hi

Lecture 17, 14.03.2017

Einstein's Theory of General Relativity

housekeeping

Question about anything?

I'll make a movie for you: Marie Curie movie anyone?

duh. I've now got a complete set of dates:

Book Review 1 is due Saturday.

FakeFacebook is due April Fools Day. tee hee Blog read-reflect project will start soon.

See calendar cartoon:





PINNED POST





Chip Brock created a poll. 18 mins

The Curie movie. This is stupid...sorry. I've now pinned down two rooms' availabilities through the week of March 27. This will be the last poll, I promise. Okay, I lied. There will be a pizza poll, but that's different, right?

sday, March 15 at 7pm
day, March 20 at 7pm
nesday, March 22 at 6:30pm
sday, March 23 at 6:30pm
day, March 27 at 7pm
3 More Options
Comment A Share

real electrons

HV transmission lines feed substations?

138,000 V is common (BWL for example) Assume that arc is at 138,000V, so electrons have that energy

...which would be the Kinetic Energy



an exercise in "electron volts"

What's the rest energy?

What's the rest mass?

What's the speed of the electrons?

What's the momentum of one of the electrons?

What's the relativistic mass of one of the electrons?

What's the total energy of one of the electrons?



This will be on video and figure into homework

https://qstbb.pa.msu.edu/storage/11 Special Relativity/spark.mp4

next few weeks, v2

S M T W Th F Sa

2/26	2/27	2/28	3/1	3/2	3/3
	midte	rm			
3/5		Spring Break			
3/12	3/13	3/14	3/15	3/16	3/17





Honors Project

has begun. First milestone was last Friday.

Read the Second of two sets of instructions:

MinervaInstructions2 2017.pdf in

www.pa.msu.edu/~brock/file sharing/QSandBB/2017homework/honors project 2017/

MasteringAstronomy

free and use of the textbook:

- The Essential Cosmic Perspective, Bennett, Megan Donahue, Schneider, Mark Voit
 - http://www.pearsonmylabandmastering.com/ northamerica/masteringastronomy/
 - Course ID is ISP220SP17
 - "code" is WSSPCT-SNELL-NAMEN-WEIGH-METIS-NJORD







$m_{H} = 100 \text{ kg} \times Y$ 100 kg

Energy/momentum relations:

object in its own "rest mass"... m "relativistic mass"... $m_R = m\gamma$ frame "Energy"... $E_T = m\gamma c^2$ the total Energy of a moving object Kinetic Energy... $K = mc^2(\gamma - 1)$ the energy due to motion momentum for Relativistic momentum... $p = m\gamma u$ each component of space Energy-momentum relation... $E_T^2 = (mc^2)^2 + (pc)^2$

the mass of an

useful expression

the mass of a moving object

"rest Energy"... $E_m = mc^2$

the mass-energy of an object in its own frame

an alternative,

looky here...

 $E_T^2 = (mc^2)^2 + (pc)^2$

two things to worry about



massless objects...





Energy and momentum are related for

E = pc

What about the negative solution?



9

three things are always, always constant С $s^2 = (ct)^2 - x^2$ S $m^2 c^4 = E_T^2 - p^2 c^2$ mc^2

Einstein preferred "Invariant Theory" to "Relativity"

General Relativity



"'rinciple of equivalence"

what's worse

Masses appearing different from different frames?

How do you deal with Newton's Universal **Gravitational formula?**

distance?



Worrying about Gravity led Einstein to

think hard about

SPACE and TIME

moving coordinate systems

accelerating

BTW, his birthday is today: Pi-day



the general theory of relativity

What's the "special" in "special" relativity?

the physics of inertial frames

What about the most obvious accelerating condition?

stupid
elevator
trick, #1

gravitational attraction



gravitational force

stupid
elevator
trick, #2

gravitational attraction



force up to create an acceleration of 1g

Here comes a Relativity-like statement:

There is no mechanical or electromagnetic experiment he can perform that would tell him that he was

