

Relativistic Quantum Mechanics.

Schrodinger's original attempt at QM was relativistic - and it failed.

Instead, he defined $\psi(\vec{x}, t)$ as a 1 particle wavefunction satisfying

$$\frac{\partial \psi(\vec{x}, t)}{\partial t} = -\frac{i}{\hbar} \nabla^2 \psi(\vec{x}, t)$$

It has definite \vec{p} and E as a free particle and solutions are

$$\psi(\vec{x}, t) = N e^{i(\vec{p} \cdot \vec{x} - Et)/\hbar}$$

Using the wave number, $\vec{k} = \frac{2\pi}{\lambda} \hat{k} = \frac{|\vec{p}|}{\hbar} \hat{k} = \frac{\vec{p}}{\hbar}$

related to the de Broglie wavelength $\lambda = \frac{2\pi\hbar}{|\vec{p}|}$

$$E = h\nu = \hbar\omega$$

So,

$$\psi(\vec{x}, t) = N e^{i(\vec{k} \cdot \vec{x} - \omega t)}$$

"First Quantization" \equiv

$$\vec{p} \rightarrow -i\hbar \vec{\nabla}$$

$$E \rightarrow i\hbar \frac{\partial}{\partial t}$$

So, the NR Hamiltonian

$$H = \frac{\vec{p}^2}{2m} \rightarrow \frac{(-i\hbar\vec{\nabla})^2}{2m} = -\frac{\hbar^2}{2m}\vec{\nabla}^2$$

and

$$-\frac{\hbar^2}{2m}\vec{\nabla}^2\psi(\vec{x},t) = -i\hbar\frac{\partial}{\partial t}\psi(\vec{x},t)$$

The free particle solutions give

$$\frac{\vec{p}^2}{2m}\psi = E\psi \quad E \text{ is all kinetic.}$$

The obvious relativistic extensions

(Schrodinger, Gordon, Klein, Fock, Kudar, deOndur, Van Duzen) in 1926:

$$H^2 = \vec{p}^2 c^2 + m^2 c^4 \quad \text{and "1st Quantized"}$$

BUT: remember Born's interpretation of ψ :

$$P = |\psi|^2 d^3x \quad \equiv \quad \text{probability of finding } \psi \text{ within the 3-volume element } d^3x$$

$$0 \leq P \leq 1$$

Probability flux density:

$$\vec{j} = -\frac{i\hbar}{2m} (\psi^* \vec{\nabla} \psi - \psi \vec{\nabla} \psi^*)$$

$$\psi^* \rightarrow \partial_t \psi^* = i\hbar \frac{\partial \psi^*}{\partial t}$$

$$\psi \rightarrow \partial_t \psi = -i\hbar \frac{\partial \psi}{\partial t}$$

Subtract the results:

$$\begin{aligned} \psi^* \partial_t \psi - \psi \partial_t \psi^* &= i\hbar \psi^* \frac{\partial \psi}{\partial t} + i\hbar \psi \frac{\partial \psi^*}{\partial t} \\ &= i\hbar \frac{\partial}{\partial t} (\psi^* \psi) \end{aligned}$$

$$\text{Sub } \mathcal{H} = -\frac{\hbar^2 \vec{\nabla}^2}{2m}$$

$$\vec{\nabla} \cdot \left[-\frac{\hbar^2}{2m} (\psi^* \vec{\nabla} \psi - \psi \vec{\nabla} \psi^*) \right] = i\hbar \frac{\partial}{\partial t} (\psi^* \psi)$$

which looks like a continuity equation

$$\frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot \vec{j} = 0$$

$$\rho = \psi^* \psi \quad \text{prob density} \quad \Rightarrow \quad \frac{\partial}{\partial t} \int d^3x \rho = 0$$

from Gauss' Theorem \Rightarrow
 ρ is constant of the Motion.

