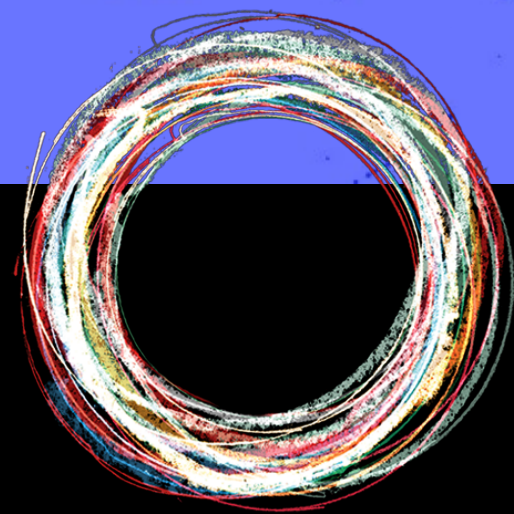


# LECTURE 7

## FEYNMAN RULES



PHY 493/803, 2017

# Recap / Up Next

Last time:

Bound States

Hydrogen

'Onium

Mesons/Baryons

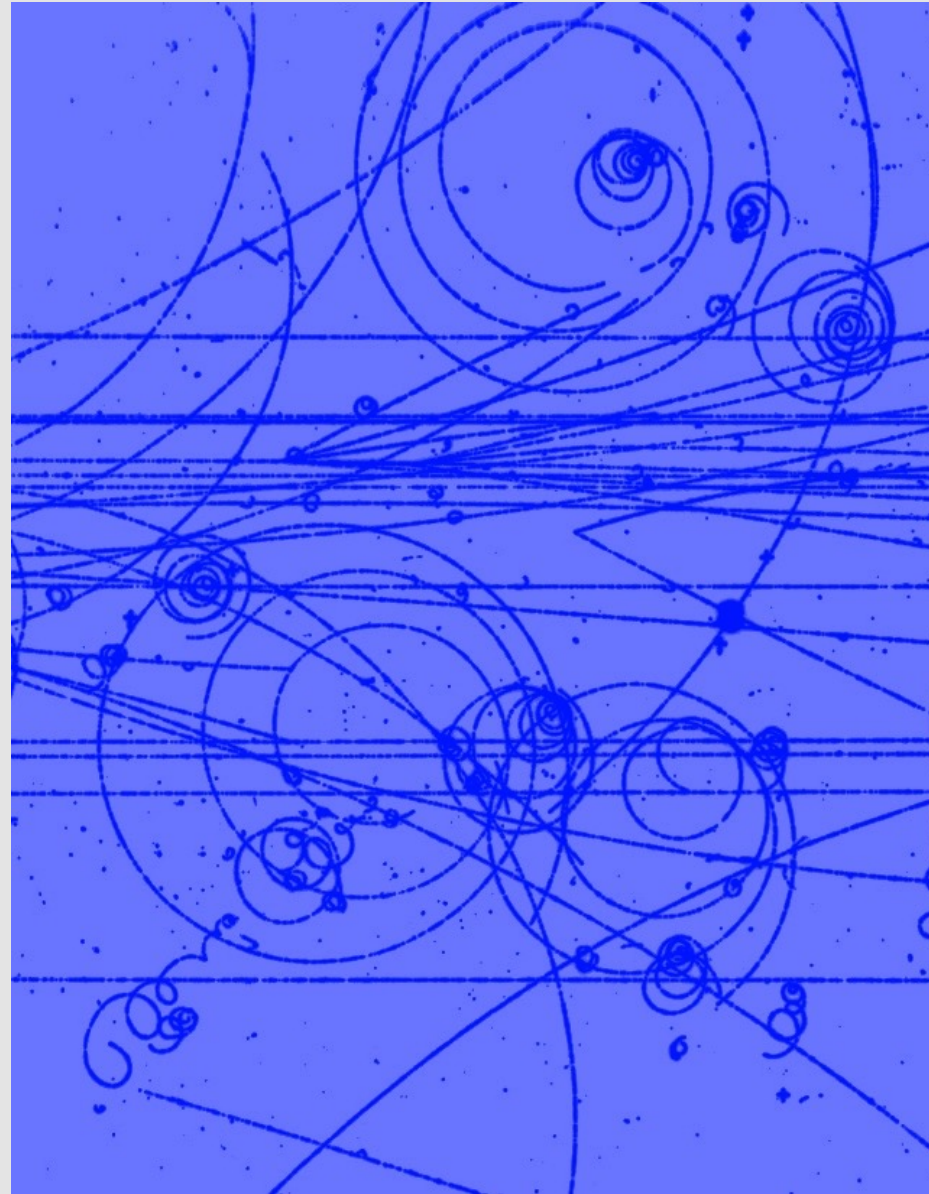
This time:

Feynman Calculus

Decays/Scattering

The Golden Rule

Feynman Rules



# Observables

---

To learn anything about particle interactions, we have to observe something.

There are two basic categories:

A) Bound states & their spectra (this was last chapter)

# Observables

To learn anything about particle interactions, we have to observe something.

There are two basic categories:

A) Bound states & their spectra (this was last chapter)

B)

**When particles are  
“Left Alone”, they can:**

- 1) Do nothing
- 2) Decay
- 3) Eventually find  
another particle and  
go to category C.



# Observables

To learn anything about particle interactions, we have to observe something.

There are two basic categories:

A) Bound states & their spectra (this was last chapter)

**B)**

**When particles are  
“Left Alone”, they can:**

- 1) Do nothing
- 2) Decay
- 3) Eventually find another particle and go to category C.

**C)**

**When particles encounter  
another particle, they can:**

- 1) Do nothing
- 2) Scatter off the other particle.
- 3) Annihilate on the other particle

# Observables


To learn anything about particle interactions, we have to observe something.

There are two basic categories:

A) Bound states & their spectra (this was last chapter)

B)

**When particles are  
“Left Alone”, they can:**

- 1) Do nothing
  - 2) Decay
  - 3) Eventually find another particle and go to category C.
- 

C)

**When particles encounter  
another particle, they can:**

- 1) Do nothing
- 2) Scatter off the other particle.
- 3) Annihilate on the other particle

Decay rate ( $\Gamma$ ): Describes how quickly particles disappear.

Lifetime ( $\tau$ ): Describes how long particles stick around

$$\tau = 1/\Gamma$$

# Decay rates & rules

---

Generally, all particles are unstable and can decay. Rough rules:

- 1) The final state cannot have more total mass, to conserve energy.
- 2) If there is not a lower energy/mass state, the particle cannot decay.
- 3) The decay must satisfy all normal conservation rules. Eg, charge conservation.

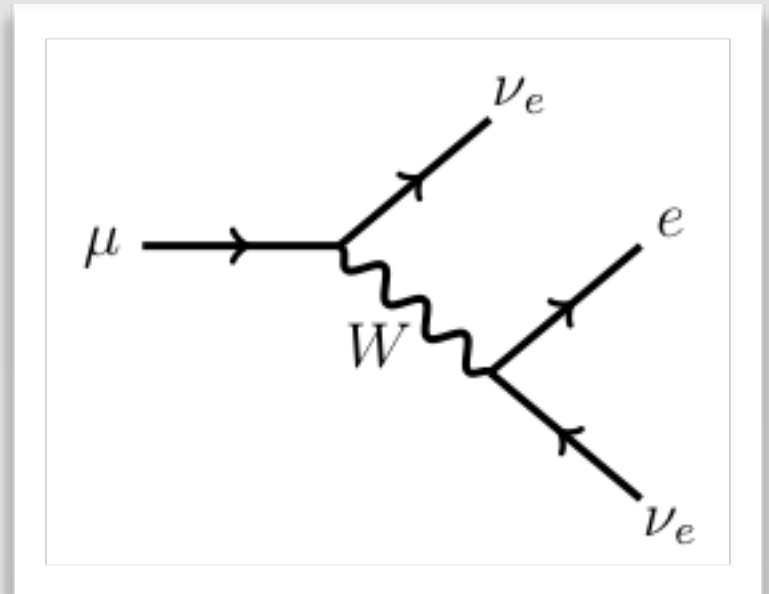
# Decay rates & rules

Generally, all particles are unstable and can decay. Rough rules:

- 1) The final state cannot have more total mass, to conserve energy.
- 2) If there is not a lower energy/mass state, the particle cannot decay.
- 3) The decay must satisfy all normal conservation rules. Eg, charge conservation.

Consider muon decay:

- The final state mass is much smaller than the initial state mass: 0.5 MeV vs 106 MeV
- The W boson must be virtual ( $M_W \sim 80 \text{ GeV}$ )
- The final state momentum is:
  - Shared amongst 3 particles
  - Determined by the difference in mass between initial and final states.





# Decay rate and Lifetime

Particles have no concept of history

The probability a particle will decay in a fixed period is constant.

A population of decaying particles can be described as a function of time.

$$\begin{aligned}\Delta N(t) &= N(t_a) - N(t_b) \\ &= p N(t_b) (t_a - t_b)\end{aligned}$$

# Decay rate and Lifetime

Particles have no concept of history

The probability a particle will decay in a fixed period is constant.

A population of decaying particles can be described as a function of time.

$$\begin{aligned}\Delta N(t) &= N(t_a) - N(t_b) \\ &= p N(t_b) (t_a - t_b)\end{aligned}$$

Written differentially, we have:

$$dN = -\Gamma N dt$$

$$N(t) = N_0 e^{-\Gamma t} = N_0 e^{-t/\tau}$$

$$\Gamma = 1/\tau$$

# Summing Decay Rates

What if a particle has more than one path to decay?

The total decay rate just becomes the sum of individual decay rates.

$$W^+ \rightarrow e^+ + \nu_e$$

$$W^+ \rightarrow \mu^+ + \nu_\mu$$

$$W^+ \rightarrow \tau^+ + \nu_\tau$$

$$W^+ \rightarrow u + \bar{d}$$

$$W^+ \rightarrow c + \bar{s}$$

# Summing Decay Rates

What if a particle has more than one path to decay?

The total decay rate just becomes the sum of individual decay rates.

$$W^+ \rightarrow e^+ + \nu_e$$

$$W^+ \rightarrow \mu^+ + \nu_\mu$$

$$W^+ \rightarrow \tau^+ + \nu_\tau$$

$$W^+ \rightarrow u + \bar{d}$$

$$W^+ \rightarrow c + \bar{s}$$

$$dN_W = -(\Gamma_{e\nu} + \Gamma_{c\bar{s}} + \cdots) N_W dt$$

$$\Gamma_{tot} = \sum \Gamma_i$$



# Branching Fraction/Ratio

The fraction of a certain particle's decays to a given final state is referred to as the Branching Fraction or Branching Ratio

It's trivially calculated!

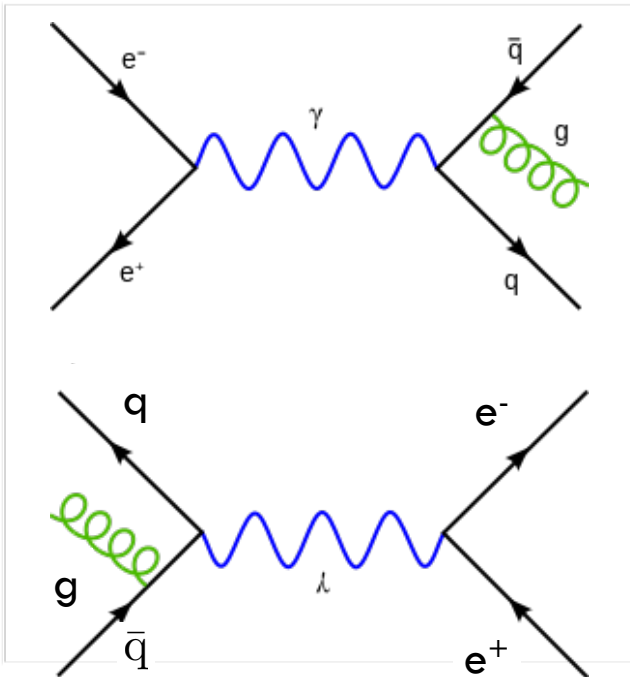
$$BR(W \rightarrow e\nu) = \frac{\Gamma_{W \rightarrow e\nu}}{\Gamma_{\text{tot}}}$$

# Decay & Production

We will see that particle decay and particle production are intrinsically related

You would have guessed this from the idea of CPT and also from how we were able to rotate Feynman diagrams

The decay rate for a process is also referred to as the “width” of the production rate “peak”.

