

Friday April 1

## Chapter 5. Photons: Covariant Theory

5.1. The classical fields ✓

5.2. Covariant quantization ✓

5.3. The photon propagator ✓

## Chapter 6. The S-Matrix Expansion

6.1. Natural Dimensions and Units ✓

6.2. The S-matrix expansion ✓

6.3. Wick's theorem ✓

## Chapter 7. Feynman Diagrams and Rules in QED

7.1. Feynman diagrams in configuration space ✓

7.2. Feynman diagrams in momentum space

7.3. Feynman rules for QED

7.4. Leptons

## Chapter 8. QED Processes in Lowest Order

8.1. The cross section

8.2. Spin sums

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8.4-7. Examples

8.8-9. Bremsstrahlung

## SECTION 7.2.

### *FEYNMAN DIAGRAMS IN MOMENTUM SPACE*

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First, let's consider an example:

ee scattering;

⇒ the Møller cross section.

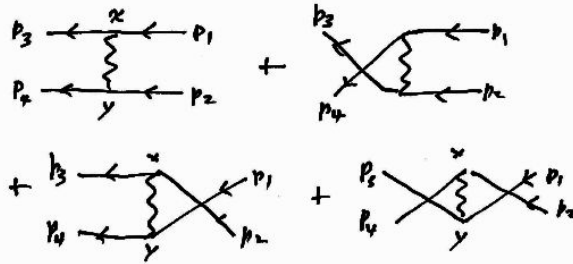
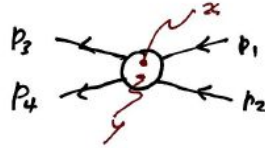
The lowest order term is 2nd order in the interaction,

$$S^{(2)} = (1/2!) (ie)^2 \iint d^4x d^4y \\ \langle f | T \bar{\psi} \gamma_\mu \psi A^\mu(x) \\ \bar{\psi} \gamma_\nu \psi A^\nu(y) | i \rangle$$

where  $|i\rangle = |e_1; e_2\rangle$  and  $|f\rangle = |e_3; e_4\rangle$

$$S^{(2)} = (1/2!) (ie)^2 \iint d^4x d^4y \\ \langle f | T \bar{\psi} \gamma_\mu \psi A^\mu(x) \bar{\psi} \gamma_\nu \psi A^\nu(y) | i \rangle$$

Apply Wick's theorem and the coordinate space Feynman rules.



Wick's theorem gives us 4 terms.

However,  $S_{22} = S_{11}$  and  $S_{21} = S_{12}$  because we integrate over  $d^4x$  and  $d^4y$ .

(E.g., in  $S_{22}$  change the variables of integration from  $x, y$  to  $y, x'$ ; then drop the primes; the result is  $S_{11}$ .)

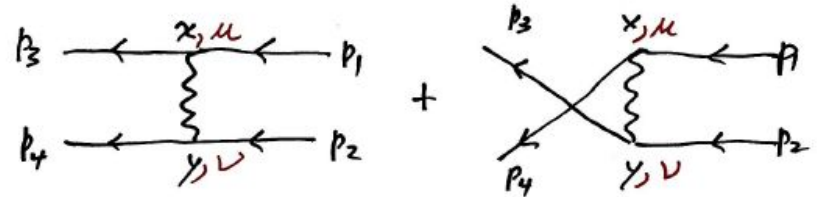
So,  $S = (S_{11} + S_{12}) \times 2$ ;  
the 2 will cancel the  $1/2!$ .

Note: The  $1/n!$  from the exponential series, will always cancel  $n!$  permutations of the vertex positions  $\{x_1, x_2, x_3, \dots, x_n\}$ .

This leads to a Feynman rule:

(1) Draw all the *topologically distinct* diagrams with the specified external lines.

For ee scattering there are two Feynman diagrams,



The corresponding S-matrix elements are

$$S_t = \iint d^4x d^4y e^{ik_3 \cdot x} e^{ik_4 \cdot y} e^{-ik_1 \cdot x} e^{-ik_2 \cdot y} \\ \prod_{a=1}^4 \left( \frac{2m}{2E_a \Omega} \right)^{1/2} \bar{u}_3 ie\gamma_\mu u_1 \bar{u}_4 ie\gamma_\nu u_2 D_F^{\mu\nu}(x-y)$$

$$S_u = \iint d^4x d^4y e^{ik_3 \cdot y} e^{ik_4 \cdot x} e^{-ik_1 \cdot x} e^{-ik_2 \cdot y} \\ \prod_{a=1}^4 \left( \frac{2m}{2E_a \Omega} \right)^{1/2} \bar{u}_3 ie\gamma_\nu u_2 \bar{u}_4 ie\gamma_\mu u_1 D_F^{\mu\nu}(x-y)$$

Transform S to momentum space.

$$D^{\mu\nu}(\mathbf{x}-\mathbf{y}) = (2\pi)^{-4} \int d^4k (-g^{\mu\nu}/k^2) e^{ik \cdot (\mathbf{x}-\mathbf{y})}$$

For the case  $S_t$ : the integral  $d^4x$  gives

$$\int d^4x e^{i\vec{p}_3 \cdot \mathbf{x}} e^{-i\vec{p}_1 \cdot \mathbf{x}} e^{i\vec{k} \cdot \mathbf{x}} = (2\pi)^4 \delta^4(\vec{p}_3 - \vec{p}_1 + \vec{k})$$

$$\therefore k^\mu = p_1^\mu - p_3^\mu$$

Comment:

That would be for infinite spacetime.