Covariant notation:

$$\phi(\vec{x}, t) \rightarrow \phi(x^\mu)$$

for free solution $\phi(x^\mu) = N e^{-i p \cdot x / \hbar}$

Add dot product

$$p \cdot x = p^\mu x_\mu = \frac{E}{c} ct - \vec{p} \cdot \vec{x}$$

$$p^\mu = (E/c, \vec{p})$$

$$x^\mu = (ct, \vec{x})$$

want $\partial_\nu^2 \phi$

$$\vec{\nabla} \phi(x^\mu) = \vec{\nabla} N e^{-i p \cdot x / \hbar} = N \vec{\nabla} e^{\frac{i}{\hbar} (\vec{p} \cdot \vec{x} - Et)}$$

$$= \frac{i}{\hbar} \vec{p} \phi(x^\mu)$$

$$\vec{\nabla}^2 \phi(x^\mu) = \left(\frac{i}{\hbar} \vec{p} \right)^2 \phi(x^\mu) = -\frac{\vec{p}^2}{\hbar^2} \phi(x^\mu)$$

$$\frac{\partial \phi(x^\mu)}{\partial t} = -\frac{i}{\hbar} E \phi(x^\mu)$$

$$\frac{\partial^2 \phi(x^\mu)}{\partial t^2} = \left(-\frac{i}{\hbar} E \right) \phi(x^\mu) = -\frac{E^2}{\hbar^2} \phi(x^\mu)$$

$$\text{So, } \left(\vec{\nabla}^2 - \frac{1}{c^2} \frac{\partial^2}{\partial t^2} \right) \phi(x^\mu) = \left(-\frac{\vec{p}^2}{\hbar^2} + \frac{E^2}{c^2 \hbar^2} \right) \phi(x^\mu)$$

$$= \frac{1}{\hbar^2} \left(\frac{E^2}{c^2} - \vec{p}^2 \right) \phi(x^\mu)$$

$$= \frac{1}{\hbar^2} (m^2 c^2) \phi(x^\mu)$$

So, this fits since:

$$\mathcal{H}^2 \phi = [c^2 (-i\hbar \vec{\nabla}^2)^2 + m^2 c^4] \phi$$

$$= (-c^2 \hbar^2 \vec{\nabla}^2 + m^2 c^4) \phi$$

$$= (c^2 \hbar^2 \frac{\vec{p}^2}{\hbar^2} + m^2 c^4) \phi$$

$$= (c^2 \vec{p}^2 + m^2 c^4) \phi$$

$$\mathcal{H}^2 \phi = \underbrace{E^2}_{\text{relativistic eigenvalue}} \phi$$

relativistic eigenvalue
equation for free
particle.

going backwards a bit,

$$\left(\vec{\nabla}^2 - \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \frac{m^2 c^2}{\hbar^2} \right) \phi(x^\mu) = 0$$

is the relativistic 2nd order wave equation

Klein Gordon Equation.

NOTHING WRONG HERE! all relativistic wavefunctions
must solve this equation

Remember: $\partial^\mu = (\partial^0, -\vec{\nabla}) = \left(\frac{1}{c} \frac{\partial}{\partial t}, -\vec{\nabla}\right)$

$$\hbar \partial^\mu = (\hbar \partial^0, -\hbar \vec{\nabla})$$

no, $\partial^\mu \partial_\mu \equiv \square = \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \vec{\nabla} \cdot \vec{\nabla}$

and K.G.

$$\left(\partial_\mu \partial^\mu + \frac{m^2 c^2}{\hbar^2}\right) \phi(x^\mu) = 0$$

$$\left(\square + \frac{m^2 c^2}{\hbar^2}\right) \phi(x^\mu) = 0$$

Okay, great. How about normalizability?

Make the continuity equation in the standard way

$$\phi^* \left(\vec{\nabla}^2 - \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \frac{m^2 c^2}{\hbar^2} \right) \phi - \phi \left(\vec{\nabla}^2 - \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \frac{m^2 c^2}{\hbar^2} \right) \phi^* = 0$$

⋮

$$\vec{\nabla} \cdot \left[\phi^* \vec{\nabla} \phi - \phi \vec{\nabla} \phi^* \right] - \frac{1}{c^2} \frac{\partial}{\partial t} \left[\phi^* \frac{\partial \phi}{\partial t} - \phi \frac{\partial \phi^*}{\partial t} \right] = 0$$

$$\uparrow$$

$$\vec{\nabla} \cdot \vec{j}$$

↑
looks like
before

$$\uparrow$$

$$\frac{1}{c^2} \frac{\partial}{\partial t} \rho_0 = 0$$

ρ_0 is not positive-definite

⇒ probability density isn't!

