#### three more historical things

- 1 Feynman Diagrams
- 2 spin
- 3 beta decay
- 4 exchange forces



#### "Feynman Diagrams"

#### actually, a shorthand for calculations

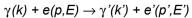
each diagram tells us how to skip about 20 pages of a calculation!

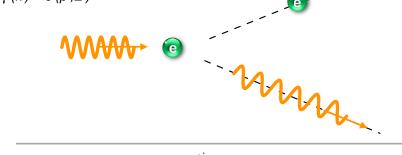
also, a marvelous "visualization" of any spacetime process

#### **Compton Scattering**

#### remember?

x-ray + electron -> x-ray + electron



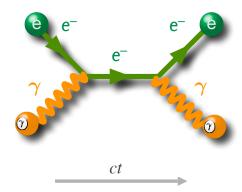


time

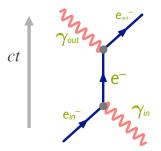
thought of as:

xray absorbed by electron...
then emitted later by electron

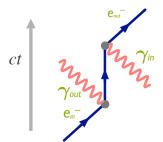
#### we draw them in a general frame



#### compton, revisited (it's $e_{\gamma} \rightarrow e_{\gamma}$ )



In order to include all possibilities, it's necessary to also take into account:



the important thing is that the initial and the final states are indistinguishable between these two diagrams

As that's all we can measure, we must include all ways of producing the same initial and final states, regardless of the direction of time.

# here's what I think of you:





### the deal:

you come to class, do the work:

You'll learn some physics

and see science differently

and you'll do okay.

## today:

#### think of it as Brian Greene

without the turtleneck or Star Wars bar characters a little stuff, and a little entertainment.

#### three more historical things

- Feynman Diagrams
- 2 spin
- 3 beta decay
- 4 exchange forces



#### "spin"

we actually define individual quanta by 3 quantities:

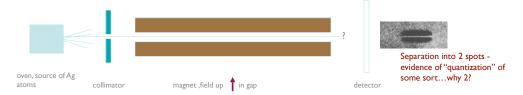
mass, electrical charge, spin and other "quantum numbers"

but, spin has a history

#### Schrodinger could not explain:

#### From out of the blue: The Stern Gerlach Effect

In 1921, Otto Stern and Walter Gerlach managed to prepare a molecular beam of Silver atoms, which they passed through a magnetic field...



IF electrons were really orbiting the nucleus classically, then the magnetic field would spread the beam out depending on the orientation of the circulating currents - blur

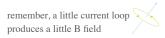
IF electrons distributed themselves according to the Schrödinger model, then there would be many discrete spots at the detector...corresponding to all different directions of all of the many orbital angular momenta in the Silver atom

**BUT** this 2-line business was disconcerting

#### my head is spinning

#### Two young Dutch theoreticians had the nutty idea

In 1925, George Uhlenbeck and Sam Goudsmit speculated that the electron is like a spinning charge...



how about a little spinning sphere of charge? ... also a little B field





That's what they said an electron was like: BUT, there's no such current, and the electron has no size - this little "magnetic moment" is an inherent quality of an electron

But, since the electron has no apparent size - "spinning" is a weird notion!! ...a metaphor

Their proposal was that to explain the Stern-Gerlach experiment, this spin-angular momentum, "S", is also quantized as ±1/2h - two states, "up" and "down"

It spins in only two orientations and there is a quantum number associated with them This quantum "number" is called "Spin" and is a property of the electron (and others).

What's happening in the Stern-Gerlach experiment?

Silver happens to be an atom with a single valence electron in its most outer shell

The magnet was selecting those of the Silver atoms that had spin projection = +1/2 from those that had -1/2...two spots

#### another gift from Dirac:

#### Spin turned out to be required

when quantum mechanics was combined with Special Relativity...it naturally emerged as a natural part of that theory...quite amazing.

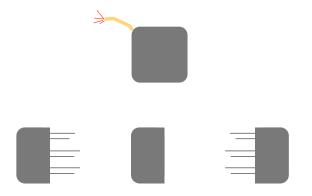


#### beta decay

was an embarrassment

#### strange firecracker

Suppose we have a firecracker exploding into two pieces



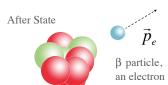
well, beta decay appeared to be the middle situation...it appeared to violate momentum and energy conservation

#### trouble with beta decay

#### Studies of the beta decay of various nuclei took a turn toward absurdity

The general reaction of interest was  $_{Z}N^{A}$   $\rightarrow$   $_{Z+1}N^{A}$  + e<sup>-</sup> nucleus adds a proton and emits an electron (so Q is conserved)





 $\vec{p}_N$  = essentially still 0!

Momentum, and hence, energy... appeared to be imbalanced between the before and after states...

#### •By the late 1920's and early 1930's, measurement precision became guite good

all experiments agreed: it behaved like a firecracker exploding into 1 moving piece

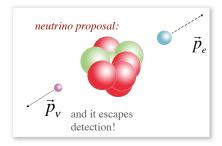
...not conserving momentum or energy! all searches for unseen radiation, like gamma rays, failed conservation of energy was seriously called into question by, of course, Bohr "I have come upon a desperate way out. To wit, the possibility that there could exist in the nucleus electrically **neutral particles** which I shall call neutrons[sic]...the mass...should not be larger than 0.01 times the proton...the continuous beta spectrum would then be understandable from the assumption that...a [neutrino] is emitted along with the electron...I admit that my way out may not seem very probable...But only he who dares wins...unfortunately I cannot appear personally in Tubingen since a ball which takes place in Zurich makes my presence here indispensable."

Pauli, distressed at the crisis and unwilling to part with energy conservation-like Bohr suggested!-in 1930 made a bold proposal, in an off-hand way

the idea hung around for a few years -but the discovery of the actual neutron gave Enrico Fermi an idea.

the neutrino...little neutron

#### prediction of an elementary particle



#### **Enrico Fermi**

## an unusual mixture of theoretician and experimentalist

...doesn't happen very often - he was simply the best At an early age, he rewrote the statistical description of spin-1/2 particles now called "fermions"

# Discovered the first artificially produced nucleon excitation (called the "delta") and managed to create the first controlled nuclear fission reaction at Chicago

the beginning of the Manhattan Project

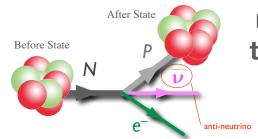


## In 1933 formulated the first workable model for beta decay

and heralded the beginning of the study of the "Weak Interaction"

1901-1954
(actually in a cafeteria in Ann Arbor, 1935)

Enrico Fermi



#### Fermi proposed:

the neutron is unstable and transmutes into a proton, electron, and neutrino

Dirac's *creation* and *annihilation*, idea put to work

He recognized that the strength of the reaction - related to the lifetime of the state which decays under this reaction - is very small compared to the electromagnetic interaction

the electromagnetic interaction has a relative strength of  $10^{-2}$  the beta decay reaction has a strength of  $\sim 10^{-13}$ 

It's called the **Weak Interaction** for that reason...there are a whole class of interactions which fall into this category any reaction involving a neutrino is weak.

# neutrons aren't what they were cracked up to be

So, the neutron would be a bit more massive than the proton

The neutrino carries away energy and momentum, balancing the electron energy, which can be a spectrum (since the proton is involved)

which Fermi predicted in a paper so audacious that it was originally rejected by *Zeitschrift fur Physik*...and so he had to publish it in Nuovo Cimento, a less prestigious Italian journal

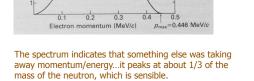
The shape of the

been a feature of concentration in

neutrino mass

searches

endpoint is sensitive

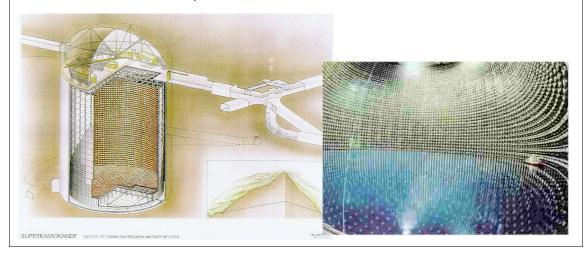


#### neutrino's putting on weight

#### since Fermi's model

neutrino has been presumed to be massless, but this has been tested for decades

now, after 50 years, it's clear that the neutrino has a tiny mass the mass difference among neutrinos is tiny fractions of an eV/c<sup>2</sup> discovered in Japan in 1998



# they're everywhere, they're everywhere

## The neutrino discovered in 1956 by Fred Reines and Clyde Cowan

inexplicably, the Nobel Prize only awarded 1995 Pauli did live to see the results

and he was right, they do not interact...very much it takes lightyears of lead to stop a neutrino, on average

#### The neutrino is ubiquitous in the universe

the nuclear reactions in the sun produce 10<sup>38</sup> neutrinos per second

the big bang is responsible for there being roughly 300 neutrinos/cm<sup>3</sup> …a bath of neutrinos, everywhere in the universe

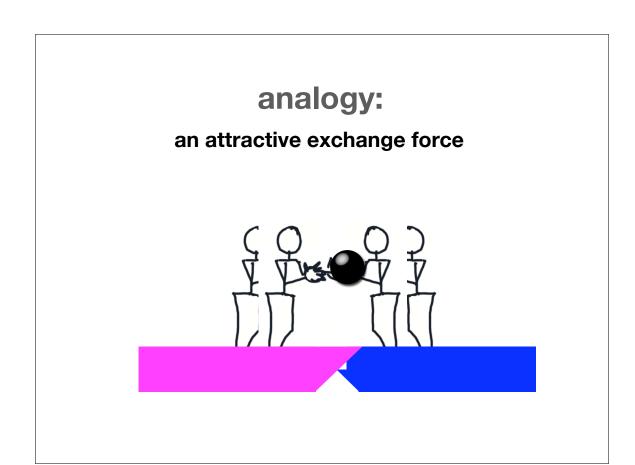
the potassium 40 in our bodies beta decays: we are emitters of  $^{\sim}3$  x  $10^8$  V per day



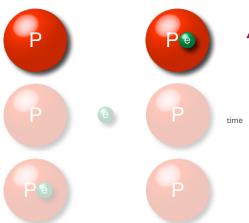
#### think about it.

what keeps the many protons large nuclei from repelling one another?

# analogy: a repulsive exchange force



# Heisenberg had an original idea before the discovery of the neutron in 1932



this is wrong.

#### exchange forces

# After the discovery of the neutron in 1932: Heisenberg proposed a family relationship:

proton and neutron are very, very close in mass

He proposed that they can be thought of as

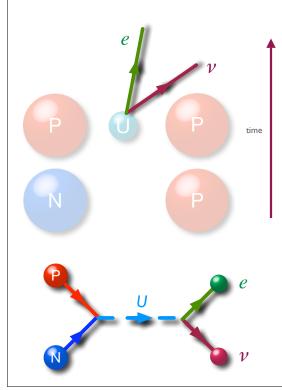
two sides of the same coin: a "nucleon"

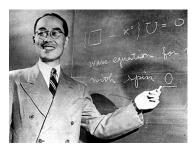
one changing into another through the exchange of a third intermediate

#### proposed a quantum number, "Isospin" of 1/2

I of proton = +1/2I of neutron = -1/2changing p into n...requires the exchange of I = 1

#### a model for beta decay





#### Hideki Yukawa in 1934

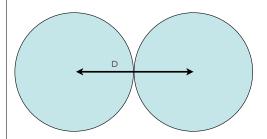
after the discovery of the neutron motivated by Fermi's theory of beta decay

Imagining the exchange particle being uncertain in position inside of the nucleon...and traveling at near the speed of light, he estimated the energy and hence, mass

#### range of strong force:

#### about the size of the proton...a rough concept

...can't travel faster than light, so the time across the proton would be: D/c Yukawa's particle would violate energy in passing from one p to another



From uncertainty

$$\Delta E \approx \frac{\hbar}{D/c}$$
 
$$m \approx \frac{\Delta E}{c^2} \approx 0.1 \times m_{proton}$$

So the hunt was on for a particle of a mass of about 0.1 times that of the proton ( $\sim 200 \times m_{electron}$ ), or about 100 MeV/c<sup>2</sup>...that would cause protons to stick together

#### Look, Ma...I'm on Top of the World

#### Cloud chambers: sophisticated & well-traveled

Anderson and others took devices to the top of Pike's Peak, the Pyrenees...

By 1937, strange things were beginning to be seen... tracks that looked like electrons...but would not slow down in Anderson-like plates

They guessed that they had masses of around 200 x melectron

#### So...they were not protons and not electrons

Could it be the Yukawa particle?

No...during WWII, Italians Conversi and Piccioni, working in Rome in basements at night fearing for their lives, built wholly electronic devices based on the Geiger-Mueller tubes

They measured the lifetime of these penetrating particles: ~2 x 10<sup>-6</sup> s - too long-lived to be the Yukawa particle

#### Named "mesotron" ('in the middle')

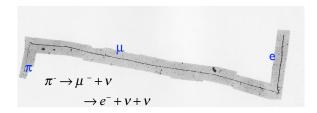
now called the "muon"

#### muons and pions and electrons, Oh, my!

#### **Robert Marshak figured it out:**

Other emulsion experiments observed, short, stopping tracks that were the Yukawa particle (dubbed the "pion")

they are unstable and decay into the muon which in turn, is unstable, decaying into an electron and 2 neutrinos

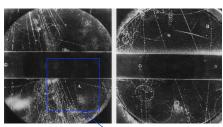


#### then, things went nuts...

#### Cosmic rays continued to produce surprises

In 1946 - in Manchester, pictures showed the presence of "Vee's"

the apparent production of a neutral particle that decayed into two charged particles



the reaction is the decay of what is now called the Kaon, or neutral version,  $K^0\!\!:$ 

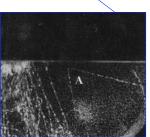
it has a mass of 0.496 MeV/c<sup>2</sup>

 $K^0 \rightarrow \pi^+ \pi^-$ 

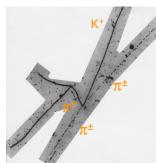
another V was found that decayed into a nucleon and a pion... called the "lambda"

it has a mass of I.II5  $\mbox{GeV}/\mbox{c}^2$ 

To add insult to injury... another Kaon was found that decayed into 3 pions



These were all "strange" things...so the particles were dubbed "Strange Particles" and the name stuck...thus beginning the tradition of naming the most fundamental bits of matter by whimsical and silly names



#### **Eventually, forces categorized**

#### By how strongly the forces act between particles

you can pick up one magnet with another magnet so, you'd say that the magnetic force > the gravitational force, okay?

