

So,

$$* \quad \mathcal{J}^{(2)}(e\gamma \rightarrow e\gamma) = \frac{(-ie)^2}{2} \int d^4x_1 \int d^4x_2 \langle e\gamma | \delta_{ij}^{\mu\nu} \delta_{2m} \{ \bar{\Psi}_i^-(x_1) \Psi_m^+(x_2) A_{\mu}(x_1) A_{\nu}(x_2) \}$$

$$* \quad \times \langle 0 | T [\Psi_j(x_1) \bar{\Psi}_k(x_2)] | 0 \rangle + \bar{\Psi}_2^-(x_1) \Psi_j^+(x_1) A_{\mu}^+(x_1) A_{\nu}^-(x_2) \times \langle 0 | T [\bar{\Psi}_m(x_1) \Psi_i(x_2)] | 0 \rangle \} | e\gamma \rangle$$

We need to deal with the contracted terms.

eg, what is $\langle 0 | T [\Psi_j(x_1) \bar{\Psi}_k(x_2)] | 0 \rangle$

we know it's

$$T [\Psi_j(x_1) \bar{\Psi}_k(x_2)] = \Psi_j(x_1) \bar{\Psi}_k(x_2) \theta(x_{10} - x_{20}) - \bar{\Psi}_k(x_2) \Psi_j(x_1) \theta(x_{20} - x_{10})$$

But, let's see it happen naturally. —

look at 1st term $x_{10} > x_{20}$

$$\langle 0 | T (\Psi_j(x_1) \bar{\Psi}_k(x_2)) | 0 \rangle = \langle 0 | T (\underbrace{\Psi_j^+ \bar{\Psi}_k^+ + \Psi_j^+ \bar{\Psi}_k^- + \Psi_j^- \bar{\Psi}_k^+ + \Psi_j^- \bar{\Psi}_k^-}_{\text{already time ordered}}) | 0 \rangle$$

$$\propto \langle 0 | \underbrace{ab}_{0} + \underbrace{aa^{\dagger}}_{0} + \underbrace{b^{\dagger}b}_{0} + \underbrace{b^{\dagger}a^{\dagger}}_{0} | 0 \rangle$$

$$= \langle 0 | \Psi_j^+(x_1) \bar{\Psi}_k^-(x_2) | 0 \rangle \quad x_{10} > x_{20}$$

$$x_2 > x_1$$

$$\langle 0 | T(\psi_j(x_1) \bar{\psi}_\ell(x_2)) | 0 \rangle = \langle 0 | T(\psi_j^+ \bar{\psi}_\ell^+ + \psi_j^+ \bar{\psi}_\ell^- + \psi_j^- \bar{\psi}_\ell^+ + \psi_j^- \bar{\psi}_\ell^-) | 0 \rangle$$

$$= - \langle 0 | \bar{\psi}_\ell^+ \psi_j^+ + \bar{\psi}_\ell^- \psi_j^+ + \bar{\psi}_\ell^+ \psi_j^- + \bar{\psi}_\ell^- \psi_j^- | 0 \rangle$$

$$\propto \langle 0 | \underbrace{b a}_{0} + \underbrace{a^\dagger a}_{0} + \underbrace{b b^\dagger}_{0} + \underbrace{a^\dagger b^\dagger}_{0} | 0 \rangle$$

$$= - \langle 0 | \bar{\psi}_\ell^+(x_2) \psi_j^-(x_1) | 0 \rangle \quad x_2 > x_1$$

Look explicitly at the first one $\langle 0 | \psi_j^+(x_1) \bar{\psi}_\ell^-(x_2) | 0 \rangle$

$$\psi_j^+(x_1) = \sum_{i=1}^2 \int d^3k \, a^{(i)}(k) u_j^{(i)}(k) e^{-i k \cdot x_1}$$

$$\bar{\psi}_\ell^-(x_2) = \sum_{m=1}^2 \int d^3k' \, \bar{u}_\ell^{(m)}(k') a^{\dagger(m)}(k') e^{i k' \cdot x_2}$$

$$\langle 0 | \psi_j^+(x_1) \bar{\psi}_\ell^-(x_2) | 0 \rangle = \sum_i \sum_m \int \frac{d^3k}{(2\pi)^3 2E} \int \frac{d^3k'}{(2\pi)^3 2E'} u_j^{(i)}(k) \bar{u}_\ell^{(m)}(k') e^{-i(k \cdot x_1 - k' \cdot x_2)} \langle 0 | a^{(i)}(k) a^{\dagger(m)}(k') | 0 \rangle$$

