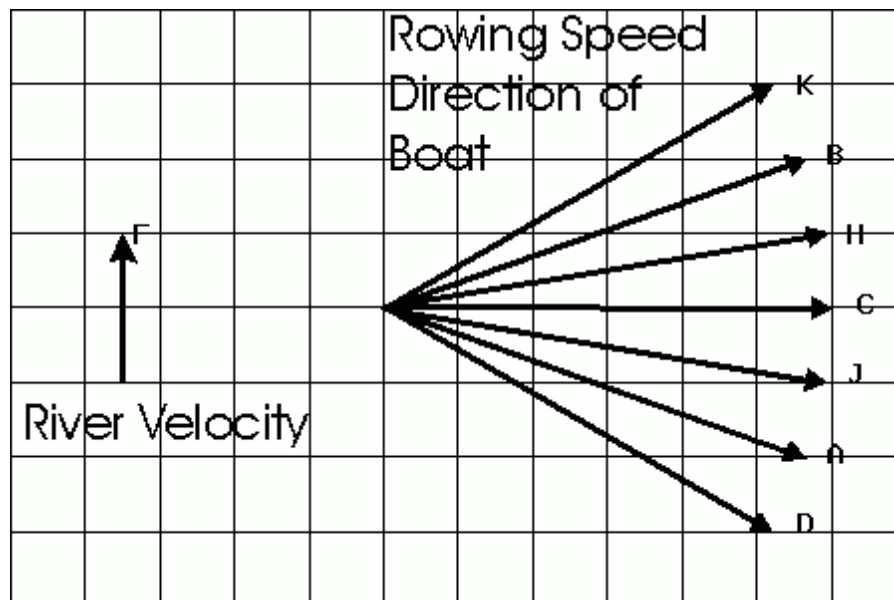


-
- LON-CAPA #4 and Mastering Physics Chapter 7 due next Tuesday
 - ◆ help room hours (Strosacker Help Room, 1248 BPS):
M: 5-8 PM
W: 5-8 PM
F: 2-6 PM
 - Register for Mastering Physics
 - ◆ >95% of you have
 - Register your iclicker on LON-CAPA
 - ◆ >98% of you have
 - ◆ I'll post the scores up to last Thursday's lecture on the web
 - First exam: Feb 6 in Life Sciences A133

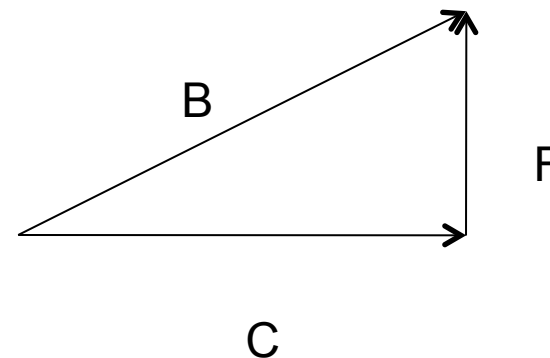
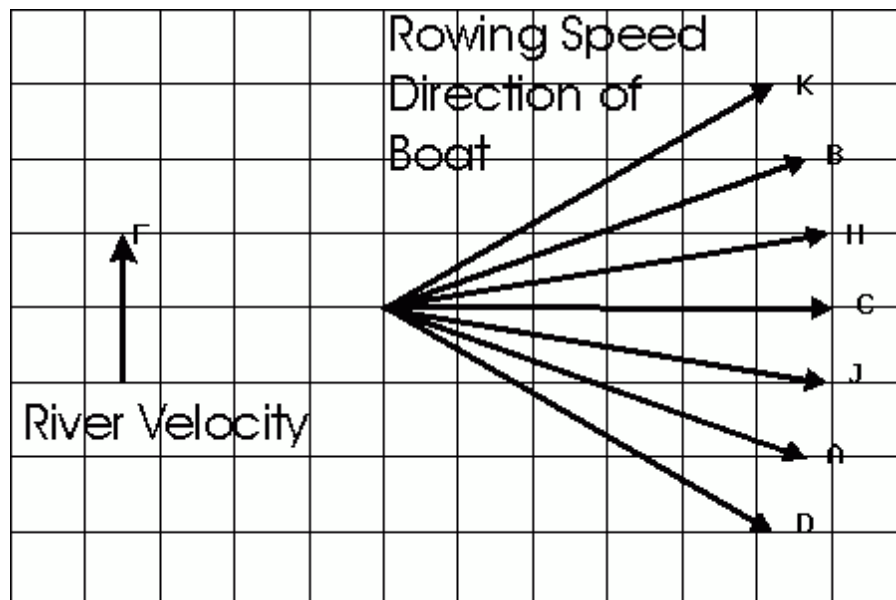
LON-CAPA problem

- A river is to be crossed by a boy using a row boat
 - He has the choice of rowing in directions A,B, C, D,...; in any direction, he rows at a constant speed with respect to the water
 - So this is a vector addition problem
- If he rows with velocity in the direction of C, where does he end up?



LON-CAPA problem

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- He has the choice of rowing in directions A,B, C, D,...; in any direction, he rows at a constant speed with respect to the water
- So this is a vector addition problem
- If he rows with velocity in the direction of C, where does he end up?
- Add two vectors



Universal law of gravitation

- Newton's law of gravity describes gravitational forces on the surface of the Earth, and off the surface of the Earth

$$F = G \frac{m_1 m_2}{d^2}$$

- A tremendous advance; the law of gravity, along with the 3 laws of motion, meant that the motion of planets, and motions on Earth, could be completely described within a mathematical framework
- Some philosophers started to talk about a 'clockwork universe'

- But how is Newton's law of gravity incomplete?

Universal law of gravitation

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- A tremendous advance; the law of gravity, along with the 3 laws of motion, meant that the motion of planets, and motions on Earth, could be completely described within a mathematical framework