What about the eigenvalues of \mathcal{H} ?

$$\mathcal{H}\varphi(x^\mu) = \pm E\varphi(x^\mu) \quad \Rightarrow \text{negative energies!}$$

TOO MUCH FOR SCHRÖDINGER!

Play with it... put back the plane wave solutions.

$$\begin{aligned} \partial^\mu \varphi &= \partial^\mu N e^{-i p \cdot x / \hbar} = -i \frac{p^\mu}{\hbar} N e^{-i p \cdot x / \hbar} \\ &= -i \frac{p^\mu}{\hbar} \varphi \end{aligned}$$

$$\text{So, if } j^\mu = (p, \vec{j}) = \frac{\hbar i}{2m} (\varphi^* \partial^\mu \varphi - \varphi \partial^\mu \varphi^*)$$

$$\text{and } j^\mu = \frac{\hbar i}{2m} \left[-i \frac{p^\mu}{\hbar} \varphi^* \varphi - \varphi \varphi^* \left(+i \frac{p^\mu}{\hbar} \right) \right]$$

$$j^\mu = \frac{1}{2m} p^\mu |\varphi|^2$$

So, the $-E$ solutions require $p < 0$.

In 1928 Dirac took an entirely new approach:

$$E \text{ \& } p \text{ problems relate to } \frac{\partial^2}{\partial t^2} \leftarrow \text{2nd order}$$

Sought a 1^{st} order equation to avoid $-E \text{ \& } -p$ issues

Assumed: $H = \sqrt{c^2 p^2 + m^2 c^4} = c \vec{\alpha} \cdot \vec{p} + \beta m c^2$

need to find $\vec{\alpha}$ and β .

$$H^2 \psi = (c^2 p^2 + m^2 c^4) \psi$$

$$= (c \vec{\alpha} \cdot \vec{p} + \beta m c)^2 \psi$$

So $c^2 p^2 + m^2 c^4 = \vec{\alpha} \cdot \vec{p} \vec{\alpha} \cdot \vec{p} c^2 + c \vec{\alpha} \cdot \vec{p} m c^2 \beta$
 $+ \beta m c^2 c \vec{\alpha} \cdot \vec{p} + \beta^2 m^2 c^4$

conclude: ① $p^2 = \vec{\alpha} \cdot \vec{p} \vec{\alpha} \cdot \vec{p}$

② $\beta^2 = 1$

③ $\beta \vec{\alpha} + \vec{\alpha} \beta = 0 \Rightarrow$ matrices

$$p^2 = \vec{\alpha} \cdot \vec{p} \vec{\alpha} \cdot \vec{p} = (\alpha^1 p^1 + \alpha^2 p^2 + \alpha^3 p^3)(\alpha^1 p^1 + \alpha^2 p^2 + \alpha^3 p^3)$$

$$= \alpha^1 p^1 \alpha^1 p^1 + \alpha^2 p^2 \alpha^2 p^2 + \alpha^3 p^3 \alpha^3 p^3$$

$$+ \alpha^1 p^1 \alpha^2 p^2 + \alpha^2 p^2 \alpha^1 p^1 + \text{etc.}$$

$$= \alpha^1 \alpha^1 p^1{}^2 + \alpha^2 \alpha^2 p^2{}^2 + \alpha^3 \alpha^3 p^3{}^2$$

$$+ (\alpha^1 \alpha^2 + \alpha^2 \alpha^1) p^1 p^2 + \text{etc.}$$

get rid of cross terms \Rightarrow impose $\alpha^i \alpha^j + \alpha^j \alpha^i = 2\delta^{ij}$

$\equiv \{\alpha^i, \alpha^j\}$
 \nearrow
 anti commutator

③ satisfied by $\{\beta, \alpha^i\} = 0$

NEW ALGEBRA - anti commuting - not so long after the newness of Heisenberg et al. new commuting algebra.