↓ use anti commutator
by adding

$$\langle 0 | a^\dagger a | 0 \rangle (= 0)$$

$$(2\pi)^3 2E \delta(k - k') \delta_{im}$$

do $\int d^3k'$ integration

$$= \sum_i \int dK \quad u_j^{(i)}(k) \bar{u}_2^{(i)}(k) e^{-ik \cdot (x_1 - x_2)}$$

Use completeness -- to project out the + energy particles

$$= \int dK \quad (\not{k} + m)_{j\ell} e^{-ik \cdot (x_1 - x_2)} \quad x_{10} > x_{20}$$

The second one

$$\begin{aligned} \langle 0 | \bar{\Psi}_2^+(x_2) \Psi_j^-(x_1) | 0 \rangle &= \sum_i \sum_m \int dK \int dK' \bar{v}_2^{(i)}(k) v_j^{(m)}(k) e^{-i(k \cdot x_2 - k' \cdot x_1)} \\ &\quad \times \langle 0 | b^{(i)}(k) b^{(m)\dagger}(k') | 0 \rangle \end{aligned}$$

since $\Psi_j^-(x_1) = \sum_{i=1}^2 \int dK' b^{(i)}(k') v_j^{(i)}(k') e^{ik' \cdot x_1}$

$$\bar{\Psi}_2^+(x_2) = \sum_{m=1}^2 \int dK \bar{v}_2^{(m)}(k) b^{(m)\dagger}(k) e^{-ik \cdot x_2}$$

In the same fashion

$$\begin{aligned} \langle 0 | \bar{\Psi}_2^+(x_2) \Psi_j^-(x_1) | 0 \rangle &= \sum_i \int dK v_j^{(i)}(k) \bar{v}_2^{(i)}(k) e^{ik \cdot (x_1 - x_2)} \\ &= \int dK (\not{k} - m)_{j\ell} e^{ik \cdot (x_1 - x_2)} \end{aligned}$$

So, going all the way back -- remembering the negative sign,

$$\langle 0|T(\psi_1(x_1)\bar{\psi}_2(x_2))|0\rangle$$

$$= \int dK (K+m)_{j\ell} e^{-i\hbar \cdot (x_1 - x_2)} \theta(x_{10} - x_{20})$$

$$- \int dK (K-m)_{j\ell} e^{i\hbar \cdot (x_1 - x_2)} \theta(x_{20} - x_{10})$$

Look at the second term...

$$- \int dK (\gamma^0 \hbar_0 - \vec{\gamma} \cdot \vec{\hbar} - m)_{j\ell} e^{i\hbar \cdot (x_1 - x_2)}$$

$$\text{note } i\hbar \cdot (x_1 - x_2) = iE(x_{10} - x_{20}) - \vec{\hbar} \cdot (\vec{x}_1 - \vec{x}_2) \\ = (-i)(-E)(x_{10} - x_{20}) - \vec{\hbar} \cdot (\vec{x}_1 - \vec{x}_2)$$

$$\text{change } \vec{\hbar} \rightarrow -\vec{\hbar} \equiv \vec{p}$$

$$- \int dP [(-\gamma^0)(-E) + \vec{\gamma} \cdot \vec{p} - m]_{j\ell} e^{-i(-E)(x_{10} - x_{20})} e^{i\vec{p} \cdot (\vec{x}_1 - \vec{x}_2)}$$

$$\text{if } p^\mu = (-E, \vec{p}) \text{ then}$$

$$\int dP \underbrace{[\gamma^0 p_0 - \vec{\gamma} \cdot \vec{p} + m]_{j\ell}}_{(p+m)} e^{\underbrace{-ip^\mu \cdot (x_{10} - x_{20}) + i\vec{p} \cdot (\vec{x}_1 - \vec{x}_2)}_{-ip \cdot (x_1 - x_2)}} e$$

Of course p^μ is a dummy, so can call it k^μ with understanding that in 2nd term $k^0 = -E$

$$\langle 0|T(\psi_1(x_1)\bar{\psi}_2(x_2))|0\rangle = \int dK (K+m)_{j\ell} e^{-i\hbar \cdot (x_1 - x_2)}$$

for $x_{10} > x_{20}$ then $k^0 = E$; for $x_{20} > x_{10}$ then $k^0 = -E$

what we are...