Instead, we normalize plane waves in a finite volume  $\Omega$ , so then the result should be

$$\Omega \delta_K(\vec{k}, \vec{p}_1 - \vec{p}_3) 2\pi \delta(E_3 - E_1 + K)$$

These are the same in the limit  $\Omega \rightarrow \infty$ ,  
but ...

The integral  $d^4y$ :

$$\int d^4y e^{i\vec{p}_4 \cdot \mathbf{y}} e^{-i\vec{p}_2 \cdot \mathbf{y}} e^{-i\vec{k} \cdot \mathbf{y}} = (2\pi)^4 \delta^4(p_4 - p_2 - k)$$

$$\therefore k^\mu = p_4^\mu - p_2^\mu$$

$$\therefore p_3^\mu + p_4^\mu = p_1^\mu + p_2^\mu$$

The transformation from coordinate space to momentum space gave us some delta functions for 4-momentum conservation. This leads to another Feynman rule:

**(2) 4-momentum is conserved at every vertex.**

$$S_t = \int \frac{d^4k}{(2\pi)^4} (2\pi)^8 \delta^4(k - p_1 + p_3) \delta^4(k + p_2 - p_4)$$

$$\frac{1}{i!} \left( \frac{2m}{2E_k \Omega} \right)^{1/2} \bar{u}_3 i\epsilon \gamma_\mu u_1 \bar{u}_4 i\epsilon \gamma_\nu u_2 \frac{-g^{\mu\nu}}{k^2}$$

$$= (2\pi)^4 \delta^4(p_f - p_i) \frac{1}{i!} \left( \frac{2m}{2E_k \Omega} \right)^{1/2} \mathcal{M}_t$$

which defines the new matrix element  $\mathcal{M}$ .  
(The other factors are the same in  $S_t$  and  $S_u$ )

Result  $M_t = e^2 (\bar{u}_3 \gamma_\mu u_1) (\bar{u}_4 \gamma^\mu u_2) / (p_1 - p_3)^2$

Similarly, exercise,

$$M_u = -e^2 (\bar{u}_3 \gamma_\mu u_2) (\bar{u}_4 \gamma^\mu u_1) / (p_1 - p_4)^2$$

Recall the Mandelstam variables,

$$s = (p_1 + p_2)^2 = 2m^2 + 2p_1 \cdot p_2 = (p_3 + p_4)^2 = 2m^2 + 2p_3 \cdot p_4$$

$$t = (p_1 - p_3)^2 = 2m^2 - 2p_1 \cdot p_3 = (p_4 - p_2)^2 = 2m^2 - 2p_2 \cdot p_4$$

$$u = (p_1 - p_4)^2 = 2m^2 - 2p_1 \cdot p_4 = (p_3 - p_2)^2 = 2m^2 - 2p_2 \cdot p_3$$

The Feynman rules in momentum space are rules for calculating the matrix element  $\mathcal{M}$ .

The S-matrix has some normalization factors that are not included in  $\mathcal{M}$ .

These normalization factors are

$$\prod_i (2\Omega E_i)^{-1/2} \quad \text{and} \quad \prod_f (2\Omega E_f)^{-1/2}$$

$$\text{and } \prod_e (2m)^{1/2}.$$

These factors are not part of  $\mathcal{M}$ .

The matrix element for ee (Møller) scattering

$$\mathcal{M}_t = e^2 (\bar{u}_3 \gamma_\mu u_1) (\bar{u}_4 \gamma^\mu u_2) / t$$

$$\mathcal{M}_u = -e^2 (\bar{u}_3 \gamma_\mu u_2) (\bar{u}_4 \gamma^\mu u_1) / u$$

$$\mathcal{M} = \mathcal{M}_t + \mathcal{M}_u$$

Note these other Feynman rules:

(3) A spinor for every external electron and positron.

(4) A polarization vector for every external photon.

(5) A propagator  $S_F(p)$  for every (internal) electron line.

(6) A propagator  $D_F^{\mu\nu}(q)$  for every (internal) photon line.

(7) A minus sign for exchanging 2 electrons.

To complete the calculation of the Møller cross section, we will need:

- $|\mathcal{M}|^2$  ;
- and the average over initial spins and sum over final spins, i.e.,  
$$\frac{1}{2} \sum_{\lambda_1} \frac{1}{2} \sum_{\lambda_2} \sum_{\lambda_3} \sum_{\lambda_4} [\dots]$$

*Let's go ahead and calculate that now.*

Homework Problem X:  
Plot the Moller cross section.

The cross section.

The transition probability is the square of the S-matrix element,

$$P = |S|^2.$$

The transition rate is the probability per unit time,

$$w = P/(2T)$$

(evolution from  $-T$  to  $T$ )

The cross section is the rate divided by the incident flux, and the incident flux is

$$\Phi = \text{density} \times \text{velocity} = v_{\text{rel}} / \Omega ;$$

(1 particle in the volume  $\Omega$ )

Thus  $d\sigma = w / \Phi =$

$$\frac{|S|^2 \Omega}{2T v_{\text{rel}}}$$