

Phase-space

288a

Let's deal with phase space separately, and then make use of it later.

Consider generally a two body process.

We could be speaking of a decay $a \rightarrow 3+4$ for which we would calculate the decay rate / particle

$$d\omega = \frac{(2\pi)^4 \delta(a \rightarrow 3+4) |T|^2 dP_3 dP_4}{2E_a}$$

and then the "total width"

$$\Gamma = \int \frac{d\omega}{ds} ds$$

or the cross section for $1+2 \rightarrow 3+4$

$$d\sigma = \frac{(2\pi)^4 \delta(1+2 \rightarrow 3+4) |T|^2 dP_3 dP_4}{(\text{flux, normalization})}$$

Common is the so-called phase space factor, which I'll represent as

$$d_2\omega(\vec{P}_3, \vec{P}_4) = (2\pi)^4 \delta(P_1 + P_2 - P_3 - P_4) d\vec{P}_3 d\vec{P}_4$$

We're going to assume that particle 3 will be measured and particle 4 will be ignored, so we need to "integrate 4 away".

Let's define

$$d_2\rho(\vec{p}_3) = d_2\rho(\Omega_3, p_3) \equiv \int d_2\rho(\vec{p}_3, \vec{p}_4) d^3\vec{p}_4$$

Then,

$$d_2\rho(\vec{p}_3) = \frac{1}{(2\pi)^2} \frac{\int d\Omega_3 p_3^2 dp_3}{4E_3 E_4} d^3\vec{p}_4 \delta(\vec{p}_1 + \vec{p}_2 - \vec{p}_3 - \vec{p}_4) \delta(E_1 + E_2 - E_3 - E_4)$$

Obviously, the $\delta^3()$ function imposes overall 3 momentum conservation

$$\vec{p}_1 + \vec{p}_2 = \vec{p}_3 + \vec{p}_4$$

$$\vec{p}_4 = \vec{p}_1 + \vec{p}_2 - \vec{p}_3$$

which acts as a constraint. Then,

$$d_2\rho(\vec{p}_3) = \frac{1}{(2\pi)^2} \frac{\int d\Omega_3 p_3^2 dp_3}{4E_3 E_4} \delta(E_1 + E_2 - E_3 - E_4) w/$$

Since $E_4^2 = p_4^2 + m_4^2$ and $\vec{p}_4 = \vec{p}_1 + \vec{p}_2 - \vec{p}_3$

the δ function is

$$\delta(E_1 + E_2 - E_3 - \sqrt{(\vec{p}_1 + \vec{p}_2 + \vec{p}_3)^2 + m_4^2})$$

can

$$E_a \equiv E_1 + E_2 \quad \text{and} \quad \left. \right\}$$

$$\vec{p}_a \equiv \vec{p}_1 + \vec{p}_2$$

makes the following

amenable to 2 body

decay $a \rightarrow 3+4$ as

well as scattering.

then

$$1+2 \rightarrow 3+4$$

From the standard δ function manipulation

$$\delta[f(x)] = \left| \frac{\partial f}{\partial x} \right|_{x=x_r}^{-1} \delta(x-x_r) \quad \text{where } f(x_r) = 0$$

Here,

$$f(E_3) = E_a - E_3 - E_4 \Rightarrow \left| \frac{\partial f}{\partial E_3} \right| = \left| -1 - \frac{\partial E_4}{\partial E_3} \right|$$

$$E_{3r} = E_a - E_4 = 1 + \frac{\partial E_4}{\partial E_3}$$

Now $E_4^2 = (\vec{p}_a - \vec{p}_3)^2 + m_4^2$

$$2E_4 \frac{\partial E_4}{\partial E_3} = 2P_3 \frac{\partial P_3}{\partial E_3} - 2\vec{p}_a \cdot \hat{p}_3 \frac{\partial \vec{p}_3}{\partial E_3}$$

since $P_3^2 = E_3^2 - m_3^2$

$$2P_3 \frac{\partial P_3}{\partial E_3} = 2E_3, \quad \frac{\partial \vec{p}_3}{\partial E_3} = \frac{E_3}{P_3}$$