So,

$$H\psi = \frac{i}{\hbar} \frac{\partial \psi}{\partial t}$$

$$H = \alpha \vec{\alpha} \cdot \vec{p} + \beta mc^2$$

$$\vec{p} \rightarrow -i\hbar \vec{\nabla}$$

} 1st quantization

$$-i\hbar c \vec{\alpha} \cdot \vec{\nabla} \psi + \beta mc^2 \psi = \frac{i}{\hbar} \frac{\partial \psi}{\partial t}$$

The DIRAC EQUATION. $\vec{\alpha}, \beta$ are matrices \rightarrow operators

Need a basis for $\vec{\alpha} \& \beta$ to operate within.

$$-i\hbar c (\vec{\alpha} \cdot \vec{\nabla})_{ab} \psi_b + mc^2 \beta_{ab} \psi_b = \frac{i}{\hbar} \frac{\partial \psi_b}{\partial t}$$

matrix indices $\Rightarrow \psi$ is a column matrix.
a SPINOR.

Convention:

$$\beta = \begin{pmatrix} 1 & & & 0 \\ & 1 & & \\ & & \dots & \\ 0 & & & -1 \\ & & & & -1 \\ & & & & & -1 \\ & & & & & & -1 \end{pmatrix}$$

\uparrow N \downarrow
 \updownarrow $N/2$ \updownarrow $N/2$

$\leftarrow N \rightarrow$

diagonal & traceless

$N=4$ was chosen - and is the smallest, unique choice

$$\beta = \begin{pmatrix} +1 & & 0 \\ & +1 & \\ 0 & & -1 \\ & & & -1 \end{pmatrix}$$

which connects to the Pauli Matrices!

$$\vec{\alpha} = \begin{pmatrix} 0 & \vec{\sigma} \\ \vec{\sigma} & 0 \end{pmatrix} \quad \text{e.g.} \quad \alpha_3 = \begin{pmatrix} 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & -1 \\ 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \end{pmatrix}, \text{ etc}$$

Then ①, ②, and ③ are all satisfied!

So:
$$\psi = \begin{pmatrix} \psi_1 \\ \psi_2 \\ \psi_3 \\ \psi_4 \end{pmatrix}$$

Done with \hbar & c : $\hbar = c = 1$.

$$i \frac{\partial \psi}{\partial t} = -i \vec{\alpha} \cdot \vec{\nabla} \psi + \beta m \psi \quad \checkmark$$

How about probability?

$$i \frac{\partial \psi}{\partial t} + i \vec{\alpha} \cdot \vec{\nabla} \psi - \beta m \psi = 0$$

$$\frac{\partial \psi}{\partial t} + \vec{\alpha} \cdot \vec{\nabla} \psi + i \beta m \psi = 0$$

so $\rho = \psi^\dagger \psi > 0$ definite!

$$\vec{j} = \psi^\dagger \vec{L} \psi$$

Symmetries \rightarrow constants of the motion.

$\frac{\partial A}{\partial t} = i [H, A] = 0 \Rightarrow A$ is a constant of the motion & in a group theoretical sense, A is a generator of a conserved quantum number.

1. ENERGY

$$[H, H] = [\vec{\alpha} \cdot \vec{p} + \beta m, \vec{\alpha} \cdot \vec{p} + \beta m] = 0 \quad \checkmark$$