- But how is Newton's law of gravity incomplete?
- Newton assumed (because he had no evidence one way or the other) that the force of gravity was transmitted instantaneously in a manner which could not be described
 - ◆ 'action-at-a-distance'
- It was left to Einstein (over 200 years later) to describe gravity as being due to the curvature of space
 - ◆ and that gravitational disturbances travelled at the speed of light

iclicker question

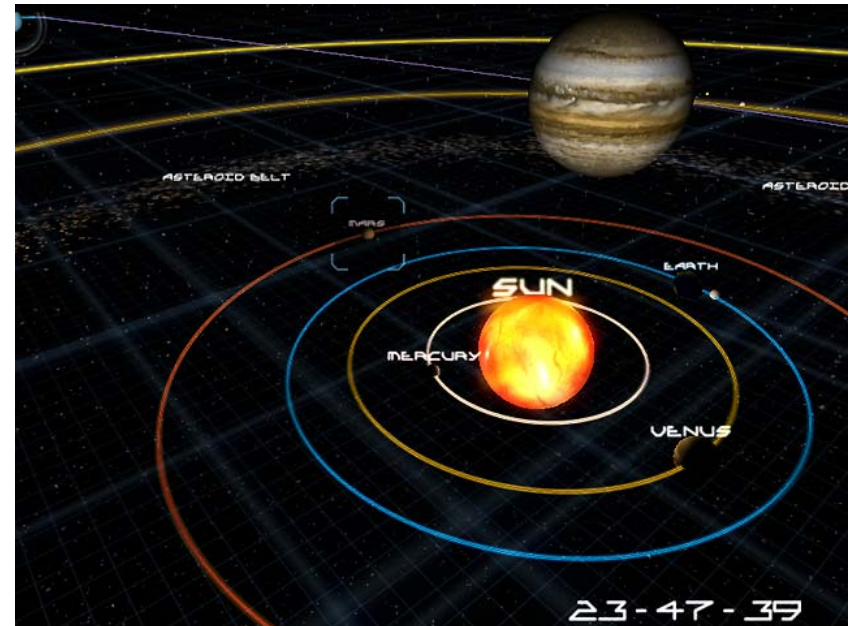
- The gravitational constant G in Newton's law of gravity
 - a) Produces the correct units of force in Newton's equation
 - b) Indicates the strength of gravity
 - c) Changes the proportion form of the law of gravity to an exact equation
 - d) All of these
 - e) None of these

iclicker question

- The gravitational constant G in Newton's law of gravity
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 - d) **All of these**
 - e) None of these

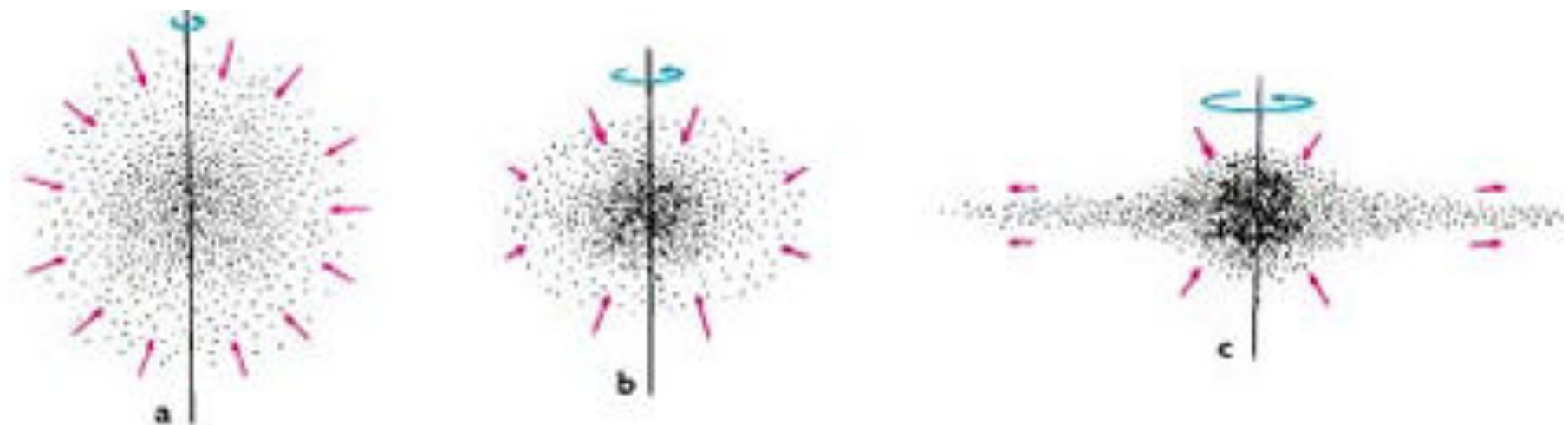
Gravity

- What's responsible for keeping the solar system together?
 - ◆ gravity
- What would happen if the sun were replaced by a black hole of the same mass?
 - ◆ nothing, except that it would get dark
 - ◆ the gravitational force of the sun already acts as if its originating from a point in the center of the sun
- What's responsible for keeping the Milky Way together?
 - ◆ gravity, except there's not enough visible matter
 - ◆ most of the universe appears to be composed of *dark matter*



Origin of the solar system

- Remember our sun was created after the universe was about 10 billion years old
- A slightly rotating ball of interstellar gas (enriched in heavy elements due to earlier supernovae) contracts due to gravitational attraction and speeds up to conserve angular momentum
- The increased momentum causes them to sweep in wider paths around the rotational axis, producing an overall disk shape
- The planets condense out of eddies in the cooling disk



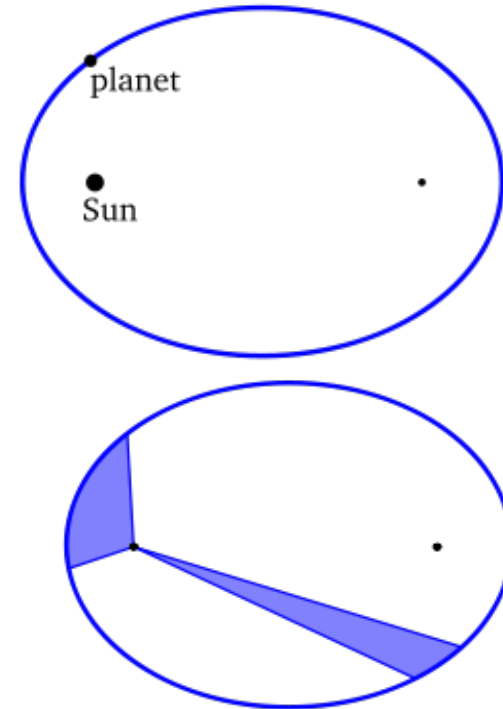
Johannes Kepler

- 1571-1630
- He took the detailed observations of Tycho Brahe on the motions of the planets and was able to formulate 3 laws that describe the motions of the planets



Kepler's 3 laws

- Every planet has an elliptical orbit with the Sun at one focus of the ellipse
- A line joining the planet and the Sun sweeps out equal areas in equal times
 - ◆ so the planet must move fastest when it's closest to the Sun
- The square of the period of the orbit of a planet is proportional to the cube of the radius (semi-major axis)