I showed the non-zero terms in the Wick expansion for $\bar{\Psi}(x) \gamma^\mu \Psi(x) A_\mu(x)$ to second order to include a variety of stand-alone, familiar processes (Compton scattering, Møller scattering, and Bhabha scattering) as well as some less familiar, "add-on" processes (electron/positron self energy, photon vacuum polarization) and some non-interesting, unmeasurable processes (annihilation & vacuum fluctuation).

We are working specifically on $e\gamma \rightarrow e\gamma$

$$* \quad J^{(2)}(e\gamma \rightarrow e\gamma) = \frac{(-ie)^2}{2} \int d^4x_1 \int d^4x_2 \langle e\gamma | \gamma_{ij}^\mu \gamma_{lm}^\nu$$

$$\left\{ \bar{\Psi}_i^-(x_1) \Psi_m^+(x_2) A_\mu(x_1) A_\nu(x_2) \langle 0 | T [\Psi_j(x_1) \bar{\Psi}_l(x_2)] | 0 \rangle \right.$$

$$* \quad \left. - \bar{\Psi}_l^-(x_2) \Psi_j^+(x_1) A_\mu(x_1) A_\nu(x_2) \langle 0 | T [\bar{\Psi}_i(x_1) \Psi_m(x_2)] | 0 \rangle \right\} | e\gamma \rangle$$

we're concentrating on the contracted term, which I showed to be

$$\langle 0 | T (\bar{\Psi}(x_1) \Psi(x_2)) | 0 \rangle = \int d^4k (k + m) e^{-ik \cdot (x_1 - x_2)}$$

$$\text{with } k^0 = E \quad x_{10} > x_{20}$$

$$k^0 = -E \quad x_{20} > x_{10}$$

look at the special integral

$$I = \frac{1}{(2\pi)^4} \int_{-\infty}^{\infty} d^4k \left(\frac{i}{k-m} \right)_{jl} e^{-ik(x_1-x_2)} \quad \text{no constraint on } k^\mu$$

multiply by $\frac{k+m}{k+m}$

denominator

$$\begin{aligned} (k+m)(k-m) &= \delta^\nu k_\nu \delta^\mu k_\mu - \delta^\nu k_\nu m + m \delta^\mu k_\mu - m^2 \\ &= \delta^\nu k_\nu \delta^\mu k_\mu - m^2 \end{aligned}$$

note $A_\nu B_\mu \delta^\mu \delta^\nu = B_\mu A_\nu (-\delta^\mu \delta^\nu + 2g^{\mu\nu})$

$$= -B_\mu A_\nu \delta^\mu \delta^\nu + 2A \cdot B$$

or

$$\begin{aligned} A B &= -B A + 2A \cdot B \\ \{A, B\} &= 2A \cdot B \end{aligned}$$

so

$$\begin{aligned} \{k, k\} &= 2k \cdot k \\ k k &= k \cdot k = k^2 \end{aligned}$$

and denominator becomes

$$\begin{aligned} k^2 - m^2 &= k_0 k_0 - \vec{k} \cdot \vec{k} - m^2 \\ &= (k_0 - \sqrt{\vec{k} \cdot \vec{k} + m^2})(k_0 + \sqrt{\vec{k} \cdot \vec{k} + m^2}) \end{aligned}$$

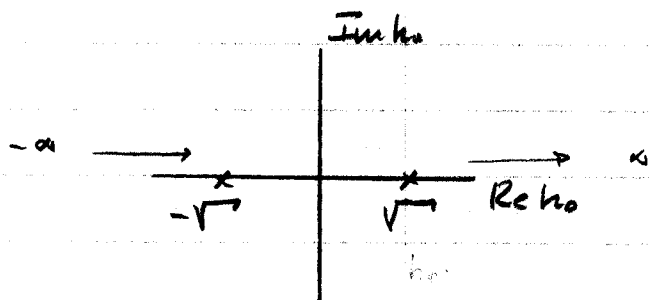
and I get

$$I = \frac{i}{(2\pi)^4} \int_{-a}^a d^3k \int_{-a}^a dk^0 \left[\frac{(k+m) e^{-ik \cdot (x_1 - x_2)}}{(k_0 - \sqrt{\vec{k} \cdot \vec{k} + m^2})(k_0 + \sqrt{\vec{k} \cdot \vec{k} + m^2})} \right]$$

