} = \frac{\Omega}{(2\pi)^3} \int d^3k$$

$$= i\hbar \delta^3(\vec{x}-\vec{x}') \text{ provided that } N^2 \omega 2\Omega = \hbar$$

$$N = \sqrt{\frac{\hbar}{2\omega\Omega}}$$

## The Hamiltonian

$$H = p\dot{q} - L$$

rewritten in terms of  $q, p$

$$H = \int \Pi(\vec{x}) \dot{\phi}(\vec{x}) d^3x - L$$

rewritten in terms of  $\phi(\vec{x}), \Pi(\vec{x})$

### Homework problem.

(A) Write H in terms of  $\Pi(\vec{x})$  and  $\phi(\vec{x})$ .

(B) Write H in terms of  $a_{\vec{k}}$  and  $a_{\vec{k}}^\dagger$ .

### Homework problem.

Determine the Green's function for the free scalar field;  $\langle 0 | T \phi(\vec{x}) \phi(\vec{y}) | 0 \rangle$ .

Next: A real scalar field  $\phi$  with a source  $\rho$ .

To make it simpler, set  $\hbar = 1$  and  $c = 1$ . (natural units)  
At the end of a calculation we can restore the factors of  $\hbar$  and  $c$  by dimensional analysis (i.e., simple units analysis).

$$\mathcal{L} = \frac{1}{2} \left( \frac{\partial \phi}{\partial t} \right)^2 - \frac{1}{2} (\nabla \phi)^2 - \frac{1}{2} m^2 \phi^2 + \rho \phi$$

$$\text{Field equation } \frac{\partial}{\partial t} \left( \frac{\delta \mathcal{L}}{\delta \dot{\phi}} \right) - \frac{\delta \mathcal{L}}{\delta \phi} = 0$$

$$\hookrightarrow \ddot{\phi} - \nabla^2 \phi + m^2 \phi - \rho = 0$$

The field equation is a linear inhomogeneous equation; so  $\phi(x,t) = \phi_{\text{particular}}(x,t) + \phi_{\text{homogeneous}}(x,t)$ .

The particular solution comes from the source; e.g., it could be a mean field produced by a static source; or, waves radiated by a time dependent source.

The homogeneous solution consists of harmonic waves.

### The particular solution for a static source

Consider  $\rho = \rho(\mathbf{x})$ , independent of  $t$

$$-\nabla^2 \phi_0 + m^2 \phi_0 = \rho(\vec{x})$$

We need the Green's function of  $-\nabla^2 + m^2$ ; i.e.,

$$(-\nabla^2 + m^2) G(\vec{x} - \vec{y}) = \delta^3(\vec{x} - \vec{y})$$

Then

$$\phi_0(\vec{x}) = \int G(\vec{x} - \vec{y}) \rho(\vec{y}) d^3y$$

## The Green's function of $-\nabla^2 + m^2$

$$(-\nabla^2 + m^2) G(\vec{x}) = \delta^3(\vec{x})$$

$$G(\vec{x}) = \int \frac{d^3k}{(2\pi)^3} \frac{e^{i\vec{k}\cdot\vec{x}}}{k^2 + m^2}$$

$$= \frac{1}{(2\pi)^2} \cdot 2\pi \int_0^\infty \frac{k^2 dk}{k^2 + m^2} \underbrace{\int_{-1}^1 d\cos\theta e^{ik\zeta \cos\theta}}_{\frac{1}{ik\zeta} (e^{ik\zeta} - e^{-ik\zeta})}$$

$$= \frac{1}{4\pi^2 i\zeta} \int_{-\infty}^{\infty} \frac{k dk}{(k-im)(k+im)} e^{ik\zeta}$$

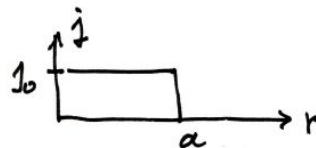
$\zeta > 0$  so close the contour above

$$= \frac{1}{4\pi^2 i\zeta} \frac{2\pi i (im)}{2im} e^{-m\zeta}$$

$$= \frac{e^{-m|\vec{x}|}}{4\pi|\vec{x}|} \quad (\text{w/ } \hbar=1 \text{ and } c=1)$$