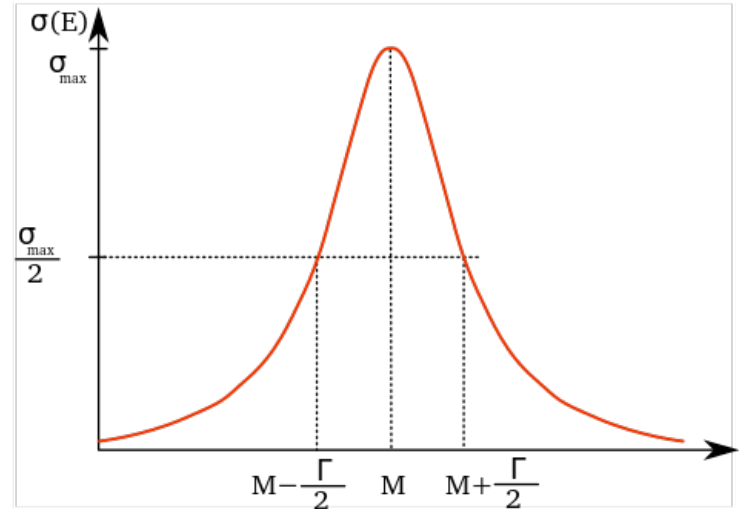
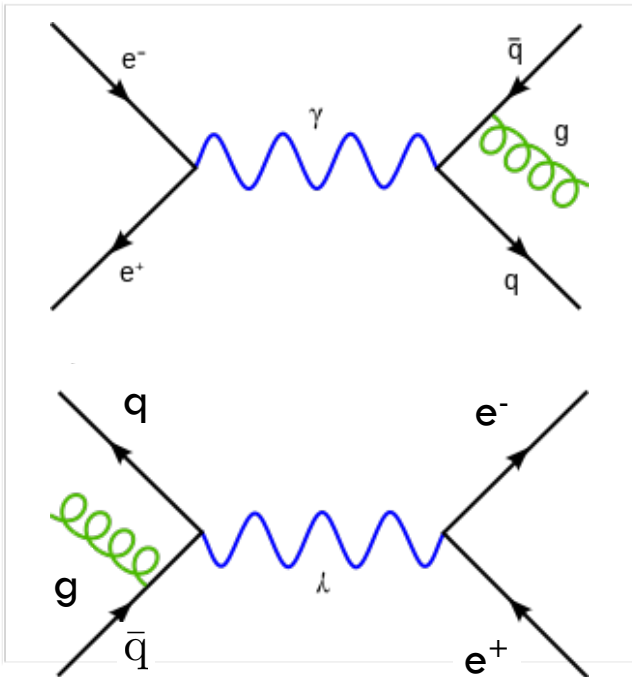


# Decay & Production

We will see that particle decay and particle production are intrinsically related

You would have guessed this from the idea of CPT and also from how we were able to rotate Feynman diagrams

The decay rate for a process is also referred to as the “width” of the production rate “peak”.



$$P(E) = \frac{1}{2\pi} \frac{\Gamma}{(E - M)^2 + \Gamma^2/4}$$

# Observables

To learn anything about particle interactions, we have to observe something.

There are two basic categories:

A) Bound states & their spectra (this was last chapter)


B)

**When particles are  
“Left Alone”, they can:**

- 1) Do nothing
- 2) Decay
- 3) Eventually find another particle and go to category C.

C)

**When particles encounter  
another particle, they can:**

- 1) Do nothing
  - 2) Scatter off the other particle.
  - 3) Annihilate on the other particle
- 

Cross section ( $\sigma$ ): Describes the probability for an interaction to occur.

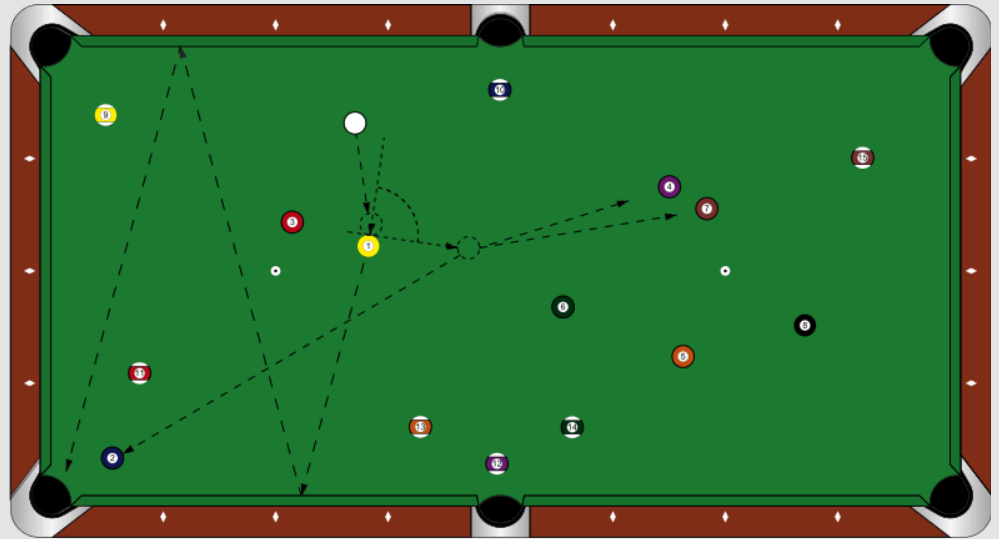
Differential cross section ( $d\sigma/dX$ ): Describes the probability for an interaction with a particular final state.



# Cross Section

We need to connect the idea of a unit area (classical) and the probability for an interaction to occur.

Classically, the cross section is inherently related to the size of an object.

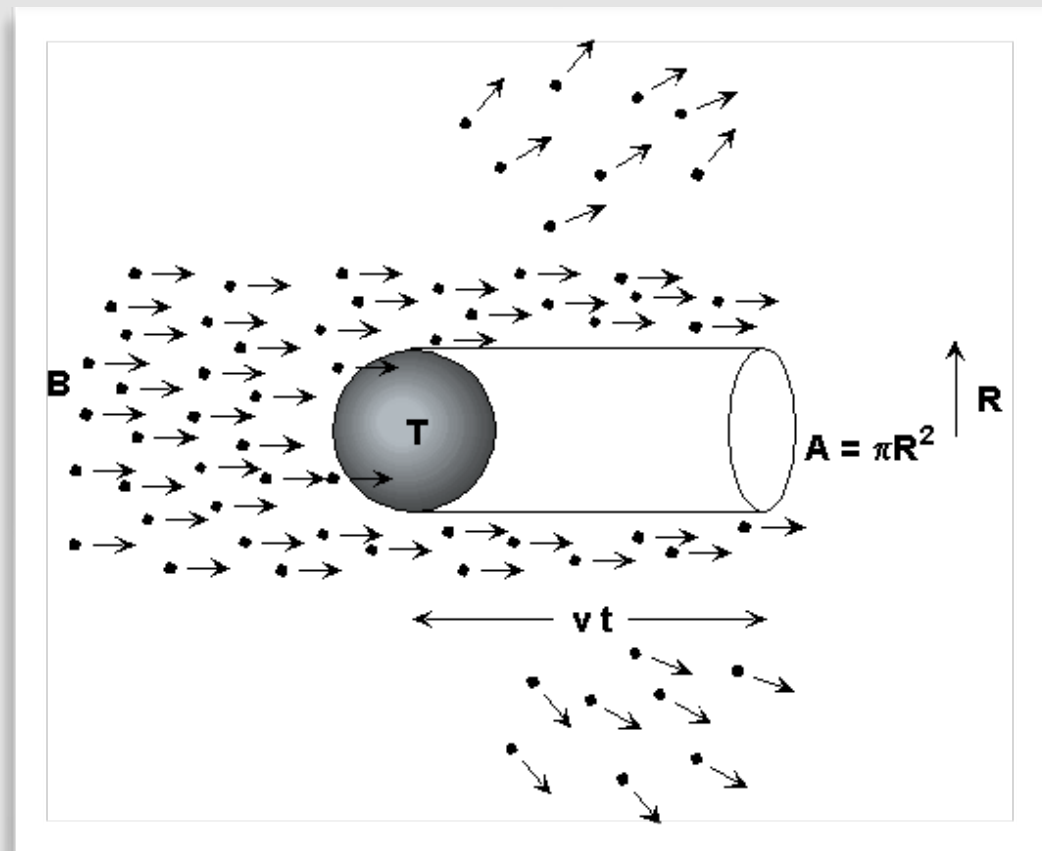


# Cross Section

We need to connect the idea of a unit area (classical) and the probability for an interaction to occur.

Classically, the cross section is inherently related to the size of an object.

In the classical sense, scattering occurs when the particles in a beam overlap their path with the target's cross section



# Hard Sphere Scattering

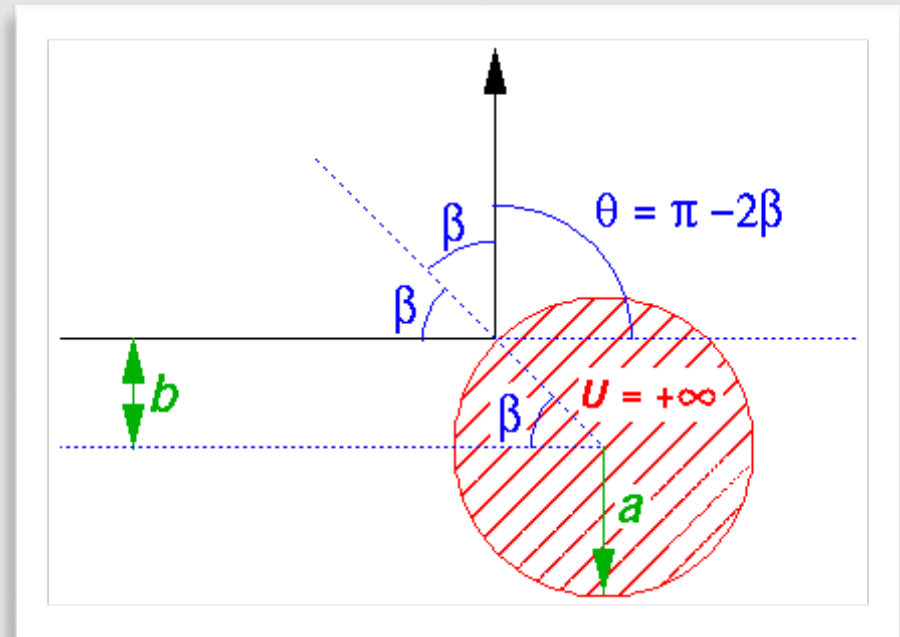
We can start with the classical hard sphere calculation

$$b = R \sin(\beta)$$

$$\theta = \pi - 2\beta$$

$$\sin(\beta) = \cos(\theta/2)$$

$$b = R \cos(\theta/2)$$



# Hard Sphere Scattering

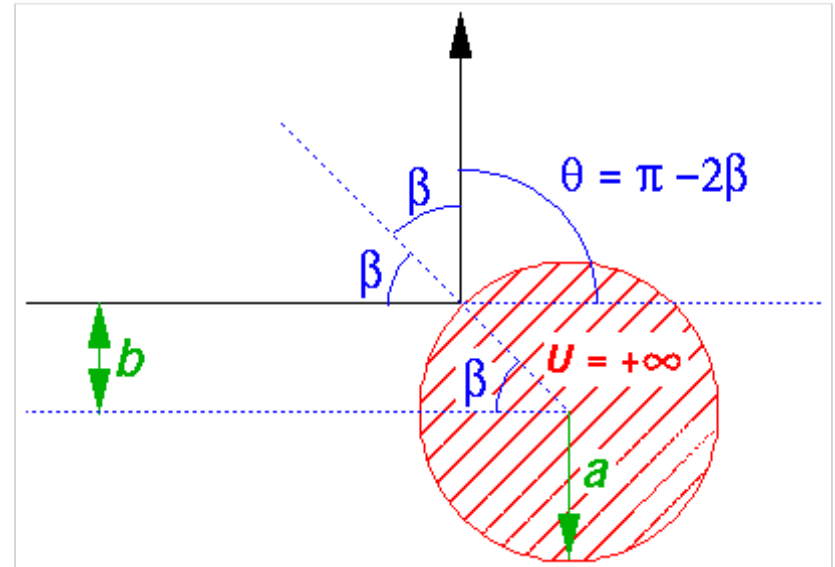
We can start with the classical hard sphere calculation