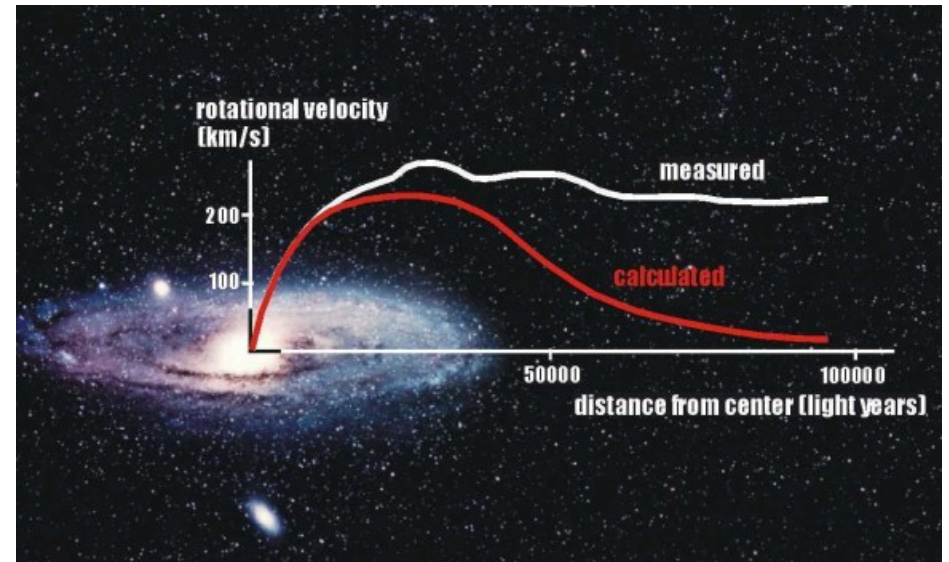


$$\left(\frac{P}{2\pi}\right)^2 = \frac{a^3}{GM_{sun}}$$

Empirical observations from Kepler; can be derived using Newton's law of gravitation

Rotation of the galaxy and dark matter

- Remember earlier I said that you would expect stars further from the center of the galaxy to rotate more slowly than ones at the center
- The fact that this does not happen is evidence of the presence of dark matter around our galaxy (10X as much dark matter as regular matter)



$$\left(\frac{P}{2\pi}\right)^2 = \frac{a^3}{GM_{sun}}$$

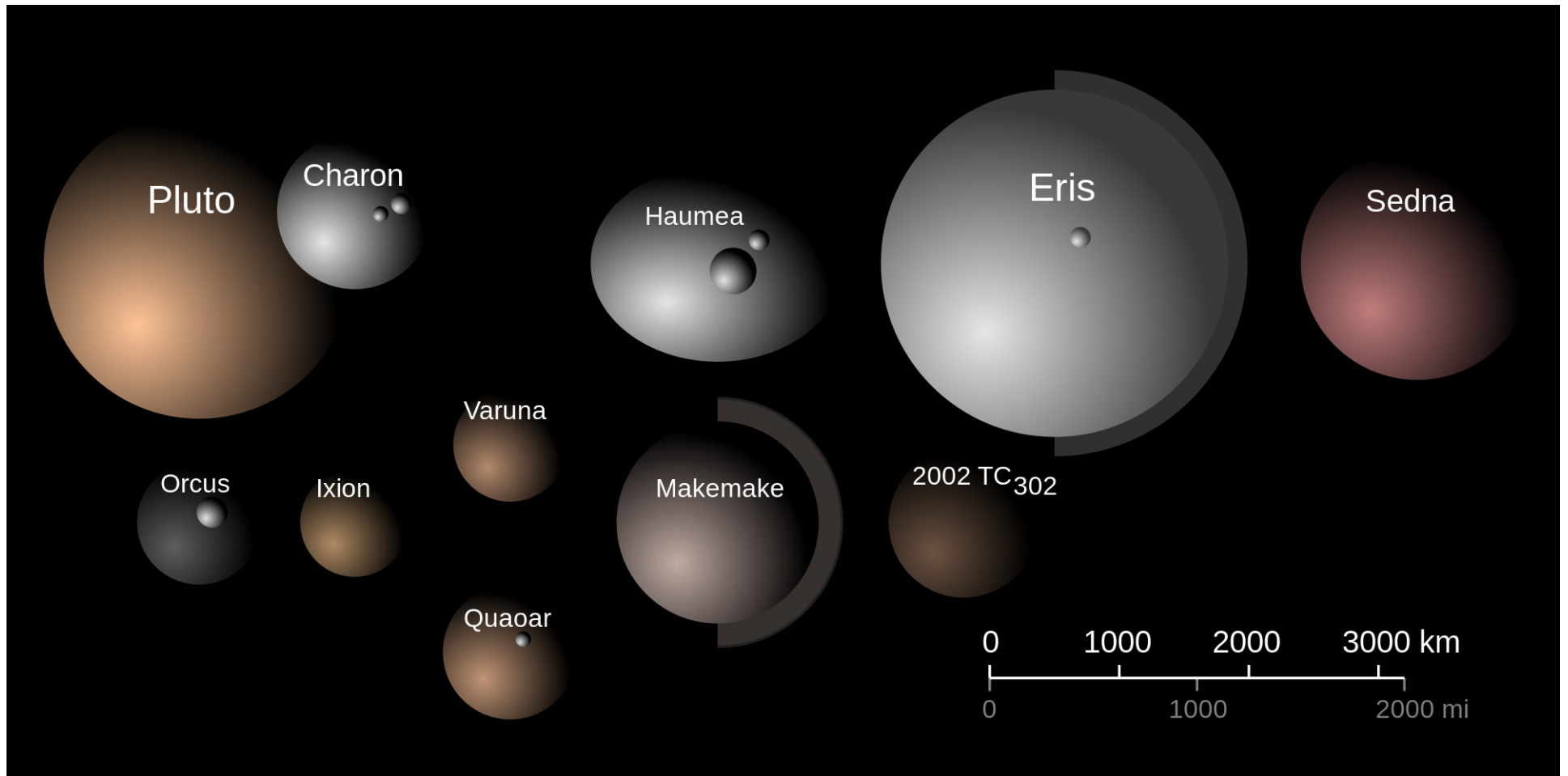
Empirical observations from Kepler; can be derived using Newton's law of gravitation