so $2E_4 \frac{\partial E_4}{\partial E_3} = 2P_3 \frac{E_3}{P_3} - 2\vec{p}_a \cdot \hat{p}_3 \frac{E_3}{P_3}$

$$\frac{\partial E_4}{\partial E_3} = \frac{E_3}{E_4} - \frac{\vec{p}_a \cdot \vec{p}_3}{E_4 P_3} \frac{E_3}{P_3} = \frac{E_3 P_3^2 - \vec{p}_a \cdot \vec{p}_3 E_3}{E_4 P_3^2}$$

so $\frac{\partial f}{\partial E_3} = 1 + \frac{\partial E_4}{\partial E_3} = \frac{E_4 P_3^2 + E_3 P_3^2 - \vec{p}_a \cdot \vec{p}_3 E_3}{E_4 P_3^2}$

$$= \frac{E_a P_3^2 - \vec{p}_a \cdot \vec{p}_3 E_3}{E_4 P_3^2}$$

$$\text{so, } d_2\rho(\vec{p}_3) = \frac{1}{(2\pi)^2} \frac{d\Omega_3}{4E_3 E_4} \frac{p_3^2 dp_3}{E_a p_3^2 - \vec{p}_a \cdot \vec{p}_3 E_{3r}} \left(\frac{E_4 p_3^2}{E_a p_3^2 - \vec{p}_a \cdot \vec{p}_3 E_{3r}} \right) \delta(E_3 - E_{3r})$$

$$\text{Now } d^3 \vec{p}_3 = p_3^2 d\Omega_3 dp_3$$

$$p_3 = \sqrt{E_3^2 - m_3^2}$$

$$dp_3 = \frac{1}{2} \cdot \frac{2E_3 dE_3}{P_3} \Rightarrow p_3 dp_3 = E_3 dE_3$$

need integral over E_3

$$\text{so, } d_2\rho(\vec{p}_3) = \frac{1}{(2\pi)^2} \frac{d\Omega_3}{4E_3 E_4} \frac{p_3 E_3 dE_3}{E_a p_3^2 - \vec{p}_a \cdot \vec{p}_3 E_{3r}} \left(\frac{E_4 p_3^2}{E_a p_3^2 - \vec{p}_a \cdot \vec{p}_3 E_{3r}} \right) \delta(E_3 - E_{3r})$$

What's E_{3r} ?

$$E_{3r} = E_a - \sqrt{(\vec{p}_a - \vec{p}_3)^2 + m_q^2} \quad \text{so,}$$

$$(\vec{p}_a - \vec{p}_3)^2 + m_q^2 = (E_a - E_{3r})^2 = E_a^2 + E_{3r}^2 - 2E_a E_{3r}$$

$$\vec{p}_a^2 + p_3^2 - 2\vec{p}_a \cdot \vec{p}_3 + m_q^2 =$$

$$\underbrace{\vec{p}_a^2 - E_a^2}_{-m_a^2} + \underbrace{p_3^2 - E_{3r}^2}_{-m_3^2} - 2\vec{p}_a \cdot \vec{p}_3 + m_q^2 = -2E_a E_{3r}$$

and

$$E_{3r} = \frac{m_a^2 + m_3^2 - m_q^2 + 2\vec{p}_a \cdot \vec{p}_3}{2E_a}$$

and, define

$$d_2\rho(\vec{p}_3) = \int d_2\rho(\vec{p}_3) d\vec{p}_3 = \int d_2\rho(\vec{p}_3) \frac{E_3}{P_3} dE_3$$

$$= \frac{1}{(2\pi)^2} \frac{d\Omega_3}{4} \frac{p_3^3}{E_a p_3^2 - \vec{p}_a \cdot \vec{p}_3 E_{3r}} \quad \text{when } E_{3r}$$

now we can choose a frame.

choose the frame in which $\vec{p}_a = 0$ ($\Rightarrow \vec{p}_1 = -\vec{p}_2$)
 if a scattering problem, - the "center of momentum
 frame" - the rest frame for a decay problem).