$$b = R \sin(\beta)$$

$$\theta = \pi - 2\beta$$

$$\sin(\beta) = \cos(\theta/2)$$

$$b = R \cos(\theta/2)$$





# Hard Sphere Scattering

We can start with the classical hard sphere calculation

$$b = R \sin(\beta)$$

$$\theta = \pi - 2\beta$$

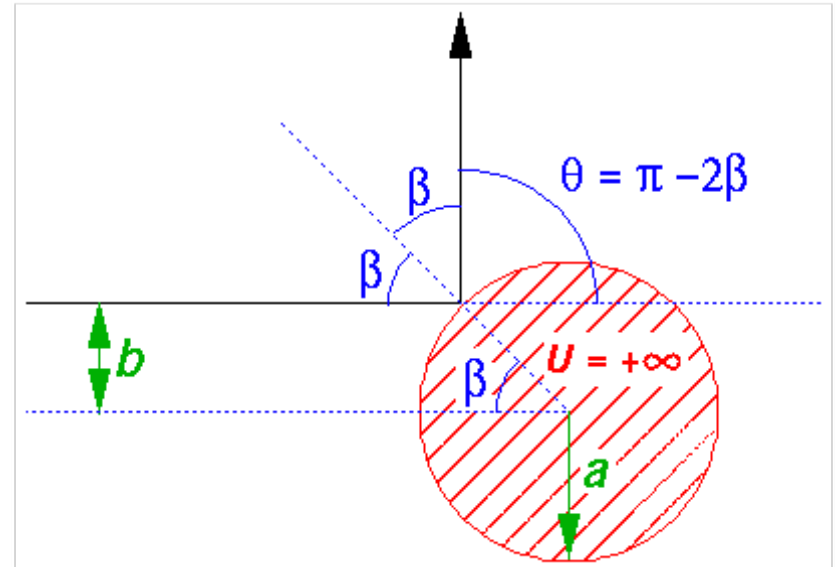
$$\sin(\beta) = \cos(\theta/2)$$

$$b = R \cos(\theta/2)$$

$$d\sigma = 2\pi b db$$

$$d\Omega = \sin(\theta) d\theta d\phi$$

$$\frac{d\sigma}{d\Omega} = \frac{b}{\sin(\theta)} \frac{db}{d\theta}$$



# Hard Sphere Scattering

We can start with the classical hard sphere calculation

$$b = R \sin(\beta)$$

$$\theta = \pi - 2\beta$$

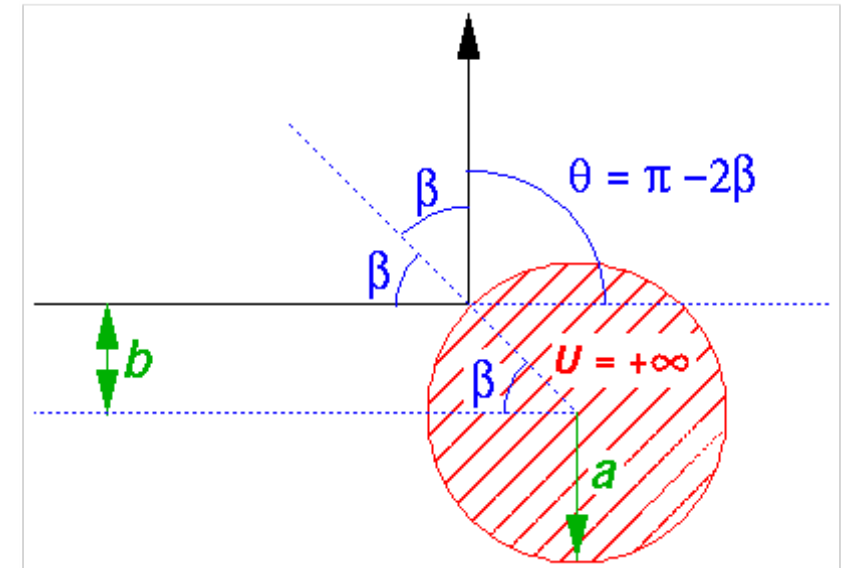
$$\sin(\beta) = \cos(\theta/2)$$

$$b = R \cos(\theta/2)$$

$$d\sigma = 2\pi b db$$

$$d\Omega = \sin(\theta) d\theta d\phi$$

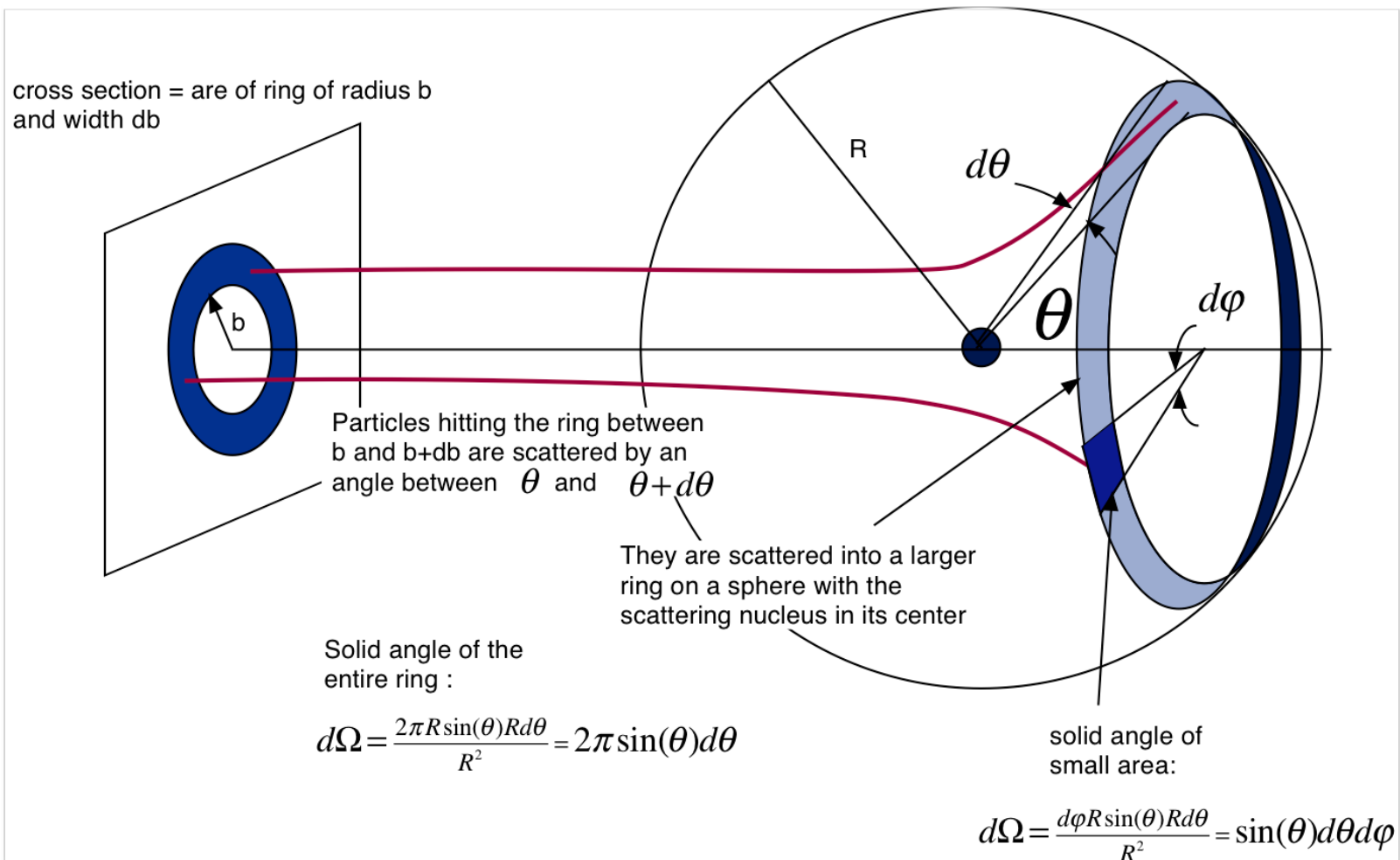
$$\frac{d\sigma}{d\Omega} = \frac{b}{\sin(\theta)} \frac{db}{d\theta}$$



$$\frac{db}{d\theta} = -\frac{R}{2} \sin\left(\frac{\theta}{2}\right) \quad \frac{d\sigma}{d\Omega} = \frac{R^2}{4}$$

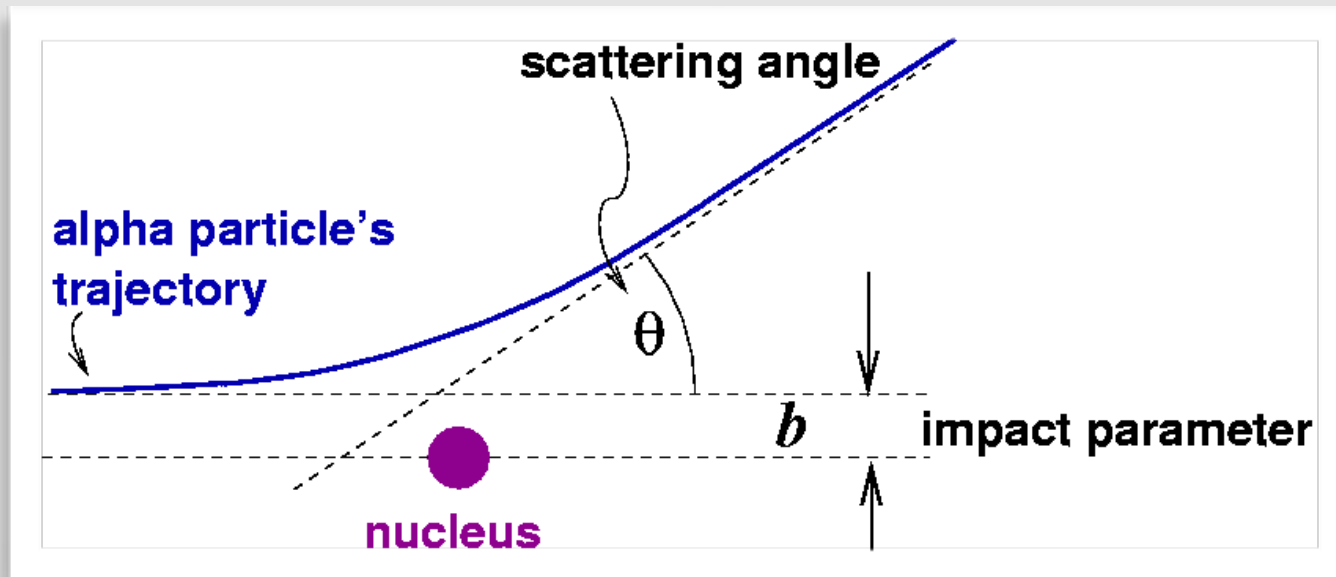
# Reminder: Rutherford Scattering

Rutherford scattering describes classical scattering off a central potential



# Reminder: Rutherford Scattering

Rutherford scattering describes classical scattering off a central potential



$$b = \frac{q_1 q_2}{2E} \cot(\theta/2) \quad \frac{d\sigma}{d\Omega} = \left( \frac{q_1 q_2}{4E \sin^2(\theta/2)} \right)^2$$

# The QFT Cross Section

---

In quantum field theory, the interaction cross section doesn't correspond to the physical size of an object.

Many particles are interpreted as point objects (zero size!)

We now refer to the probability of particles interacting with each other.

We've seen a form of this in the context of spin & angular momentum:

# The QFT Cross Section

In quantum field theory, the interaction cross section doesn't correspond to the physical size of an object.

Many particles are interpreted as point objects (zero size!)

We now refer to the probability of particles interacting with each other.

We've seen a form of this in the context of spin & angular momentum:

$$|1, 1\rangle = \sqrt{\frac{1}{3}} |3/2, 1/2\rangle + \sqrt{\frac{2}{3}} |1/2, 1/2\rangle$$

# The QFT Cross Section

In quantum field theory, the interaction cross section doesn't correspond to the physical size of an object.

Many particles are interpreted as point objects (zero size!)

We now refer to the probability of particles interacting with each other.

We've seen a form of this in the context of spin & angular momentum:

$$|1, 1\rangle = \sqrt{\frac{1}{3}} |3/2, 1/2\rangle + \sqrt{\frac{2}{3}} |1/2, 1/2\rangle$$

Be careful with this particular analogy, though. But we can consider:

The cross section for the  $|3/2, 1/2\rangle$  state is  $1/3$  of the total cross section.

The cross section for anything but  $|3/2, 1/2\rangle$  or  $|1/2, 1/2\rangle$  is 0.