#### The four fundamental forces appear to be:

the **strong force** (strongest in Nature)

nuclear binding, originally thought to be propagated by the pion... the strongest force in nature

#### the electromagnetic force

between particles of electric charge - always involving a photon

the weak force

short decays like beta decay, any reaction involving neutrinos and other reactions that we'll come to realize involve new "photons"

the gravitational force (weakest in Nature)

particles of mass/energy...namely all particles

In all of nature, and for nearly the entire time of the universe, there appear to be only these 4

#### each force:

#### propagated by a quantum cousin of the photon:

strong force: the "gluon"

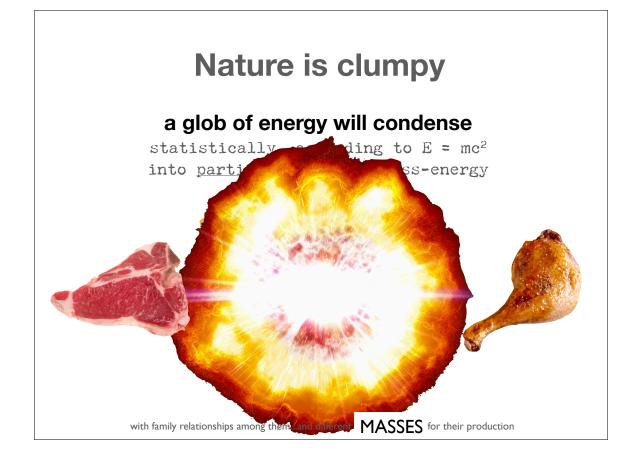
weak force: the "W"

electromagnetic force: the photon

gravitational force: the "graviton" (speculated)

#### and a bonus: the weak and electromagnetic forces

are one in the same at very high energy densities theory predicted: the "Z," directly related to photon



# the clumps? a menagerie of "particles".

#### subsequently, hundreds of them

produced artificially in accelerators sometimes, they don't live very long...



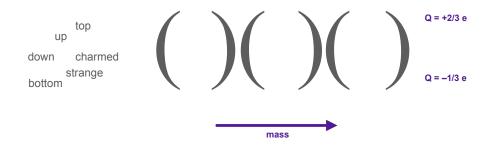


	LIGHT UNF			STRANGE		BOT	
	$f^{G}(J^{PC})$	B = 0)	$f^{G}(J^{PC})$	(S = ±1, 0	$I(J^D)$	(B =	$I^G(J^{PC})$
- π <sup>±</sup>	1-(0-)	<ul> <li>π<sub>2</sub>(1670)</li> </ul>	1-(2-+)	• K±	1/2(0-)	• B±	1/2(0-)
π <sup>0</sup>	1-(0-+)	• ¢(1680)	0-(1)	• K <sup>0</sup>	1/2(0-)	• B <sup>0</sup>	1/2(0-)
• η	0+(0-+)	<ul> <li>ρ<sub>3</sub>(1690)</li> </ul>	1+(3)	• K <sup>0</sup>	1/2(0-)	<ul> <li>B±/B<sup>0</sup> ADM</li> </ul>	
f <sub>0</sub> (600)	0+(0++)	<ul> <li>ρ(1700)</li> </ul>	1+(1)	• K,	1/2(0-)	<ul> <li>B±/B<sup>0</sup>/B<sup>0</sup>,/b</li> </ul>	-baryon AD-
ρ(770)	1+(1)	$a_2(1700)$	1-(2++)	K*(800)	1/2(0+)	MIXTURE	
• ω(782)	0-(1)	<ul> <li>f<sub>0</sub>(1710)</li> </ul>	0+(0++)	<ul> <li>K*(892)</li> </ul>	1/2(1-)	V <sub>cb</sub> and V <sub>ub</sub> ( Elements	-rivi Matrix
• n/(958)	0+(0-+)	$\eta(1760)$	0+(0-+)	<ul> <li>K<sub>1</sub>(1270)</li> </ul>	$1/2(1^+)$	• B*	1/2(1-)
• f <sub>0</sub> (980)	0+(0++)	<ul> <li>π(1800)</li> </ul>	1-(0-+)	<ul> <li>K<sub>1</sub>(1400)</li> </ul>	$1/2(1^+)$	B* <sub>J</sub> (5732)	?(??)
• a <sub>0</sub> (980)	1-(0++)	f <sub>2</sub> (1810)	0+(2++)	• K*(1410)	1/2(1-)	BOTTOM,	STRANCE
• ¢(1020)	0-(1) 0-(1+-)	X(1835)	? <sup>?</sup> (? - +)	<ul> <li>K<sub>0</sub>(1430)</li> </ul>	1/2(0+)	(B = ±1,	
<ul> <li>h<sub>1</sub>(1170)</li> <li>b<sub>1</sub>(1235)</li> </ul>	1+(1+-)	<ul> <li>φ<sub>3</sub>(1850)</li> </ul>	0+(2-+)	<ul> <li>K<sub>2</sub>(1430)</li> </ul>	1/2(2+)	• B <sup>0</sup>	0(0-)
• a <sub>1</sub> (1260)	1-(1++)	$\eta_2(1870)$ $\rho(1900)$	1+(1)	K(1460)	1/2(0-)	B;	0(1-)
• f <sub>2</sub> (1270)	0+(2++)	f <sub>2</sub> (1910)	0+(2++)	K <sub>2</sub> (1580)	1/2(2-)	B*, (5850)	?(??)
f <sub>1</sub> (1285)	0+(1++)	• f <sub>2</sub> (1950)	0+(2++)	K(1630)	1/2(??)	D <sub>2</sub> (3030)	:(: )
η(1295)	0+(0-+)	ρ <sub>3</sub> (1990)	1+(3)	K₁(1650) • K*(1680)	1/2(1 <sup>+</sup> ) 1/2(1 <sup>-</sup> )	воттом,	
<ul><li>π(1300)</li></ul>	1-(0-+)	• f <sub>2</sub> (2010)	0+(2++)	• K <sub>2</sub> (1770)	1/2(2-)	(B = C	= ±1)
<ul> <li>∂₂(1320)</li> </ul>	1-(2++)	f <sub>0</sub> (2020)	0+(0++)	• K <sub>3</sub> (1780)	1/2(3-)	• B <sup>±</sup> <sub>c</sub>	0(0-)
• f <sub>0</sub> (1370)	0+(0++)	<ul> <li>a<sub>4</sub>(2040)</li> </ul>	1-(4++)	• K <sub>2</sub> (1820)	1/2(2-)	6	7
$h_1(1380)$	7-(1+-)	<ul> <li>f<sub>4</sub>(2050)</li> </ul>	0+(4++)	K(1830)	1/2(0-)	• η <sub>ε</sub> (15)	0+10-
$\pi_1(1400)$	1-(1-+)	$\pi_2(2100)$	1-(2-+)	K*(1950)	1/2(0+)	• J/ψ(15)	0-(1-
• η(1405)	0+(0-+)	$f_0(2100)$	0+(0++)	K5(1980)	1/2(2+)	<ul> <li>χ<sub>c0</sub>(1P)</li> </ul>	0+(0+
f <sub>1</sub> (1420)	0+(1++) 0-(1)	f <sub>2</sub> (2150)	0+(2++)	<ul> <li>K<sub>4</sub>*(2045)</li> </ul>	1/2(4+)	• χc1(1P)	0+(1+
• ω(1420)	0+(2++)	ρ(2150) ε (2200)	1 <sup>+</sup> (1 <sup></sup> ) 0 <sup>+</sup> (0 <sup>++</sup> )	K2(2250)	1/2(2-)	$h_c(1P)$	??(???)
f <sub>2</sub> (1430) • a <sub>0</sub> (1450)	1-(0++)	$f_0(2200)$ $f_2(2220)$	0+(2 or 4 + +)	$K_3(2320)$	$1/2(3^{+})$	<ul> <li>χ<sub>C2</sub>(1P)</li> </ul>	0+(2+
ρ(1450)	1+(1)	η(2225)	0+(0-+)	K <sub>5</sub> (2380)	1/2(5-)	<ul> <li>η<sub>c</sub>(25)</li> </ul>	0+(0-
η(1475)	0+(0-+)	ρ <sub>3</sub> (2250)	1+(3)	K <sub>4</sub> (2500)	1/2(4-)	<ul> <li>ψ(25)</li> </ul>	0-(1-
f <sub>0</sub> (1500)	0+(0++)	<ul> <li>f<sub>2</sub>(2300)</li> </ul>	0+(2++)	K(3100)	??(???)	<ul> <li>ψ(3770)</li> </ul>	0-(1- 0?(??+)
$f_1(1510)$	0+(1++)	f <sub>4</sub> (2300)	0+(4++)	CHAR	MED	• X(3872)	0+(2+
• f' <sub>2</sub> (1525)	0+(2++)	<ul> <li>f<sub>2</sub>(2340)</li> </ul>	0+(2++)	(C=	±1)	• χc2(2P) Υ(3940)	??(???)
$f_2(1565)$	0+(2++)	$\rho_5(2350)$	1+(5)	• D±	1/2(0-)	• \$\psi(4040)	0-(1-
$h_1(1595)$	0-(1+-)	a <sub>6</sub> (2450)	1-(6++)	• D <sup>0</sup>	1/2(0-)	<ul> <li>ψ(4160)</li> </ul>	0-(1-
$\pi_1(1600)$	1-(1-+)	f <sub>6</sub> (2510)	0+(6++)	• D*(2007)0	1/2(1-)	Y(4260)	??(1
a <sub>1</sub> (1640)	1-(1 + +) 0+(2 + +)	OTH	ER LIGHT	• D*(2010)±	1/2(1-)	<ul> <li>ψ(4415)</li> </ul>	0-(1-
f <sub>2</sub> (1640) • η <sub>2</sub> (1645)	0+(2-+)	Further Sta	tes	D*(2400)0	1/2(0+)		
ω(1650)	0-(1)			D*(2400)*	1/2(0+)	b	
ω(1670) ω <sub>3</sub> (1670)	0-(3)			• D <sub>1</sub> (2420) <sup>0</sup>	1/2(1 <sup>+</sup> ) 1/2(? <sup>?</sup> )	$\eta_b(15)$	0+(0-
	- (- /			$D_1(2420)^{\pm}$ $D_1(2430)^0$	1/2(1+)	• T(15)	0-(1-
				• D;(2460)0	1/2(2+)	<ul> <li>χ<sub>b0</sub>(1P)</li> <li>χ<sub>b1</sub>(1P)</li> </ul>	0+(1+
				• D <sub>2</sub> (2460)±	1/2(2+)	• χ <sub>b2</sub> (1P)	0+(2+
				D*(2640)±	1/2(??)	• T(25)	0-(1-
				. ,		T(1D)	0-(2-
				CHARMED,		<ul> <li>\chi_{b0}(2P)</li> </ul>	0+(0+
				(C = S		<ul> <li>χ<sub>b1</sub>(2P)</li> </ul>	0+(1+
				• D <sub>s</sub> <sup>±</sup>	0(0-)	<ul> <li>χt2(2P)</li> </ul>	0+(2+
				• D*±	0(??)	<ul> <li>T(3S)</li> </ul>	0-(1-
				• D*(2317)±	0(0+)	• T(45)	0-(1-
				• D <sub>s1</sub> (2460)±	0(1 <sup>+</sup> ) 0(1 <sup>+</sup> )	• 7'(10860)	0-(1-
				<ul> <li>D<sub>81</sub>(2536)<sup>±</sup></li> <li>D<sub>52</sub>(2573)<sup>±</sup></li> </ul>	0(2?)	• T(11020)	0-(1-