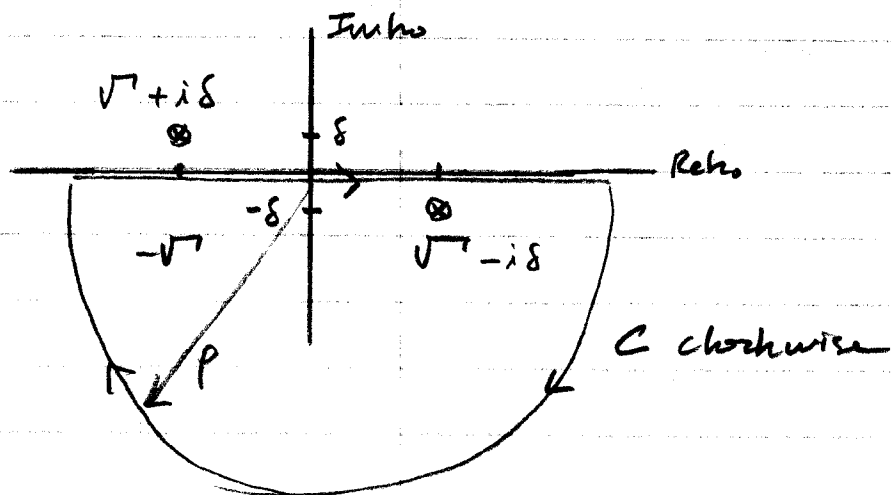
which has poles at $k_0 = \pm \sqrt{\vec{k} \cdot \vec{k} + m^2}$

$$\text{So, } I_0 = \int_{-a}^a dk_0 \frac{(k+m) e^{-ik^0(x_{10} - x_{20})}}{(k_0 - \sqrt{\vec{k} \cdot \vec{k} + m^2})(k_0 + \sqrt{\vec{k} \cdot \vec{k} + m^2})}$$

map this onto the complex k_0 plane.



avoid the poles in the standard way by shifting them and use Cauchy's Theorem



As $p \rightarrow \text{large}$ and $h^0 = -i|h^0|$, the exponential becomes,

$$e^{-ih^0(x_{10}-x_{20})} \rightarrow e^{-i(-i|h^0|)(x_{10}-x_{20})} \\ = e^{-|h^0|(x_{10}-x_{20})} \\ > 0$$

so as $p \rightarrow \infty$ $|h^0| \rightarrow \infty$ and exponential damps.

Then,

$$\oint_C dh^0 = -2\pi i (\sum \text{residues of enclosed poles})^R$$

$$= \int_{-a}^{+a} dh^0 \text{ (real, straight line)}$$

$$+ \int_{-a}^{+a} (curve)$$

$$-2\pi i R = \int_{-a}^{+a} f(h^0) dh^0 = I_0$$

The pole is simple, so the residue is

$$\frac{(k+m) e^{-ih^0(x_{10}-x_{20})}}{(h^0 - \sqrt{\quad} + i\delta)(h^0 + \sqrt{\quad} - i\delta)} \Big|_{h^0 = k_p = \sqrt{\quad} - i\delta}$$

$$I_0 = -2\pi i \frac{(k_p + m) e^{-i[\sqrt{k^2 - i\delta}](x_{10} - x_{20})}}{(\sqrt{k^2 - i\delta} + \sqrt{k^2 - i\delta})}$$

$\delta \rightarrow 0$

$$I_0 = -2\pi i \frac{(k + m) e^{-i\sqrt{k^2 + m^2}(x_{10} - x_{20})}}{2\sqrt{k^2 + m^2}}$$

and

$$I = \frac{1}{(2\pi)^4} \int d^3k \frac{(k + m) e^{i\vec{k} \cdot (\vec{x}_1 - \vec{x}_2) - i\sqrt{k^2 + m^2}(x_{10} - x_{20})}}{2\sqrt{k^2 + m^2}}$$

$$= \int \frac{d^3k}{(2\pi)^3} \frac{(k + m) e^{-ik \cdot (x_1 - x_2)}}{2E} \quad E = +\sqrt{k^2 + m^2}$$

$$= \int dK (k + m) e^{-ik \cdot (x_1 - x_2)}$$

which is what we had for

$$\langle 0 | T(\psi(x_1) \bar{\psi}(x_2)) | 0 \rangle \quad \text{for } x_{10} > x_{20}$$