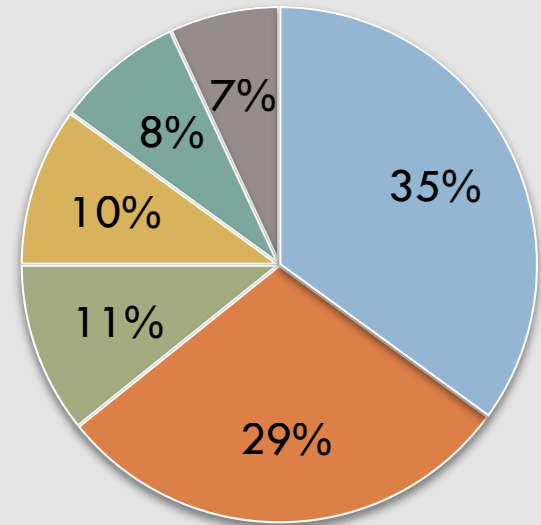
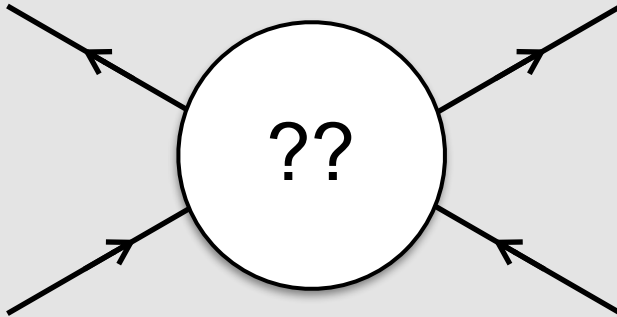
However, we don't know the total cross section!

We'll talk about that in the context of the Golden Rule.

# Total Cross Section

As we saw with the decay rate, the cross section can be subdivided into constituent pieces.

To find the constituents, we have to consider all ways to produce the final state.

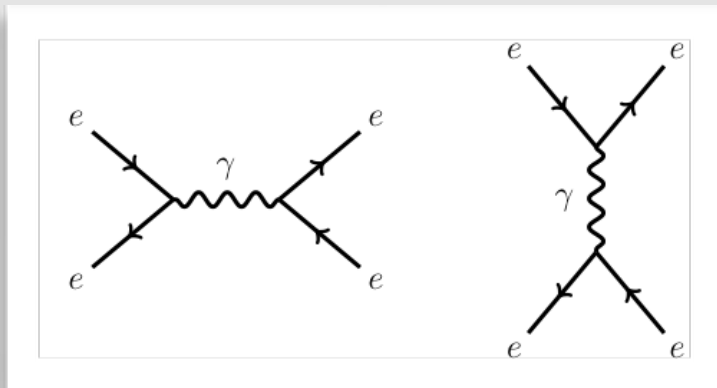
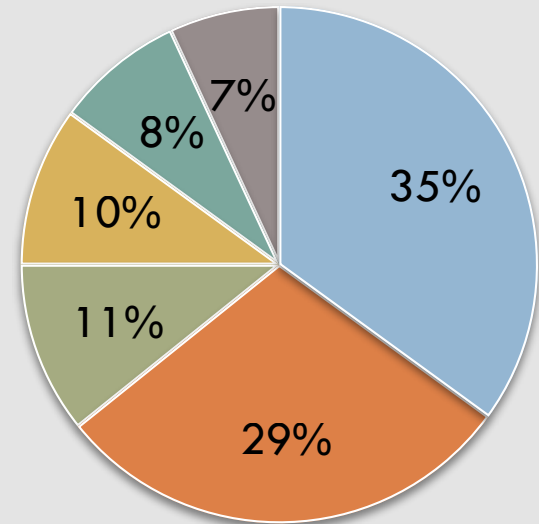
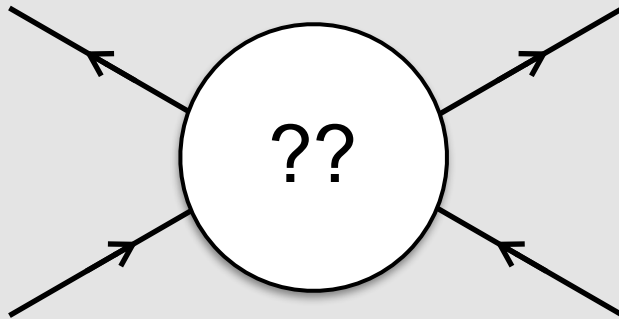




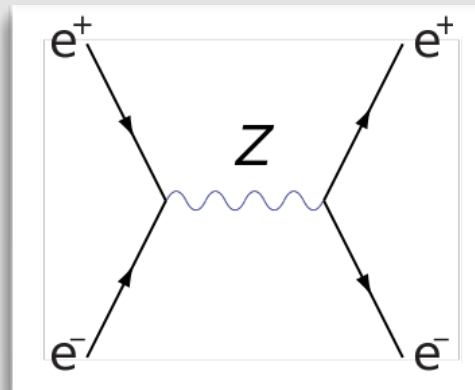
# Total Cross Section

As we saw with the decay rate, the cross section can be subdivided into constituent pieces.

To find the constituents, we have to consider all ways to produce the final state.



+



$$\sigma_{\text{tot}} = \sum \sigma_i$$

# Reminder: From Rutherford Scattering

Cross section ( $\sigma$ ) : the probability of a collision of occurring  
between two particles (beam and target)

Reaction rate (W)

$\propto \sigma$

$\propto N_{\text{target}}$  : number of target  
illuminated by the beam

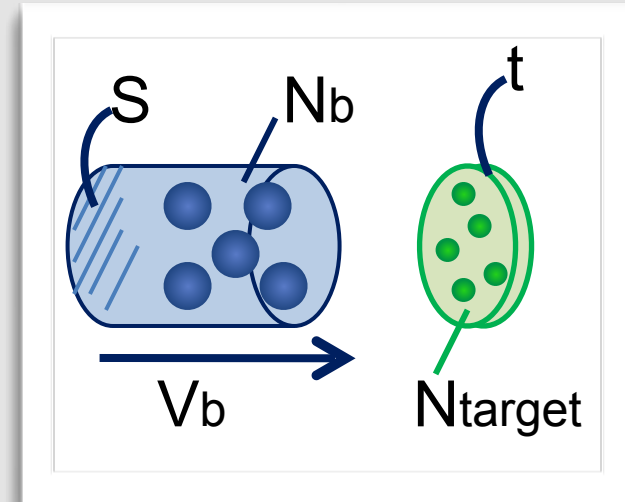
$\propto \text{Flux } J$  : beam rate per unit area

$$J = n_b \cdot V_b$$

( $n_b$  : number density of beam particles)

( $V_b$ : beam velocity)

Beam intensity  $I = J S$  ( $S$ : beam area)



# Reminder: From Rutherford Scattering

Cross section ( $\sigma$ ) : the probability of a collision of occurring  
between two particles (beam and target)

Reaction rate ( $W$ )

$$\propto \sigma$$

$$\propto N_{\text{target}} : \text{number of target illuminated by the beam}$$

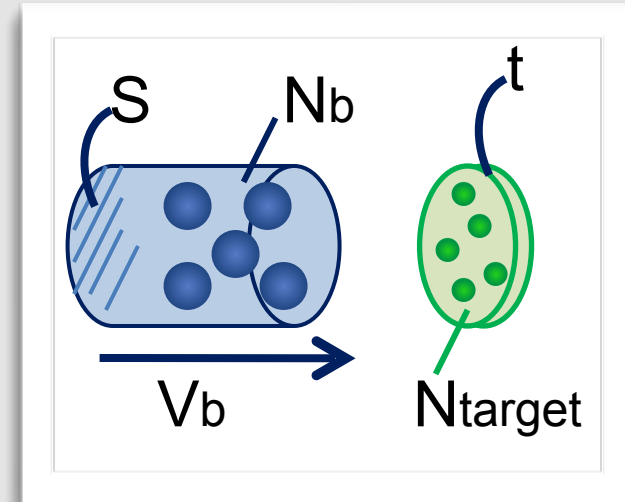
$$\propto \text{Flux } J : \text{beam rate per unit area}$$

$$J = n_b \cdot V_b$$

( $n_b$  : number density of beam particles)

( $V_b$  : beam velocity)

Beam intensity  $I = J S$  ( $S$ : beam area)



$$W = \sigma \cdot N_{\text{target}} \cdot J$$

$$= \sigma \cdot N_{\text{target}} \cdot I/S$$

$$= \sigma \cdot (n_t \cdot V) \cdot I/S$$

$$= \sigma \cdot (n_t \cdot t) \cdot I$$

$$= \sigma \cdot \rho \cdot (N_A/M_A) \cdot t \cdot I$$

( $n_t$ : number of target particles/unit volume)

( $t$ : thickness of target)

( $\rho$ : target density)

( $N_A$ : Avogadro's constant,  $M_A$ : mass)

# The Golden Rule

Fermi's Golden Rule states:

The transition rate for a system depends on two fundamental quantities

- 1) The “Matrix Element” or “Transition Amplitude”
- 2) The final state phase space

$$W = \frac{2\pi}{\hbar} |\mathcal{M}_{if}|^2 \rho(E)$$

# The Golden Rule

Fermi's Golden Rule states:

The transition rate for a system depends on two fundamental quantities

- 1) The “Matrix Element” or “Transition Amplitude”
- 2) The final state phase space

$$W = \frac{2\pi}{\hbar} |\mathcal{M}_{if}|^2 \rho(E)$$



**The transition rate  
(decays or  
interactions)**

# The Golden Rule

Fermi's Golden Rule states:

The transition rate for a system depends on two fundamental quantities

- 1) The “Matrix Element” or “Transition Amplitude”
- 2) The final state phase space

$$W = \frac{2\pi}{\hbar} |\mathcal{M}_{if}|^2 \rho(E)$$

**The transition rate  
(decays or  
interactions)**

**The Matrix Element.  
We may not actually  
know this.**

# The Golden Rule

Fermi's Golden Rule states:

The transition rate for a system depends on two fundamental quantities

- 1) The “Matrix Element” or “Transition Amplitude”
- 2) The final state phase space

$$W = \frac{2\pi}{\hbar} |\mathcal{M}_{if}|^2 \rho(E)$$


**The transition rate  
(decays or  
interactions)**

**The Matrix Element.  
We may not actually  
know this.**

**Phase Space or  
Density of States**

# Density of States

## Density of states\*

Also known as the available phase space or just phase space

Describes how many equivalent ways a final state can be configured

1) Assume we have a particle with quantized momentum confined to volume  $V$

$$p = \hbar k = h/\lambda$$



Smallest element of phase space in any coordinate is  $\mathbf{h}$ !

\*Some people like to use the particle-in-a-box example. It's fine, but I find it confusing for our purposes here.



# Density of States

## Density of states\*

Also known as the available phase space or just phase space

Describes how many equivalent ways a final state can be configured

1) Assume we have a particle with quantized momentum confined to volume  $V$

$$p = \hbar k = h/\lambda$$



Smallest element of phase space in any coordinate is  $\hbar$ !

2) The number of equivalent states,  $N_i$ , can be calculated by dividing the total phase space volume by the elemental phase space volume:

$$N_i = \frac{1}{(2\pi\hbar)^3} \int dx dy dz dp_x dp_y dp_z = \frac{V}{(2\pi\hbar)^3} \int d^3p$$

\*Some people like to use the particle-in-a-box example. It's fine, but I find it confusing for our purposes here.

# Density of States

## Density of states

Also known as the available phase space or just phase space

Describes how many equivalent ways a final state can be configured

3) Simplify life by calculating over a unit volume ( $V=1$  of your favorite unit). Also consider the option to have  $n$  particles in the final state:

$$N_n = \frac{1}{(2\pi\hbar)^{3n}} \int \prod_{i=1}^n d^3 p_i$$

# Density of States

## Density of states

Also known as the available phase space or just phase space

Describes how many equivalent ways a final state can be configured

3) Simplify life by calculating over a unit volume ( $V=1$  of your favorite unit). Also consider the option to have  $n$  particles in the final state:

$$N_n = \frac{1}{(2\pi\hbar)^{3n}} \int \prod_{i=1}^n d^3 p_i$$

4) Impose conservation of momentum in the final state, which means there is one less degree of freedom in the distribution of momenta:

$$\begin{aligned} N_n &= \frac{1}{(2\pi\hbar)^{3n}} \int \prod_{i=1}^n d^3 p_i \delta \left[ p_n - \left( p_0 - \sum_{i=1}^{n-1} p_i \right) \right] \\ &= \frac{1}{(2\pi\hbar)^{3n}} \int \prod_{i=1}^{n-1} d^3 p_i \end{aligned}$$

# Dirac Delta Function

I hope this is review. If not, go study up on this!