		****	4/1000)		****			****	Σ+		****	=0		****
P	$\rho_{11}$	****	∆(1232)	$P_{33}$	***	Λ Λ(1405)	$P_{01}$	****	Σ0	$P_{11}$	****	=-	$\rho_{11}$	****
n ***(******)	P <sub>11</sub>	****	$\Delta(1600)$	$P_{33}$	****		$S_{01}$	****	Σ-	$P_{11}$	****	= =(1530)	P <sub>11</sub>	****
N(1440)	$P_{11}$		$\Delta(1620)$	$S_{31}$		A(1520)	$D_{03}$	***		$P_{11}$	****		$P_{13}$	
N(1520)	$D_{13}$	****	$\Delta(1700)$	$D_{33}$	****	A(1600)	$P_{01}$		Σ(1385)	$P_{13}$		≡(1620)		
N(1535)	$S_{11}$	****	$\Delta(1750)$	$P_{31}$	*	A(1670)	$S_{01}$	****	Σ(1480) Σ(1560)			±(1690)	_	***
N(1650)	$S_{11}$		$\Delta(1900)$	$S_{31}$	**	A(1690)	$D_{03}$			_		±(1820)	$D_{13}$	
(1675)	$D_{15}$	****	$\Delta(1905)$	$F_{35}$	****	A(1800)	$S_{01}$	***	Σ(1580)	$D_{13}$		±(1950)		***
(1680)	$F_{15}$	****	$\Delta(1910)$	$P_{31}$	****	A(1810)	$P_{01}$	***	$\Sigma(1620)$	$S_{11}$	***	±(2030)		***
(1700)	$D_{13}$	***	$\Delta(1920)$	$P_{33}$	***	A(1820)	$F_{05}$	****	$\Sigma(1660)$	$P_{11}$		$\Xi(2120)$		*
(1710)	$P_{11}$	***	$\Delta(1930)$	$D_{35}$	***	A(1830)	$D_{05}$	****	$\Sigma(1670)$	$D_{13}$	****	$\Xi(2250)$		**
(1720)	$P_{13}$	****	$\Delta(1940)$	$D_{33}$	*	A(1890)	$P_{03}$	****	$\Sigma(1690)$		**	±(2370)		**
(1900)	$P_{13}$	**	$\Delta(1950)$	$F_{37}$	****	A(2000)			$\Sigma(1750)$	$S_{11}$	***	±(2500)		*
(1990)	F <sub>17</sub>	**	$\Delta(2000)$	F35	**	A(2020)	F <sub>07</sub>	*	Σ(1770)	$P_{11}$				
(2000)	$F_{15}$	**	$\Delta(2150)$	$S_{31}$	*	A(2100)	G <sub>07</sub>	****	$\Sigma(1775)$	$D_{15}$	****	Ω-		****
(2080)	$D_{13}$	**	∆(2200)	$G_{37}$	*	A(2110)	F <sub>05</sub>	***	$\Sigma(1840)$	$P_{13}$	*	$\Omega(2250)^{-}$		***
(2090)	$S_{11}$	*	$\Delta(2300)$	H <sub>39</sub>	**	A(2325)	$D_{03}$	*	$\Sigma(1880)$	$P_{11}$	**	$\Omega(2380)^{-}$		**
(2100)	$P_{11}$		$\Delta(2350)$	$D_{35}$	*	A(2350)	Hoo	***	Σ(1915)	$F_{15}$	****	Ω(2470)-		**
(2190)	$G_{17}$	****	∆(2390)	F37	*	A(2585)		**	Σ(1940)	$D_{13}$	***			
(2200)	$D_{15}$	**	∆(2400)	G20	**				$\Sigma(2000)$	$S_{11}$	*	$A_c^+$		****
(2220)	H19	****	$\Delta(2420)$	$H_{3,11}$	****				Σ(2030)	F <sub>17</sub>	****	$\Lambda_c(2593)^+$		***
((2250)	$G_{19}$	****	$\Delta(2750)$	/3.13	**				Σ(2070)	F <sub>15</sub>	*	$A_c(2625)^+$		***
(2600)	1.11	***	Δ(2950)	K <sub>3.15</sub>	**				Σ(2080)	$P_{13}$	**	$\Lambda_c(2765)^+$		*
(2700)	K <sub>1.13</sub>	**	L(2300)	143,15					Σ(2100)	G17	*	$\Lambda_c(2880)^+$		**
	,.,		Θ(1540)+						Σ(2250)		***	$\Sigma_c(2455)$		****
			0(1510)						Σ(2455)		**	$\Sigma_c(2520)$		***
									E(2620)		**	$\Sigma_c(2800)$		***
									Σ(3000)		*	Ξ+		***
			1						Σ(3170)			=0		***
									2(0110)			='+		***
			1									=0		***
			1									= <sub>c</sub> ≡ <sub>c</sub> (2645)		***
												$\Xi_c(2045)$ $\Xi_c(2790)$		***
														***
												$\Xi_c(2815)$ $\Omega_c^0$		***
												12°C		
												Ξ+,		*
												100		
												=0. =-		

# all can be described by patterns of 6 "quarks"

#### called "generations"



# rest of matter described by patterns of 3 "electrons" and neutrinos

#### called "generations"

#### why?

are the quark patters like the lepton patterns?

are the masses arranged?

are the charges arranged?

are there 3 generations?

dunno

#### jargon

"Fermions": n/2 spins; "Bosons": integer spins

#### particles which interact via strong force:

"Hadrons"

which can be spin n/2 ("baryons") [proton, neutron,  $\Delta$ ,  $\Sigma$ ,  $\Xi$ ,  $\Omega$ , ...] or spin 0, 1 ("mesons") [pion, kaon,  $\eta$ ,  $\rho$ ,...]

#### particles which do not interact via the strong force:

"Leptons"

which are all spin 1/2 [e,  $\mu$ ,  $\tau$ ,  $\nu$ e,  $\nu_{\mu}$ ,  $\nu_{\tau}$ ]

#### And, fields which propagate the 4 forces:

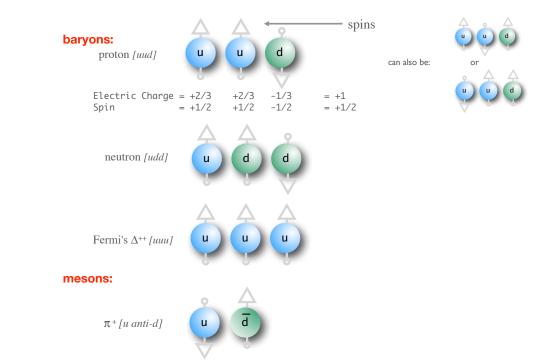
Electromagnetic: photon

Weak: "W boson"
Strong: "gluon"

Gravity: "graviton"... speculated

#### And a surprise...conceived in 1967.

#### How to build a proton: (originally) hypothetical building blocks of hadrons



#### 1969: Discovery of partons, Kendall, Freedman, and Taylor

Lo and behold: quarks were found rattling around inside of the protons and neutrons

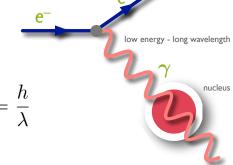
an electron beam emits a photon with a wavelength inversely proportional to its energy

> scattering from whole nuclei, required photons of wavelengths ~ nuclear sizes

INCREASE the energy of the electrons - smaller sizes are resolvable

and an amazing thing happened:

individual protons and neutrons were broken up,



and the pattern of Rutherford scattering emerged!

Something was inside and was much, much smaller than the proton, and the higher energy photons dug in and scattered directly from these point-like objects

called PARTONS by Feynman

They indeed turned out to be quarks.

#### 

#### **Connection with electrodynamics**

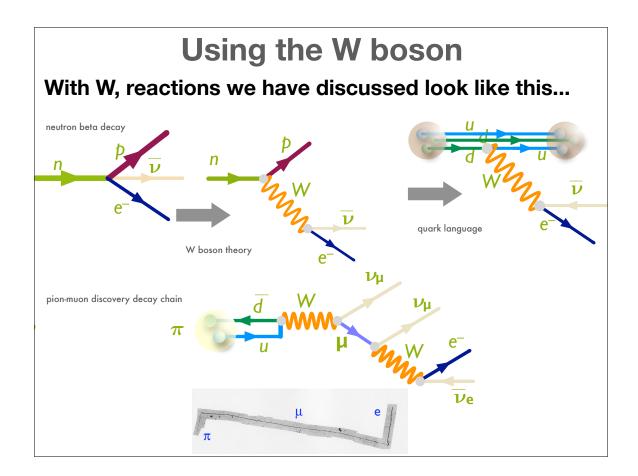
# In the late 1950's, Feynman and Gell-Mann put this idea on a mathematical footing for the Weak Interactions

...that the weak force also was propagated by a spin-1 boson but it was different from the photon.

#### It had to be:

- 1.electrically charged (it changes  $n \rightarrow p$ , so Q goes from 0 to 1)
- 2.very massive (the weak force is propagated over short distances)
- 3.capable of changing isospin...ie, it can change one particle into another within isospin families (within other conservation requirements)

dubbed the "W Boson"



#### so: the force carriers:

connect the quarks and leptons to one another

according to strengths

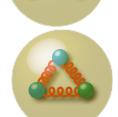
strong interaction: the gluon

#### so, what's a proton?

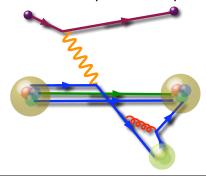
#### A collection of three quarks...u, u, d... What keeps the positive u quarks from repelling one another?

the strong force...the one associated with Yukawa's particle: but not the pion, a new "photon" that transmits the strong force called the "gluon"

the gluons themselves are moving and spitting out new particles...quarks...







We scatter from all of these quarks with very high-energy particles...and produce the gluons all the time

#### the issue:

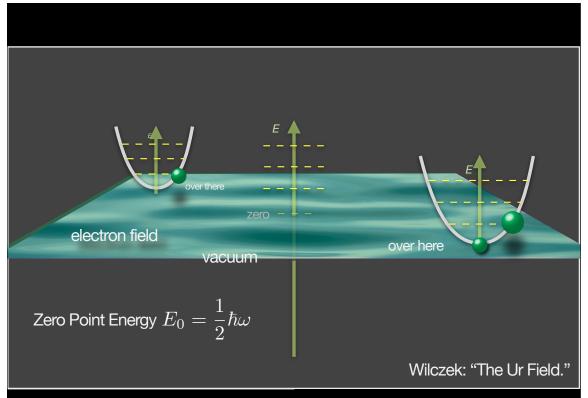
mass.

## question:

what's a "particle"?

## harmonic excitations of a





**UI-**  $|O(\partial)\Gamma|$   $|O(\partial)\Gamma|$  combining form primitive; original; earliest : *urtext.* ORIGIN from German.

# identity!





# 1967



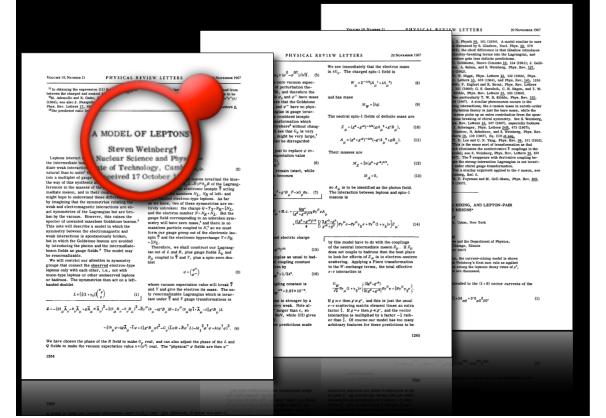
http://www.mustangdreams.com/mdfastback.htm

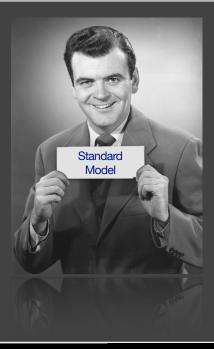


http://nobelprize.org/nobel\_prizes/physics/laureates/1979/weinberg-autobio.html



ttp://hacks.mit.edu/Hacks/by\_year/2006





# Standard Model

standard | 'standərd |

model |'mädl|

noun

1 a level of quality or attainment 2 ...a simplified descr

 ${\bf 2}$  ...a simplified description, esp. a mathematical one, of a system or process, to assist calculations and predictions

a model of leptons and quarks....

$$\begin{pmatrix} \nu_e \\ e \end{pmatrix}_L \begin{pmatrix} \nu_\mu \\ \mu \end{pmatrix}_L \begin{pmatrix} \nu_\tau \\ \tau \end{pmatrix}_L \nu_{eR}, e_R, \nu_{\mu R}, \mu_R, \nu_{\tau R}, \tau_R$$

$$\begin{pmatrix} u \\ d \end{pmatrix}_L \begin{pmatrix} c \\ s \end{pmatrix}_L \begin{pmatrix} t \\ b \end{pmatrix}_L u_R, d_R, c_R, s_R, t_R, b_R$$

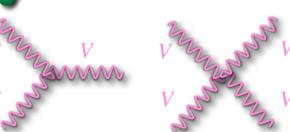
and forces....

$$W^{\pm}, Z^0, \gamma, g$$

the weak, electromagnetic, and strong interactions

and dynamics....





