

Chapter 5. Photons: Covariant Theory

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SECTION 5.2.

COVARIANT QUANTIZATION

Review the covariant equations for the classical electromagnetic field

$$F = \text{curl } A ; \quad F^{\mu\nu} = \partial^\nu A^\mu - \partial^\mu A^\nu$$

In the Lorentz gauge ($\partial_\mu A^\mu = 0$) we can use the Lagrangian density

$$L = -\frac{1}{2} (\partial_\nu A_\mu)(\partial^\nu A^\mu) - s_\mu A^\mu$$

Now apply canonical quantization to this Lagrangian density. Then calculate

$$[A^\mu(x), A^\nu(y)] = i D^{\mu\nu}(x-y)$$

and

$$\langle 0 | T A^\mu(x) A^\nu(y) | 0 \rangle = i D_F^{\mu\nu}(x-y)$$

$$L = -\frac{1}{2} (\partial_\nu A_\mu)(\partial^\nu A^\mu) - s_\mu A^\mu$$

Recall the real scalar field,

$$L_\phi = \frac{1}{2}(\partial_\nu \phi)(\partial^\nu \phi) - \frac{1}{2} m^2 \phi^2$$

$$[\phi(x), \phi(y)] = i \Delta(x-y)$$

$$\langle 0 | T \phi(x) \phi(y) | 0 \rangle = i \Delta_F(x-y)$$

Note that the free e.m. field has

$$L = -\frac{1}{2}(\partial_\nu A^0)(\partial^\nu A^0) + \frac{1}{2}(\partial_\nu A^i)(\partial^\nu A^i)$$

(sum i = 1 2 3)

■ $A^i(x)$ is just like $\phi(x)$ with $m=0$

$$\therefore [A^i(x), A^i(y)] = i \Delta(x-y)$$

$$\therefore \langle 0 | T A^i(x) A^i(y) | 0 \rangle = i \Delta_F(x-y)$$

(no sum on i)

Different i and j are independent

$$\therefore [A^i(x), A^j(y)] = i \delta_{ij} \Delta(x-y)$$

$$\& \quad \langle 0 | T A^i(x) A^j(y) | 0 \rangle = i \delta_{ij} \Delta_F(x-y)$$

■ $A^0(x)$ is a little different (the sign)

$$\Pi^0 = \partial L / \partial(\partial A^0 / \partial t) = -\partial A^0 / \partial t$$

(compare $\Pi_\phi = \partial \phi / \partial t$)

So the commutator is

$$[A^0(x), A^0(y)] = -i \Delta(x-y)$$

Commutator result

$$[A^\mu(x), A^\nu(y)] = -i g^{\mu\nu} \Delta(x-y)$$

$$= i D^{\mu\nu}(x-y)$$

$$\therefore D^{\mu\nu}(x-y) = -g^{\mu\nu} \Delta(x-y)$$

↑ (with $m=0$) 2

■ As for the propagator,

same thing $\Rightarrow$

$$\langle 0 | T A^\mu(x) A^\nu(y) | 0 \rangle = -i g^{\mu\nu} \Delta_F(x-y)$$

$$= i D_F^{\mu\nu}(x-y)$$

$$D_F^{\mu\nu}(x-y) = -g^{\mu\nu} \Delta_F(x-y)$$

(with $m = 0$)

The Fourier integral,

$$D_F^{\mu\nu}(x-y) = \int -\frac{d^4k}{(2\pi)^4} D_F^{\mu\nu}(k) e^{-i k \cdot x}$$

$$D_F^{\mu\nu}(k) = \frac{-g^{\mu\nu}}{k^2 + i\epsilon}$$

Expansion in plane waves

$$A^\mu(x) = \sum_{\vec{k}} \sum_{r=0}^3 \left(\frac{\hbar c^2}{252\omega} \right)^{1/2} \epsilon_r^\mu(\vec{k})$$

$$\left\{ a_r(\vec{k}) e^{-i\vec{k} \cdot \vec{x}} + a_r^\dagger(\vec{k}) e^{i\vec{k} \cdot \vec{x}} \right\}$$

where $k^0 = |\vec{k}| = \omega$.

The four-vector polarization vectors are defined like this:

$$\epsilon_0^\mu(\vec{k}) = \eta^\mu \equiv (1, 0, 0, 0) \quad \text{pure time like}$$

$$\epsilon_i^\mu(\vec{k}) = (0, \hat{\epsilon}_i(\vec{k})) \quad \text{pure space like } (i=1,2,3)$$

$$\hat{\epsilon}_3(\vec{k}) = \vec{k}/\omega \quad \text{longitudinal}$$

$$\hat{\epsilon}_1, \hat{\epsilon}_2, \hat{\epsilon}_3 \quad \text{form an orthogonal triad.}$$

transverse



The canonical commutation relation

$$[A^\mu(x), A^\nu(y)] = -i g^{\mu\nu} \Delta(x-y) \text{ implies}$$

$$[a_r(\vec{k}), a_s^\dagger(\vec{k}')] = \delta_{rs} \delta_{\vec{k}\vec{k}'} \zeta_r \quad \text{where } \zeta_r = \begin{cases} -1 & \text{for } 0 = r \\ +1 & \text{for } 123 \end{cases}$$

The constraint $\partial_\mu \mathbf{A}^\mu = 0$;
i.e., Lorentz gauge quantization

◦ We cannot set $\partial_\mu \mathbf{A}^\mu = 0$ as an operator equation.