Note if we had had scalar fields, one can go all the way back to the original T product and would find

$$\int dP e^{-ip \cdot (x_1 - x_2)} = \langle 0 | T(\phi(x_1) \phi^\dagger(x_2)) | 0 \rangle \equiv i\Delta^{(+)}(x_1 - x_2)$$

$x_{10} > x_{20}$

called the Positive Frequency propagator - one of a family

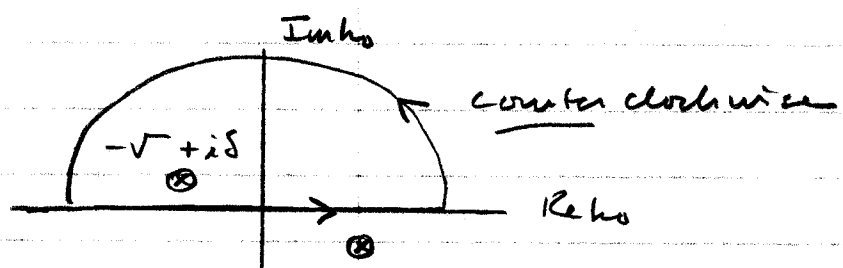
Likewise, $\langle 0 | T(\psi^+(x_2) \psi(x_1)) | 0 \rangle \equiv -i \Delta^{(-)}(x_1 - x_2)$
 $= \int dP e^{+iP \cdot (x_1 - x_2)} \quad \text{for } x_{20} > x_{10}$

In general the Feynman scalar Propagator is:

$$i \Delta_F(x_1 - x_2) = i \Delta^{(+)}(x_1 - x_2) \theta(x_{10} - x_{20}) - i \Delta^{(-)}(x_1 - x_2) \theta(x_{20} - x_{10})$$

We've just seen that the Dirac propagator for $x_{20} > x_{10}$ looks like that of the scalar, but for the projection operator (as I hinted at before)

Back to Dirac fields - what about $x_{20} > x_{10}$?



so $I_0 = +2\pi i (\sum R) \Big|_{L^0 = -\sqrt{k \cdot k} + i\delta}$

so $I = \frac{(2\pi i)}{(2\pi)^4} \int \frac{d^3k}{2(-\sqrt{k \cdot k} + m)} e^{ik \cdot (x_1 - x_2)} (i \not{k} + m) e^{-i(-\sqrt{k \cdot k} + m)(x_{10} - x_{20})}$

call $k_0 = -\sqrt{k \cdot k} + m \equiv -E$

and $I = - \int \frac{d^3k}{(2\pi)^3 (=2E)} (i \not{k} + m) e^{-ik \cdot (x_1 - x_2)}$

which is what we had for $\langle 0 | T(\psi(x_1) \bar{\psi}(x_2)) | 0 \rangle$
when $x_2 > x_1$

So, we have a single way of writing $\psi(x_1) \bar{\psi}(x_2)$
regardless of the time ordering.

$$\langle 0 | T(\psi_j(x_1) \bar{\psi}_k(x_2)) | 0 \rangle$$

$$= \frac{i}{(2\pi)^4} \int d^4k \frac{(k\!\!\!/ + m)_{jk} e^{-ik \cdot (x_1 - x_2)}}{[k_0 - \sqrt{\vec{k} \cdot \vec{k} + m^2} + i\delta][k_0 + \sqrt{\vec{k} \cdot \vec{k} + m^2} - i\delta]}$$

$$\text{denominator} = k_0^2 - (\vec{k} \cdot \vec{k} + m^2) + \underbrace{2i\delta \sqrt{\vec{k} \cdot \vec{k} + m^2}}_{\text{Call this } i\eta} + \delta^2$$

$$= k \cdot k + m^2 + i\eta$$

and

$$\langle 0 | T(\psi_j(x_1) \bar{\psi}_k(x_2)) | 0 \rangle = \int \frac{d^4k}{(2\pi)^4} \frac{(k\!\!\!/ + m)_{jk} e^{-ik \cdot (x_1 - x_2)}}{(k^2 + m^2 + i\eta)}$$