$$\text{Then: } E_{3r} = \frac{m_a^2 + m_3^2 - m_q^2}{2m_a}$$

$$\begin{aligned} \text{Since, } p_3^2 &= E_{3r}^2 - m_3^2 \\ &= \frac{m_a^4 + m_3^4 + m_q^4 + 2m_a^2m_3^2 - 2m_a^2m_q^2 - 2m_3^2m_q^2}{4m_a^2} \\ &\quad - m_3^2 \\ &= \frac{m_a^4 + m_3^4 + m_q^4 - 2m_a^2m_3^2 - 2m_a^2m_q^2 - 2m_3^2m_q^2}{4m_a^2} \end{aligned}$$

This numerator happens a lot in kinematics and
 is given a name -

$$\text{The triangle function: } \lambda(a, b, c) = a^2 + b^2 + c^2 - 2ab - 2ac - 2bc$$

So here we have

$$p_3^2 = \frac{\lambda(m_a^2, m_3^2, m_q^2)}{4m_a^2}$$

There is more jargon. The center of mass energy squared is called

$$S = (\vec{P}_1 + \vec{P}_2)^2 = (\vec{P}_3 + \vec{P}_4)^2 = \vec{p}_a^2 \quad (\text{4 vectors})$$

$$S = m_a^2$$

so, you sometimes see (jargon),

$$E_3 = \frac{s + m_3^2 - m_4^2}{2\sqrt{s}} \quad \left. \right\} \text{cm frame.}$$

and you could also find

$$E_4 = \frac{s + m_4^2 - m_3^2}{2\sqrt{s}}$$

$$\text{also, } \vec{P}_3^2 = \frac{[s - (m_3 + m_4)^2][s - (m_3 - m_4)^2]}{4s} = \vec{p}_4^2$$

and

$$|\vec{p}_3| = \frac{1}{2s} \lambda^t(s, m_3^2, m_4^2)$$

In this frame,

$$d_2\rho(\vec{p}_3) = \frac{1}{(2\pi)^2} \frac{\vec{p}_3^3}{4} \frac{1}{m_2 P_3^2} d\vec{p}_3$$

$$d_2\rho(\vec{p}_3) = \frac{\vec{p}_3}{16\pi^2 \sqrt{s}} d\vec{p}_3$$

$$\text{and } d\omega = \frac{|T|^2}{2E_a} d\varphi(\mathbf{r}_3)$$

$$d\omega = \frac{|T|^2}{32\pi^2} \left(\frac{\rho_3}{M_a^2} \right) d\Omega_3$$

note $|T|^2$ will have θ_3
angular dependence, so
this is as far as we can
go

When integrated, this is the "total width" of the unstable A.

$$\Gamma = \int \frac{dw}{d\omega} d\omega$$

The decay rate per particle is

$$\rho = -\frac{dN_A}{dt} \frac{1}{N_A}$$

so,

$$N_A(t) = N_A(0) e^{-\rho t}$$

and we refer to $\rho^{-1} \equiv \tau$ as the "lifetime"
(more later)

Now look at 2 body scattering, $1+2 \rightarrow 3+4$

$$d\sigma = |T|^2 d\Omega_2$$

(flux normalization)

$$\begin{aligned} (\text{flux normalization}) &= |\vec{v}_1 - \vec{v}_2| / 2E_1 2E_2 \\ &= |\vec{p}_1 - \vec{p}_2| / 4E_1 E_2 \\ &= \left| \frac{\vec{p}_1}{E_1} - \frac{\vec{p}_2}{E_2} \right| / 4E_1 E_2 \\ &= 4 |\vec{p}_1 E_2 - \vec{p}_2 E_1| \end{aligned}$$

$$= 4 \sqrt{E_1^2 \vec{p}_1^2 + E_2^2 \vec{p}_2^2 - 2E_1 E_2 \vec{p}_1 \cdot \vec{p}_2}$$

which is general.