#### astonishing level of understanding

$$Z^{0} \to e + e; \ \mu + \mu, \ \tau + \tau, \ \nu + \nu \\ \to u + u, \ d + d, \ s + s, \ c + c, \ b + b \\ W^{\pm} \to e + \nu, \ \mu + \nu, \ \tau + \nu \\ \to u + d, \ u + s, \ u + b \\ + g, + 2g... + \sum_{n=1}^{\infty} g$$

$$d \rightarrow u, \ell, \nu$$

$$s \rightarrow d, u, \ell, \nu$$

$$c \rightarrow d, u, \ell, \nu$$

$$b \rightarrow d, u, \ell, \nu$$

$$t \rightarrow W + b$$

Including precise measurements of:

- masses
- branching fractions,
- mixing
- production cross sections

 $\nu \rightarrow \nu'$ 

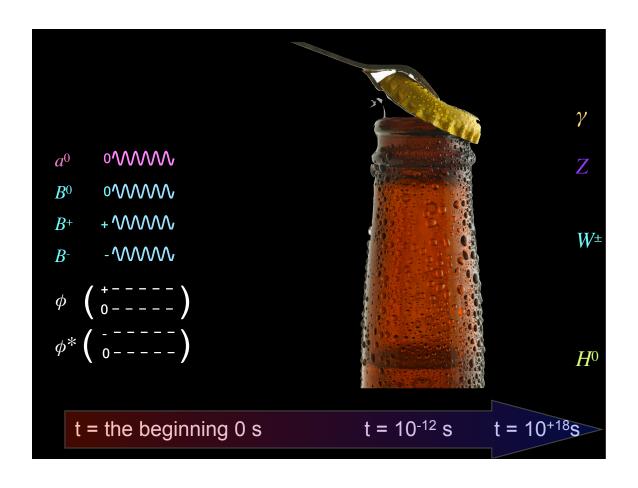
Odd man out: neutrino sector

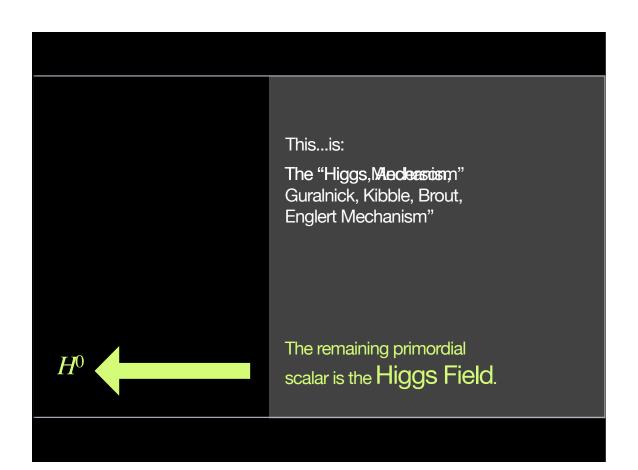
# SM is an effective theory

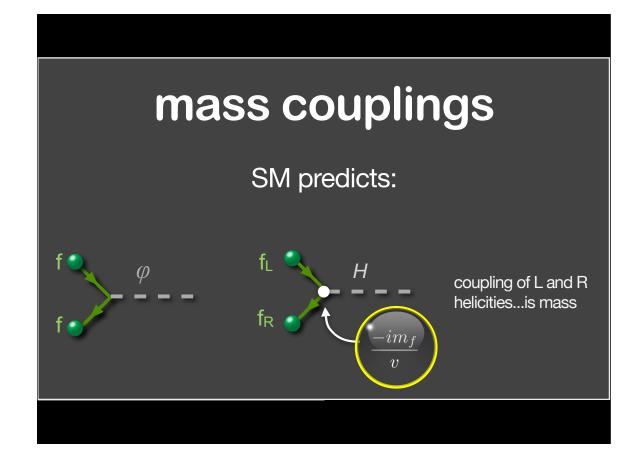
pretty damn good.

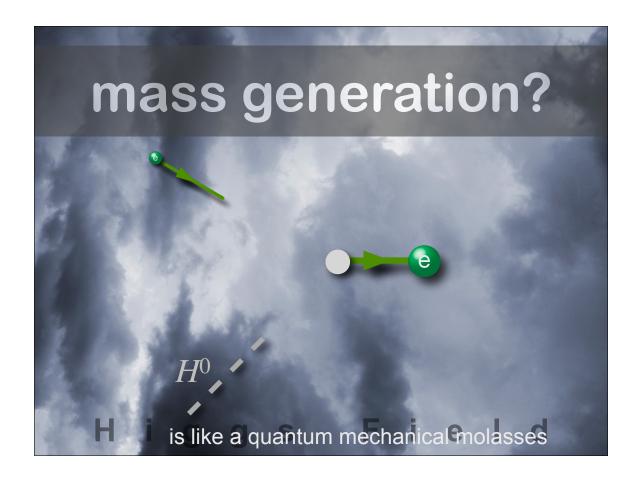
Quantity	Value	Standard Model	Pull
$m_t$ [GeV]	$172.7 \pm 2.9 \pm 0.6$	$172.7 \pm 2.8$	0.0
$M_W$ [GeV]	$80.450 \pm 0.058$	$80.376 \pm 0.017$	1.3
arw [Gev]	$80.392 \pm 0.039$	00.370 ± 0.017	0.4
$M_Z$ [GeV]	$91.1876 \pm 0.0021$	$91.1874 \pm 0.0021$	0.1
$\Gamma_Z$ [GeV]	$2.4952 \pm 0.0023$	$2.4968 \pm 0.0011$	-0.7
Γ(had) [GeV]	$1.7444 \pm 0.0020$	$1.7434 \pm 0.0010$	
Γ(inv) [MeV]	$499.0 \pm 1.5$	$501.65 \pm 0.11$	
$\Gamma(\ell^+\ell^-)$ [MeV]	$83.984 \pm 0.086$	$83.996 \pm 0.021$	
σ <sub>had</sub> [nb]	$41.541 \pm 0.037$	$41.467 \pm 0.009$	2.0
$R_e$	$20.804 \pm 0.050$	$20.756 \pm 0.011$	1.0
$R_{\mu}$	$20.785 \pm 0.033$	$20.756 \pm 0.011$	0.9
$R_{\tau}$	$20.764 \pm 0.045$	$20.801 \pm 0.011$	-0.8
$R_h$	$0.21629 \pm 0.00066$	$0.21578 \pm 0.00010$	0.8
R.	$0.1721 \pm 0.0030$	$0.17230 \pm 0.00004$	-0.1
$A_{FB}^{(0,e)}$	$0.0145 \pm 0.0025$	$0.01622 \pm 0.00025$	-0.7
$A_{FB}^{(0,\mu)}$	$0.0169 \pm 0.0013$		0.5
.(0.x)	$0.0188 \pm 0.0017$		1.5
$A_{FB}^{(0,b)}$ $A_{FB}^{(0,b)}$	$0.0992 \pm 0.0016$	$0.1031 \pm 0.0008$	-2.4
$A_{FB}^{(0,c)}$ $A_{FB}^{(0,c)}$	$0.0707 \pm 0.0035$	$0.0737 \pm 0.0006$	-0.8
$A_{FB}^{(0,s)}$	$0.0976 \pm 0.0114$	$0.1032 \pm 0.0008$	-0.5
$\bar{s}_{\ell}^{2}(A_{FB}^{(0,q)})$	$0.2324 \pm 0.0012$	$0.23152 \pm 0.00014$	0.7
-£(FB)	$0.2238 \pm 0.0050$		-1.5
$A_c$	$0.15138 \pm 0.00216$	$0.1471 \pm 0.0011$	2.0
	$0.1544 \pm 0.0060$		1.2
	$0.1498 \pm 0.0049$		0.6
$A_{\mu}$	$0.142 \pm 0.015$		-0.3
$A_{\tau}$	$0.136 \pm 0.015$		-0.7
	$0.1439 \pm 0.0043$		-0.7
$A_b$	$0.923 \pm 0.020$	$0.9347 \pm 0.0001$	-0.6
$A_c$	$0.670 \pm 0.027$	$0.6678 \pm 0.0005$	0.1
$A_s$	$0.895 \pm 0.091$	$0.9356 \pm 0.0001$	-0.4
$g_{\vec{k}}^2$	$0.30005 \pm 0.00137$	$0.30378 \pm 0.00021$	-2.7
$g_R^2$	$0.03076 \pm 0.00110$	$0.03006 \pm 0.00003$	0.6
$q_V^{\nu e}$	$-0.040 \pm 0.015$	$-0.0396 \pm 0.0003$	0.0
$g_A^{\nu e}$	$-0.507 \pm 0.014$	$-0.5064 \pm 0.0001$	0.0
$A_{PV}$	$-1.31 \pm 0.17$	$-1.53 \pm 0.02$	1.3
$Q_W(Cs)$	$-72.62 \pm 0.46$	$-73.17 \pm 0.03$	1.2
$Q_W(Tl)$	$-116.6 \pm 3.7$	$-116.78 \pm 0.05$	0.1
$\frac{\Gamma(b \rightarrow s\gamma)}{\Gamma(b \rightarrow Xe\nu)}$	$3.35^{+0.50}_{-0.44} \times 10^{-3}$	$(3.22 \pm 0.09) \times 10^{-3}$	0.3
$\frac{1}{2}(g_{\mu} - 2 - \frac{\alpha}{\pi})$	$4511.07 \pm 0.82$	$4509.82 \pm 0.10$	1.5
$\tau_T$ [fs]	290:89 ± 0:58 1	$291.87 \pm 1.76$	-0.4
$\tau_{\tau}$ [fs]	290789 E-0788 1	$^{1.05}$ $291.87 \pm 1.76$	-0.4
	$4511.07 \pm 0.82$	$4509.82 \pm 0.10$	

The Review of Particle Physics W.-M. Yao et al., Journal of Physics, G 33, 1 (2006)









#### Masses come from the Higgs Mechanism. An analogy:

a room full of people, randomly oriented talking...

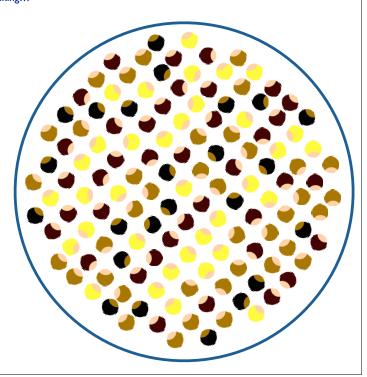
the noise (energy level) in the room is constant and forms a background (ground state) energy which is largely ignored by each member in his individual conversation

There is no ordering to the orientation of the people - a highly symmetric configuration

The room is the vacuum. The people are collectively a **higgs field.** 

...the ground state energy level is unimportant and tuned out

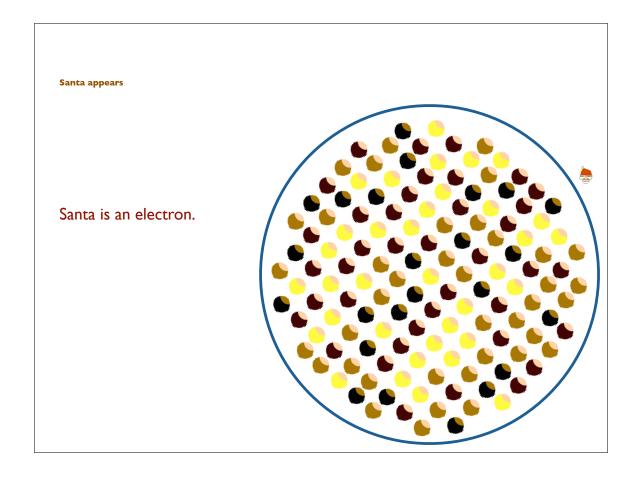
(after David Miller)



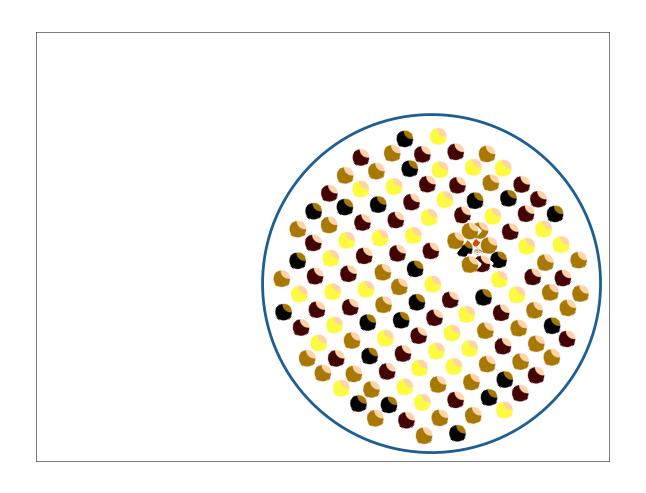
sleigh bells ring from the NE corner of the room...

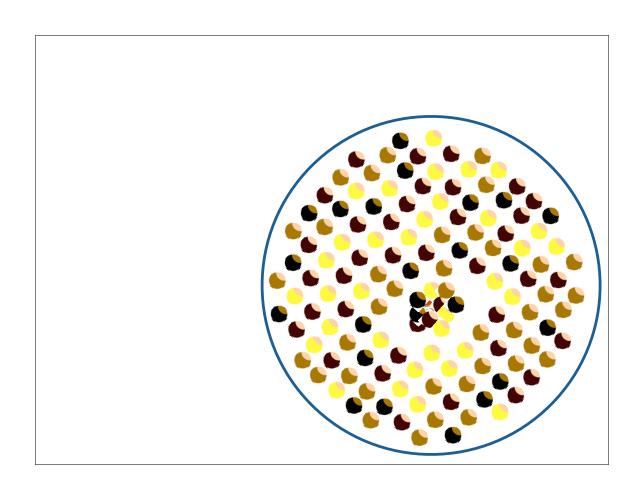
everyone turns to the sound and is immediately silent - so, the energy level in the room is lowered when the randomness is removed

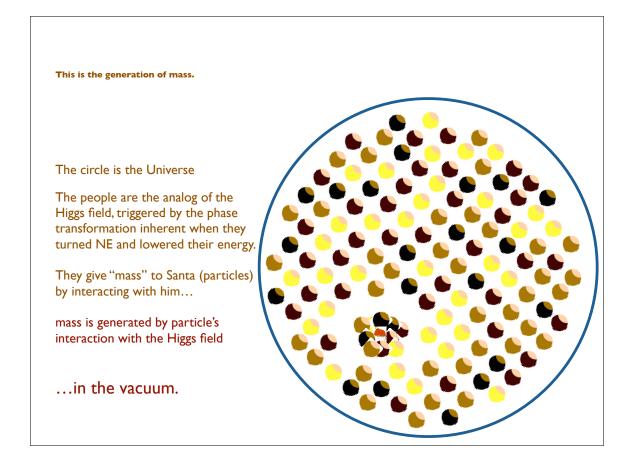
The earlier symmetry is "broken"

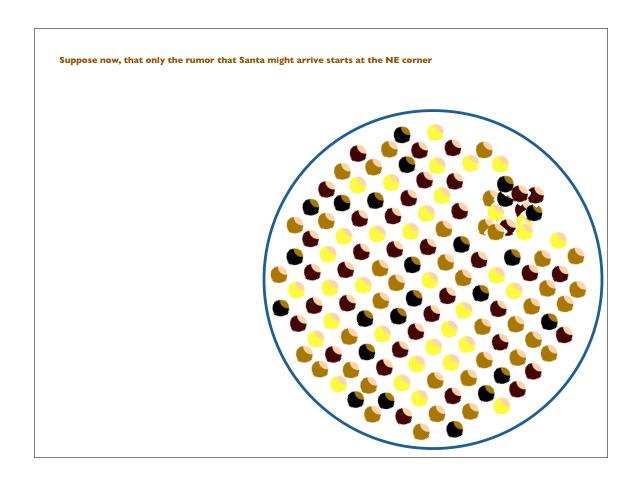


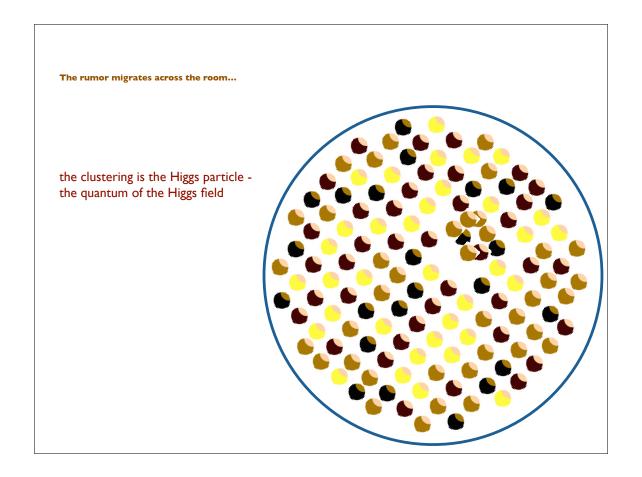
his progress is impeded as if his inertia has increased ie, as if his mass has increased by virtue of his interaction with the people around him.