Planets

| | | | | | | | | | |
|--|----------------------|-----------------------------------|--------------------------------|-------------------------------|-----------------------------------|-------------------------|-------------------------|-----------------------------------|-----------------------------------|
| Table of the largest objects in the Solar System – Wikipedia, the free encyclopedia | | | | | | | | | |
| http://en.wikipedia.org/wiki/Attributes_of_the_largest_solar_system_bodies | | | | | | | | | |
| Resummation...tal at MSU MTA SZTAKI: ... Dictionary CSCNotesLis...las < TWiki PatVancouve...las < TWiki PhysicsAnaly...las < TWiki Quick guide...nda mo | | | | | | | | | |
| <div>Community portal</div> <div>Recent changes</div> <div>Contact Wikipedia</div> <div>Donate to Wikipedia</div> <div>Help</div> <div>toolbox</div> <div>What links here</div> <div>Related changes</div> <div>Upload file</div> <div>Special pages</div> <div>Printable version</div> <div>Permanent link</div> <div>Cite this page</div> <div>languages</div> <div>Alemannisch</div> <div>Deutsch</div> <div>Español</div> <div>Italiano</div> <div>Lëtzebuergesch</div> <div>Română</div> <div>Suomi</div> <div>ייִדיש</div> <div>中文</div> | | | | | | | | | |
| Planets | | | | | | | | | |
| | | Mercury | Venus | Earth | Mars | Jupiter | Saturn | Uranus | Neptune |
| Astronomical symbol | | ♿ | ♀ | ♁ | ♂ | ♃ | ♄ | ♅ | ♆ |
| Mean distance from Sun | km | 57,909,175 | 108,208,930 | 149,597,870 | 227,936,640 | 778,412,010 | 1,426,725,400 | 2,870,972,200 | 4,498,252,900 |
| | AU | 0.38709893 | 0.72333199 | 1 | 1.52366231 | 5.20336301 | 9.53707032 | 19.19126393 | 30.06896348 |
| Mean radius | km | 2,439.64 | 6,051.59 | 6,378.15 | 3,397.00 | 71,492.68 | 60,267.14 | 25,557.25 | 24,766.36 |
| | :E ^[1] | 0.3825 | 0.9488 | 1 | 0.53226 | 11.209 | 9.449 | 4.007 | 3.883 |
| Surface area | km² | 75,000,000 | 460,000,000 | 510,000,000 | 140,000,000 | 64,000,000,000 | 44,000,000,000 | 8,100,000,000 | 7,700,000,000 |
| | :E ^[1] | 0.1471 | 0.9010 | 1 | 0.2745 | 125.5 | 86.27 | 15.88 | 15.10 |
| Volume | km³ | 6.083×10 ¹⁰ | 9.28×10 ¹¹ | 1.083×10 ¹² | 1.6318×10 ¹¹ | 1.431×10 ¹⁵ | 8.27×10 ¹⁴ | 6.834×10 ¹³ | 6.254×10 ¹³ |
| | :E ^[1] | 0.056 | 0.87 | 1 | 0.151 | 1,321.3 | 763.59 | 63.086 | 57.74 |
| Mass | kg | 3.302×10 ²³ | 4.8690×10 ²⁴ | 5.9742×10 ²⁴ | 6.4191×10 ²³ | 1.8987×10 ²⁷ | 5.6851×10 ²⁶ | 8.6849×10 ²⁵ | 1.0244×10 ²⁶ |
| | :E ^[1] | 0.055 | 0.815 | 1 | 0.107 | 318 | 95 | 14 | 17 |
| Density | g/cm³ | 5.43 | 5.24 | 5.515 | 3.940 | 1.33 | 0.70 | 1.30 | 1.76 |
| Equatorial gravity | m/s² | 3.70 | 8.87 | 9.81 | 3.71 | 23.12 | 8.96 | 8.69 | 11.00 |
| Escape velocity | km/s | 4.25 | 10.36 | 11.18 | 5.02 | 59.54 | 35.49 | 21.29 | 23.71 |
| Rotation period | days ^[2] | 58.646225 | -243.0187 ^[3] | 0.99726968 | 1.02595675 | 0.41354 | 0.44401 | -0.71833 ^[3] | 0.67125 |
| Orbital period | years ^[2] | 0.2408467 | 0.61519726 | 1.0000174 | 1.8808476 | 11.862615 | 29.447498 | 84.016846 | 164.79132 |
| Mean orbital speed | km/s | 47.8725 | 35.0214 | 29.7859 | 24.1309 | 13.0697 | 9.6724 | 6.8352 | 5.4778 |
| Eccentricity | | 0.20563069 | 0.00677323 | 0.01671022 | 0.09341233 | 0.04839266 | 0.05415060 | 0.04716771 | 0.00858587 |
| Inclination | deg. | 7.00487 | 3.39471 | 0.00005 | 1.85061 | 1.30530 | 2.48446 | 0.76986 | 1.76917 |
| Axial tilt ^[4] | deg. | 0.0 | 177.3 | 23.45 | 25.19 | 3.12 | 26.73 | 97.86 | 29.58 |
| Mean surface temp. | K | 440 | 730 | 288-293 | 186-268 | 152 | 134 ^[5] | 76 ^[5] | 72 ^[5] |
| Mean air temp. ^[6] | K | | | 288 | | 165 | 135 | 76 | 73 |
| Atmospheric composition | | He Na ⁺ P ⁺ | CO ₂ N ₂ | N ₂ O ₂ | CO ₂ N ₂ Ar | H ₂ He | H ₂ He | H ₂ He CH ₄ | H ₂ He CH ₄ |
| Number of known moons | | 0 | 0 | 1 | 2 | 63 | 60 | 27 | 13 |
| Rings? | | No | No | No | No | Yes | Yes | Yes | Yes |
| Planetary discriminant ^[7] | | 9.1×10 ⁴ | 1.35×10 ⁶ | 1.7×10 ⁶ | 1.8×10 ⁵ | 6.25×10 ⁵ | 1.9×10 ⁵ | 2.9×10 ⁴ | 2.4×10 ⁴ |
| Dwarf planets | | | | | | | | | |
| | | Ceres | Pluto | Makemake | Eris | | | | |

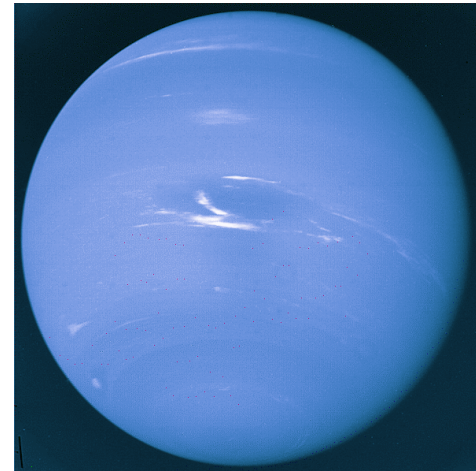
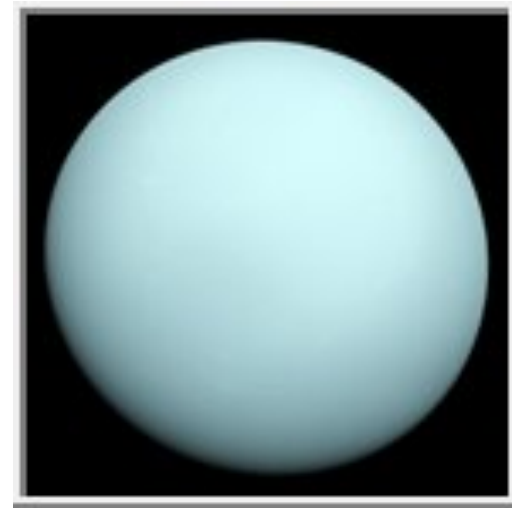
Dwarf planets