1. being attracted by the Earth due to gravity or

2. being pulled and accelerated g with no gravitational field anywhere

this is called the

"Equivalence Principle"

stupid elevator trick, #1 gravitational attraction

stupid elevator trick, #2 gravitational attraction



identical



some subtly to the Equivalence Provide Antiple

force down at g



gravitational force

EP says that if these are the same

the laws of physics will be identical

So, anything that happens in L happens in R and visa versa

 $F = G \frac{Mm(\text{grav})}{R^2}$

 $m(\text{grav}) \equiv m(\text{inertial})$



space

force up at g



inertial force

F = m(inertial)a



relative to couch people:

at rest

constant speed up





straight, slanted

CP observe light beam: horizontal

boy observes light beam: horizontal

horizontal

accelerated, up



horizontal

what's "straight"?

around a gravitating mass, the curve path is still:

"shortest distance between two points"

in practice: the path that a beam of light would take

BUT: light travels differently shaped paths between relatively accelerated frames

Then the Equivalence Principle requires:

light should also curve in the presence of gravity

light paths

map the shape of space



not just light

acceleration messes with geometry

straight is not straight

and Einstein knew that this was problematic $C = 2\pi R$

experimentally: you could show that

Now, start it rotating.

fast...so special relativistic effects are apparent.

The ruler on the radius?

The rulers on the circumference?

 $C \neq 2\pi R$

The rules of Euclid's geometry – flat geometry – don't hold for an accelerating object.

Hold this thought: accelerated motion seems to change regular Geometry.



acceleration warps space

from the Equivalence Principle



gravity should warp space

light beam

what about time?

use a clock

accelerating.... so **B** moving away from **A** A and B are different inertial frames at each time



receives at say 5 ticks per second

1 second

B would say that A's clock has longer between ticks:

Runs Slower

sends at say 10 ticks per second

1 second



same idea as last time, slight different take



Equivalence Principle would require that:

B would say that A's clock has longer between ticks:

Runs Slower



"red shift" longer between ticks?

like the wavelength of the light is longer at B than A



acceleration warps time

from the Equivalence Principle



gravity should warp time

I was sitting in a chair in the patent office at Bern when all of a sudden a thought occurred to me. If a person falls freely, he will not feel his own weight. Albert Einstein

reminiscence from 1907 later he called this the "happiest thought of my life"

free-fall.