## Example

Suppose  $\rho(x) = \rho_0 \theta(a - r)$ .



$$\phi_0(\vec{x}) = \int \frac{e^{-m|\vec{x}-\vec{y}|}}{4\pi|\vec{x}-\vec{y}|} \rho(\vec{y}) d^3y$$

$$\phi_0(r) = \frac{\rho_0}{4\pi} \int \frac{e^{-m|\vec{x}-\vec{y}|}}{|\vec{x}-\vec{y}|} \theta(a-y) d^3y$$

Limiting cases —

- Large  $r$

$$\phi_0(r) \sim \frac{\rho_0}{4\pi} \frac{e^{-mr}}{r} \cdot \frac{4}{3}\pi a^3$$

- Small  $r$

$$\phi_0(r) \sim \frac{\rho_0}{4\pi} \int \frac{e^{-my}}{y} \theta(a-y) d^3y$$

$$= \frac{\rho_0}{m^2} \left\{ 1 - (1+ma)e^{-ma} \right\} \quad \text{Still } \hbar=1 \text{ and } c=1.$$

The interaction Lagrangian density

$$\mathcal{L}_{\text{interaction}} = g \Psi_{\alpha\rho}^{\dagger} \Psi_{\alpha\rho} \phi$$

- This  $\mathcal{L}_{\text{int}}$  acts as a source for  $\phi$ , with

$$\rho(\mathbf{x}, t) = g \Psi_{\alpha\rho}^{\dagger} \Psi_{\alpha\rho}.$$

- It also acts as a potential for  $\Psi$ :

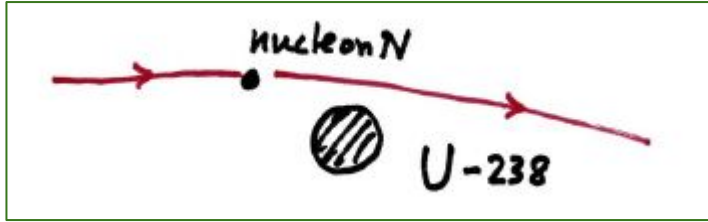
$$V_{\text{int}}(\mathbf{x}, t) = -g \phi(\mathbf{x}, t).$$

∴ The field equations;  
i.e., Lagrange's equations,

$$-\frac{\hbar^2}{2m} \nabla^2 \psi + (-g\phi) \psi = i\hbar \frac{\partial \psi}{\partial t}$$

$$\frac{\partial^2 \phi}{\partial t^2} - \nabla^2 \phi + m^2 \phi = g \psi^{\dagger} \psi$$

Calculate the potential energy for a nucleon (N) attracted to a heavy isotope (Z,A)



First step -- calculate the mean field created by the nucleons in the heavy isotope.

$$-\nabla^2 \phi + m^2 \phi = \langle g \psi^\dagger \psi \rangle$$

$$\langle \psi^\dagger \psi \rangle_{U-238} = \text{density of nucleons}$$

$$= \sum_{n=1}^{238} |\psi_n(\vec{r})|^2$$

$$\approx \frac{A}{\frac{4}{3}\pi R^3} \theta(R-r) \text{ where } R = r_0 A^{1/3}$$

$$\phi(\vec{r}) = \int G(\vec{r}-\vec{y}) n(\vec{y}) d^3y$$

Second step -- calculate the potential energy for the presence of the extra nucleon.

$$V(\vec{x}) = -g \phi = -g \int G(\vec{x}-\vec{y}) n(\vec{y}) d^3y$$

$$V(\vec{x}) = \frac{-3g}{4\pi r_0^3 A} \int \frac{e^{-m|\vec{x}-\vec{y}|}}{4\pi|\vec{x}-\vec{y}|} \theta(r_0 A^{1/3} - |\vec{y}|) d^3y$$

$$\hbar = 1 \text{ and } c = 1.$$

Rewrite this for numerical calculation...