Consider specific cases:



i) anti-colinear beams: $\vec{p}_1 \cdot \vec{p}_2 = -p_1 p_2$ but $|\vec{p}_1| \neq |\vec{p}_2|$ necessarily

$$(\text{flux} \cdot \text{norm}) = 4 \sqrt{E_2^2 p_1^2 + E_1^2 p_2^2 + 2 E_1 E_2 p_1 p_2}$$

simplifying $2 E_1 E_2 p_1 p_2 = (E_1 E_2 + p_1 p_2)^2 - E_1^2 E_2^2 - p_1^2 p_2^2$

:

$$(\text{flux} \cdot \text{norm}) = 4 \sqrt{-E_1^2 m_2^2 + p_1^2 m_2^2 + (E_1 E_2 + p_1 p_2)^2}$$

write using invariants $p_1^\mu \cdot p_2^\mu = E_1 E_2 - \vec{p}_1 \cdot \vec{p}_2$

$$= E_1 E_2 + |\vec{p}_1| |\vec{p}_2|$$

so

$$(\text{flux} \cdot \text{norm}) = 4 \sqrt{(p_1 \cdot p_2)^2 - m_1 m_2}$$

ii) center of momentum frame:

$$|\vec{p}_1| = |\vec{p}_2| \equiv p$$

$$|\vec{p}_3| = |\vec{p}_4| \equiv p'$$

$$\begin{aligned} \text{then, } (p_1 \cdot p_2)^2 - m_1^2 m_2^2 &= (E_1 E_2 - \vec{p}_1 \cdot \vec{p}_2)^2 - m_1^2 m_2^2 \\ &= (E_1 E_2 + |\vec{p}|^2)^2 \\ &= p^2 (E_1 + E_2)^2 \end{aligned}$$

so

$$(\text{flux} \cdot \text{norm}) = 4p(E_1 + E_2)$$

since, for anti colinear beams,

$$S = (P_1 + P_2)^2 \quad (\text{4 vectors})$$

$$= P_1^2 + P_2^2 + 2P_1 \cdot P_2$$

$$= m_1^2 + m_2^2 + 2E_1 E_2 - 2\vec{P}_1 \cdot \vec{P}_2 \quad \text{General}$$

$$= m_1^2 + m_2^2 + 2E_1 E_2 + 2|\vec{p}|^2$$

anti colinear &
center of momentum

$$= P_1^2 + m_1^2 + P_2^2 + m_2^2 + 2E_1 E_2$$

$$= E_1^2 + E_2^2 + 2E_1 E_2$$

$$= (E_1 + E_2)^2$$

obviously, if beams are same species (particle-particle
or particle-antiparticle) then $E_L = E_R = E$

$$S = (2E)^2$$

anticolinear, center of mass, same mass.

and one typically refers to an accelerator's
capability in terms of "root S", here

$$\sqrt{S} = 2E$$

@ Fermilab $E \approx p = 1 \text{ TeV}$

$$\sqrt{S} = 2 \text{ TeV}$$

max: @ LEP $E \approx p = 105 \text{ GeV}$

$$\sqrt{S} = 210 \text{ GeV}$$

@ LHC $E \approx p = 7 \text{ TeV}$

$$\sqrt{S} = 14 \text{ TeV}$$

(iii) For scattering in the stationary target frame (measure particle 3)

$$d\sigma(\vec{p}_3, \vec{p}_4) = \frac{1}{\text{flux.norm}} \frac{(2\pi)^4 \delta^4(p_1 + p_2 - p_3 - p_4)}{(2\pi)^3 2E_3} \frac{\sum |T|^2 d^3 p_3}{(2\pi)^3 2E_4} d^3 p_4$$

do standard trick. - $\frac{d^3 p_4}{2E_4} \rightarrow d^4 p_4 \delta(p_4^2 - m_4^2) \theta(p_{43})$
and integrate, enforcing a constraint on 4-momentum conservation.

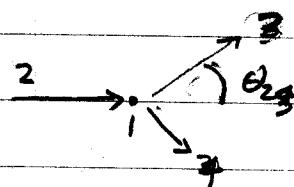
$$d\sigma(\vec{p}_3) = \frac{1}{\text{flux.norm}} \frac{1}{(2\pi)^2} \delta(p_4^2 - m_4^2) \theta(p_{43}) \frac{d^3 p_4}{2E_3} \Big|_{P_4 = P_1 + P_2 - P_3}$$