The Dirac delta function is simply the derivative of the Heaviside step function

$$\theta(x) = \begin{cases} 0, & (x < 0) \\ 1, & (x > 0) \end{cases} \longrightarrow \delta(x) = \begin{cases} 0, & (x \neq 0) \\ \infty, & (x = 0) \end{cases}$$

# Dirac Delta Function

I hope this is review. If not, go study up on this!

The Dirac delta function is simply the derivative of the Heaviside step function

$$\theta(x) = \begin{cases} 0, & (x < 0) \\ 1, & (x > 0) \end{cases} \longrightarrow \delta(x) = \begin{cases} 0, & (x \neq 0) \\ \infty, & (x = 0) \end{cases}$$

The integral of the delta function has a valuable property

The delta function selects out the zeros of its argument

$$\int_{-\infty}^{+\infty} \delta(x) = 1 \longrightarrow \int_{-\infty}^{+\infty} f(x) \delta(x - a) = f(a)$$

# Density of States

## Density of states

Also known as the available phase space or just phase space

Describes how many equivalent ways a final state can be configured

3) Simplify life by calculating over a unit volume ( $V=1$  of your favorite unit). Also consider the option to have  $n$  particles in the final state:

$$N_n = \frac{1}{(2\pi\hbar)^{3n}} \int \prod_{i=1}^n d^3 p_i$$

4) Impose conservation of momentum in the final state, which means there is one less degree of freedom in the distribution of momenta:

$$\begin{aligned} N_n &= \frac{1}{(2\pi\hbar)^{3n}} \int \prod_{i=1}^n d^3 p_i \delta \left[ p_n - \left( p_0 - \sum_{i=1}^{n-1} p_i \right) \right] \\ &= \frac{1}{(2\pi\hbar)^{3n}} \int \prod_{i=1}^{n-1} d^3 p_i \end{aligned}$$

# Phase Space Problem!

---

What we did is just fine, but it's not Lorentz Invariant

But we can fix that by taking into account the what we've learned

Express our density of states as a 4-momentum

# Phase Space Problem!

What we did is just fine, but it's not Lorentz Invariant

But we can fix that by taking into account the what we've learned

Express our density of states as a 4-momentum

1) Simplify life by calculating over a unit volume ( $V=1$  of your favorite unit). Also consider the option to have  $n$  particles in the final state:

$$N_n = \frac{1}{(2\pi\hbar)^{3n}} \int \prod_{i=1}^n d^3 p_i \longrightarrow N_n = \frac{1}{(2\pi\hbar)^{4n}} \int \prod_{i=1}^n d^4 p_i$$



# Phase Space Problem!

What we did is just fine, but it's not Lorentz Invariant

But we can fix that by taking into account the what we've learned

Express our density of states as a 4-momentum

1) Simplify life by calculating over a unit volume ( $V=1$  of your favorite unit). Also consider the option to have  $n$  particles in the final state:

$$N_n = \frac{1}{(2\pi\hbar)^{3n}} \int \prod_{i=1}^n d^3 p_i \longrightarrow N_n = \frac{1}{(2\pi\hbar)^{4n}} \int \prod_{i=1}^n d^4 p_i$$

2) Impose Lorentz invariance by fixing the Lorentz invariant inner product for each final state particle:

$$N_n = \frac{1}{(2\pi\hbar)^{4n}} \int \prod_{i=1}^n d^4 p_i \delta \left[ (p^\mu p_\mu)_i - m_i^2 c^2 \right]$$

# Lorentz Invariant Phase Space

---

While we're at it, we may as well introduce “sanity checks”

- 1) Energy & momentum are conserved
- 2) There can be no negative energy states

# Lorentz Invariant Phase Space

While we're at it, we may as well introduce "sanity checks"

- 1) Energy & momentum are conserved
- 2) There can be no negative energy states

3) Introduce a delta function for 4-momenta conservation (energy & momentum):

$$N_n = \frac{1}{(2\pi\hbar)^{4n}} \int \delta^4(p_i - \sum p_f) \prod_{j=1}^n d^4p_j \delta[(p^\mu p_\mu)_j - m_j^2 c^2]$$

# Lorentz Invariant Phase Space

While we're at it, we may as well introduce "sanity checks"

- 1) Energy & momentum are conserved
- 2) There can be no negative energy states

3) Introduce a delta function for 4-momenta conservation (energy & momentum):

$$N_n = \frac{1}{(2\pi\hbar)^{4n}} \int \delta^4(p_i - \sum p_f) \prod_{j=1}^n d^4p_j \delta [(p^\mu p_\mu)_j - m_j^2 c^2]$$

4) Introduce a Heaviside function to remove negative final state energies:

$$N_n = \frac{1}{(2\pi\hbar)^{4n}} \int \delta^4(p_i - \sum p_f) \prod_{j=1}^n d^4p_j \delta [(p^\mu p_\mu)_j - m_j^2 c^2] \theta(E_j)$$

# Lorentz Invariant Phase Space

---

We're almost there!! One more thing to do...

Recall, our integrals were not truly over continuous momentum space.

Momentum is quantized, so we're actually integrating integer  $k$  space!

# Lorentz Invariant Phase Space

We're almost there!! One more thing to do...

Recall, our integrals were not truly over continuous momentum space.

Momentum is quantized, so we're actually integrating integer  $k$  space!

5) Fourier transform from  $p$ -space back to  $k$ -space (or vice-versa):

$$\phi(p) = \frac{1}{2\pi} \int \psi(k) e^{ipk} dk$$



Every delta function gets a  $2\pi$  in the integral.

# Lorentz Invariant Phase Space

We're almost there!! One more thing to do...

Recall, our integrals were not truly over continuous momentum space.

Momentum is quantized, so we're actually integrating integer  $k$  space!

5) Fourier transform from  $p$ -space back to  $k$ -space (or vice-versa):

$$\phi(p) = \frac{1}{2\pi} \int \psi(k) e^{ipk} dk$$



Every delta function gets a  $2\pi$  in the integral.

6) We done! We call the energy derivative the “differential Lorentz Invariant phase Space” or just dLIPS.

$$\rho(E) = \frac{\partial N_n}{\partial E} = \frac{(2\pi)^{n+4}}{(2\pi\hbar)^{4n}} \delta^4(p_i - \sum p_f) \prod_{j=1}^n d^4 p_j \delta [(p^\mu p_\mu)_j - m_j^2 c^2] \theta(E_j)$$

# The Golden Rule

Fermi's Golden Rule states:

The transition rate for a system depends on two fundamental quantities

- 1) The “Matrix Element” or “Transition Amplitude”
- 2) The final state phase space

$$W = \frac{2\pi}{\hbar} |\mathcal{M}_{if}|^2 \rho(E)$$


**The transition rate  
(decays or  
interactions)**

**The Matrix Element.  
We may not actually  
know this.**

**Phase Space or  
Density of States**



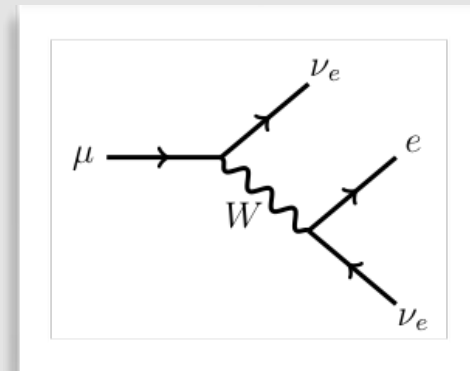
# Golden Rule for Decays

We can now build the form of the decay rate

Assume we have a particle at rest decaying to  $n$  particles:  $1 \rightarrow 2, 3, 4 \dots n$

Here we assume the form for the decay rate and insert elements as needed

$$\Gamma = \frac{S}{2\hbar m_1} \int |\mathcal{M}|^2 \rho(E)$$



The mysterious  $S$  factor takes into account final state particle interchange

If the final state has  $N$  identical particles, we include a factor of  $N!$

Thus  $S = 1/N!$

# Golden Rule for Decays

We can now build the form of the decay rate

Assume we have a particle at rest decaying to  $n$  particles:  $1 \rightarrow 2, 3, 4 \dots n$

Here we assume the form for the decay rate and insert elements as needed

$$\Gamma = \frac{S}{2\hbar m_1} \int |\mathcal{M}|^2 \rho(E)$$

$$\Gamma = \frac{S}{2\hbar m_1} \int |\mathcal{M}|^2 (2\pi)^4 \delta^4(p_1 - p_2 - p_3 \cdots - p_n) \\ \times \prod_{j=2}^n 2\pi \delta(p_j^2 - m_j^2 c^2) \theta(E_j) \frac{d^4 p_i}{(2\pi)^4}$$

Following the book or in-class derivation for 2-particle decays ( $1 \rightarrow 2, 3$ ):

$$\Gamma = \frac{S |p|}{8\pi \hbar m_1^2 c} |\mathcal{M}|^2$$

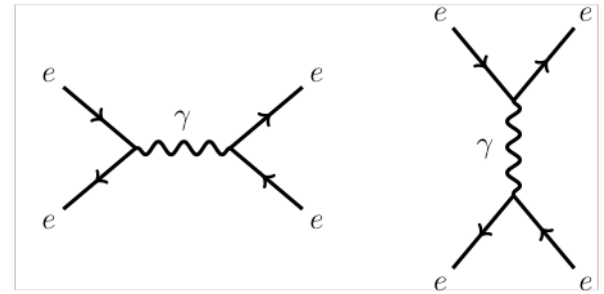
# Golden Rule for Scattering

We can also build the form for particle scattering

Suppose we have two particle scattering:  $1, 2 \rightarrow 3, 4, \dots, n$

Don't worry about the inertial frame just yet.

$$\sigma = \frac{S \hbar^2}{4\sqrt{(p_1 \cdot p_2)^2 - (m_1 m_2 c^2)^2}} \int |\mathcal{M}|^2 \rho(E)$$



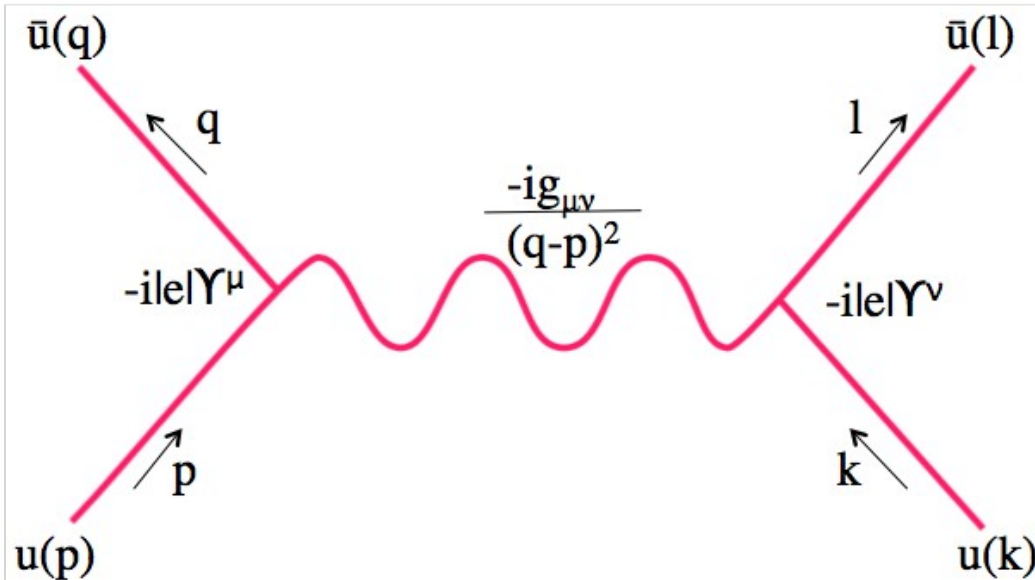
Following the book's derivation for 2-body scattering ( $1, 2 \rightarrow 3, 4$ ):

$$\frac{d\sigma}{d\Omega} = \left( \frac{\hbar c}{8\pi} \right)^2 \frac{S |\mathcal{M}|^2}{(E_1 + E_2)^2} \frac{|p_f|}{|p_i|}$$

# Feynman Rules

We've been completely ignoring the dynamical part of the equation  
All of the specific interaction dynamics are bound up in the **Matrix Element**  
We need to now build up the rules for calculating it!

The Feynman diagram will be our guide in most cases  
We will assign kinematical features to each part of the diagram,  
depending on what type of particle it represents.