...from the outer reaches of the solar system; they take 100's of years (or more) to orbit the sun



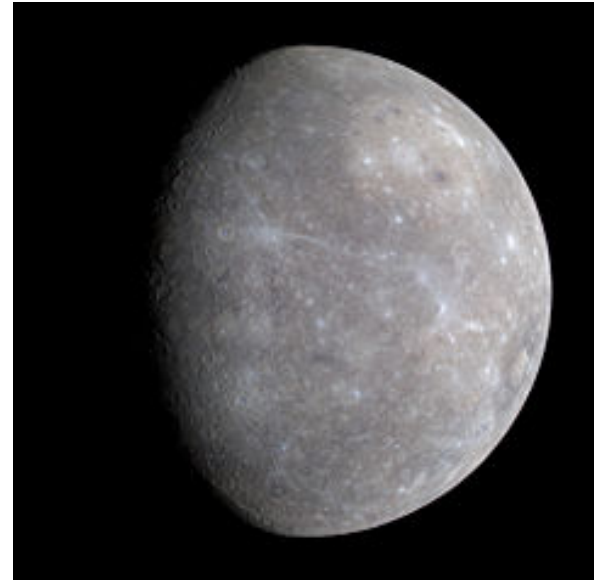
Uranus

- First planet discovered in modern times
 - ◆ not visible to naked eye
- The largest gravitational force in the solar system is due to the Sun (most of the mass)
- But the other planets in the Solar System tug on each other and cause the planets to wobble in their orbits
- If you calculate the effects of all of the other planets on Uranus' wobble, it's not enough
- Either the universal law of gravitation doesn't work at these large distances or there's an 8th planet
- There is an 8th planet (Neptune) and it was where they calculated it should be



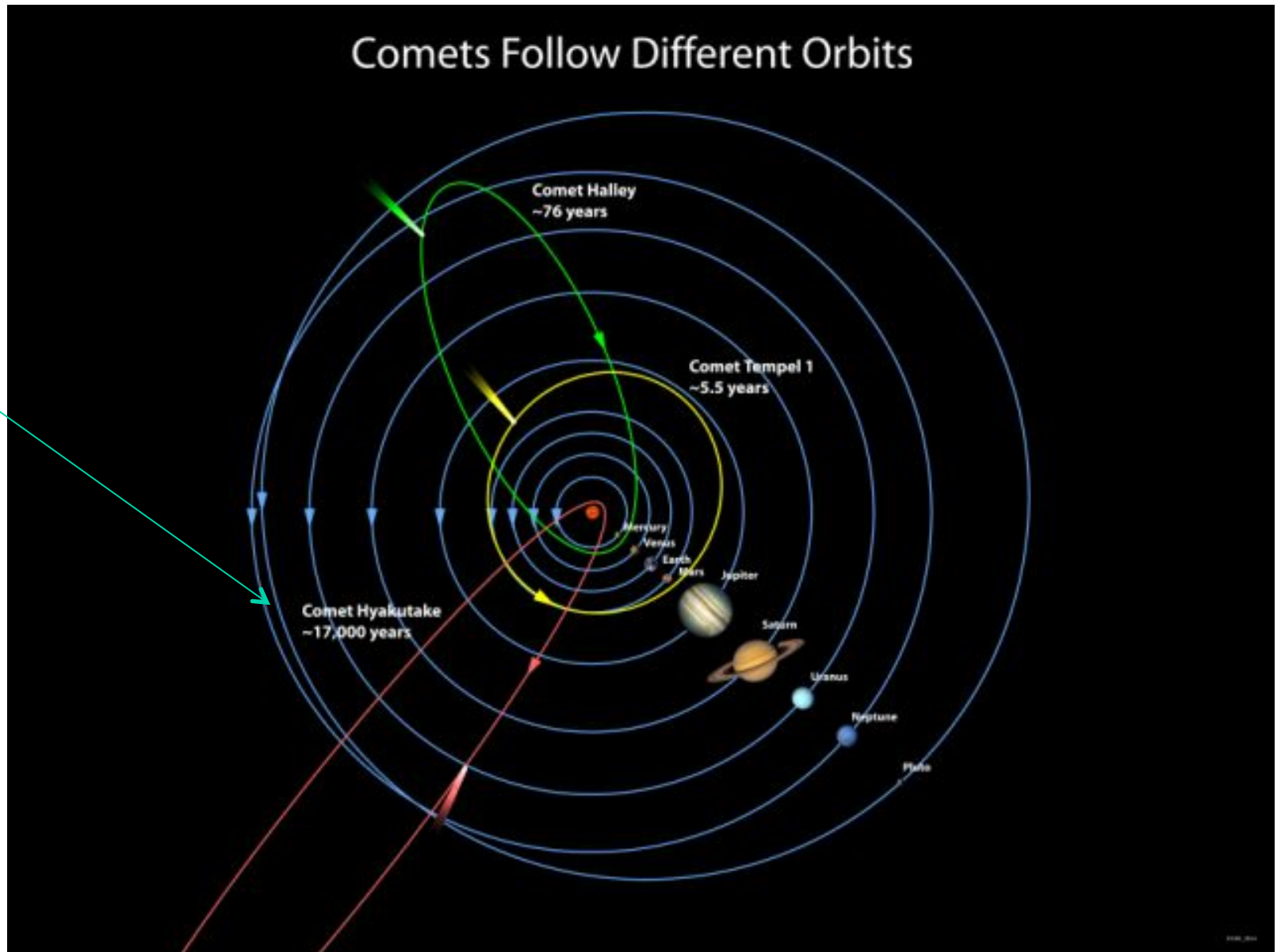
Mercury

- Closest planet to the sun
- Tugs on Mercury perturb orbit
- Not covered by Newtonian physics
- Another planet closer to the sun?
 - ◆ Vulcan, Spock's planet
- ...or effects of general relativity
 - ◆ hint: it's not Vulcan



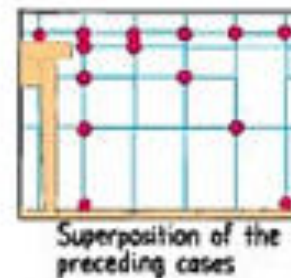
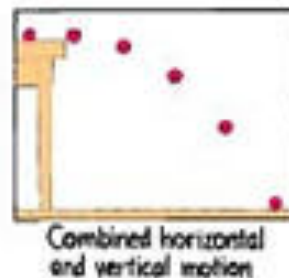
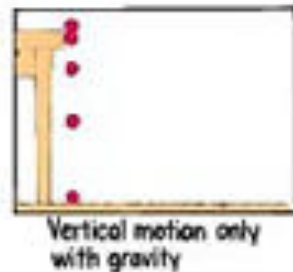
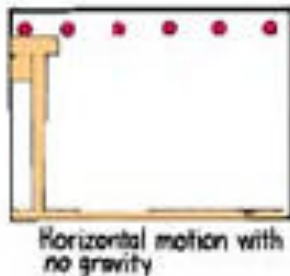
Orbits

