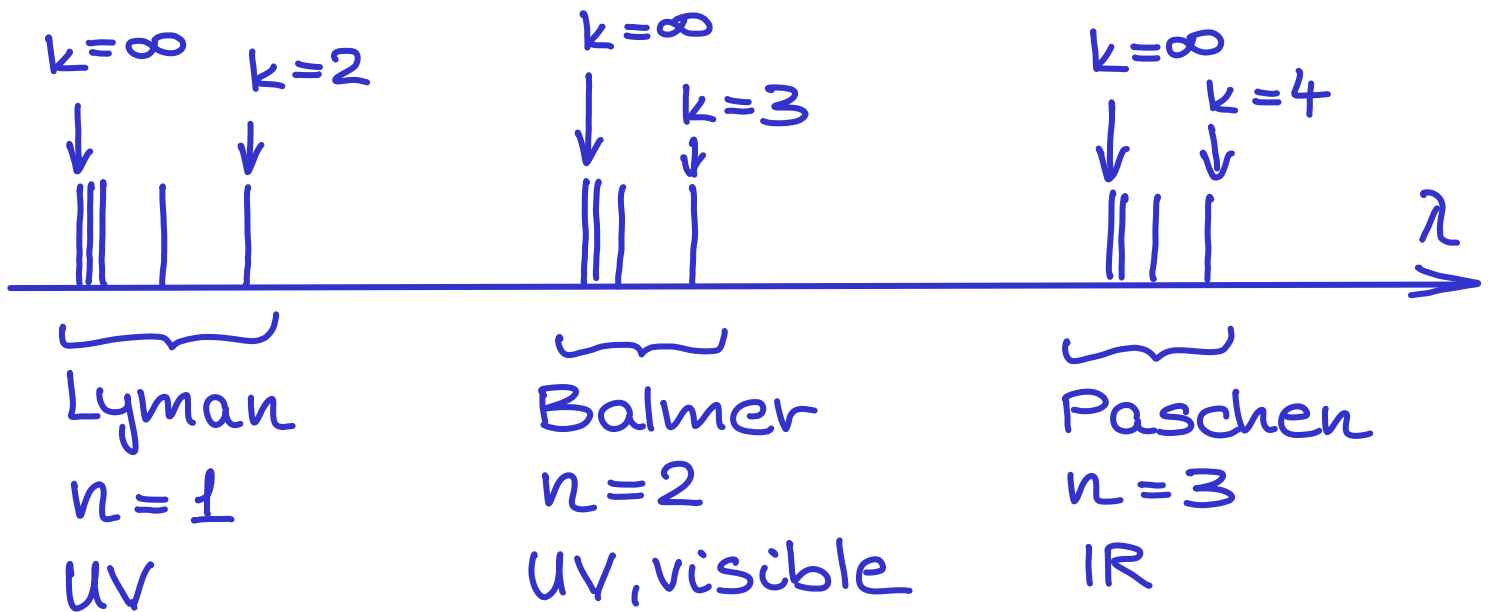


Series limits

$$\frac{1}{\lambda} = R_H \left(\frac{1}{n^2} - \frac{1}{k^2} \right); R_H = 1.096776 \cdot 10^7 \frac{1}{m}$$



$$\frac{1}{\lambda_{\text{lim}}} = R_H \left(\frac{1}{n^2} - \frac{1}{\infty^2} \right) = \frac{R_H}{n^2}$$

$$\lambda_{\text{lim}} = \frac{n^2}{R_H} \left. \begin{array}{l} \text{shortest wavelength} \\ \text{in the series} \end{array} \right\}$$

Longest wavelength in a series:

$$\begin{aligned} \frac{1}{\lambda_{\text{long}}} &= R_H \left(\frac{1}{n^2} - \frac{1}{(n+1)^2} \right) = R_H \frac{(n+1)^2 - n^2}{n^2 \cdot (n+1)^2} \\ &= R_H \frac{2n+1}{n^2 \cdot (n+1)^2} \end{aligned}$$

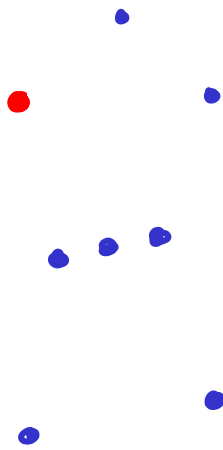
Series limits 2.