2. MOMENTUM

$$[H, p_i] = [\vec{\alpha} \cdot \vec{p} + \beta m, p_i]$$

$$= [\alpha_j p_j + \beta m, p_i]$$

but $[p_i, p_j] = 0$ & α & β commute with p

$$\Rightarrow [H, p_i] = 0 \quad \checkmark$$

3. ORBITAL ANGULAR MOMENTUM

$$\vec{L} = \vec{r} \times \vec{p} \quad \rightarrow \quad L_i = \epsilon_{ijh} r_j p_h$$

$$i[H, L_i] = i[\alpha_n p_n + \beta m, \epsilon_{ijh} r_j p_h]$$

$$[r_j, p_h] = i \delta_{jh} \quad \text{and other stuff commutes}$$

$$i[H, L_i] = i[\alpha_n p_n, \epsilon_{ijh} r_j p_h]$$

$$= i(\alpha_n p_n \epsilon_{ijh} r_j p_h - \epsilon_{ijh} r_j p_h \alpha_n p_n)$$

$$= i \alpha_n [p_n, r_j] \epsilon_{ijk} p_k$$

$$= -i \alpha_n i \delta_{nj} \epsilon_{ijk} p_k$$

$$= \alpha_j \epsilon_{ijk} p_k = \epsilon_{ijk} \alpha_j p_k$$

$$\text{so } i[H, \vec{L}] = \vec{\alpha} \times \vec{p} = \frac{d\vec{L}}{dt} \neq 0 \quad \times$$

L is not a symmetry \Rightarrow no conserved
angular momentum

In regular Q.M. $\psi \sim f(r) \cdot \underbrace{Y_\ell^m(\theta, \phi)}_{\text{eigenstates of } L}$

not in Dirac theory

Search for something that's half way related...

$$\text{try } \vec{L} \equiv \vec{\alpha} \times \vec{\alpha}$$

$$i [H, \vec{\alpha} \times \vec{\alpha}]$$

Some useful identities:

$$\{\beta, \alpha_i\} = 0 \Rightarrow \beta \alpha_i = -\alpha_i \beta$$

$$\{\alpha_i, \alpha_j\} = 2\delta_{ij} \Rightarrow \alpha_i \alpha_j = 2\delta_{ij} - \alpha_j \alpha_i$$

$$i [H, (\vec{\alpha} \times \vec{\alpha})_j] = i [H, \epsilon_{jkl} \alpha_k \alpha_l]$$

$$= i [\alpha_i p_i + \beta m, \epsilon_{jkl} \alpha_k \alpha_l]$$

$$= i \{ (\alpha_i p_i + \beta m) \epsilon_{jkl} \alpha_k \alpha_l - \epsilon_{jkl} \alpha_k \alpha_l (\alpha_i p_i + \beta m) \}$$

$$= i \{ \alpha_i p_i \epsilon_{jkl} \alpha_k \alpha_l + \beta m \epsilon_{jkl} \alpha_k \alpha_l - \epsilon_{jkl} \alpha_k \alpha_l \alpha_i p_i - \epsilon_{jkl} \alpha_k \alpha_l \beta m \}$$

$$= i \{ \alpha_i p_i \epsilon_{jkl} \alpha_k \alpha_l + \beta m \epsilon_{jkl} \alpha_k \alpha_l - \epsilon_{jkl} \alpha_k \alpha_l \alpha_i p_i - \epsilon_{jkl} \alpha_k \alpha_l \beta m \}$$

$$= i \{ \alpha_i p_i \epsilon_{jkl} \alpha_k \alpha_l + \beta m \epsilon_{jkl} \alpha_k \alpha_l - \epsilon_{jkl} \alpha_k \alpha_l \alpha_i p_i - \epsilon_{jkl} \alpha_k \alpha_l \beta m \}$$

$$= i \{ \alpha_i p_i \epsilon_{jkl} \alpha_k \alpha_l + \beta m \epsilon_{jkl} \alpha_k \alpha_l - \epsilon_{jkl} \alpha_k \alpha_l \alpha_i p_i - \epsilon_{jkl} \alpha_k \alpha_l \beta m \}$$