Look at argument:

$$\begin{aligned} p_4^2 - m_4^2 &= (p_1 + p_2 - p_3)^2 - m_4^2 \\ &= m_1^2 + m_2^2 + m_3^2 + 2p_1 \cdot p_2 - 2p_1 \cdot p_3 - 2p_2 \cdot p_3 - m_4^2 \\ &= (\cancel{m_1^2} - \cancel{m_4^2}) + (\cancel{m_2^2} + \cancel{m_3^2}) \\ &\quad \text{III} \qquad \text{III} \\ &= M_{12}^+ + M_{23}^+ \end{aligned}$$

in the lab. $p_i = (m_i, \vec{p}_i)$

$$p_i = (E_i, \vec{p}_i)$$



now,

$$\begin{aligned} p_4^2 - m_4^2 &= M_{12}^- + M_{23}^+ + 2m_1 E_2 - 2m_1 E_3 - 2E_2 E_3 + 2\vec{p}_2 \cdot \vec{p}_3 \\ &= M_{12}^- + M_{23}^+ + 2m_1 E_2 - 2E_3 (m_1 + E_2) + 2\vec{p}_2 \cdot \vec{p}_3 \cos \theta_{23} \\ &\equiv f(E_3) \end{aligned}$$

use standard S function theorem

we'll convert to dE_3 - always can do this.

$$d^3P_3 = P_3^2 dP_3 d\Omega_3 \quad P_3 = \sqrt{E_3^2 - m_3^2}$$

$$dP_3 = \frac{1}{2} \frac{2E_3 dE_3}{P_3}$$

$$P_3 dP_3 = E_3 dE_3$$

call the root E_r , w $f(E_r) = 0 \rightarrow$ ugly

$$E_r = (M_{1\bar{2}} + M_{2\bar{3}} + 2m_1 E_2 + 2P_2 P_3 \cos \theta_{23}) / 2(m_1 + E_2)$$

$$\frac{df}{dE_3} = 2P_2 \cos \theta_{23} \frac{dP_3}{dE_3} - 2(m_1 + E_2)$$

"
 E_3/P_3

$$= 2P_2 \cos \theta_{23} E_3/P_3 - 2(m_1 + E_2)$$

so,

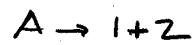
$$d\sigma(\Omega_3) = \frac{1}{\text{flux.norm}} \frac{1}{(2\pi)^2} \int \frac{\delta(E_3 - E_r) P_3 E_3 dE_3}{2(m_1 + E_2 - \frac{P_2}{P_3} E_3 \cos \theta_{23})} \frac{d\Omega_3}{2P_3}$$

$$\begin{aligned} \text{flux.norm} &= 4\sqrt{(P_1 P_2)^2 - m_1^2 m_2^2} \quad (\text{4d}) \\ &= 4\sqrt{m_1^2 E_2^2 - m_1^2 m_2^2} \\ &= 4m_1 \sqrt{E_2^2 - m_2^2} \\ &= 4m_1 |\vec{P}_2| \end{aligned}$$

so,

$$\frac{d\sigma}{d\Omega_3} = \frac{1}{4m_1 |\vec{P}_2|} \frac{1}{(2\pi)^2} \frac{P_3}{4[m_1 + E_2 - \frac{P_2}{P_3} E_3 \cos \theta_{23}]} \left. \sum |T|^2 \right|_{E_3 = E_r}$$

Unstable particles -



One imagines $N_A(0)$ particles at $t=0$ with some average lifetime τ , such that

$$N_A(t) = N_A(0) e^{-t/\tau}$$

In accordance with the Uncertainty Principle, there is an uncertainty in the time each state exists which in turn leads to an uncertainty in the energy level according to $\Delta E \Delta t \geq \hbar$

The number of particles at any time would be related to $|Y_A(t)|^2$. For free propagation

$$Y_A(t) \propto Y_A(0) e^{-iEt}$$

That the state is unstable means,

$$|Y_A(t)|^2 \propto |Y_A(0)|^2 e^{-t/\tau}$$

Notice one gets this if one makes the energy imaginary

$$E = E_0 - \frac{i\Gamma}{2}$$

$$Y_A(t) \sim Y_A(0) e^{-iE_0 t} e^{-\Gamma/2 t}$$

(go)