$$\lambda_{\text{longest}} = \frac{n^2 \cdot (n+1)^2}{R_H \cdot (2n+1)}$$

n	n^2	$\lambda_{\text{lim}} = \frac{n^2}{R_H}$	$\frac{n^2(n+1)^2}{2n+1}$	λ_{longest}
1	1	91.2 nm	$\frac{4}{3} = 1.333$	121.6 nm
2	4	364.7 nm	$\frac{36}{5} = 7.2$	656.5 nm
3	9	820.6 nm	$\frac{144}{7} = 20.57$	1876 nm
4	16	1459 nm	$\frac{400}{9} = 44.44$	4052 nm
5	25	2279 nm	$\frac{900}{11} = 81.81$	7460 nm

visible spectrum: 400-700 nm

Betelgeuse



$$\lambda_{\max} = 800 \text{ nm} \\ (\text{infrared})$$

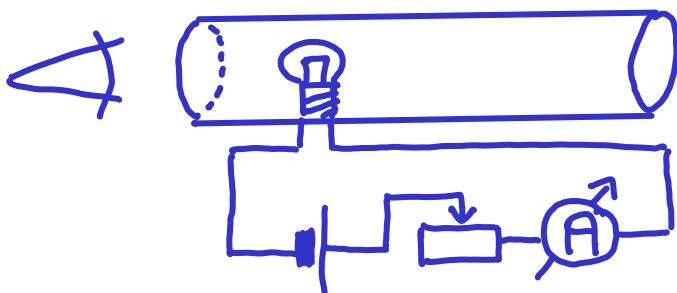
Wien's displacement law:

$$\lambda_{\max} \cdot T = b ; b = 2.898 \cdot 10^{-3} \text{ m} \cdot \text{K}$$

$$T = \frac{b}{\lambda_{\max}} = \frac{2.898 \cdot 10^{-3}}{800 \cdot 10^{-9}} = 3622 \text{ K}$$

$$\text{Sun: } \lambda_{\max} = 500 \text{ nm} \\ \rightarrow T = 5800 \text{ K}$$

Bolometer:



Cosmic background radiation

$$\lambda_{\max} = 1.07 \text{ mm} \longleftrightarrow T = 2.7 \text{ K}$$

$$f_{\max} = \frac{c}{\lambda_{\max}} = \frac{3 \cdot 10^8 \text{ m/s}}{1.07 \cdot 10^{-3} \text{ m}} = 280 \text{ GHz}$$

Penzias, Wilson (1978, Nobel)

COBE : Cosmic Background
Explorer : $T = 2.73 \text{ K}$

WMAP : Wilkinson Microwave
Anisotropy Probe :
 $T = 2.725 \text{ K}$

Efficiency of an incandescent bulb

$$P = \sigma \epsilon A T^4 \quad T = 2330K \quad \text{Tungsten}$$

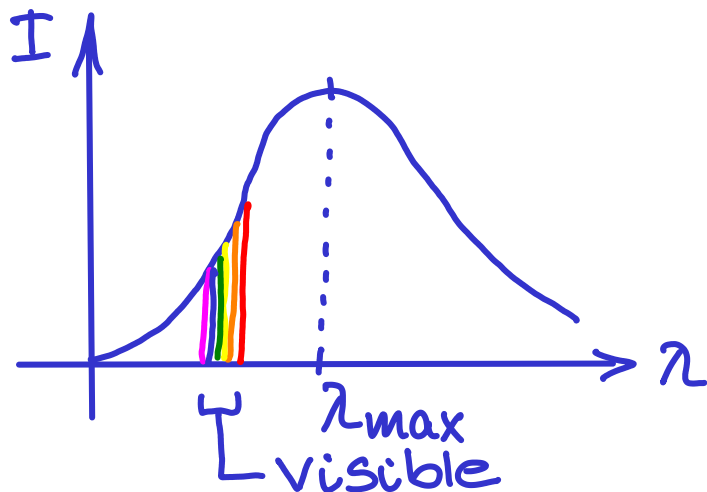
$$\frac{P}{A} = \sigma \epsilon T^4$$

$$\epsilon = 1$$

$$\sigma = 5.6704 \cdot 10^{-8} \frac{W}{m^2 K^4}$$

$$\frac{P}{A} = 5.67 \cdot 10^{-8} \cdot 2330^4 = 1.67 \cdot 10^6 \frac{W}{m^2}$$

$$(100W \text{ lightbulb} \Rightarrow A \approx 60 \text{ mm}^2)$$



$$\lambda_{\max} = \frac{b}{T} \approx 1240 \text{ nm}$$

$$\lambda_{\text{blue}} = 400 \text{ nm}$$

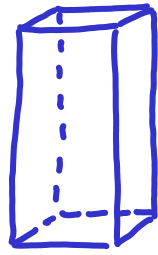
$$\lambda_{\text{red}} = 700 \text{ nm}$$

