Reconstruction in DeepCore and PINGU





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Avviso!

- The results shown in this presentation are preliminary
 - Contents under pressure
 - Listeners assume all liability for damages, incidental or otherwise
 - This material may be hazardous to your health or scientific reputation
 - I was never here. You do not know me.
 - The Secretary will disavow all knowledge of this talk
 - The first rule of Fight Club is that you don't talk about Fight Club
 - Burn before reading
 - These slides will self-destruct in 5 seconds

The Problem (as I see it)

- Our traditional techniques were developed originally for small detectors and high energies
- They assume infinite tracks, starting outside the detector and passing completely through
 - Five parameters: x, y, t, θ, φ
 - No attempt to discriminate particle type at best, compare separate likelihoods
 - Assume minimum ionizing, and measure energy separately ex post facto
 - Rely on high photon statistics to counteract any problems with description of photon transport
 - Use only first arriving photon (basically, enforce causality), and ignore dogs that don't bark
- For (most of) these physics topics, we need to do better
 - Eight parameters (or 10?): x, y, z, t, $\theta_{(\mu)}$, $\varphi_{(\mu)}$, Ec, L_{μ} , (θ_c , φ_c)?
 - Accurate track vs. no-track discrimination very important
 - Can we play statistical games to separate v from \bar{v} , v_{τ} from v_{e} , and/or NC from v_{e} ?

The Story So Far

- In IceCube, we have suffered from unrecognized numerical problems in our photon transport models for many years
 - Affected both reconstruction and simulation
 - Separate from uncertainties in optical properties of ice
- Fixing the obvious problems in our reconstruction algorithms did not yield expected improvements (often made things worse)
 - Likelihood space fundamentally inaccurate better minimizers gave worse results, more correct descriptions gave worse results, etc.
 - Kudos in particular to Jakob v. Santen, Marek Kowalski, and Nathan Whitehorn for figuring out what was wrong





SPE as configured in IC79 L2





SPE as configured in IC79 L2

From Chang Hyon's Uppsala talk Binning artifacts, and older likelihoo





Where We Stand

- We have (mostly) fixed those issues in the last year
 - In the process of implementing all the ideas we had over the last decade...
- Key new techniques are GPU-enabled direct photon transport simulations, and a smooth, non-parametric spline-based descriptor of photon transport for reconstructions
- The next few slides were smuggled out of Aachen and show the current state of the art with (mostly) fully-implemented production reconstructions
 - Numerical issues solved in cascade reconstructions
 - Track reconstructions allowing non-infinite tracks, but either based on direct photorec tables (with artifacts)...
 - ...or with splined photon description but numerical approximations for high light levels (e.g. Gaussian amplitude distribution)

Directional Reconstruction (Cascades)

- Decent angular resolution possible with Credo and enough CPU cycles
 Klaus Wiebe, Alexandra Schulte
- Median space angular res. around 25° for reasonable NCh and NStr
 - Around 50° for
 NCh > 8
- Good enough for interesting physics



Remarks: Cut1 is nstr>1 and nch>8, Cut2 is nstr>4 and nch>21, switch red and yellow in legend of right plot x axis: angular resolution, y axis: abundance (left), cumulative events (right)

Energy Reconstruction (Cascades)

• With appropriate

seeds, energy resolution comparable to that at higher energies

 Some oddities seen in DOM responses, improvements still possible?



Current low-en cascade reco. status: Monopod seeded with Credo Resolution: ~0.26

> Vertex.energy* f_{had} + Particle.energy Vertex.energy**f*had

Donglian Xu

 $E_0 = 0.399, m = 0.130, f_0 = 0.467$

Ref: Marek Kowalski's PhD dissertation



Directional Reconstruction (Muons)

Median zenith error [degrees]

Events

0.2

30

25

20

- Advanced fits with some quality cuts now achieve decent ang. resolution at oscillation E's
 - 12°-15° in zenith at 25 GeV
- Signal loss around 50%, but we are not statistically limited

Results for only zenith

Meike de With



Meike de With | IceCube Collaboration Meeting | October 3, 2012 | 14

Directional Reconstruction (Muons)

fit

SANTA zenith

- An alternative approach to fitting, focusing on zenith directly
 - Avoids direct dependence on ice modeling, but at the cost of lower efficiency

Juan Pablo Yañez, Jürgen Brunner

Zenith error as a function of E



 Maintains resolution better than 10° down to 20 GeV *neutrino* energy, for the selected events



Track Length Reconstruction

- Reconstructed neutrino energy depends
- largely on track length (min. ionizing tracks)
- With correct trajectory, can achieve decent length reco. (~30% if vertex cascade small)
 - Better for vertical events, worse for larger cascades

Millipede/finiteReco combination

Better end point estimation of finiteReco
Better vertex reconstruction by millipede

finiteReco end - millipede vtx, track, vert. up sample

Length = finiteReco end - millipede vtx.