β terms cancel

$$= i \{ \alpha_i p_i \epsilon_{jkl} \alpha_k \alpha_l - \epsilon_{jkl} \alpha_k \alpha_l \alpha_i p_i \}$$

$$= i \epsilon_{jkl} p_i \{ \alpha_i \alpha_k \alpha_l - \alpha_k \alpha_l \alpha_i \}$$

$$\begin{aligned}
 \left\{ \begin{array}{l} \\ \end{array} \right\} &= \alpha_i \alpha_h \alpha_k + \alpha_h \alpha_i \alpha_k - 2 \delta_{ih} \alpha_k \\
 &= \alpha_i \alpha_h \alpha_k - \alpha_i \alpha_h \alpha_k + 2 \delta_{hi} \alpha_k - 2 \delta_{ih} \alpha_k \\
 &= 2 \delta_{hi} \alpha_k - 2 \delta_{ih} \alpha_k
 \end{aligned}$$

$$\begin{aligned}
 i[H, (\vec{\alpha} \times \vec{\alpha})_j] &= 2i \epsilon_{jhl} p_i \delta_{hi} \alpha_k - 2i \epsilon_{jhl} p_i \delta_{ih} \alpha_k \\
 &= 2i \epsilon_{jhl} p_h \alpha_k - 2i \epsilon_{jhl} p_h \alpha_k \\
 &= 4i \epsilon_{jhl} p_h \alpha_k = 4i (\vec{p} \times \vec{\alpha})_j
 \end{aligned}$$

$$i[H, (\vec{\alpha} \times \vec{\alpha})_j] = -4i (\vec{\alpha} \times \vec{p})_j = \frac{d}{dt} (\vec{\alpha} \times \vec{\alpha})_j$$

\uparrow
 sorta like $\frac{d\vec{L}_j}{dt} = (\vec{\alpha} \times \vec{p})_j$

so,

$$i[H, \vec{L} + \frac{1}{4i} (\vec{\alpha} \times \vec{\alpha})] = 0$$

$\vec{L} + \frac{1}{4i} (\vec{\alpha} \times \vec{\alpha})$ is a constant of the motion ?

Look at this

$$(\vec{\alpha} \times \vec{\alpha})_i = \epsilon_{ijk} \alpha_j \alpha_k = \epsilon_{ijk} \begin{pmatrix} 0 & \sigma_j \\ \sigma_j & 0 \end{pmatrix} \begin{pmatrix} 0 & \sigma_k \\ \sigma_k & 0 \end{pmatrix}$$

$$= \epsilon_{ijk} \begin{pmatrix} \sigma_j \sigma_k & 0 \\ 0 & \sigma_j \sigma_k \end{pmatrix}$$

$$= \epsilon_{ijk} \begin{pmatrix} i \epsilon_{jkl} \sigma_l & 0 \\ 0 & i \epsilon_{jkl} \sigma_l \end{pmatrix}$$

↑
from commutation
relations for $\sigma \rightarrow$
The Lie Algebra of
 $SU(2)$

$$= i \epsilon_{ijk} \epsilon_{jkl} \begin{pmatrix} \sigma_l & 0 \\ 0 & \sigma_l \end{pmatrix}$$

$$\equiv i \underbrace{\epsilon_{ijk} \epsilon_{jkl}}_{2 \delta_{il}} \Sigma_l \quad \Sigma_l \equiv \begin{pmatrix} \sigma_l & 0 \\ 0 & \sigma_l \end{pmatrix}$$

$$(\vec{\alpha} \times \vec{\alpha})_i = 2i \delta_{il} \Sigma_l = 2i \Sigma_i$$

or. $\frac{1}{4i} (\vec{\alpha} \times \vec{\alpha})_i = \frac{1}{2} \Sigma_i$

↑
"spin" - ish

$$S_i \equiv \frac{1}{2} \Sigma_i \quad \text{The spin operator.}$$

The constant of the motion is

$$\vec{J} = \vec{L} + \vec{S} \quad \text{total angular momentum.}$$

SPIN comes for FREE in the Dirac Equation!