usually written

$$= \int \frac{d^4k}{(2\pi)^4} \left(\frac{i}{k\!\!\!/ - m} \right) e^{-ik \cdot (x_1 - x_2)}$$

$$\equiv iS_F(x_1 - x_2)$$

the Feynman fermion propagator

Notice

$$S_{F,jk}(x_1-x_2) = (i\not{\partial} + m)_{jk} \Delta^{(+)}(x_1-x_2) \theta(x_{10}-x_{20}) \\ - (i\not{\partial} + m)_{jk} \Delta^{(+)}(x_1-x_2) \theta(x_{20}-x_{10})$$

Remember what these are...

look at $S_F(x_1-x_2) = \frac{-i}{(2\pi)^4} \int d^4k \frac{(i\not{k} + m) e^{-ik \cdot (x_1-x_2)}}{k^2 - m^2}$

operate with the Dirac Equation operator.

$$(i\not{\partial} + m) S_F(x_1-x_2) = \frac{-i \cdot i}{(2\pi)^4} \int d^4k \frac{(i(-i\not{k}) + m) e^{-ik \cdot (x_1-x_2)}}{k^2 - m^2} \\ = \frac{1}{(2\pi)^4} \int d^4k \frac{k^2 - m^2}{k^2 - m^2} e^{-ik \cdot (x_1-x_2)} \\ = \frac{1}{(2\pi)^4} \int d^4k e^{-ik \cdot (x_1-x_2)} \\ = \delta(x_1-x_2)$$

proving that $S_F(x_1-x_2)$ is the Green's Function for the Dirac Equation.

We will need the same thing for the photon. Again

$$\langle 0 | T(A_\mu(x_1) A_\nu(x_2)) | 0 \rangle = \langle 0 |$$

$$\langle 0 | A_\mu^+(x_1) A_\nu^-(x_2) | 0 \rangle \theta(x_{10} - x_{20}) + \langle 0 | A_\nu^-(x_2) A_\mu^+(x_1) | 0 \rangle \theta(x_{20} - x_{10})$$

proceed the same way. -

first, for $x_{10} > x_{20}$

$$\langle 0 | T() | 0 \rangle = \int dK \int dK' \sum_{\lambda=0}^3 \sum_{\lambda'=0}^3 \epsilon_{\mu(\lambda)}(k) \epsilon_{\nu(\lambda')}(k') \cdot e^{-ik \cdot x_1} e^{ik' \cdot x_2} \langle 0 | a_{(\lambda)}(k) a_{(\lambda')}^+(k') | 0 \rangle$$

Remember, for photons the commutation relations are a little odd

$$[a_{(\lambda)}(k), a_{(\lambda')}^+(k')] = -g_{\lambda\lambda'} 2k_0 (2\pi)^3 \delta(\vec{k} - \vec{k}')$$

$$iD_{F\mu\nu}(x_1 - x_2) = \int dK \sum_{\lambda=0}^3 (-g_{\lambda\lambda'}) \epsilon_{\mu(\lambda)}(k) \epsilon_{\nu(\lambda')}(k) e^{-ik \cdot (x_1 - x_2)}$$

Remember the completeness relation

$$\sum_{\lambda=0}^3 g_{\lambda\lambda'} \epsilon_{\mu(\lambda)}(k) \epsilon_{\nu(\lambda')}(k) = g_{\mu\nu} \quad (\text{I wrote } \sum_{\lambda} \epsilon_{\lambda} \epsilon_{\lambda} = -g_{\mu\nu})$$

same.

With the same logic as before, we get a similar expression for $x_{20} > x_{10}$.