Andrii Terliuk





Better performance with better tails, but lower efficiency of the method.



- At Penn State, we've been working on a full likelihood description of events, with the following assumptions:
 - One event in the detector at a time
 - Cascade vertex aligned with muon track, if one exists
 - Muons are minimum-ionizing, no stochastics
 - (Assume no physics hits expected more than 300 m from light sources, only noise)
 - Poisson likelihood based on spline-fitted light tables for muon track segments and cascades, full waveform information used, all DOMs included
 - Based on several years of work by Pat Toale, Sven Lafèbre, Chang Hyon Ha, Mike Prikockis, and most recently Matt Dunkman and Ryan Eagan

First Step:Validate Likelihood

- MC truth will not, in general, have the best likelihood
 - With a perfect likelihood description, the true values will be distributed with characteristic spread of 1σ around the extremum of the likelihood space
 - The corresponding errors in the parameters of interest represents the (optimal) resolution of the detector
- This presumes that the optimum associated with the MC truth is the global optimum
 - In the past, this has not been true
 - Manual scans of slices of likelihood space suggest this is now the case (for at least 90% of events), but it remains to be proven definitively



Intrinsic Resolutions

- In the following plots, we attempt to measure the intrinsic resolution (error between truth and likelihood optimum nearest the true values)
 - We run the full 8-parameter reconstruction, using the MC truth as the seed value
 - Not a perfect measure minimizer can get stuck or not find the likelihood optimum, and conversely the optimum may not be global

• Nonetheless, this is a meaningful measure of detector performance

- A detector incapable of measuring parameters would have a shallow likelihood space, and the 1σ range would be wide
- We have verified that the minimizer is exploring the local space (50-100 steps)
- If optima are global, sufficient application of computing power will allow us to approach arbitrarily close to the optimal resolution – a matter of event selection

Event Sample

- Charged current muon neutrino events, 10 GeV $< E_{\nu} < 100$ GeV
 - Hadronic cascade at interaction vertex included, but using generic nugen simulation, not full GEANT
- Neutrino interaction vertex within DeepCore
 - Radial distance less than 150 m of DeepCore central string (a bit larger than the detector radius)
 - Vertical position -200 m < z < -400 m (DeepCore DOMs range from -150 m to -500 m)
- More than 8 hits remaining after standard hit cleaning
- Recently squished a bug and reimplemented for speed, low statistics available at present

Zenith Angle Resolution

• Zenith angle is the relevant one for oscillation studies

- Comparing to true muon direction, not the neutrino (kinematics significant at these energies)
- Nearly unbiased, RMS resolution < 10°
- Azimuthal resolution worse, due to asymmetric spacing (72 m between strings, 7 m between DOMs)



Energy Dependence

- In pre-bugfix version, median zenith resolution around 10° at lowest muon energy (1-10 GeV), down to <5° at 25 GeV
 - Neutrino energies a factor of 2 higher, on average (but with large spread)
 - Need more statistics to confirm with current version, but resolution looks good
 - Angular errors will likely be dominated by kinematics of neutrino interaction



Track Length Resolution

• Unbiased, RMS resolution 11 m (equivalent to ~2.3 GeV)

• Note: track segment granularity is 15 m!



Cascade Energy Resolution

• A tail to overestimated energy, RMS resolution 15 GeV

- Relation to event parameters (position, Björken-y, etc.) unknown at present
- Apparently the dominant component of the energy resolution
- Can we improve resolution, or select events with better reconstructions?



PINGU Reconstruction Studies

- Uses one candidate geometry, of several under discussion
 - 20 strings of 60 DOMs
 - 26 m inter-string spacing, 5 m DOM spacing
 - Full GENIE simulation, including hadronic vertex alignment and GEANT tracking of all charged particles

• Again, a study of intrinsic resolution theoretically possible

- Assumes selection of events interacting within the volume
- No data processing pipeline, just examination of the precision of the likelihood minimum around the MC truth
- No estimate of efficiencies associated with background rejection
- So: preliminary!

PINGU Reconstruction Studies

• Errors relative to true muon direction, not neutrino direction



Outlook

- Full likelihood reconstruction looks promising
 - Intrinsic resolution is very good, and we *think* the global optima are in the right place
 - Current version running quickly (~1 second per event in DeepCore, 5-10 in PINGU)
 - Somewhat longer processing times expected with real seeds, and of course much longer if intensive scans of the likelihood space are required – but this would be an annoyance, not a fundamental problem
- Need to establish complete analysis pipeline
 - Optimization is difficult likelihood falls away very quickly in all parameters, so accurate seeds for all parameters are needed
 - May require new minimizers or other tricks
 - Event selection efficiencies are still unknown (in DeepCore: 10-20%)