is a strange state of motion

you don't notice your own weight





training in the Vomit Comet KC 135





Gravity is relative

another Equivalence Principle

Neither situation "sees" gravity

all Einstein

sidewalk boy in **Free Fall** feels no forces

everything has same relative velocity

go far away from any gravitational sources

exactly the same situation as free fall near a gravitating body

John Wheeler calls it Free Float



intertial frame

identical





modeling all of this was

arguably one of the most technically challenging piece of physics ever

Hilley 105 $\frac{d}{dt} \cdot 2\dot{f} = -\int \sigma \frac{\dot{s}}{2t} = -\int \sigma \frac{d}{dt} \left(\frac{t}{t}\right)$ $2\dot{f} = \int \sigma \cdot \frac{t}{2t} + C$ Fine Die Bewegungs gleidenzen no unstraillen Tweekles lamben ales (bei west): - d & & xi+...+ gro = - 1 > 2 y - 2 y d. thornu 1/2 12 auf Dos Reptorat ! E= c, (1- 2 2+ 2 2;) -12 5=1-12+12 , Anternyo igalam : Geradaus 25 rentrat 4 (T + x2 ×) Perike (bury us) pn pn pn : C.E : 20:-0: 4 + 2" OT = x dr aT 2'd q + da' = 2 c, E - 2 c, + C, A Datis ist (and flow S. & F. and S.7 Gl. 5) w 2'dq2: (Jo : de + C)de du ling and mad ma ? " all multich Alin er N = d + . / A. x + . + . + $\frac{d^{2}T}{dx^{2}} = \frac{1}{r}\frac{dT}{dr} + x^{2} \frac{1}{r}\frac{d}{dr}\left(\frac{1}{r}\frac{dT}{dr}\right)$ <u>K</u> <u>46</u> <u>477</u> 9 ... Bw 26,E-26,45)24dy = (fo + + C)2(22dy + d2) $AT = \frac{3}{r} \frac{dT}{dr} + i \frac{d}{r} \frac{d}{dr} \left(\frac{1}{r} \frac{dT}{dr} \right)$ $t \left[\frac{2c_0(b-c_0)}{+c_0^* A} e^{i(b-\frac{1}{2}+C_1)^2} dy^* = \left(\frac{b-\frac{1}{2}+C}{+c_0^* A} \right)^2 dx^*$ with buicksideligt Dess sich am Filde unt Du Best wicht Eyz' = 39. $\frac{2}{2x}\left(x^{2}\frac{N}{r^{2}}\right) = 2x\frac{N}{r^{2}} + x^{3} \cdot \frac{1}{r}\frac{d}{dr}\left(\frac{N}{r^{2}}\right)$ 2 x dr (+ the : O ge sign sind , were auch wide for x, = x , x = y the Cortalle $d p = \frac{\int_{-\pi} \frac{1}{2} \cdot \mathcal{C}}{\sqrt{c_0 [2(\delta \tau_0) + \frac{1}{2} \frac{1}{2}] \cdot 2^4 - (S_0 + \mathcal{C}_2)^2}} d\alpha$ W . di . [-9' + C $\frac{\partial^{2}}{\partial x^{2}} = 2 \frac{W}{r^{2}} + 5 x^{2} \frac{1}{r} \frac{d}{dr} \left(\frac{W}{r^{2}} \right) + x^{4} \frac{1}{r} \frac{d}{dr} \left(\frac{1}{r} \frac{d}{dr} \left(\frac{M}{r^{2}} \right) \right)$ ds = Vg, x + g ... + 2g, yx + ... 2g, x + ... + g. " " " B' (c'-q'- c'+ nc' + 6 1-d. c' (1-d+x d')'= c' (1-2 d 2 = x2y. 1 d (N/r2) Die group and aber , tis and weende . Al. werde Owing day de = Tacalé.ci) 9" = " (1 - 1 + n -1) - co (1 - 2 + 1) gry = Co (1 - if + 3 di) Dola whereint we Do given = $x^2 \cdot \frac{1}{r} \frac{d}{dr} \left(\frac{N}{r^2} \right) + x^2 y^2 \cdot \frac{1}{r} \frac{d}{dr} \left(\frac{1}{r} \frac{d}{dr} \frac{d}{r'} \right)$ $\frac{dt^{*} + t^{*} dy^{*}}{dt^{*}} = C_{0}^{*} \left[1 - \frac{C_{0}^{*}}{C} + \frac{d}{2} \left(2 \frac{C_{0}^{*}}{L^{*}} - 1 \right) + \frac{d^{*}}{2^{*}} \left[n - \frac{C_{0}^{*}}{C} \right] \right]$ $\Delta(x^{2}\frac{N}{r^{2}}) = 2\frac{N}{r^{2}} + \int_{r}^{N} x^{2}\frac{1}{r}\frac{d}{dr}\binom{N}{r^{2}} + x^{2}\frac{1}{r}\frac{d}{dr}\binom{1}{r}\frac{d}{dr}\binom{1}{r}$ $= c_{*}^{*} \left\{ -L + \frac{d}{2} \left(2L + i \right) = \frac{d_{*}}{2} \left(n + i \right) \right\}$ $\frac{t^{*} dy}{dt^{*}} = \left(\frac{2}{L} \right)^{2} c_{*}^{4} \left(i - \frac{d}{2} + n \cdot \frac{d}{2^{*}} \right)^{*} - \left(\frac{\partial}{\delta} \right)^{*} c_{*}^{*} \left(i - \frac{d}{2} + n \cdot \frac{d}{2^{*}} \right)^{*} = \left(\frac{\partial}{\delta} \right)^{*} c_{*}^{*} \left(i - \frac{d}{2} + n \cdot \frac{d}{2^{*}} \right)^{*} = \left(\frac{\partial}{\delta} \right)^{*} c_{*}^{*} \left(i - \frac{d}{2} + n \cdot \frac{d}{2^{*}} \right)^{*} = \left(\frac{\partial}{\delta} \right)^{*} c_{*}^{*} \left(i - \frac{d}{2} + n \cdot \frac{d}{2^{*}} \right)^{*} = \left(\frac{\partial}{\delta} \right)^{*} c_{*}^{*} \left(i - \frac{d}{2^{*}} + n \cdot \frac{d}{2^{*}} \right)^{*} = \left(\frac{\partial}{\delta} \right)^{*} c_{*}^{*} \left(i - \frac{d}{2^{*}} + n \cdot \frac{d}{2^{*}} \right)^{*} = \left(\frac{\partial}{\delta} \right)^{*} c_{*}^{*} \left(i - \frac{d}{2^{*}} + n \cdot \frac{d}{2^{*}} \right)^{*} = \left(\frac{\partial}{\delta} \right)^{*} c_{*}^{*} \left(i - \frac{\partial}{2^{*}} + n \cdot \frac{d}{2^{*}} \right)^{*} = \left(\frac{\partial}{\delta} \right)^{*} c_{*}^{*} \left(i - \frac{\partial}{2^{*}} + n \cdot \frac{d}{2^{*}} \right)^{*} = \left(\frac{\partial}{\partial} \right)^{*} \left(\frac{\partial}{\partial} \right)^{*} = \left(\frac{\partial}{\partial} \right)^{*} \left(\frac$ the ALL PAR IN LANK KARANI Sie Ti con $\frac{d}{dt} \frac{d}{dt} = -\frac{i}{2} \frac{\partial g_{vv}}{\partial t} \frac{\partial g_{vv}}{\partial x} \frac{dt}{dt} \frac{dt}{dt} \frac{i}{dt} = -\frac{i}{2} \frac{\partial g_{vv}}{\partial t} \frac{\partial t}{\partial t}$ $\frac{3}{r}\frac{dT}{dr} + \frac{1}{r}\frac{d}{r}\left(\frac{i}{r}\frac{dT}{dr}\right) + 2\frac{N}{V^2} + x^2 \frac{dT}{r}\frac{1}{dr}\frac{d}{r}\left(\frac{y}{r}\right)$ de (1+x2) (1-2.)(2-2.)(2-2')(2-2') $\frac{2\pi}{V_{2,1_{1}}} \left\{ 1 + \frac{1}{2} \left(\frac{1}{2} + \frac{1}{2} \right) \left(\alpha + \frac{2}{2} + \frac{1}{2} \right) \right\}$ H $\int_{-\infty}^{\infty} \frac{d^{-1}}{dt} dx^{-1} + x^{+} dy^{-1} \left(1 - 2\frac{d}{2} - \mu \cos \frac{d^{2}}{2x} \right) = \int_{0}^{\infty} x^{+} dy^{-1} \frac{1}{2} - E +$ 2'+2" - S- & / C = - S. - #2 -+ # # 2 d2' (() (1-2' + w))= dy = 3 go 2' $+2, +2, +2'+2'' = \frac{A}{2z(c_{0}-b)}, 2, +2, \frac{A}{2z(c_{0}-b)} + \frac{A}{C}$ + 1/2 (1 1/4 - 1/16 y de (x/de) - x de (y/de) = d de yx-xy = de 60 6 (1-2 + 100 1,2, = الم المد المعد - 2 0 2"+ A (21+1) 4 2 3- [A" (n+1 ay = B. 1-12-14-1 arl) = - 1 1 6 di (ENCO) 1-2'+ (+2) 12'-+ 1 71 1=-1" - 3 + 1 + 8 + - 5 + bed Di grossen Hayde + 2, 12 plats . Si bethe Klam 2 and 2' 1 = 3 # Inhorment mon 2'+2 2 A E (1+04)= ((1-2) , 2,+22 = 2(1+1) 1 = 3 ! $2_{1} 2_{1} = \frac{1}{E} \left[k \left(\frac{1}{E} \right) + (n+1) A \right] - \frac{1}{E} A = \frac{1}{E} \left[\left(\frac{1}{E} \right) \left(\frac{1}{E} (n-1) A \right) \right]$ + (+ # + + 1 - 27)/+1 $\frac{1}{2}\left[\int_{\mathbb{R}^{d}} \frac{1}{2} \int_{\mathbb{R}^{d}} \frac{1}{2$