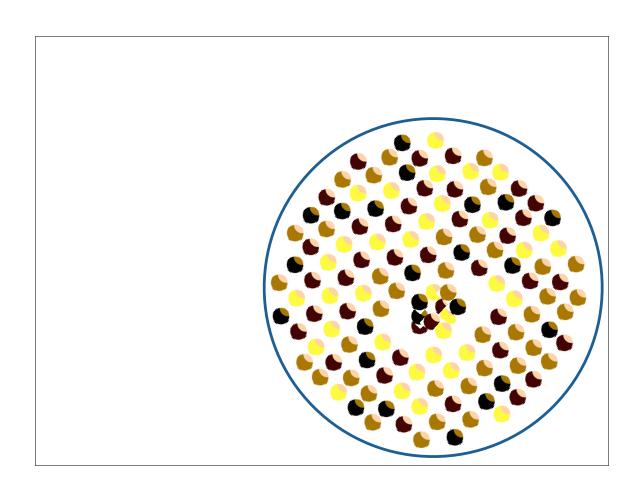












Finding the Higgs Boson is of fundamental importance to confirming the Standard Model and the veracity of the phase transition

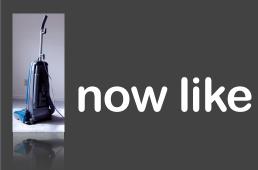


## relativistic quantum field theory

says

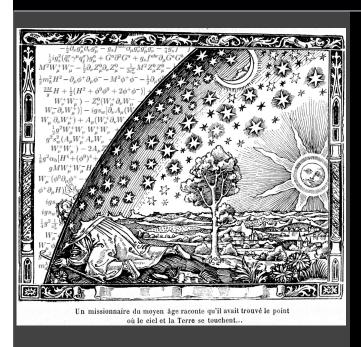


is seething.



that drawer you all have at home





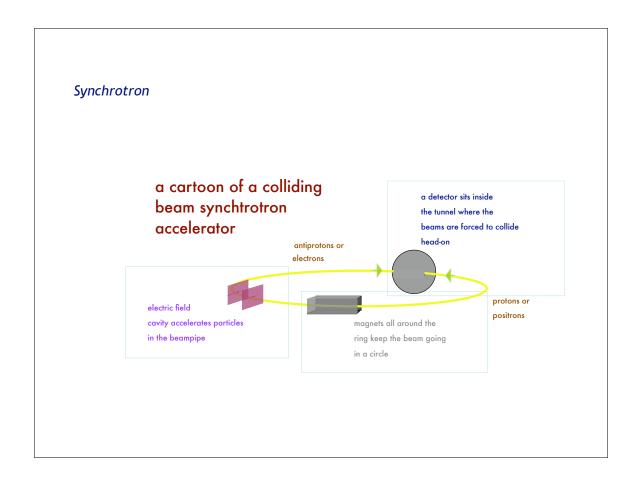
# the deeper...

the Standard Model fails...

be there dragons?

after: Camille Flammarion, L'Atmosphere: Météorologie Populaire (Paris, 1888), p. 163.

## accelerators



#### what is Fermilab?



### it's many things to me...

## it's a dedicated scientific community made up of:

- 1200 physicists, engineers, and staff
- >1000 faculty, post docs, and students
- from > 80 US & ~20 foreign institutions

#### it's an amazing scientific instrument

#### consisting of:

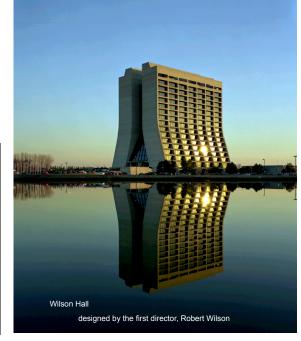
- A time machine
- A particle accelerator for antirotating beams of protons and antiprotons
- hand-made vehicles to explore the current and the very early universe
- A source of high energy/intensity beams of kaons and neutrinos

#### it's a beautiful single-purpose DOE national lab

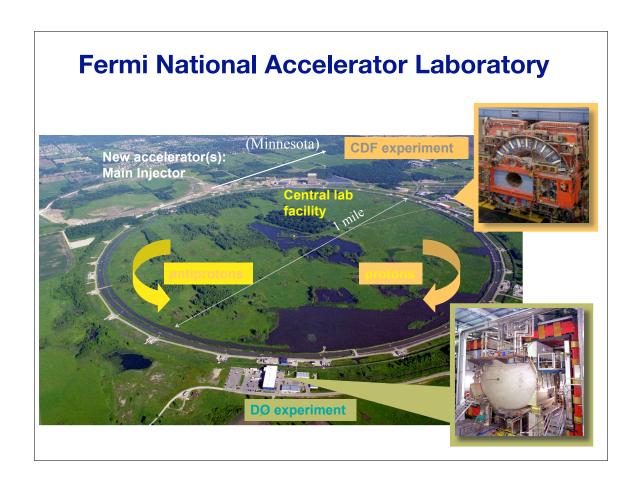
#### located at:

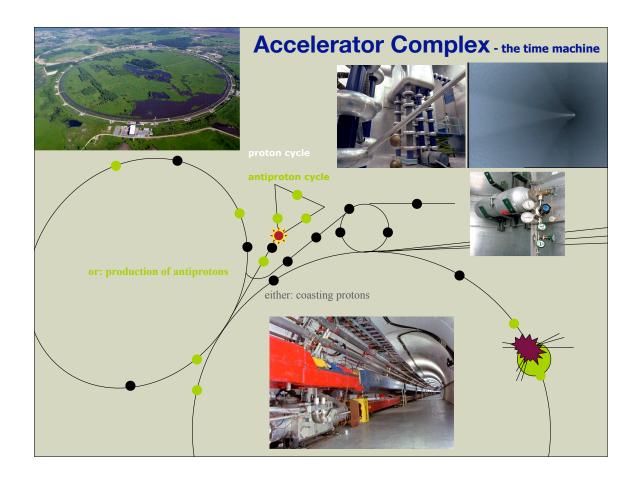
- real space: 60 mi west of Chicago
- cyberspace: www.fnal.gov

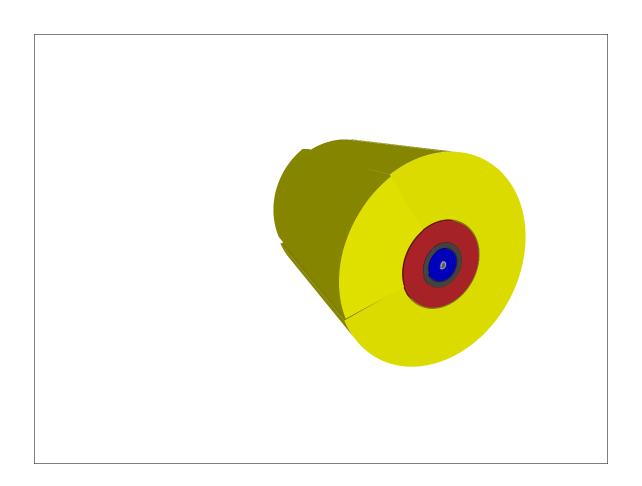
## an inspiring place to work

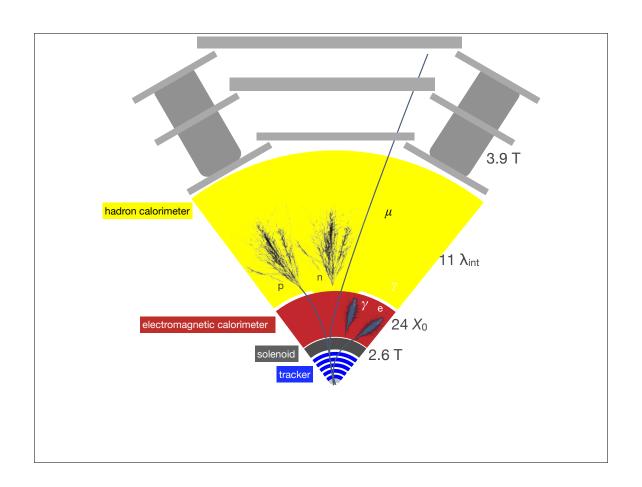












## the DØ detector: inside



## the DØ detector: outside







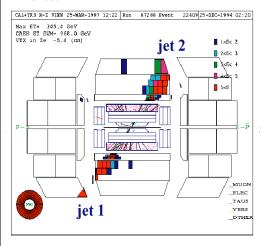


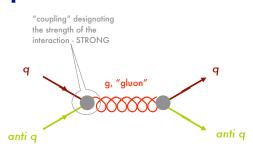


## most violent elementary particle collision produced on earth

Rutherford Scattering of one quark in the proton off of another quark from the antiproton

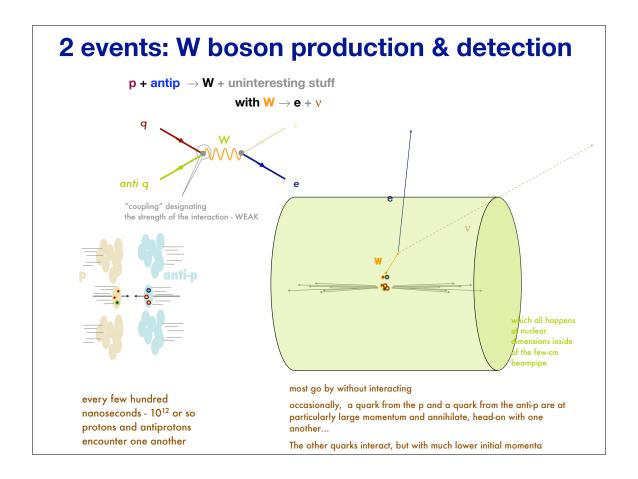
with the exchange of a "gluon" a photon-like particle that transmits only the STRONG force.

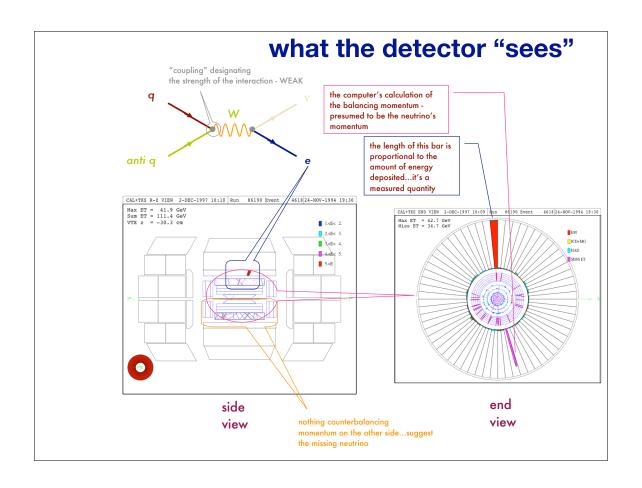


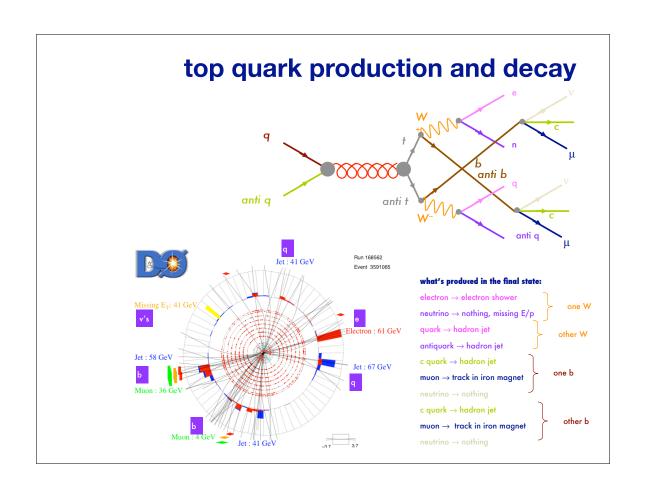


It required that the quarks annihilated within 10<sup>-19</sup> m of one another or 1/10,000 the size of a proton