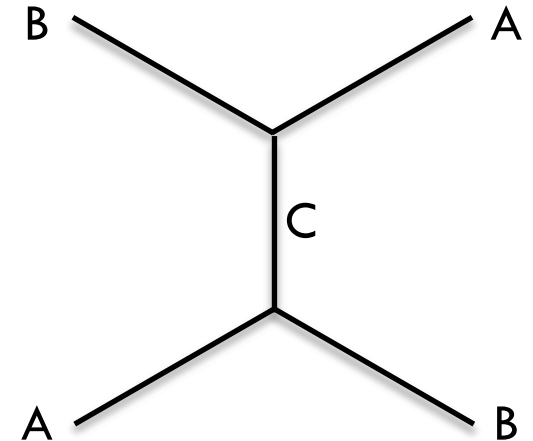
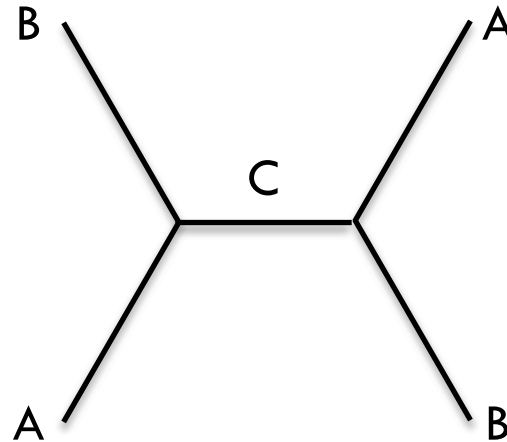
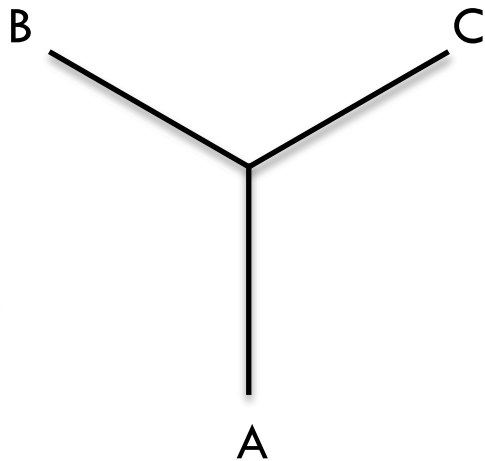


The Feynman diagram will give rise to the matrix element.

# ABC Theory

To exercise the Feynman rules, we'll use a toy theory: ABC theory

- There are 3 spinless particles: A, B, C
- Each particle is its own antiparticle (No arrows!)
- Only 1 vertex exists and includes all three particles (ABC). Eg, (AAA) is forbidden.
- If  $m_A > m_B + m_C$ , then A can decay to B and C.

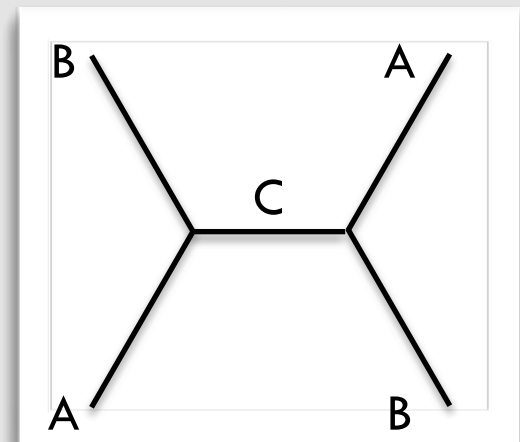


# Feynman Rules for ABC Theory

The Feynman rules provide the recipe for constructing an amplitude (Matrix Element  $\mathcal{M}$ ) from a Feynman diagram.

## Rule 1:

Draw the Feynman diagram with the minimum number of vertices. There may be more than 1.



# Feynman Rules for ABC Theory

The Feynman rules provide the recipe for constructing an amplitude (Matrix Element  $\mathcal{M}$ ) from a Feynman diagram.

## Rule 1:

Draw the Feynman diagram with the minimum number of vertices. There may be more than 1.

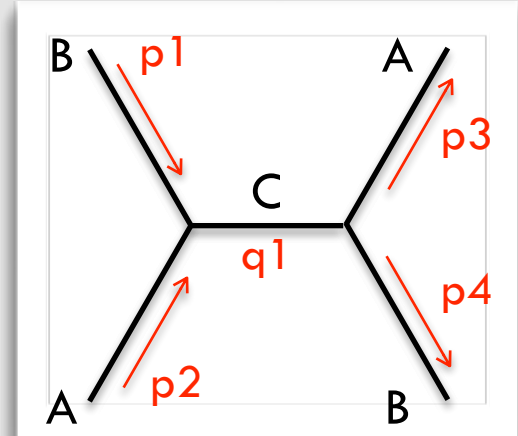
## Rule 2:

Label the four-momentum of each line (with arrows), enforcing four-momentum conservation at each vertex.

$p_1, p_2, \dots$  external momenta: need arrows

$q_1, q_2, \dots$  internal momenta: arrow is arbitrary

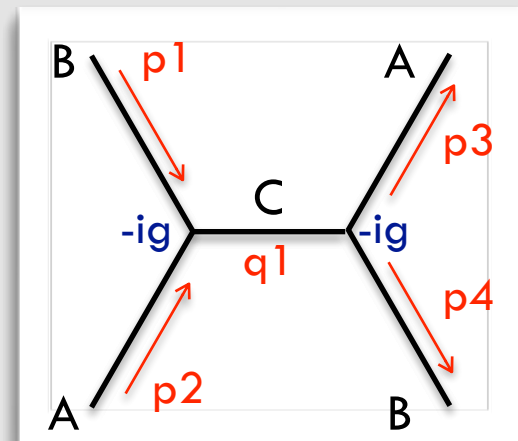
We'll keep track of arrows into/out of vertices.



# Feynman Rules for ABC Theory

## Rule 3:

Each vertex contributes a factor of  $(-ig)$ , where  $g$  is referred to as the coupling constant. It specifies the strength of the ABC interaction.





# Feynman Rules for ABC Theory

## Rule 3:

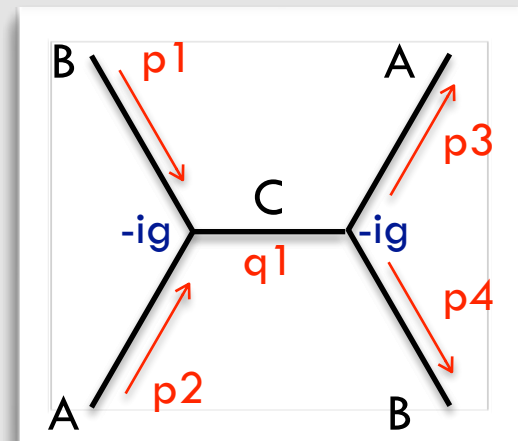
Each vertex contributes a factor of  $(-ig)$ , where  $g$  is referred to as the coupling constant. It specifies the strength of the ABC interaction.

## Rule 4:

Each internal line, or propagator, with mass  $m$  and four-momentum  $q$  gets a factor of:

$$\frac{i}{q^2 - m^2 c^2}$$

**Note:**  $q^2$  doesn't have to equal  $m^2$ . These are virtual particles!



# Feynman Rules for ABC Theory

## Rule 3:

Each vertex contributes a factor of  $(-ig)$ , where  $g$  is referred to as the coupling constant. It specifies the strength of the ABC interaction.

## Rule 4:

Each internal line, or propagator, with mass  $m$  and four-momentum  $q$  gets a factor of:

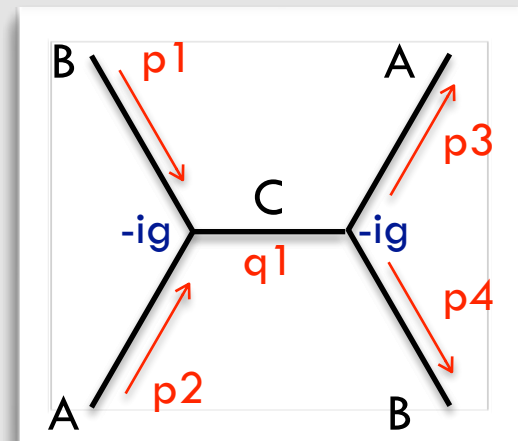
$$\frac{i}{q^2 - m^2 c^2}$$

**Note:**  $q^2$  doesn't have to equal  $m^2$ . These are virtual particles!

## Rule 5:

Each vertex contributes a delta function to conserve energy and momentum. The  $k_i$  are the momenta coming into the vertex:

$$(2\pi)^4 \delta^4(k_1 + k_2 + k_3)$$



# Feynman Rules for ABC Theory

## Rule 6:

Build up the proto-Matrix Element from the previous factors & add an  $i$ :

$$\mathcal{M} = i (\text{vertex factors}) (\text{propagator factors}) (\text{momentum conservation})$$

# Feynman Rules for ABC Theory

## Rule 6:

Build up the proto-Matrix Element from the previous factors & add an  $i$ :

$$\mathcal{M} = i (\text{vertex factors}) (\text{propagator factors}) (\text{momentum conservation})$$

## Rule 7:

Integrate over the internal momenta:

$$\frac{1}{(2\pi)^4} d^4 q_i$$

This connects initial/final state momenta via the delta functions.

# Feynman Rules for ABC Theory

## Rule 6:

Build up the proto-Matrix Element from the previous factors & add an  $i$ :

$$\mathcal{M} = i (\text{vertex factors}) (\text{propagator factors}) (\text{momentum conservation})$$

## Rule 7:

Integrate over the internal momenta:

$$\frac{1}{(2\pi)^4} d^4 q_i$$

This connects initial/final state momenta via the delta functions.

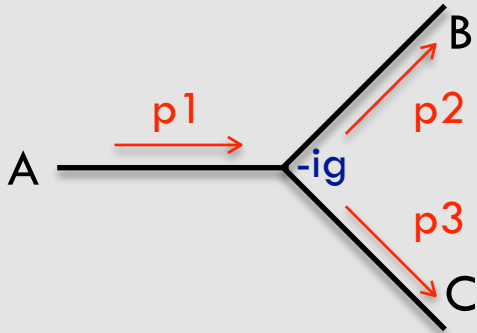
## Rule 8:

Drop the extra delta function. We do this because the Matrix Element gets squared, and that would double-count the delta function. **Don't** worry, it gets put back in the Golden Rule Equation!

**The result of step 8 is the matrix element!**

# Decay of the A

We are now in a position to calculate something! The easiest is the lifetime and/or decay rate of  $A \rightarrow BC$ .



**We've already done the kinematics!**

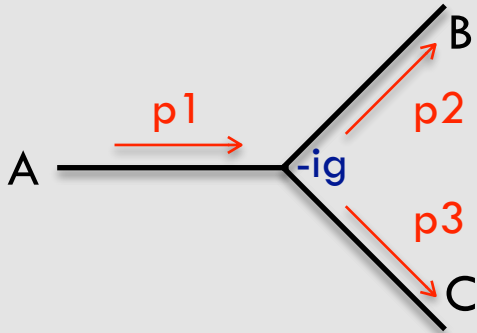
$$\Gamma = \frac{S |p|}{8\pi m_A^2} |\mathcal{M}|^2$$

$\hbar = c = 1$

S Factor: The final state particles are not the same, so  $S=1$

# Decay of the A

We are now in a position to calculate something! The easiest is the lifetime and/or decay rate of  $A \rightarrow BC$ .



**We've already done the kinematics!**

$$\Gamma = \frac{S |p|}{8\pi m_A^2} |\mathcal{M}|^2$$

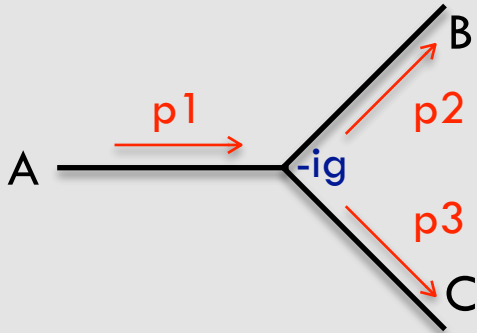
$\hbar = c = 1$

S Factor: The final state particles are not the same, so  $S=1$

Rule 3: Only one vertex, so we get a factor of  $(-ig)$

# Decay of the A

We are now in a position to calculate something! The easiest is the lifetime and/or decay rate of  $A \rightarrow BC$ .