From 1979
to 1999
Pluto was
closer to
the Sun
than Neptune



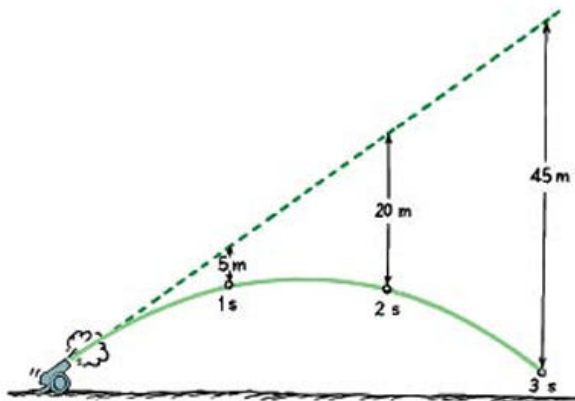
Vertical and horizontal motions

- When both vertical and horizontal motions are present, they can be treated completely independently
- For example, below is shown a ball rolling off of a table with a constant horizontal velocity
- The constant horizontal velocity continues (ignoring any air resistance) while there is a vertical acceleration due to gravity

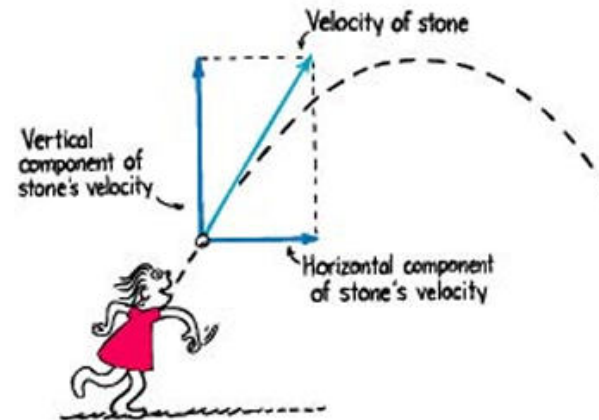


Projectile motion

- With no gravity, the cannon ball would follow a straight line
- Because of the acceleration due to gravity, it follows a parabolic path



- If I throw a stone, it will also follow a parabolic path



simulation

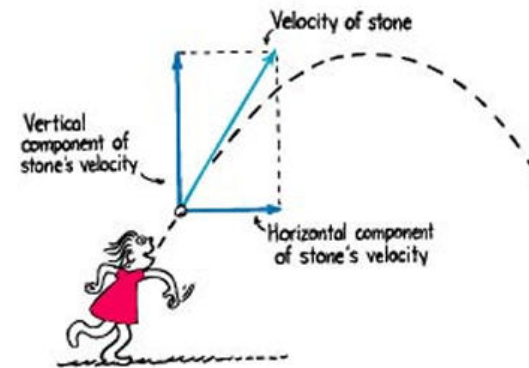
Projectile Motion

- So we have motion in both the x and y directions
- And the two motions are independent so we can write down two separate equations for the x and y motions

$$x = x_0 + v_0^x t + \frac{1}{2} a_x t^2$$

$$y = y_0 + v_0^y t + \frac{1}{2} a_y t^2$$

- I can simplify somewhat since there is no acceleration in the x direction and I can write the acceleration in the y direction as -g

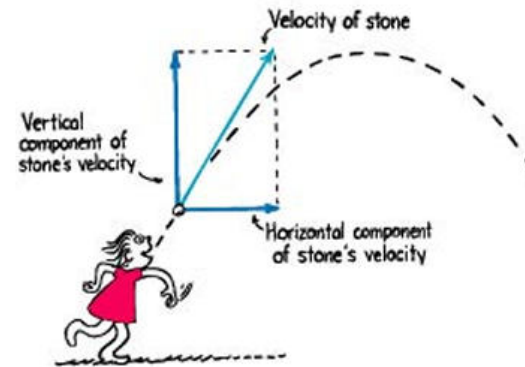


$$x = x_0 + v_0^x t$$

$$y = y_0 + v_0^y t - \frac{1}{2} g t^2$$

Projectile Motion

- Let's start simple
- I throw the ball horizontally with a speed of 20 m/s
- How long before it hits the ground?
- How far has it travelled?



$$x = x_0 + v_0^x t$$

$$y = y_0 + v_0^y t - \frac{1}{2} g t^2$$

Projectile Motion

- Assume that I release it 2 m from the ground
- $y_0 = 2\text{m}$, $v_0^y = 0 \text{ m/s}$

$$y = y_0 - \frac{1}{2}gt^2$$

$$0 = 2\text{m} - \frac{1}{2}(9.83\text{m/s}^2)t^2$$

$$t^2 = \frac{4\text{m}}{9.83\text{m/s}^2} = 0.407\text{s}^2$$

$$t = 0.64\text{s}$$

$$x = x_0 + 20\text{m/s}(0.64\text{s}) = x_0 + 12.8\text{m}$$



$$x = x_0 + v_0^x t$$

$$y = y_0 + v_0^y t - \frac{1}{2}gt^2$$

Projectile Motion

- Suppose I throw it at 20 m/s at an angle of 45°
- Let's again start with the vertical motion
 - ◆ how long before it hits the ground?