the question is

Could gravity be an illusion?

A circumstance relative only to your state of motion?

Could gravity be "transformed away"

by the change of a reference frame?

Maybe gravity is not a force at all?

there should be observable consequences

and Einstein knew it

and calculated them - half a decade of Newton-like concentration

35

what we've found:



gravitating bodies...masses:

warp both space and time.

They warp: spacetime

Einstein had to learn that geometry & <u>energy-mass</u>

interact & that space and time respond

That took him 5 years after his happy thought to figure

out



he had to go back to school...privately with his buddy Marcel Grossman





tests of general relativity

There are a handful of "classic tests"

of these ideas:

that space and time are warped by gravitation



40

light beam what about time?

use a clock



В

Gravitational Red Shift is built so B moving away from A A and B ulection of the set of inertial frames at each ifiyou get where you want to go you just confirmed General Relativity

force up at g



receives at say 5 ticks per second

between ticks: per second

1 second

"Advance of the Perihelion of the Orbit of Mercury"

Vulcan?

Mercury

misbehaves

"advance of the perihelion"

Einstein calculated it including the sun's warping of space





1916: Got precisely the right amount.

Had heart palpitations when the result appeared on his paper...

point of closest approach of the orbit advances by 43 seconds of arc per century



the mother of all experiments

the "solar eclipse" experiment





Sep 23, 2013 - Check out this fun science fair project idea for 1st grade: a solar eclipse model that demonstrates total partial and annular solar eclipses

"Solar Eclipse Experiment"

light

obeys the strong Equivalence Principle

the laser pointer...for real

Not totally surprising...light has energy...behaves like mass - it should bend

The star is actually at A

Zariel. 14. K. 13. ch gecharter Hers Kollege! time surfache theoretische Ufer legung macht die Annahme plausikel, dass Lichtstrahlen in einem Geavitation felde eme Deviation uphren. 15 Lechrobahl An Tonnenrande misste diere Ablenkung 0,84 betrayer and wie 1 abuchmen To ware deshall von geösstem Intresse, bis gu wie grosse Somen whe fills Firsteene bei turendung der stinketen Vergrösserungen ber Tage (ohne Some of insternis) gerehen werden komun · Same of instances gereken werden 1911 calculation – initially wrong, only the E=mc² component...

In 1915 he changed his 1911 calculation to include the warping of space...worth x2

The deflection should be about 1/4 milli-degree



eclipse experiment May 29,1919

Sir Arthur Eddington led 2 teams:

Gulf of Guinea

& Brazil





there was some cloudiness!

Eddington had 10 seconds to get a photograph







1/16 plates had usable data

Eddington announced the result

November 6, 1919 at the Royal Astronomical Society meeting

Einstein woke up in Berlin the next morning and was famous.

eclipse announcement at scientific meeting, 11/06/19:

instant celebrity, 11/07/19

cover of December 14, 1919 issue of Berliner Illustrirte



caption: "A new great figure in world history: Albert Einstein, whose investigations signify a complete revision of our concepts of Nature, and are on a par with the insights of a Copernicus, a Kepler, and a Newton."