Proof
$$[\partial_\mu A^\mu(x), A^\nu(y)] = -i \partial^\nu \Delta(x-y)$$

 and it is not 0.

◦ The Gupta-Bleuler formalism:

(Comment: it's not a theory;
 it's not a model; it's a *formalism*.)

Apply the constraint, $\partial_\mu \mathbf{A}^\mu = 0$, to the states of the Hilbert space; require

$$\partial_\mu A^{(+)\mu} |\Psi\rangle = 0,$$

for any physical state $|\Psi\rangle$.

◦ Theorem

$$[a_3(\mathbf{k}) - a_0(\mathbf{k})] |\Psi\rangle = 0 \text{ for all } \mathbf{k}$$

Proof

$$\begin{aligned} \partial_\mu A^{(+)\mu} &= \partial_\mu \sum_{\vec{k}} \sum_{r=0}^3 \frac{1}{\sqrt{2\pi\omega}} \epsilon_r^\mu(\vec{k}) a_r(\vec{k}) e^{-ik \cdot x} \\ &= \sum_{\vec{k}} \sum_r \frac{1}{\sqrt{2\pi\omega}} (-ik_\mu) \epsilon_r^\mu(\vec{k}) a_r(\vec{k}) e^{-ik \cdot x} \\ &= \sum_{\vec{k}} \sum_{r=0,3} \frac{(-i)}{\sqrt{2\pi\omega}} [k^0 \epsilon_r^0 - \vec{k} \cdot \vec{\epsilon}_r] a_r(\vec{k}) e^{-ik \cdot x} \\ &= \sum_{\vec{k}} \frac{(-i)}{\sqrt{2\pi\omega}} \underbrace{\{ k^0 a_0(\vec{k}) - \vec{k} \cdot \vec{\epsilon}_3 a_3(\vec{k}) \}}_{= \omega (a_0(\vec{k}) - a_3(\vec{k}))} e^{-ik \cdot x} \end{aligned}$$

$$\partial_\mu A^{(+)\mu} |\Psi\rangle = 0 \quad \text{implies}$$

$$[a_0(\vec{k}) - a_3(\vec{k})] |\Psi\rangle = 0 \quad \text{all } \vec{k}$$

$$[a_3(\mathbf{k}) - a_0(\mathbf{k})] |\Psi\rangle = 0, \quad (\star)$$

for any physical state; arb $\mathbf{k}$.

What does it mean?

◦ Recall the free vacuum $|0\rangle$; it has

$$a_r(\mathbf{k}) |0\rangle = 0 \quad \text{for } r = 0, 1, 2, 3$$

so $|0\rangle$ obeys $(\star)$

◦ Now consider $a_r^\dagger(\mathbf{k}) |0\rangle$.

For $r = 1, 2$, it obeys $(\star)$.

Creating any transverse photons will yield a physical state.

For $r = 3, 0$, the state does not obey $(\star)$.
Creating single longitudinal or scalar photons will yield an unphysical state; or, a physical state requires creating longitudinal and scalar together.

◦ Example. Consider

$$|\psi\rangle = [a_3^\dagger(\mathbf{p}) - a_0^\dagger(\mathbf{q})] |0\rangle$$

It has

$$a_3(\mathbf{k}) |\psi\rangle = \delta_{\mathbf{k},\mathbf{p}} |0\rangle$$

$$a_0(\mathbf{k}) |\psi\rangle = \delta_{\mathbf{k},\mathbf{q}} |0\rangle$$

So $|\psi\rangle$ is a physical state $(\star)$ provided that $\mathbf{p} = \mathbf{q}$.

(M. & Sh. problem 5.2; homework)

◦ But that is just a gauge transformation.

(M. & Sh. problem 5.3; homework)

Results of the formalism

In the Gupta-Bleuler formalism,

- we only calculate transition amplitudes between states with transverse photons;
- i.e., longitudinal and scalar photons are not included in asymptotic states;
- but longitudinal and scalar photons do exist as virtual particles;
- i.e., we use the full propagator $-g^{\mu\nu}/(k^2+i\epsilon)$.

This is all we need to proceed.

So now we could forget about the Gupta-Bleuler formalism.

But note what Mandl and Shaw say at the end of Section 5.2:

“ ...

For most purposes, the complete formalism is not required.

[footnote: if interested, read 3 other books.]

”

An alternative approach (which is required for QCD but optional for QED) is to use *functional integration* with the *Faddeev-Popov formalism*. (M.&Sh. chapters 10 - 14)