For the photon, we consider an integral like

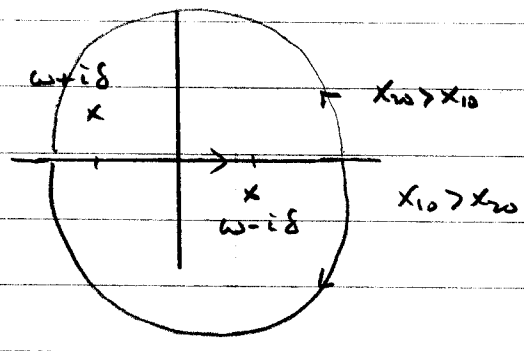
$$\frac{1}{(2\pi)^4} \int \frac{d^4k}{k^2 k_\mu} e^{-ik \cdot (x_1 - x_2)}$$

$$k^2 = k_0^2 - \vec{k} \cdot \vec{k} = (k_0 - \sqrt{\vec{k} \cdot \vec{k}})(k_0 + \sqrt{\vec{k} \cdot \vec{k}})$$

which leads to
$$I_0 = \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} \frac{e^{-i\omega(x_1 - x_2)}}{(\omega + i\epsilon)(\omega - i\epsilon)}$$

where $\omega \equiv \sqrt{k^2}$

Again,



and we get

$$D_{F\mu\nu}(x_1 - x_2) = \int \frac{d^4k}{(2\pi)^4} e^{-ik(x_1 - x_2)} D_{F\mu\nu}(k)$$

where $D_{F\mu\nu}(k) = \frac{-g_{\mu\nu}}{k^2 + i\epsilon}$ like before

This is the momentum space Feynman Propagator.

Note, this is in a particular gauge - The Feynman gauge. We can generalize as before with the gauge fixing parameter, in which we would have

$$D_{F\mu\nu}(k) = \frac{-g_{\mu\nu}}{k^2 + i\epsilon} + \frac{\xi - 1}{\xi} \frac{k^\mu k^\nu}{(k^2 + i\epsilon)^2}$$

Feynman $\xi = 1$

Landau $\xi \rightarrow \infty$ (transverse in 4d, i.e. $k^\mu D_{F\mu\nu} = 0$)

This is all covariant. - we could have done it for only the physical, transverse states - that's often done. However, it is instructive to break down the components of the polarization sum.

Define the unit vector:

$$\eta_\mu \equiv (1, \vec{0}) \quad \text{timelike and } \perp \text{ to } \epsilon_{(1,2)}^\mu$$

and define another unit vector:

$$\pi^\mu \equiv \frac{k^\mu - (k \cdot \eta) \eta^\mu}{\sqrt{(k \cdot \eta)^2 - k^2}} \quad \text{spacelike and} \\ \text{orthogonal: } \pi \cdot \epsilon = 0 \\ \pi \cdot \eta = 0$$

So, this provides a basis $(\eta_\mu, \pi^\mu, \epsilon_{(1,2)}^\mu)$ for which we can write a completeness relation

$$\eta^\mu \eta^\nu - \sum_{i=1}^2 \epsilon_{(i)}^\mu \epsilon_{(i)}^\nu - \pi^\mu \pi^\nu = g^{\mu\nu}$$

or,

$$iD_F^{\mu\nu}(k) = \frac{\sum_{i=1}^2 \epsilon_{(i)}^\mu \epsilon_{(i)}^\nu}{k^2 + i\epsilon} \quad \text{transverse piece.}$$

$$+ \frac{1}{k^2 + i\epsilon} \frac{k^\mu k^\nu - \eta^\mu \eta^\nu}{\sqrt{(k \cdot \eta)^2 - k^2}} \quad \text{Coulomb piece}$$

$$+ \frac{1}{k^2 + i\epsilon} \frac{k^\mu k^\nu - (k^\mu \eta^\nu + k^\nu \eta^\mu)(k \cdot \eta)}{\sqrt{(k \cdot \eta)^2 - k^2}} \quad \text{"residual piece"}$$

Notice, the Coulomb piece is a combination of the scalar and longitudinal pieces -- then in the frame in which $\eta = (1, 0, 0, 0)$

$$D_F^{\mu\nu}(k)_{\text{Coulomb}} = \frac{g^{\mu 0} g^{\nu 0}}{|k|^2}$$

• Couples only to $\mu, \nu = 0$ photon "leg"

→ the static Coulomb potential emerges and is, as advertised, a mixture of the two ^{separately} unphysical pieces of the renormalization. → our $\sum_{\lambda=0}^3$ sum includes the static potential.

The "residual piece is not observable due to current conservation. That is, in momentum space

$$k^\mu j_\mu(k) = 0$$

so any coupling of the $D_F^{\mu\nu}$ propagator with a conserved EM current leads to the k^μ parts of $D_F^{\mu\nu}$ residual to give no real contribution.