LIGHTS ALL ASKEW, the Times golf editor IN THE HEAVENS New York Times, Men of Science More or Less November 10, 1919 Agog Over Results of Eclipse Observations. One of the speakers at the Royal Society's meeting suggested that Euclid EINSTEIN THEORY TRIUMPHS was knocked out. Schoolboys should not rejoice prematurely, for it is pointed Stars Not Where They Seemed out that Euclid laid down the axiom or Were Calculated to be, that parallel straight lines, if produced but Nobody Need Worry. ever so far, would not meet He said nothing about light lines. A BOOK FOR 12 WISE MEN Some cynics suggest that the Einstein theory is only a scientific version of the well-known phenomenon that a coin in a No More in All the World Could Comprehend It, Said Einstein When basin of water is not on the spot where His Daring Publishers Accepted It. it seems to be and ask what is new in the refraction of light. Albert Einstein is a Swiss citizen, Special Cable to THE NEW YORK TIMES. LONDON, Nov. 9.-Efforts made to about 50 years of age. After occupying put in words intelligible to the nona position as Professor of Mathematical eclentific public the Einstein theory of light proved by the cellpse expedition Physics at the Zurich Polytechnic so far have not been very successful. The new theory was discussed at a recent School and afterward at Prague Unimeeting of the Royal Society and Royal versity, he was elected a member of Astronomical Society, Sir Joseph Thom-Emperor William's Scientific Academy son, President of the Royal Society, docinres it is not passible to put Einstein's in Berlin at the outbreak of the war. theory into really intell'gible words, yet at the same time Thomson adds : Dr. Einstein protested against the Ger-"The results of the eclipse expedition man professors' manifesto approving of demonstrating that the rays of light from the stars are bent or deflected Germany's participation in the war, and from their normal course by other aerial at its conclusion he welcomed the revobodies acting upon them and consequently the inference that light has lution. He has been living in Berlin weight form a most important concribution to the laws of gravity given us for about six years. sipce Newton haid down his principles." When he offered his last important Thompson states that the difference between theories of Newton and those of work to the publishers he warned them Einstein are infinitesimal in a popular there were not more than twelve persons sense, and as they are purely mathematical and can only be expressed in in the whole world who would understrictly scientific terms it is useless to stand it, but the publishers took the endeavor to detail them for the man in the; street.

risk.

New York Times, December 3, 1919

EINSTEIN EXPOUNDS HIS NEW THEORY

It Discards Absolute Time and Space, Recognizing Them Only as Related to Moving Systems.

IMPROVES ON NEWTON

Whose Approximations Hold for Most Motions, but Not Those of the Highest Velocity.

INSPIRED AS NEWTON WAS

But by the Fall of a Man from a Roof Instead of the Fall of an Apple.

Copyright, 1919, by The New York Times Company Special Cable to THE NEW YORK TIMES.

BERLIN, Dec. 2 .- Now that the Royal Society, at its meeting in London on Nov. 6, has put the stamp of its official "authority on Dr. Albert Einstein's muchdebated new "theory of relativity," man's conception of the universe seems likely to undergo radical changes. Indeed, there are German savants who believe that since the promulgation of Newton's theory of gravitation no discovery of such importance has been made in the world of science.

When THE NEW YORK TIMES correspondent called at his home to gather from his own lips an interpretation of what to laymen must appear the book with the seven seals. Dr. Einstein him-



now recovered from exhaustion and photogenic: 1920



Gravitational Lensing - an off-hand prediction of Einstein





Foreground objects can distort, and magnify distant background galaxies.



SDSS J1420+6019

SDSS J1627-0053

Today, the dramatic effects of light bending are observed in the form of gravitational lensing

this is the bending of light around a very massive object, like a large galaxy tool for studying dark matter: looking for Massive Astrophysical Compact Halo **ObjectS**

MACHOS

Credit: Canada-France-Hawaii Telescope Corporation 2006.

"Gravitational Lensing"



"Einstein Cross" - 4 images of a quasar







Second view of this:

"The Hafele-Keating experiment"

an atomic clock was carefully carried around the world in 1972 and carefully calibrated and compared with groundbased clocks



correct

Predicted Effect	Flying East	Flying West	
GTR (Gravitation) STR (Velocity) Total	+ 144 ± 14 ns - 184 ± 18 ns - 40 ± 23 ns	+ 179 ± 18 ns + 96 ± 18 ns + 275 ± 21 ns	
measured:	- 59 ± 10 ns	+273 ± 7 ns	

redone twice more in airplanes and rockets/satellites



J. Hafele and R. Keating

about half of their effect was due to the gravitational difference between Earth and the flight's altitude

spacetime in general relativity

Einstein got rid of gravitational forces in GR

Masses warp spacetime...