**We've already done the kinematics!**

$$\Gamma = \frac{S |p|}{8\pi m_A^2} |\mathcal{M}|^2$$

$\hbar = c = 1$

S Factor: The final state particles are not the same, so  $S=1$

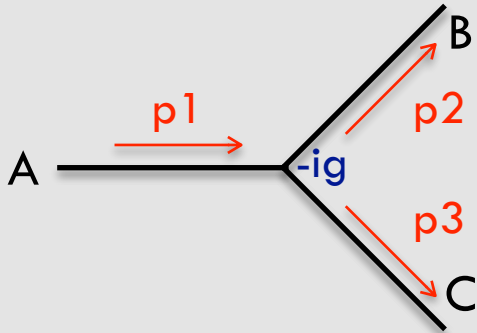
Rule 3: Only one vertex, so we get a factor of  $(-ig)$

Rule 4: No propagators to worry about



# Decay of the A

We are now in a position to calculate something! The easiest is the lifetime and/or decay rate of  $A \rightarrow BC$ .



**We've already done the kinematics!**

$$\Gamma = \frac{S |p|}{8\pi m_A^2} |\mathcal{M}|^2$$

$\hbar = c = 1$

S Factor: The final state particles are not the same, so  $S=1$

Rule 3: Only one vertex, so we get a factor of  $(-ig)$

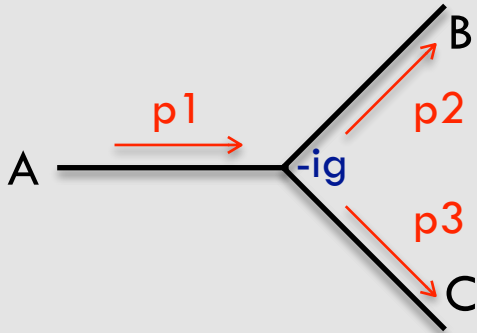
Rule 4: No propagators to worry about

Rule 5: We pick up a delta function over the vertex in/out momenta

$$(2\pi)^4 \delta^4(p_A - p_B - p_C)$$

# Decay of the A

We are now in a position to calculate something! The easiest is the lifetime and/or decay rate of  $A \rightarrow BC$ .



**We've already done the kinematics!**

$$\Gamma = \frac{S |p|}{8\pi m_A^2} |\mathcal{M}|^2$$

$\hbar = c = 1$

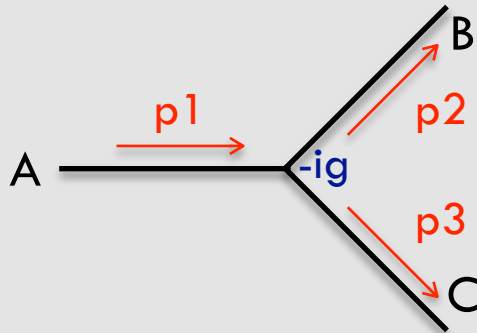
Rule 6: Our matrix element is thus:  $\mathcal{M} = i(-ig) = g$

Rule 7: No internal lines to integrate over

Rule 8: Now we drop the delta function we got from Rule 5. There was no integral, so it was unnecessary.

# Decay of the A

We are now in a position to calculate something! The easiest is the lifetime and/or decay rate of  $A \rightarrow BC$ .



**We've already done the kinematics!**

$$\Gamma = \frac{S |p|}{8\pi m_A^2} |\mathcal{M}|^2$$

$\hbar = c = 1$

Rule 6: Our matrix element is thus:  $\mathcal{M} = i(-ig) = g$

Rule 7: No internal lines to integrate over

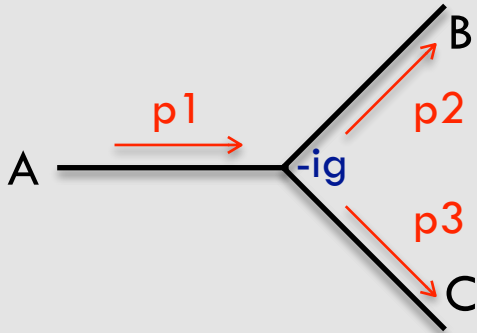
Rule 8: Now we drop the delta function we got from Rule 5. There was no integral, so it was unnecessary.

Momentum:  $|\mathbf{p}_B| = |\mathbf{p}_C| = p$  is uniquely determined by  $M_A$ ,  $M_B$ , &  $M_C$ .

**See problem 3.19:**  $|p| = \frac{1}{2m_A} \sqrt{m_A^4 + m_B^4 + m_C^4 - 2m_A^2 m_B^2 - 2m_A^2 m_C^2 - 2m_B^2 m_C^2}$

# Decay of the A

We are now in a position to calculate something! The easiest is the lifetime and/or decay rate of  $A \rightarrow BC$ .



**We've already done the kinematics!**

$$\Gamma = \frac{S |p|}{8\pi m_A^2} |\mathcal{M}|^2$$

$\hbar = c = 1$

$$\Gamma(A \rightarrow BC) = \frac{g^2 |p|}{8\pi m_A^2} \quad \tau(A \rightarrow BC) = \frac{8\pi m_A^2}{g^2 |p|}$$

See problem 3.19:  $|p| = \frac{1}{2m_A} \sqrt{m_A^4 + m_B^4 + m_C^4 - 2m_A^2 m_B^2 - 2m_A^2 m_C^2 - 2m_B^2 m_C^2}$

# Example: $\pi^0 \rightarrow \gamma\gamma$ Decays

---

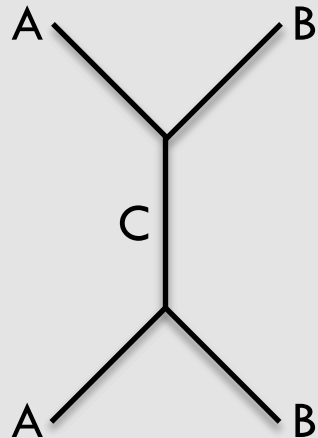
Example done in class.

# $A+A \rightarrow B+B$ Scattering

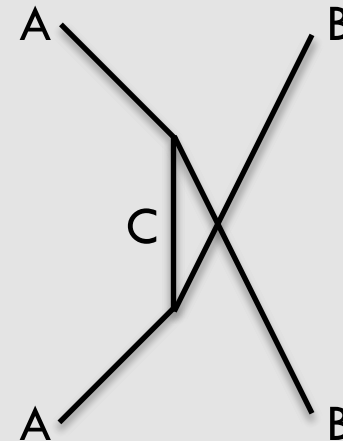
Now we can try something more complicated.

The scattering of  $A+A$  into  $B+B$  (or, equivalently  $A+A$ ) is highly relevant  
In ABC theory, this occurs via the exchange of particle C.

**We have the t-channel diagram**



**But also the u-channel diagram!**



Why two diagrams?

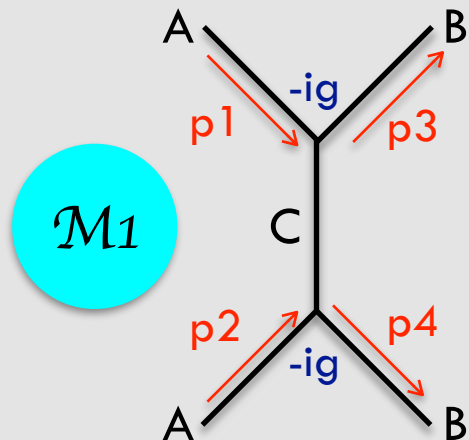
- The u-channel diagram is necessary because we cannot tell which outgoing particles connect to which vertex. So we have to include both possibilities in our matrix element and, thus, in the integrations.
- But don't worry about double-counting, as our S-factor will now be  $S=1/2!=1/2$ . By doing so, we effectively average over the two diagrams

# A+A→B+B Scattering

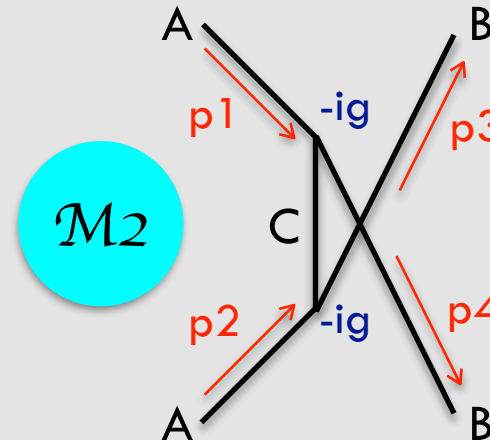
Now we can try something more complicated.

The scattering of A+A into B+B (or, equivalently A+A) is highly relevant  
In ABC theory, this occurs via the exchange of particle C.

**We have the t-channel diagram**



**But also the u-channel diagram!**



**We've already done the kinematics, just fix  $S=1/2$ :**

$$\frac{d\sigma}{d\Omega} = \left( \frac{1}{8\pi} \right)^2 \frac{|\mathcal{M}|^2}{2(E_1 + E_2)^2} \frac{|p_f|}{|p_i|} \quad \hbar = c = 1$$

# A+A→B+B Scattering

First let's calculate the t-channel matrix element.

Rule 3: Two vertices, so we get two factors of  $(-ig)$

Rule 4: One propagator gets us a factor of:

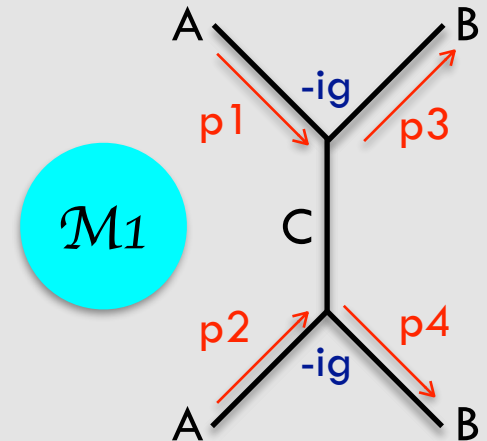
$$\frac{i}{q^2 - m_C^2}$$

Rule 5: We pick up two delta functions for the momenta in/out of the vertices

$$(2\pi)^4 \delta^4(p_1 - q - p_3)$$

$$(2\pi)^4 \delta^4(p_2 + q - p_4)$$

Rule 7: We have to integrate over the phase space of the C propagator (we'll do Rule 6 next)



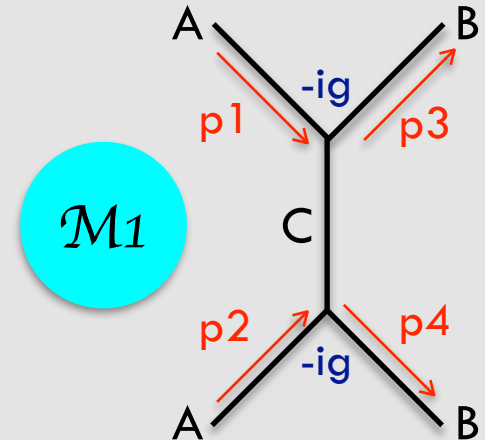


# A+A→B+B Scattering

First let's calculate the t-channel matrix element.

**Rule 3: Vertex Factors**

**Rule 5: Vertex E/p conservation**



**Rule 6:**

$$\mathcal{M}_1 = i \int \underline{(-ig)^2} \left( \frac{i}{\underline{q^2 - m_C^2}} \right) \underline{(2\pi)^4 \delta^4(p_1 - q - p_3)} \underline{(2\pi)^4 \delta^4(p_2 + q - p_4)} \underline{\frac{d^4 q}{(2\pi)^4}}$$

**Rule 4: Propagator Factor**

**Rule 7: Vertex E/p conservation**

# A+A→B+B Scattering

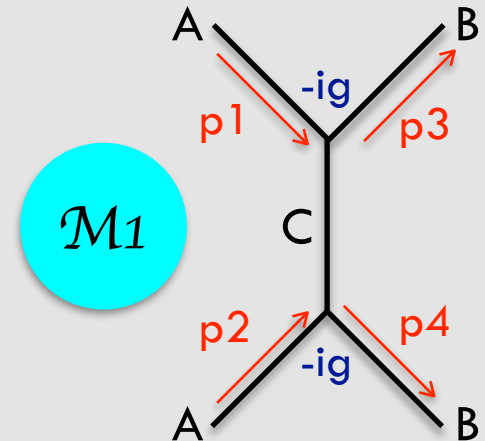
First let's calculate the t-channel matrix element.