Note

$$S^2 = \frac{1}{4} \sum_i \Sigma_i \Sigma_i = \frac{1}{4} (\Sigma_1 \Sigma_1 + \Sigma_2 \Sigma_2 + \Sigma_3 \Sigma_3)$$

$$\Sigma_i \Sigma_i = \begin{pmatrix} \sigma_i^2 & 0 \\ 0 & \sigma_i^2 \end{pmatrix} = \begin{pmatrix} \mathbb{1} & 0 \\ 0 & \mathbb{1} \end{pmatrix} = \mathbb{1}$$

So,

$$S^2 = \frac{3}{4} \mathbb{1}$$

$$\frac{1}{2} (1 + \frac{1}{2}) \Rightarrow s(s+1)$$

it really is SPIN — S^2 is the Casimir Operator for
IRR of $SU(2)$

SPIN $\frac{1}{2}$ COMES FOR FREE in the Dirac Equation!

shut UP!!

Form: $\vec{p} \cdot \vec{J} = \vec{p} \cdot \vec{L} + \vec{p} \cdot \vec{S}$

$$\vec{p} \cdot \vec{J} = \vec{p} \cdot \vec{S} \quad \leftarrow \text{component of SPIN along direction of motion} \\ \equiv \text{HELICITY}$$

So, obviously

$$i[H, \vec{p} \cdot \vec{S}] = i[H, \vec{p} \cdot \left(\frac{\vec{\alpha} \times \vec{\alpha}}{4i}\right)] = -i\vec{p} \cdot \vec{\alpha} \times \vec{p} = 0$$

and we see that $\vec{p} \cdot \vec{S}$, HELICITY is a constant of the motion.

Helicity conservation is the real relativistic analog to angular momentum conservation.

Need a covariant form for Dirac Equation

CONVENTION: $\gamma^0 \equiv \beta = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$ (4x4, right?)

$$\vec{\gamma} \equiv \beta \vec{\alpha} = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \begin{pmatrix} 0 & \vec{\sigma} \\ \vec{\sigma} & 0 \end{pmatrix} = \begin{pmatrix} 0 & \vec{\sigma} \\ -\vec{\sigma} & 0 \end{pmatrix}$$

so, $\gamma^\mu = (\gamma^0, \vec{\gamma})$

can raise and lower $\gamma_\mu = g_{\mu\nu} \gamma^\nu = (\gamma_0, -\vec{\gamma})$

The following are really useful:

$$\begin{aligned}\{\gamma_0, \gamma_i\} &= \gamma_0 \gamma_i + \gamma_i \gamma_0 \\ &= \beta \beta \alpha_i + \beta \alpha_i \beta = \beta \beta \alpha_i - \beta \beta \alpha_i = 0\end{aligned}$$

$$\begin{aligned}\{\gamma_i, \gamma_j\} &= \gamma_i \gamma_j + \gamma_j \gamma_i \\ &= \beta \alpha_i \beta \alpha_j + \beta \alpha_j \beta \alpha_i \\ &= -\beta \beta \alpha_i \alpha_j - \beta \beta \alpha_j \alpha_i \\ &= -(\alpha_i \alpha_j + \alpha_j \alpha_i) \\ &= -\{\alpha_i, \alpha_j\} = -2 \delta_{ij}\end{aligned}$$

so $\{\gamma_\mu, \gamma_\nu\} = 2g_{\mu\nu}$

Also: $\beta^\dagger = \beta = (\gamma^0)^\dagger = \gamma^0$

$$\gamma_i^\dagger = (\beta \alpha_i)^\dagger = \alpha_i^\dagger \beta^\dagger = \alpha_i^\dagger \beta = -\beta \alpha_i = -\gamma_i$$

$\Rightarrow \beta$ is Hermitian
 γ^i is anti-Hermitian

Useful:

$$\gamma^i \dagger = -\gamma^i \times \underset{\substack{\uparrow \\ \gamma^0 \gamma^0}}{\mathbb{1}}$$

$$\gamma^{i\dagger} = -\gamma^i \gamma^0 \gamma^0 = +\gamma^0 \gamma^i \gamma^0$$

$$\gamma^{0\dagger} = \gamma^0 \gamma^0 \gamma^0 = \gamma^0$$

so

$$\gamma^{\mu\dagger} = \gamma^0 \gamma^\mu \gamma^0$$

Now, the Dirac Equation becomes

$$(i\alpha_j \frac{\partial}{\partial x_j} + i \frac{\partial}{\partial t} - \beta m) \psi = 0$$