$$0 = 2m + (20m/s) \sin 45^\circ t - \frac{1}{2}(9.83m/s^2)t^2$$

$$0 = 2m + (20m/s)(0.707)t - (4.915m/s^2)t^2$$

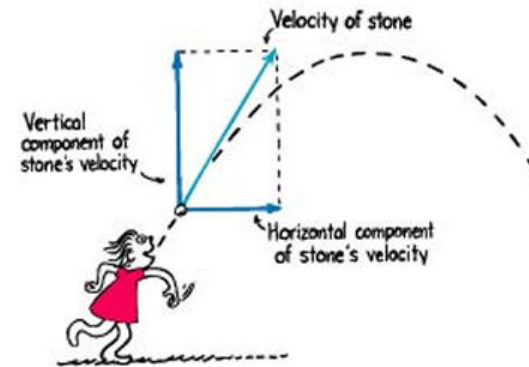
$$4.915t^2 - 14.14t - 2 = 0$$

$$t = 3.01s$$

- Now the horizontal motion

$$x = x_0 + (20m/s) \cos 45^\circ t$$

$$x = x_0 + (20m/s)(0.707)(3.01s) = x_0 + 42.6m$$



$$x = x_0 + v_0^x t$$

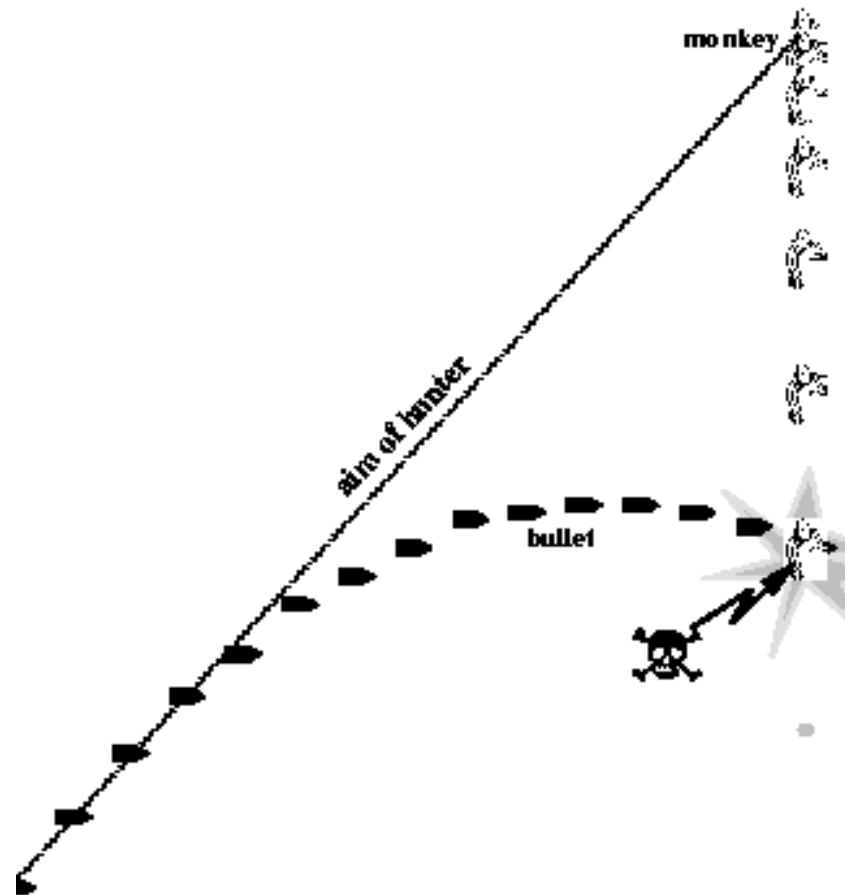
$$y = y_0 + v_0^y t - \frac{1}{2} g t^2$$

Projectile motion
Shoot the monkey

'Shoot the monkey'

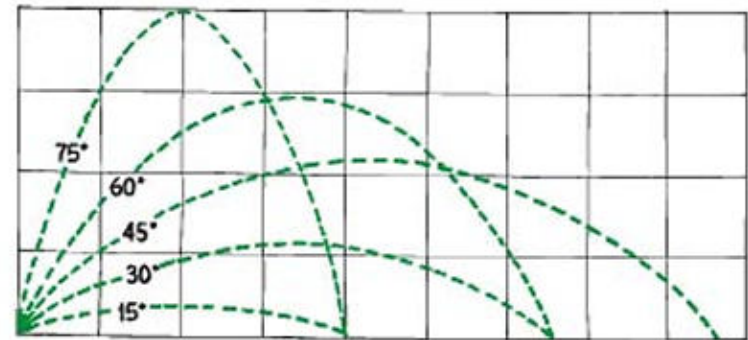
- We said that the vertical and horizontal motions for projectiles were independent
- Neglecting air resistance, there is an acceleration only in the vertical direction (due to gravity) and it is the same regardless of whether there is horizontal motion or not
- Suppose a hunter is aiming at a monkey hanging in a tree
- The monkey lets go at the same instant that the hunter pulls the trigger
- Does the monkey get hit?
- Yes, if the initial aim is correct, because the monkey and the bullet have the same acceleration

Figure 4.5: Shoot the "monkey": an illustration of motion in two dimensions.



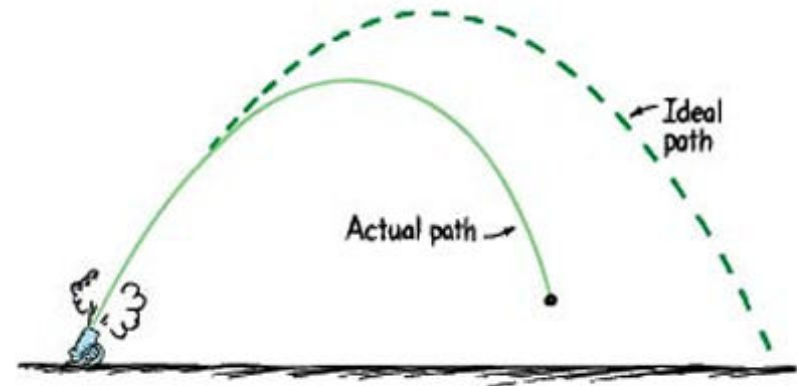
...no actual monkeys will be harmed in this demonstration

- What angle should you throw a ball in order for it to go the maximum distance, given that the initial release velocity is the same?



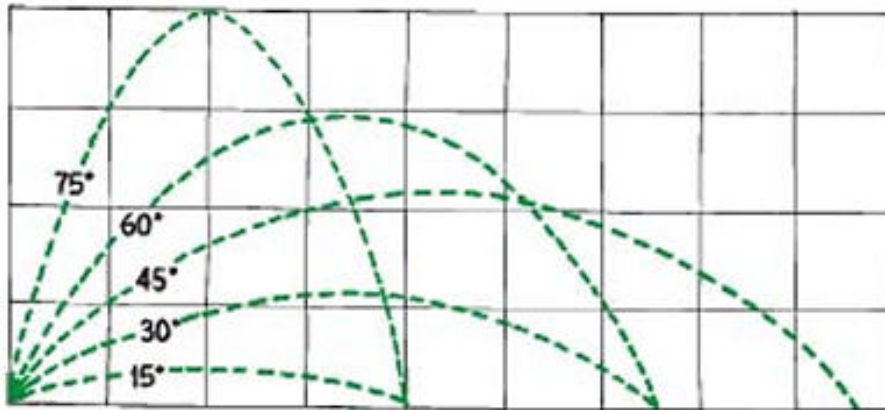
- ◆ somewhere between 0 and 90 degrees
- ◆ to be more precise 45 degrees

- What is the impact of air resistance?