Since the shortest distance between two spacetime points is a light-path, this "maps" the shape

In GR gravity is not a force, but a "topography" of spacetime that forces objects to take the shortest curved path in spacetime

Earth's orbit is then just us following the shortest distance around the sun...not a gravitational force



okay.

Spacetime might be curvy, bumpy, ..."warped"

a "non-Euclidean geometry"?

Euclid's Geometry starts with 4 terms and 4+1 postulates:

Point, Line, Plane, Space

- A straight line can be drawn between any two points 1.
- 2. A finite line can be extended infinitely in both directions
- A circle can be drawn with any center and any radius 3.
- All right angles are equal to each other 4.
- Given a line and a point not on the line, only one line 5. can be drawn through the point parallel to the line

a System of a series of proofs, each building on the previous, to a whole system of mathematics



like, actually... a lot of algebra problems before algebra was invented

Einstein's mathematics of GR

led him to have to consider non-Euclidean Geometries which were still timidly being studied by mathematicians

not so far-fetched

we live in such a geometry



what's a ''straight line''



on a sphere?

shortest distance between 2 points

http://gc.kls2.com/



Euclid's 5th Postulate

parallel lines never meeting?

only in a flat space

sum of interior
angles in a triangle
= 180°?

only in flat space

on sphere > 180°



"warping"

means that geometry

spacetime geometry

mixes with mass, energy, and pressure

General Relativity

Einstein's GR equation

complicated mathematics geometry of spacetime

$R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} = \frac{8\pi}{c^4} T_{\mu\nu}$

we'll call it: "G = T "



 \leftrightarrow

mass-energy, pressure, & momentum



Einstein grossly underestimated

the richness of his theory

he knew he'd exhausted the possible solutions to the GR equations

He was wrong...and irritable about it

wrong. Almost immediately:

from the foxhole, 1915



Karl Schwarzschild, 1873-1916

Yes. I mean *from* a foxhole.

The **first exact solution** to GR...Einstein had used some approximations for light-bending, etc.

The equations of spacetime outside of a spherical mass.

a big mass.



escape

Suppose a rocket is shot straight up... when it goes "ballistic" (no propulsion)...what happens?

It depends.



More initial velocity, the more likelihood that the rocket will escape the pull of the Earth's gravity.

This happens when the kinetic energy = potential energy

$$v_{\rm esc} = \sqrt{\frac{2GM_E}{R_E}}$$

From Earth: 11.2 kilometers per second...~25,000 mph

 $v \approx v_{esc}$

what about light?

suppose the question is not:

"What's the escape velocity from a sphere of mass M?"

BUT

"What's the radius of a mass M for which the escape velocity is = *c*?"

 $v_{\rm esc} = \sqrt{\frac{2GM_E}{R_E}} \quad \longrightarrow \quad c = \sqrt{\frac{2GM}{R_S}}$

R_S called the Schwarzchild Radius

$$R_S = \frac{2GM}{c^2}$$

It seemed to be a magic radius...



66

the

Schwarzchild Radius falls out of his solution to G.R.

it's not likely

RS is incredibly small

and density, incredibly high All of the mass of:



Impossible, right?. But, since Nature doesn't do infinity...thought to be a disaster for the theory.

inside of R_S :





everyone fretted over this for more than a decade

1932, Georges Lemaître found that a slight change of coordinate axes

changed the problem completely

Black Holes

The Schwarzchild Radius was not a flaw in the theory

simply an insult from Nature!





Einstein calculated that the normal formation of a star of gravitational accretion could never form in so small a volume...and stars get bigger not smaller, right?

1939: Robert Oppenheimer & Hartland Snyder showed how.

5¢ worth of stellar physics no charge

Hertzsprung-Russell Diagram...aka H-R Diagram



5¢ worth of stellar physics no charge

Hertzsprung-Russell Diagram...aka H-R Diagram



stars radiate energy – that's their job!

being stable is their challenge...



The "main" sequence

a balancing act

inward pressure from gravity

VS

outward pressure from radiation



A star's fate is determined by how massive it is.

gravity pulls core/atmosphere: in Radiation pressure from nuclear fusion in core: out

H begins to "burn" to He