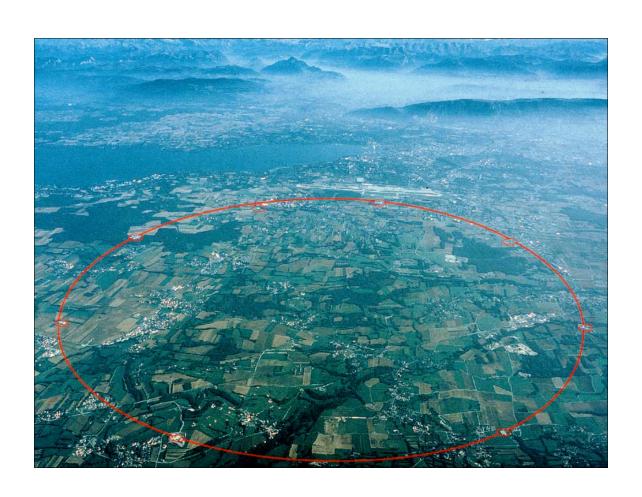
The energetics of this event is consistent with interactions in the early universe ~10-20 s after the big bang

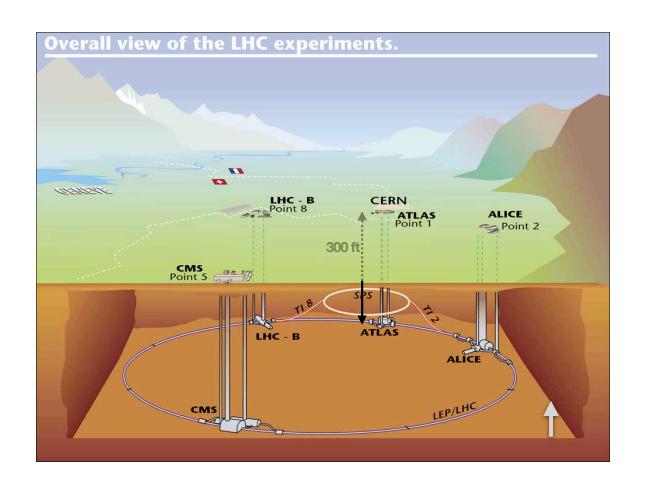


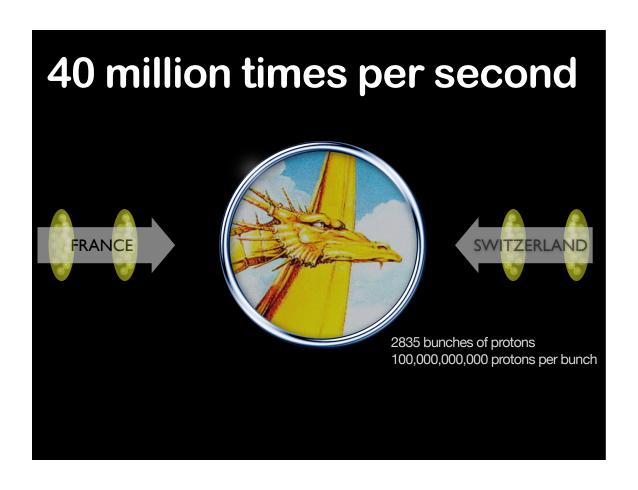




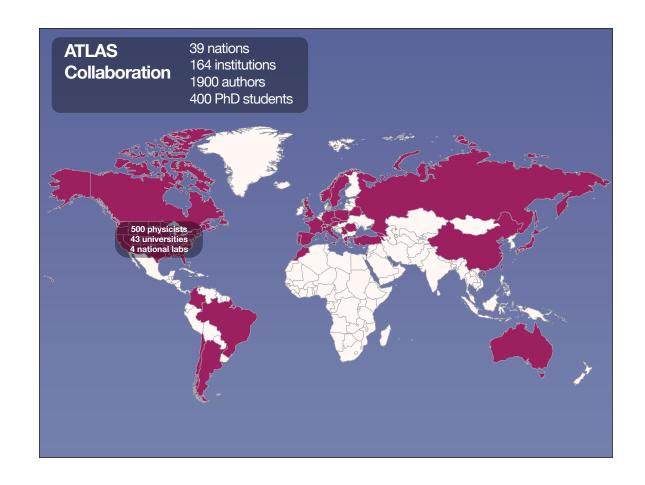
## the Large Hadron Collider

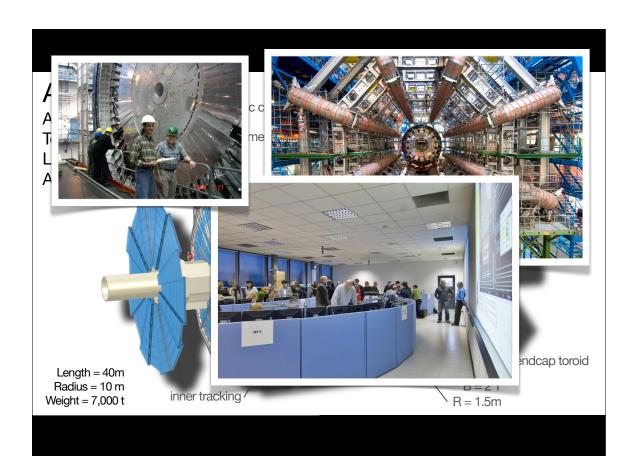




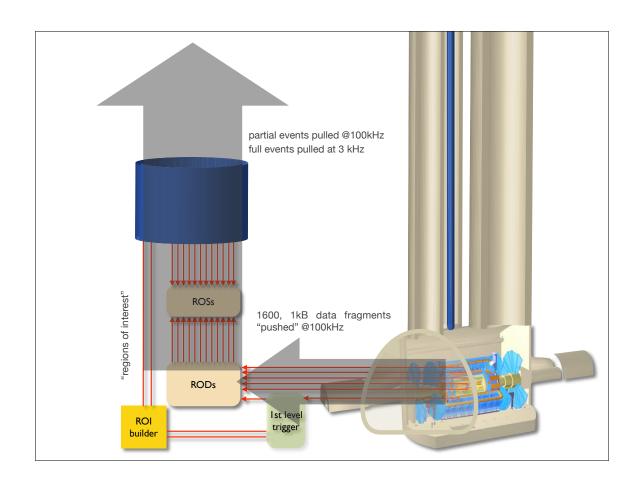


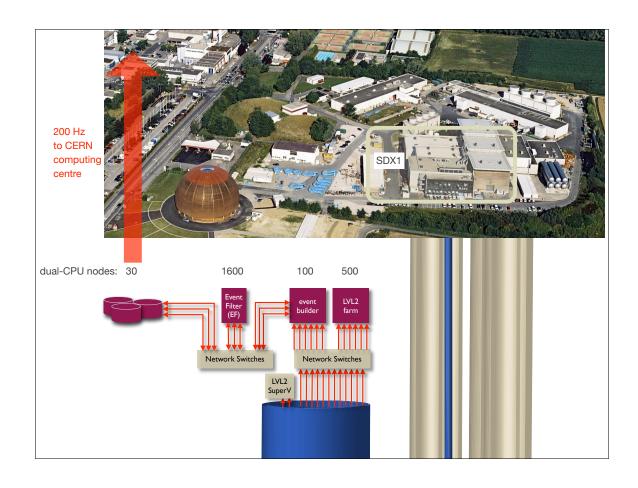










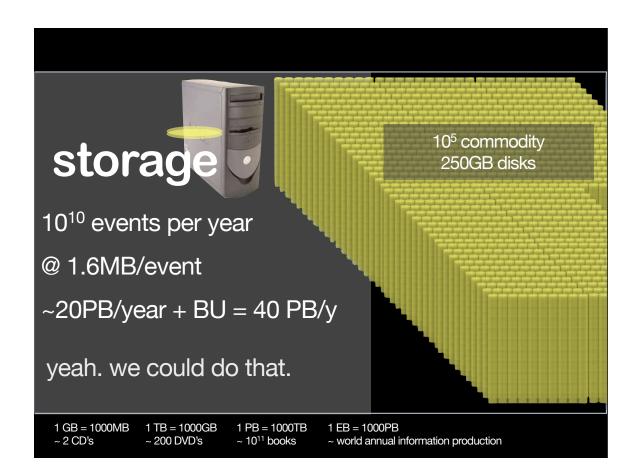


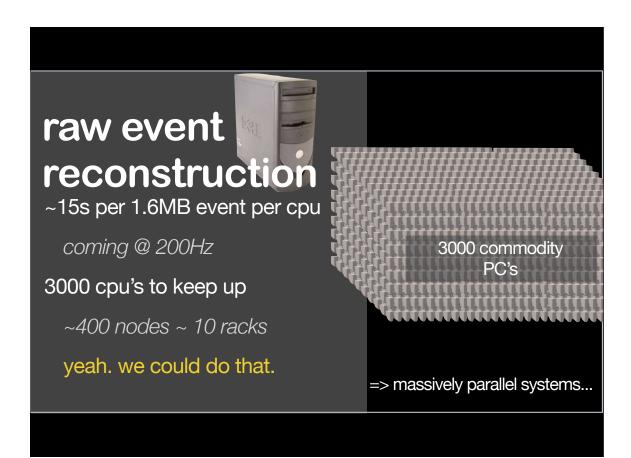
## now what?

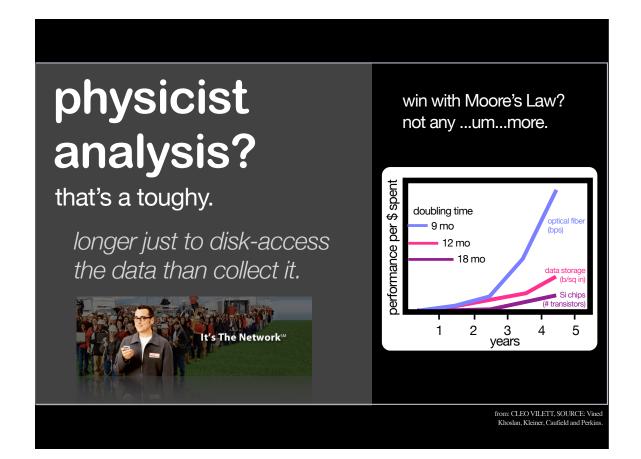
gotta store it. gotta reduce it gotta analyze it.

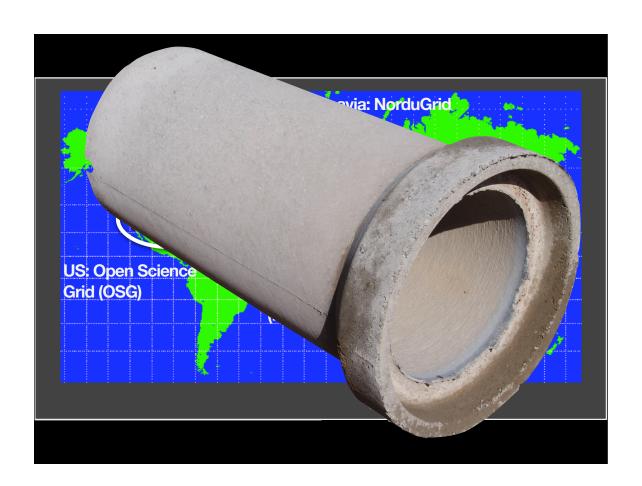
keep that up for 2 decades.