## Yukawa's theory of the nucleon-nucleon force (1935)

- (1) Nucleons interact through a scalar field  $\phi$  with mass  $m$ .
- (2) The range of the force is

$$\text{range} = \frac{\hbar}{mc} = 1 \text{ to } 2 \text{ fm}$$

$$\therefore mc^2 = \frac{\hbar c}{\text{range}} = 100 \text{ to } 200 \text{ MeV}$$

Of course Yukawa did not know about pions, which were discovered in 1947.

$$\text{mass}(\pi^\pm) = 139.6 \text{ MeV}/c^2$$

$$\text{mass}(\pi^0) = 135.0 \text{ MeV}/c^2$$

The Lagrangian density for the theory is

$$\mathcal{L} = \mathcal{L}_{\text{nucleon}} + \mathcal{L}_{\text{meson}} + \mathcal{L}_{\text{interaction}}$$

$$\text{Nucleon field} = \psi_{\alpha\rho}(\vec{x})$$

$\alpha = \text{spin index}$  and  $\rho = \text{isospin index}$

$$\psi_{\alpha\rho}(\vec{x}) = \sum_{k,s,t} \frac{1}{\sqrt{V}} e^{i\vec{k}\cdot\vec{x}} \eta_{\alpha}^{(s)} \eta_{\rho}^{(t)} a_{kst}$$

$$\text{Meson field} = \phi(\vec{x})$$

with isospin 0 to follow Yukawa

Lagrange's equations including the interaction,

$$\mathcal{L}_{\text{interaction}} = g \psi^\dagger_{\alpha\rho} \psi_{\alpha\rho} \phi$$

$$\square \square \quad -\frac{\hbar^2}{2m} \nabla^2 \psi + (-g\phi) \psi = i\hbar \frac{\partial \psi}{\partial t}$$

$$\square \square \quad \frac{\partial^2 \phi}{\partial t^2} - \nabla^2 \phi + m^2 \phi = g \psi^\dagger \psi$$

## Numerical calculations

$$r_0 = 1.25 \text{ fm}$$

$$R = r_0 A^{1/3}$$

$$mc^2 = 140 \text{ MeV}$$

pion mass

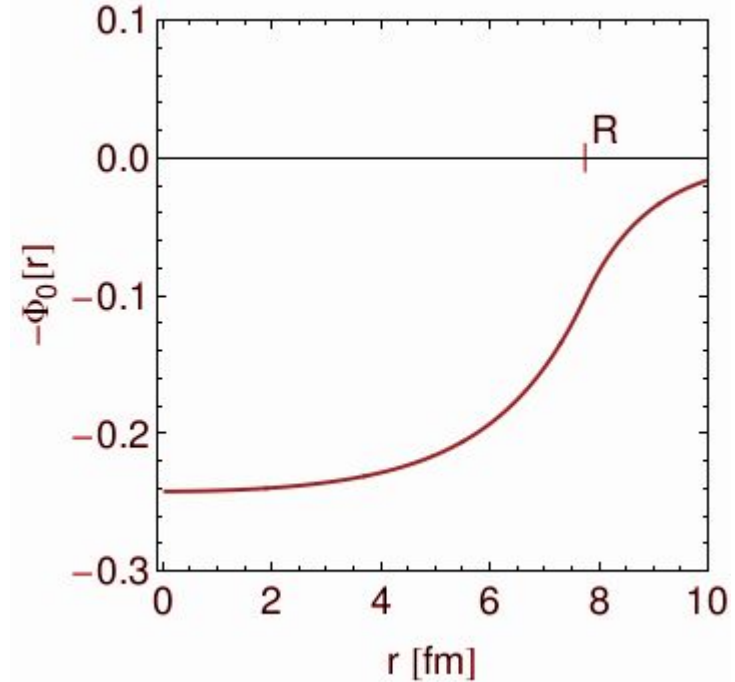
$$A = 238$$

uranium

$$g = 15$$

strong interaction

The potential energy for the extra nucleon is  
 $V(r) = -g^2 \Phi_0(r)$ .



## Homework due Wednesday, March 2

### Problem 32.

For the free real scalar field,

(A ) Write  $H$  in terms of  $\pi(\mathbf{x})$  and  $\phi(\mathbf{x})$ .

(B ) Write  $H$  in terms of  $a_{\mathbf{k}}$  and  $a_{\mathbf{k}}^\dagger$ .

### Problem 33.

(A ) Mandl and Shaw problem 3.3.

(B ) Mandl and Shaw problem 3.4.