$$(i\beta\alpha_j \frac{\partial}{\partial x_j} + i\beta \frac{\partial}{\partial t} - \beta m) \psi = 0$$

$$(i\gamma_j \frac{\partial}{\partial x_j} + i\gamma_0 \frac{\partial}{\partial t} - m) \psi = 0$$

$$(i\gamma^0 \frac{\partial}{\partial t} - i\gamma^j \frac{\partial}{\partial x_j} - m) \psi = 0$$

$$(i\gamma^\mu \frac{\partial}{\partial x^\mu} - m) \psi = 0$$

or $(i\gamma^\mu \partial_\mu - m) \psi = 0$ ✓

In momentum space: $(\gamma^\mu p_\mu - m) \psi = 0$ ✓

The conjugate equation will be important

$$\psi^\dagger (-i \overleftarrow{\partial}_\mu (\psi^\mu)^\dagger - m) = 0$$

$$\psi^\dagger (-i \overleftarrow{\partial}_\mu \gamma^0 \gamma^\mu \gamma^0 - m) = 0$$

$$\psi^\dagger (-i \overleftarrow{\partial}_\mu \gamma^0 \gamma^\mu \gamma^0 \gamma^0 - m \gamma^0) =$$

$$\psi^\dagger \gamma^0 (-i \overleftarrow{\partial}_\mu \gamma^\mu - m) = 0$$

DEFINE $\bar{\psi} \equiv \psi^\dagger \gamma^0$ so we get:

$$\bar{\psi} (i\gamma^\mu \overleftarrow{\partial}_\mu + m) = 0 \quad \text{and} \quad \bar{\psi} (\gamma^\mu p_\mu + m) = 0 \quad \checkmark$$

Feynman's shorthand notation - "slash notation"

$$\beta_\mu \gamma^\mu \equiv \not{\beta}$$

in LaTeX... `\def\slash#1{\!/\!#1}`

so that $\not{\beta} \rightarrow \text{\slash}\{\beta\}$

no charge.

Then, $(i \not{\not{x}} - m) \psi(x) = 0$

$$\bar{\psi}(x) (i \overleftarrow{\not{x}} + m) = 0$$

$$(\not{p} - m) \psi(p) = 0$$

$$\bar{\psi}(p) (\overleftarrow{\not{p}} + m) = 0$$

Watch this:

$$(i \gamma^\mu \partial_\mu - m) \psi = 0$$

$$i \gamma^\nu \partial_\nu \rightarrow$$

$$(-i \gamma^\nu \gamma^\mu \partial_\nu \partial_\mu - i \gamma^\nu \partial_\nu m) \psi = 0$$

notice $m i \gamma^\nu \partial_\nu \psi = m^2 \psi$ from the D.E.

so $(-\gamma^\nu \gamma^\mu \partial_\nu \partial_\mu - m^2) \psi = 0$

add in another one

$$(-\gamma^\nu \gamma^\mu \partial_\nu \partial_\mu - m^2 - \gamma^\mu \gamma^\nu \partial_\mu \partial_\nu - m^2) \psi = 0$$

$$(-\gamma^\nu \gamma^\mu \partial_\mu \partial_\nu - m^2 - \gamma^\mu \gamma^\nu \partial_\mu \partial_\nu - m^2) \psi =$$

$$[-(\gamma^\nu \gamma^\mu + \gamma^\mu \gamma^\nu) \partial_\mu \partial_\nu - 2m^2] \psi =$$

$$[-2g^{\mu\nu} \partial_\mu \partial_\nu - 2m^2] \psi = 0$$

look at a single component of $\psi = \begin{pmatrix} \psi_a \end{pmatrix} \dots \psi_a$.

$$(\partial^\nu \partial_\nu + m^2) \psi_a = 0$$

$$(\square + m^2) \psi_a = 0$$

a Klein-Gordon Equation!
for every ψ component.

This had to happen... arguments for K.G.E. were completely general.

SO: probability chance
still have negative energies.