The propagator integral sends either:

$$q \rightarrow p_4 - p_2$$

or:

$$q \rightarrow p_3 - p_1$$



Rule 6:

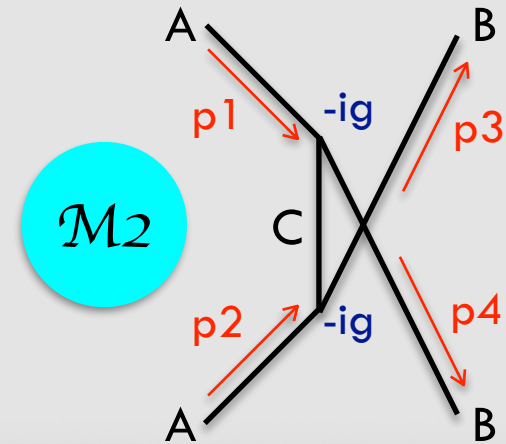
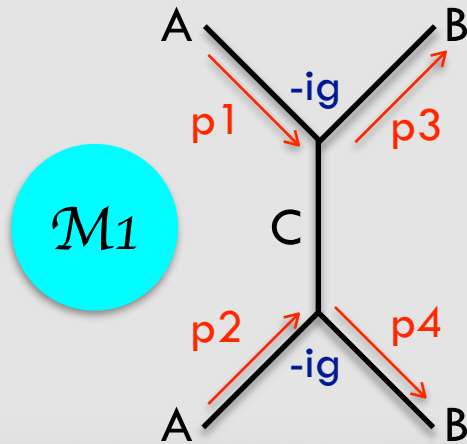
$$\mathcal{M}_1 = i \int (-ig)^2 \left( \frac{i}{q^2 - m_C^2} \right) (2\pi)^4 \delta^4(p_1 - q - p_3) (2\pi)^4 \delta^4(p_2 + q - p_4) \frac{d^4 q}{(2\pi)^4}$$

$$\mathcal{M}_1 = \frac{g^2}{(p_4 - p_2)^2 - m_C^2} (2\pi)^4 \delta^4(p_1 + p_2 - p_3 - p_4)$$

# A+A→B+B Scattering

Now we can consider the u-channel matrix element.

No work required, just swap out the relevant momenta



$$\mathcal{M}_1 = \frac{g^2}{(p_4 - p_2)^2 - m_C^2}$$

$$\mathcal{M}_2 = \frac{g^2}{(p_3 - p_2)^2 - m_C^2}$$

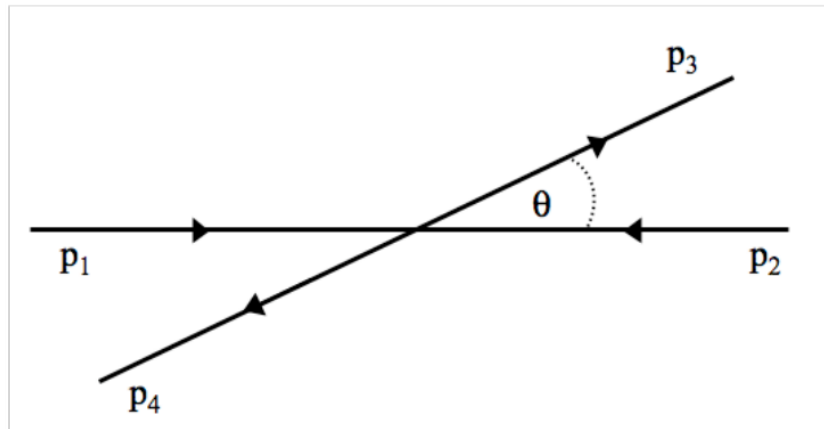
$$\mathcal{M}_{\text{tot}} = \mathcal{M}_1 + \mathcal{M}_2$$

# A+A→B+B Scattering

The A+A→B+B scattering matrix element is Lorentz Invariant  
But the evaluated quantities may depend on the inertial frame  
Let's consider the CoM frame and let  $M_A=M_B=m$  and  $M_C=0$

$$\mathcal{M}_1 = \frac{g^2}{(p_4 - p_2)^2 - m_C^2}$$

$$\mathcal{M}_2 = \frac{g^2}{(p_3 - p_2)^2 - m_C^2}$$



$$t = (p_4 - p_2)^2 = p_4^2 + p_2^2 - 2p_2 \cdot p_4 = -2p^2(1 - \cos \theta)$$

$$u = (p_3 - p_2)^2 = p_3^2 + p_2^2 - 2p_3 \cdot p_2 = -2p^2(1 + \cos \theta)$$

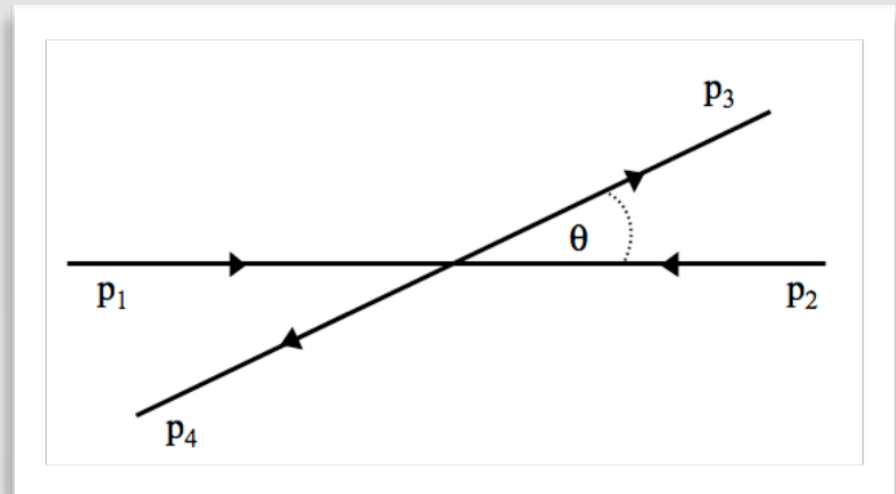
# A+A→B+B Scattering

The  $A+A \rightarrow B+B$  scattering matrix element is Lorentz Invariant  
But the evaluated quantities may depend on the inertial frame  
Let's consider the CoM frame and let  $M_A=M_B=m$  and  $M_C=0$

$$\mathcal{M}_1 = \frac{g^2}{(p_4 - p_2)^2 - m_C^2}$$

$$\mathcal{M}_2 = \frac{g^2}{(p_3 - p_2)^2 - m_C^2}$$

$$\mathcal{M} = -\frac{g^2}{p^2 \sin^2 \theta}$$



$$t = (p_4 - p_2)^2 = p_4^2 + p_2^2 - 2p_2 \cdot p_4 = -2p^2(1 - \cos \theta)$$

$$u = (p_3 - p_2)^2 = p_3^2 + p_2^2 - 2p_3 \cdot p_2 = -2p^2(1 + \cos \theta)$$

# A+A→B+B Scattering

Finally, we can evaluate the differential cross section!

We worked out the 2-body kinematics earlier

We just calculated the matrix element

$$\frac{d\sigma}{d\Omega} = \left(\frac{1}{8\pi}\right)^2 \frac{|\mathcal{M}|^2}{2(E_1 + E_2)^2} \frac{|p_f|}{|p_i|} \quad \mathcal{M} = -\frac{g^2}{p^2 \sin^2 \theta}$$

From the CoM condition we have:

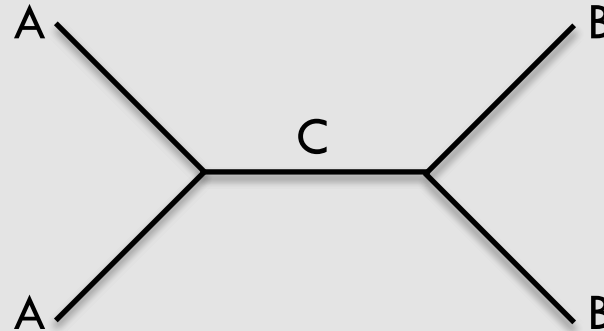
$$E_1 = E_2$$

$$|p_f| = |p_i|$$

$$\frac{d\sigma}{d\Omega} = \frac{1}{2} \left( \frac{g^2}{16\pi E |p|^2 \sin^2 \theta} \right)^2$$

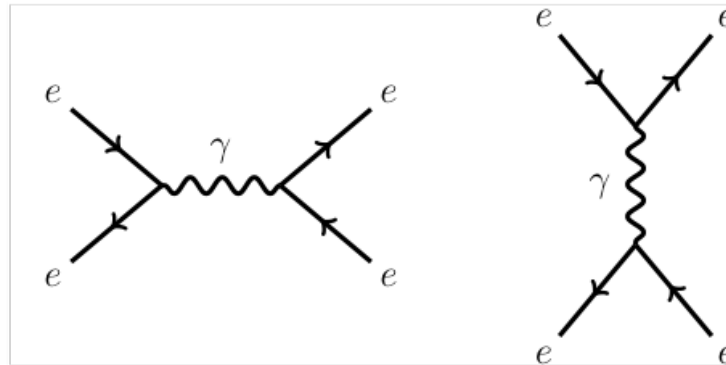
# $A+A \rightarrow B+B$ Scattering

A great question! Why didn't we consider this diagram?  
Be careful to not confuse or read too deep into ABC theory!



Answer:

There are no AAC or CBB vertices in ABC theory.  
But this diagram does exist in QED, for example.



# Another question

Earlier we said cross sections sum:

$$\sigma_{\text{tot}} = \sum \sigma_i$$

And we just said matrix elements sum:

$$\mathcal{M}_{\text{tot}} = \mathcal{M}_1 + \mathcal{M}_2$$

The first statement is clearly only true if there is no interference amongst the matrix elements!

$$|\mathcal{M}|^2 = |\mathcal{M}_1|^2 + |\mathcal{M}_2|^2 + 2|\mathcal{M}_1 \cdot \mathcal{M}_2|^2$$

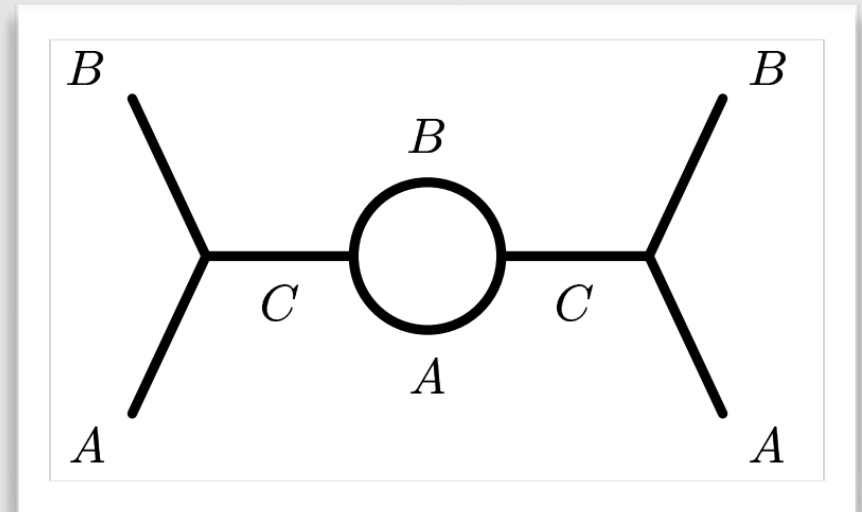
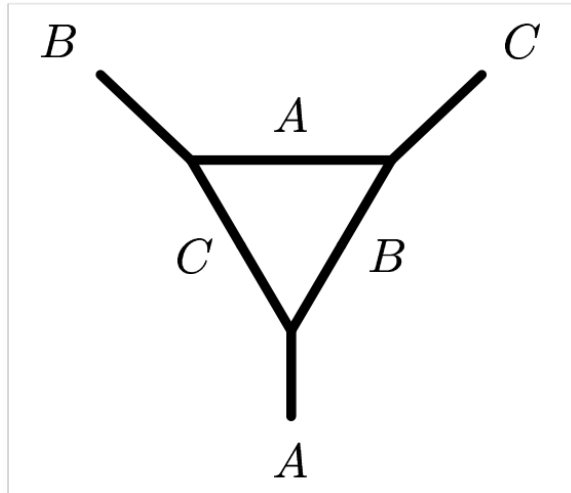


# Higher Order Diagrams

These interference effects frequently (but not always) arise due to higher order diagrams

ABC theory limits the number of such diagrams, so continue to be wary

We're not going to dig deep into this right now, but be aware they exist



# Recap / Up Next

This time:

Feynman Calculus

Decays/Scattering

The Golden Rule

Feynman Rules

Next time:

Particle Accelerators

Cyclotrons

Synchrotrons

The Large Hadron Collider