$$|Y_A(t)|^2 \sim |Y_A(0)|^2 e^{-\Gamma t}$$

Showing that $\frac{1}{\tau} = 2 \text{Im}(E) = \Gamma$

This constant is the decay rate, Γ

We always calculate in momentum space, so

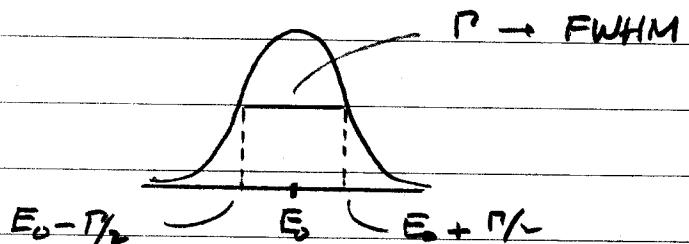
$$\begin{aligned}\hat{\psi}_A(t) &= \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} dt e^{iEt} \psi_A(t) \\ &= \frac{1}{\sqrt{2\pi}} \psi_A(0) \int dt e^{i(E-E_0)t - i\Gamma t/2} \\ &= \frac{i \psi_A(0)}{\sqrt{2\pi}} \frac{1}{(E-E_0) + i\Gamma/2}\end{aligned}$$

The probability of finding A at this time.

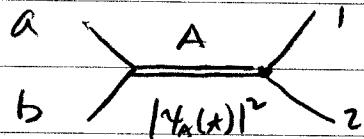
$$|\hat{\psi}_A(E)|^2 \propto \frac{|\psi_A(0)|^2}{2\pi} \left(\frac{1}{(E-E_0)^2 + \Gamma^2/4} \right)$$

(Bret-Wigner shape)

So, Γ is both the rate at which the states decay and the uncertainty in the value of the energy due to the finite lifetime



Often, an unstable state is produced in a scattering process $a + b \rightarrow A \rightarrow 1 + 2$



(said to be a resonant state)

The relativistic point of view is that the energy of the resonant peak \approx the mass of the resonance - definition.

$$E_0^2 = m_A^2 = s$$

write,

$$\frac{m_A}{m_A} \cdot \frac{1}{(E-E_0) + i\Gamma/2} = \frac{2m_A}{2m_A(E-E_0) + im_A\Gamma}$$

If $|E-m_A| \ll m_A$ - hard to think of a resonant "particle" interpretation otherwise

$$E^2 - m_A^2 = (E-m_A)(E+m_A) \\ \sim (E-m_A)2m_A$$

\Rightarrow

$$\frac{2m_A}{(E^2 - m_A^2) - im_A\Gamma}$$

which squared is proportional to,

$$\frac{1}{(s - m_A^2)^2 + m_A^2\Gamma^2}$$

Relativistic Breit-Wigner

Notice:

$$\frac{1}{E^2 - m_A^2 - i m_A \Gamma}$$

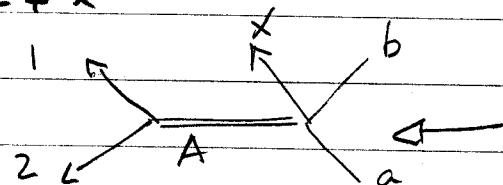
$$\sim \frac{1}{E^2 - m_A^2 - i \gamma}$$

which are propagators for the particle which is intermediate. This leads to the definition of mass as the pole of the propagator.

The total amplitude for the process

$$a + b \rightarrow A \rightarrow 1 + 2 + x$$

can be written,



$$T(a + b \rightarrow 1 + 2 + x) = T(A \rightarrow 1 + 2) \frac{1}{\rho^2 - m_A^2 - i m_A \Gamma} T(a + b \rightarrow A + x)$$

Imagine one intermediate state $a + b \rightarrow A \rightarrow 1 + 2$
then,

$$\sigma(a + b \rightarrow 1 + 2) \propto \Gamma(A \rightarrow 1 + 2) \frac{1}{(s - m_A^2)^2 + m_A^2 \Gamma^2} \Gamma(\bar{A} \rightarrow \bar{a} + \bar{b})$$

For calculational purposes, often the "narrow width approximation" is helpful.

$$\frac{1}{(s - m^2)^2 + m^2 \Gamma^2} \sim \frac{\pi}{m \Gamma} \delta(s - m^2)$$