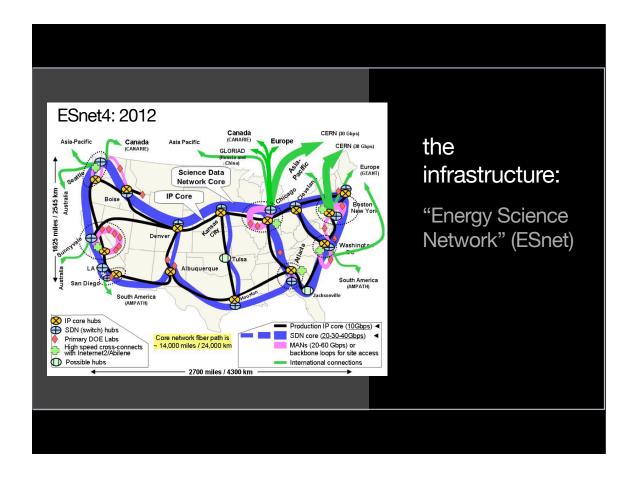


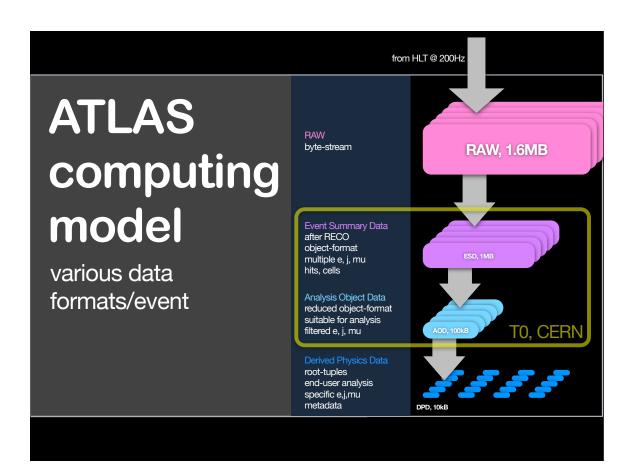


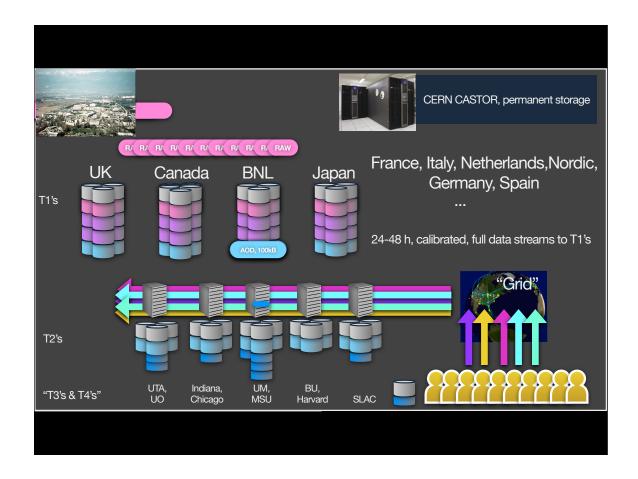


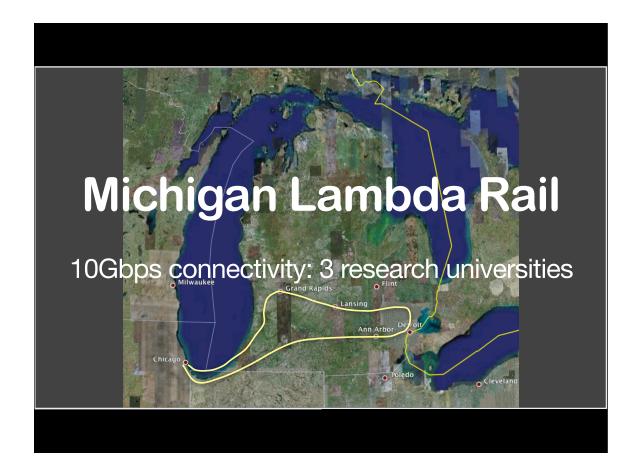


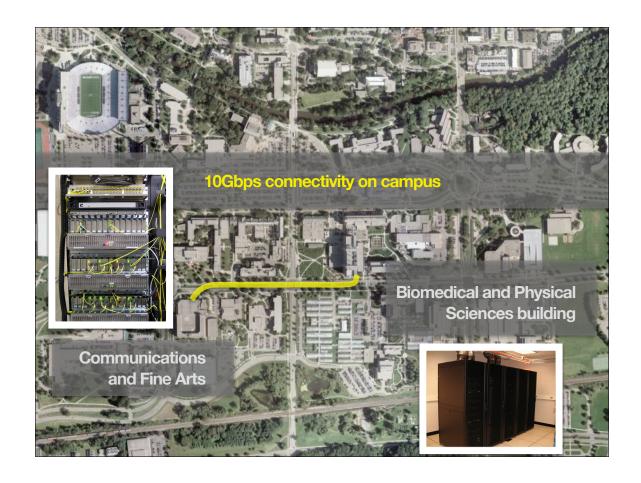


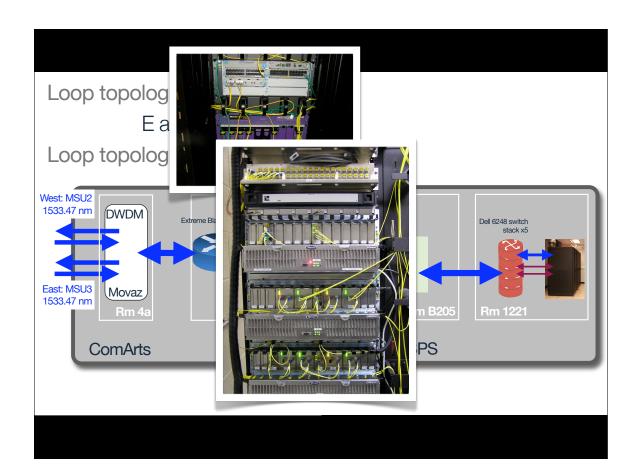


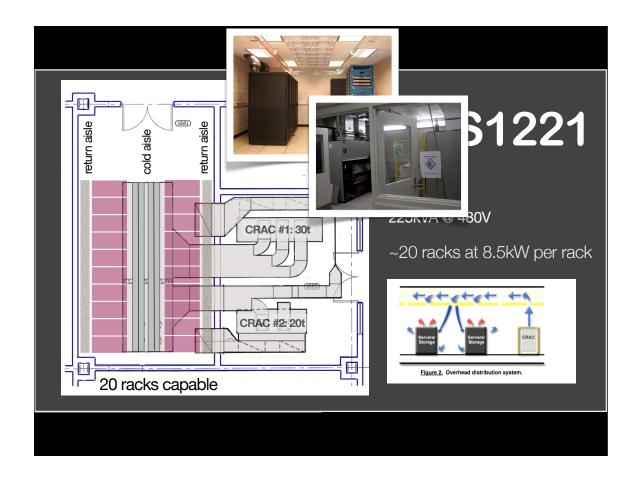












### 5 Racks running since December

#### Hardware:

#### computing:

Dell Poweredge 1950

54 nodes Intel Xeon 5355, 2.67 GHz dual, quad core

(SPECint\_2000: ~2178/cpu => ~17,424/node => 940k SPECint2000)

so...432 cpus. will double by end of summer.

I plan for about 1500 cpu's

#### storage:

PowerVault MD1000

225,000GB storage in 5 shelves

## just arrived for "T3":

40 nodes & 15TB



## will it work?

probably but not as designed

gridifying people is much harder than gridifying computers!



# This'll keep us busy for 15 years. • We have to find the Higgs Boson...or everything that's worked before crashes • We have to find new physics... or everything that's worked before crashes and wouldn't that be neat? Cosmology **Big Bang** Microwave background **Measuring the Geometry of the Universe** uh oh

## What about...The Universe?

#### The first, most remarkable thing perhaps is that it's expanding

- We've seen Hubble evidence for that, but there's a trickier and more subtle piece of evidence
- If, by extrapolating backwards, everything in the universe started at a fixed time and from a tiny, even infinitesimal size - 10 Billion years ago

#### BUT, this is a state of the whole universe, spacetime and matter/energy

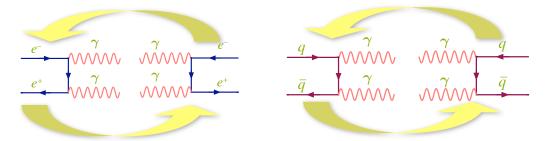
- there is not an empty universe into which stuff expanded
- rather, the spacetime fabric of the universe expands with matter/energy
- So, the conditions at that time must have been very hot.

so hot, that only elementary particles and radiation could have existed

a soup of energetic quarks, leptons, and the precursors to the photon, sibling massive particles, W and Z

#### These conditions

 fundamental interactions happen constantly...in equilibrium between radiation and matter things like:

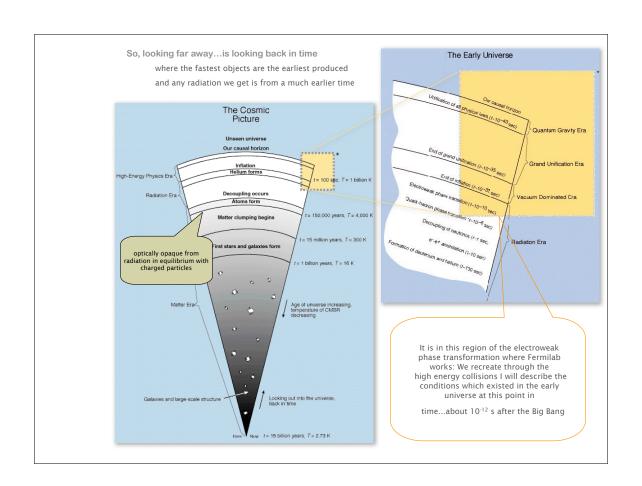


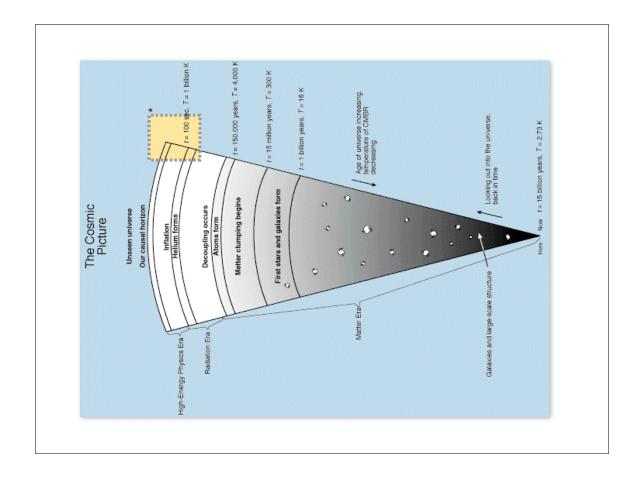
## further... is earlier



 $t_1$ 

At an early time, two objects form Light from 2 spreads out





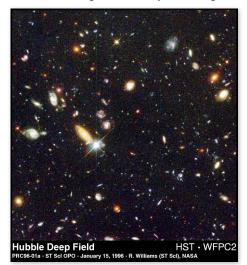
## the universe is bigger than...um.. anything

#### One of the more stunning results of the HST

- else • It was trained on a small patch of sky in the little dipper for 10 days previously cleared as not having any major, nearby galaxies that would be in the way
- a patch of sky the size of a dime, seen from 75' away • This limit corresponds to looking back in time 10 billion years

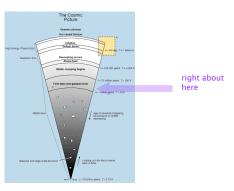
when the universe was 1/16 of its current age

to periods in which the galaxies were just starting to form, a billion y from the big bang



Roughly 1500 galaxies in this patch, up to this

100's of billions of galaxies in the universe as a

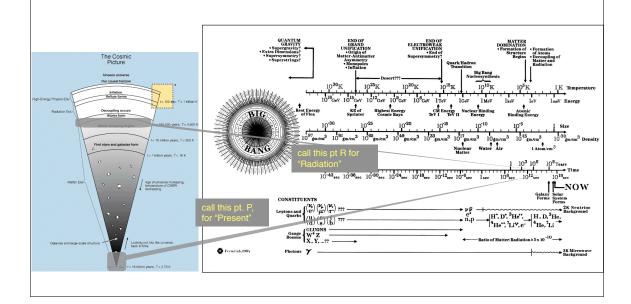


## warm glow...actually, a cool glow

#### A little earlier

At some point the universe cooled enough that the equilibrium processes stop and the ratio of matter to radiation becomes fixed

We have a snapshot of the conditions when that happened and a test of expansion:

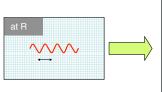


## Cosmic Microwave Background (CMB) Radiation

#### The radiation that is left has since cooled...a lot

· but it's not the kind of cooling that you might think of

It's because the "measuring stick" for wavelength has stretched through the expansion of spacetime



at P

lengths, short

so, wavelengths short,

or...frequency high,

or...energy high,

or...temperature high

lengths, longer

so, wavelengths longer,

or...frequency lower,

or...energy lower,

or...temperature lower

#### The prediction was:

the wavelengths should be about 7.35cm, which corresponds to a frequency of 4000MHz...or a temperature of ~3K

and that it should be everywhere...all

## seeing that and seeing as...

## In 1965, Arno Penzias and Robert Wilson were trying to get rid of noise in a large radio telescope (designed to study satellite communications at AT&T)

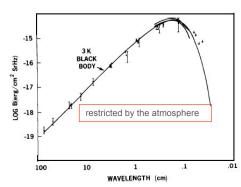
 They couldn't and began to believe that it was coming from outer space, all day, all night, from all directions

Rather quickly it was learned that this was precisely what he just been predicted at Princeton by James Peebles

#### In the 1940s, George Gamow had predicted

 that the big bang radiation would have indeed cooled and that it would have the characteristics of a black body



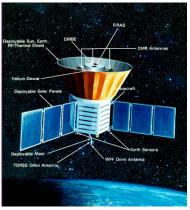


From Wilson's 1978 Nobel Lecture

## The Cosmic Background Explorer (COBE)

Specifically, a mission launched in 1989 to measure the CBM...and it's uniformity - the hot soup must start out uniform...BUT

- ...I mean, clumps do happen (us, Milky Way, etc)
- Now, it's incredibly precise...this plot has data points with error bars
   COBE measured E&M radiation as a function of frequency outside of the earth's atmosphere



showing precisely the blackbody spectrum for a temperature of 2.726K

Wavelength [cm]

10 10 1 0.1

FIRAS COBE satellite

DMR COBE satellite

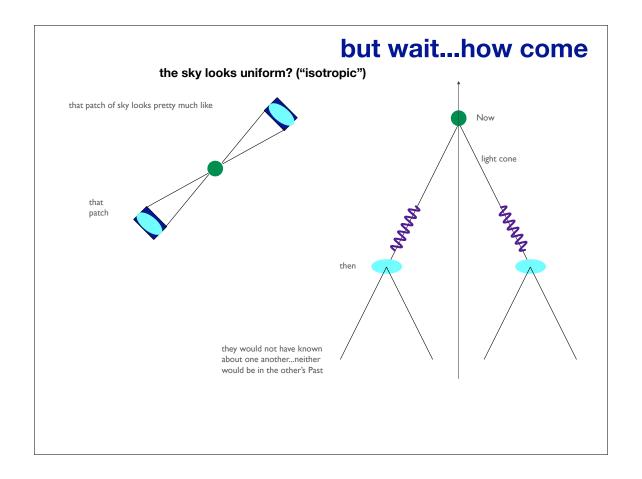
LBL - Italy White Mm & South Pole

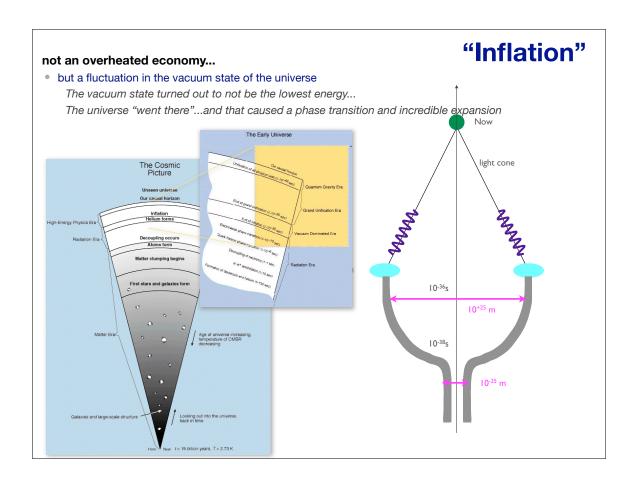
Princeton gounding rocket

Cyanogen optacl

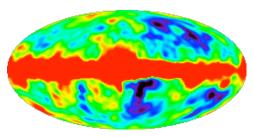
10 10 100

Frequency [GHz]

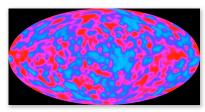








2.7279 (blue)-2.7281 (red) K x 25000Zoom



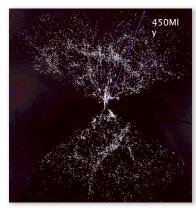
fluctuations are  $0.000030\ K$ 

This is incredible.