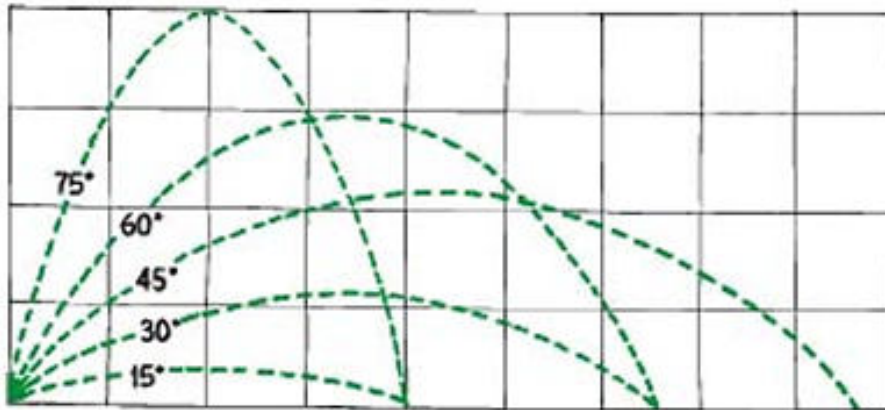
Clicker question

- For the possible paths for the projectile below, which has the largest vertical acceleration?
- A) 75°
- B) 60°
- C) 45°
- D) 30°
- E) they're all the same

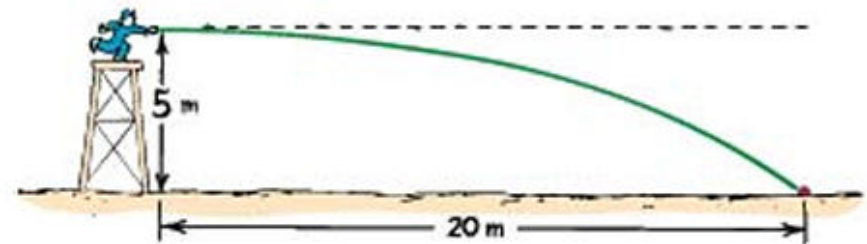


Clicker question

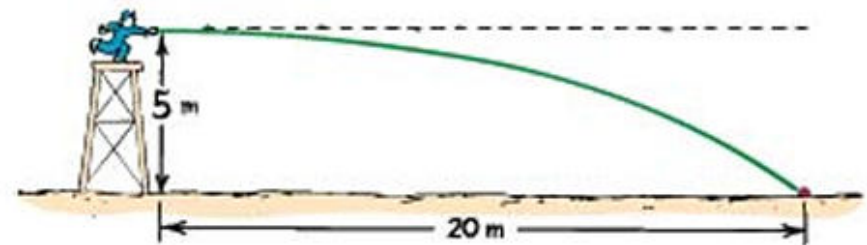
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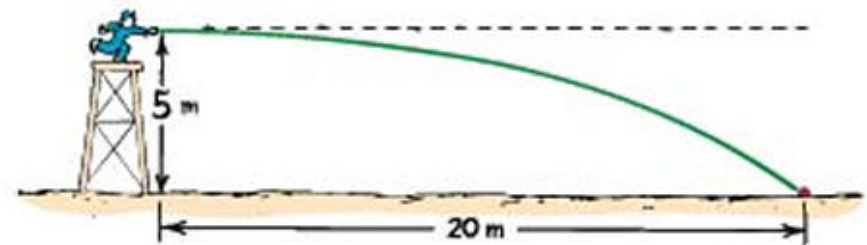
- How fast is the ball thrown?
- How would you approach the problem?



- How fast is the ball thrown?
- How would you approach the problem?
- Separate into horizontal and vertical motions
- How long does it take for the ball to drop 5 m?
- During that time it has travelled 20 m horizontally



- How fast is the ball thrown?
- How would you approach the problem?
- Separate into horizontal and vertical motions
- **How long does it take for the ball to drop 5 m?**
- During that time it has travelled 20 m horizontally



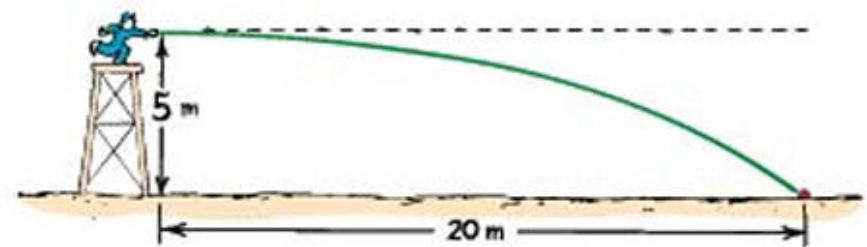
$$y = y_o + v_o^y t - \frac{1}{2} g t^2$$

$$0 = 5m - \frac{1}{2} (9.83m/s^2) t^2$$

$$t^2 = \frac{10m}{9.83m/s^2}$$

$$t = 1.01s$$

- How fast is the ball thrown?
- How would you approach the problem?
- Separate into horizontal and vertical motions
- How long does it take for the ball to drop 5 m?
- During that time it has travelled 20 m horizontally



$$t = 1.01s$$

$$x = x_o + v_o^x t$$

$$20m = v_o^x (1.01s)$$

$$v_o^x = 19.8m/s$$

iclicker question

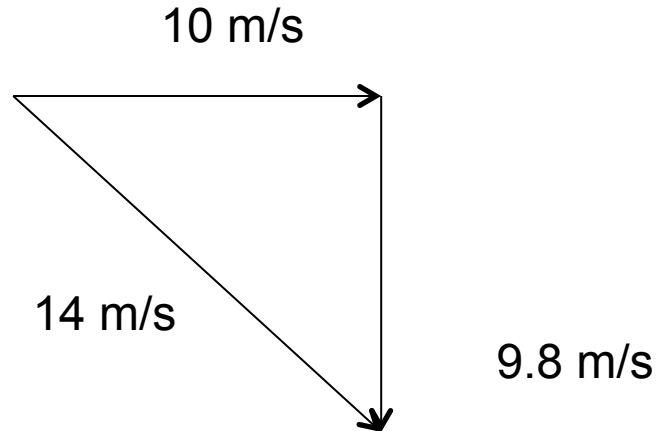
A cannonball is fired horizontally with a speed of 10 m/s from the edge of a cliff. What is its speed 1 second after being fired? Ignore air resistance.

- a) 10 m/s
- b) 14 m/s
- c) 16 m/s
- d) 20 m/s
- e) 30 m/s

iclicker question

A cannonball is fired horizontally with a speed of 10 m/s from the edge of a cliff. What is its speed 1 second after being fired?

- a) 10 m/s
- b) **14 m/s**
- c) 16 m/s
- d) 20 m/s
- e) 30 m/s



Satellites

- What happens as you throw the ball harder and harder?
- It goes farther before it hits the Earth's surface
- The Earth's surface falls off about 5 m every 8000 m
- If you can throw a ball hard enough so that it travels 8000 m in the 1 second it takes to fall 5 m, then it will keep on falling around the surface of the Earth
 - ◆ 8 km/s
 - ◆ or 29,000 km/hour