It is decisive evidence that the Big Bang model is correct and...the ripples are primordial density fluctuations consistent with that required to form galaxies.

An all-sky image (like a Mercator projection) of the sky...notice the Galactic halo across the midline...

Then this large-scale structure is digitally removed...

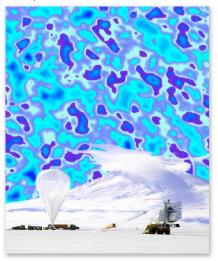


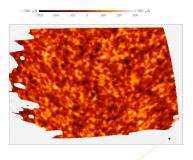
And, that's true. There is non-random structure: These filament-like strands are combinations of 11,000 galaxies (MW at the center).

### more measurements:

A balloon experiment, launched from the South Pole in 1998 has done even more precise measurements of the CMB

 BOOMERanG (Balloon Observations of Millimetric Extragalactic Radiation and Geomagnetics)





That's the moon's disk size for comparison

## so, then if it started, does it end?

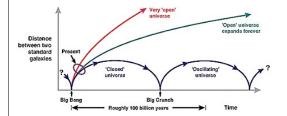
#### The evidence that the Big Bang occurred is overwhelming

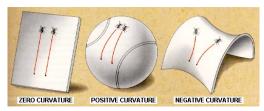
Olber's Paradox

#### What is the future of the universe?

- Thought to depend on the balance between expansion and gravitation that depends on how much mass/energy there is
- Two general ways to determine that count the stuff: how much stuff/light can we see and account for? measure spacetime: what clues do we have to the geometry of the universe?

These are the 3 extremes of geometries for spacetime: a totally closed (in which C<2 $\pi$ r) flat (C=2 $\pi$ r), and open (C>2 $\pi$ r)...parallel lines?





## for the universe: density⇔geometry

## critical density...from General Relativity

$$\rho_c = \frac{3H^2}{8\pi G}$$

$$\rho_{\text{critical}} = (0.97 \pm 0.12) \text{ x } 10^{-29} \text{ g/cc}$$

$$\Omega = \frac{\rho_{measured}}{\rho_c}$$

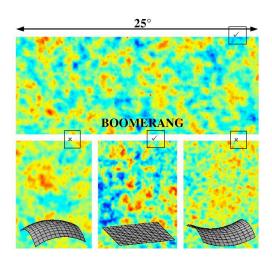
$$\Omega=1$$
  $\Omega>1$   $\Omega<1$ 

## we can now measure the Geometry of the Universe

#### By comparing the anisotropy of the CMB to computer simulations

• the geometry of the universe is determined to be very close to  $F \perp A T$  (remember, in 4 spacetime dimensions)

The characteristic quantity for flatness is  $\Omega = 1$  means flat.



Lots is in play here...Spacetime will be bent and shaped by the amount of mass in the universe and how it's distributed

This is usually expressed in terms of  $\Omega$  the deviation from flatness. If  $\Omega$  = 1, then the universe is flat.

From these measurements,

 $\Omega_{\text{geometry}} = 1$ 

But, it gets worse.

### a little short

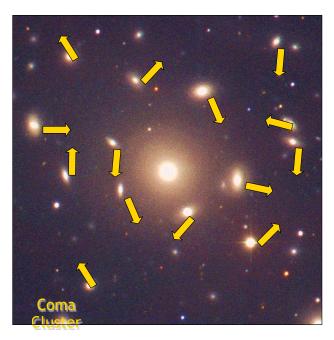
## Everything that has mass/energy would contribute to $\Omega$ So, add up everything that's visible:

- The various visible contributions to the critical density  $\Omega_{\text{luminous matter}}$  are: Visible stuff (stars, dust, gas) from telescopes  $\Omega_{\text{luminous matter}} \approx 0.01$ Estimation of other regular matter..deuterium abundance & theory  $\Omega_{\text{luminous matter}} \approx 0.05$
- But: from measurements  $\Omega_{\text{geometry}} = 1$

Meaning...that more than 94% of the mass/energy in the universe is unaccounted for by what we know as normal matter/radiation.

But, it gets worse.

## galactic boogie



Galaxies are moving...should be consistent with their masses

it's not...

## So, there is lots of stuff...

#### BUT, the universe is lacking in another way:

Measurements of the motion of galaxies and other large-scale gravitational motions suggest
there is not enough stuff that we can see with any telescope (optical, radio, microwave, infrared, etc) to account
for the gravitational pulls that would account for these large scale motions. - the Coma Cluster is 400x too light to
be consistent with its motions

There is something missing in the universe that specifically influences large-scale dynamics - The Missing Mass Problem

Our paradigm: All manifestations of energy/mass are in the form of elementary "particles"

So...We expect that there is some yet-to-be-observed elementary particle: "Dark Matter" created in large abundance in the Big Bang

weakly interacting so as to have not scattered or interacted with "regular" matter, so that it's still around

Trying to produce it in the laboratory, nor has it been possible to detect it passing through the earth - WIMP's

•The accounting suggests that there must be a contribution... $\Omega_{\text{dark matter}} \sim 0.25$ 

This means that there must be some large amount of mass/energy in the universe which will affect the way in which it expands

• So, adding it up:  $\Omega_{\text{dark matter}} + \Omega_{\text{luminous matter}} \sim 0.3$ 

• but  $\Omega_{\text{geometry}} = 1$ 

Will we end in a whimper? Or a big crunch?

That is, will the universe continue to expand, and not be slowed down by all of the mass? Or, will the mass be sufficient to slow it down, bring it to a halt, and then cause contraction?

All of the observed and dynamically required matter lead to a poor accounting of  $\Omega$ 

But, it gets worse.

## Supernovae

#### Remember, the determination of the Hubble Constant led to all of this:

 In order to make this measurement, one needs the ability to measure how far away something is and determine it's velocity

This was done with a variety of "standard candles", such as Cepheid Variable Stars

#### Well, there's a new candle on the block:

A particular kind of supernova

A star is held together through two competing forces:

the nuclear fuel in the core is "burned" in fusion reactions which tend to increase the mass of the center...and hence increase the gravitational pull

Radiation from the nuclear reactions causes a pressure out stabilizing the gravitational forces in

A supernova is a star which has depleted its nuclear fuel & gravitationally collapses very quickly

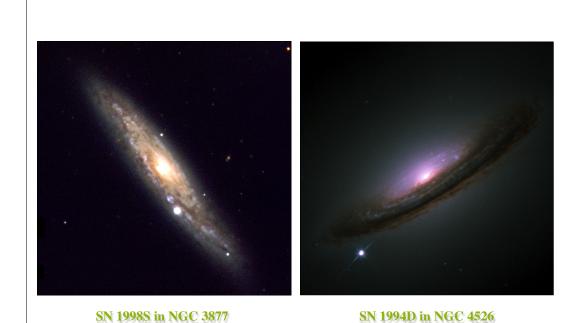
this results in an explosive release of neutrinos

- dramatically heating the outer shells to the point where they are explosively ejected
- This super-heated shell is extremely bright

Not all stars will supernova, ours won't. Only especially large ones will and they are rather spectacular when they occur nearby (there are 100's catalogued per year in other galaxies)

The bright 'star' appears, sometimes in the day (Tycho saw one and so did Kepler) and

 then the universe is bathed in a neutrino flux that's measurable on earth in elementary particle experiments, slightly preceding the visible explosion

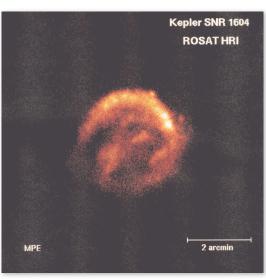


## not very often...

#### The last supernovae to occur in the Milky Way:

• Tycho's in 1572 and Kepler's in 1604





## a little excitement

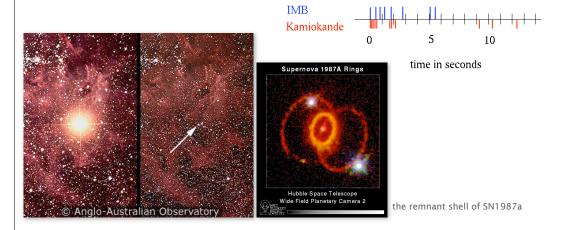
#### The first naked eye supernova in 400 years:

SN1987a - a modest Blue Giant

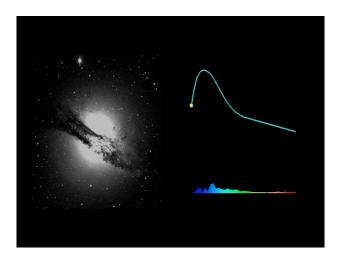
Large Megallanic Cloud (a companion galaxy 170,000 ly away, visible from the southern hemisphere)

0736 23 Feb, neutrinos observed; 0930AM, nothing visible by amateur Aussie astronomer, but by 1030...it was visible

According to the model of stellar collapse, it was expected that on earth there should be  $\sim$ 10 neutrino events, and in two very large "proton decay" experiments in Japan and Cleveland, detected 19



## smile, you're on candid camera



Three "representations" of a Supernova 1986G in the Centaurus A galaxy

Supernova Cosmology Project (P. Nugent: spectral sequence; A. Conley: image sequence) with the help of Lawrence Berkeley National Laboratory's Computer Visualization Laboratory (N. Johnston: animation) at the National Energy Research Scientific Computing Center.

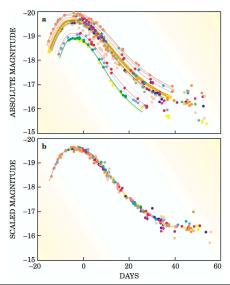
#### 1a supernovae are different

• Typically stars not massive enough by themselves to nova

but in close proximity to another star which it siphons matter from, enough to cause a supernova explosion after all

These are special as they stay bright for a period of time which is directly related to their magnitude - So, like Cepheids, measure the time, deduce the brightness

with brightness, can determine the distance.



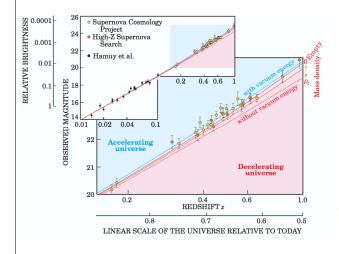
But, it gets worse.

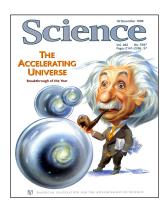
### there are lots of them

There are enough of these to have one about per second in deep galaxies

• So, by searching one gets a wide range of distances many, quite far away...at very "large redshifts"

#### An amazing thing happens:





The data require an interpretation that the Universe is Accelerating in its expansion

## um...accelerating???

#### What would cause the universe to accelerate??

- Hubble data suggested maybe a constant expansion
- CMB data demand that the geometry of the universe is flat (an open geometry could account for this)

## There has to be some "antigravity" kind of force at work to do this

- What's more, there has to be a lot of it
- The combination of all observables (only some of which have been mentioned), suggest a new contribution to  $\Omega$ , called  $\Omega_{\Lambda}$

$$\Omega_{\text{lum matter}} + \Omega_{\text{dark matter}} + \Omega_{\Lambda} = \Omega(\text{total})$$

### $\Omega_{\Lambda}$ could be due to the vacuum...

- Particles, fields in the vacuum that add a term of Negative Pressure to Einstein's equations!
- That is, it adds in essentially a quantum mechanical Cosmological Constant Term...back to the future.

Einstein's "blunder" may have been right after all!

## Called the TISSINI ENERGY PROBLEM

Understanding this is perhaps the biggest problem in science

The world is mounting huge multiple-satellite Supernovae measuring missions

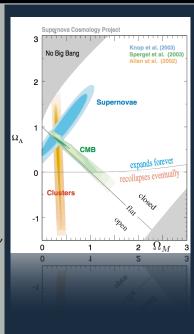
# interpreting dark energy

as a vacuum energy:

$$R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} + \Lambda g_{\mu\nu} = 8\pi G T_{\mu\nu} + \Lambda_{\text{eff}} g_{\mu\nu}$$

$$|\rho_{vac}(\text{cosmology})| < 10^{-9} \text{erg/cm}^3$$

Here, the vacuum energy absolute value matters!



## uh oh.

## THE 2 biggest problems in physical science

for High Energy Physics:

well...the vacuum belongs to us ;-) it's energy density is: 10<sup>-43</sup> erg/cc cosmology suggests: 10<sup>-9</sup> erg/cc

and any missing mass/energy is presumably quantum mechanical

that belongs to us also!

so, we are very aware of the possibility of producing such states in our collisions

we have to, for this to make any sense

It's all great fun!	