Max Planck:

$$I(\lambda, T) = \frac{2\pi c^2 h}{\lambda^5} \cdot \frac{1}{e^{\frac{hc}{\lambda k_B T}} - 1}$$

$$\eta \approx 2-6\%$$

Power emitted by the human body



$$\begin{aligned}H &= 1.88 \text{ m} \\W &= 0.315 \text{ m} \\L &= 0.26 \text{ m}\end{aligned}$$

$$\epsilon = 97\% = 0.97$$

$$\alpha = 97\% = 0.97$$

$$T_{\text{body}} = 35.0^\circ\text{C} = 308 \text{ K}$$

$$P_{\text{out}} = \epsilon \sigma A T_{\text{body}}^4$$

$$\sigma = 5.67 \cdot 10^{-8} \frac{\text{W}}{\text{m}^2 \text{K}^4}$$

$$A = 2(HW + WL + LH)$$

$$P_{\text{out}} \approx 1000 - 1500 \text{ W!}$$

$$T_{\text{env}} = 22.5^\circ\text{C} = 295.5 \text{ K}$$

$$P_{\text{in}} = \alpha \cdot P_{\text{env}} = \alpha \cdot \epsilon \cdot 1 \cdot A T_{\text{env}}^4$$

↑ black body rad.

$$P_{\text{in}} = P_{\text{out}} \cdot \left(\frac{T_{\text{env}}}{T_{\text{body}}} \right)^4$$

$$P_{\text{in}} \approx 900 - 1200 \text{ W}$$

$$P_{\text{loss}} = P_{\text{out}} - P_{\text{in}} \approx 100 - 200 \text{ W}$$

$$\lambda_{\text{max}} = \frac{b}{T_{\text{body}}} = \frac{2.898 \cdot 10^{-3} \text{ m} \cdot \text{K}}{308 \text{ K}} = \underline{9410 \text{ nm}}$$

far infrared

Radio station

$$f = 94.9 \text{ MHz (WMMQ)}$$

$$P = 46.5 \text{ kW (?)}$$

$$\text{Planck constant: } h = 6.63 \cdot 10^{-34} \text{ Js}$$

Energy of one photon at the frequency above:

$$\begin{aligned} E = hf &= 6.63 \cdot 10^{-34} \text{ Js} \cdot 94.9 \cdot 10^6 \frac{1}{\text{s}} = \\ &= 6.29 \cdot 10^{-26} \text{ J} \end{aligned}$$

Number of photons per second:

$$\frac{P}{E} = \frac{46.5 \cdot 10^3 \text{ W}}{6.29 \cdot 10^{-26} \text{ J}} = 7.39 \cdot 10^{29} \frac{1}{\text{s}}$$

Roughly a million Avogadro number of photons each second.

Photoelectric effect

Work function: $\phi = 3.25 \text{ eV}$

$$\text{Einstein: } hf = \phi + \underbrace{KE}_{=0} \Rightarrow f_{\min} = \frac{\phi}{h}$$

$$h = 6.63 \cdot 10^{-34} \text{ Js} = 4.14 \cdot 10^{-15} \text{ eVs}$$

$$f_{\min} = \frac{3.25 \text{ eV}}{4.14 \cdot 10^{-15} \text{ eVs}} = 7.85 \cdot 10^{14} \text{ Hz}$$

$$\lambda_{\max} = \frac{c}{f_{\min}} = \frac{3 \cdot 10^8 \text{ m/s}}{7.85 \cdot 10^{14} \text{ Hz}} \approx \underline{382 \text{ nm}}$$

(ultra) violet

$$hc = 1240 \text{ eVnm}$$

$$\lambda_{\max} = \frac{hc}{hf} = \frac{1240 \text{ eVnm}}{3.25 \text{ eV}} \approx 382 \text{ nm}$$

When the photon frequency is doubled, the energy of the photon doubles:

$E_{\gamma} = 6.50 \text{ eV}$. Therefore the kinetic energy of the electrons is going to be $KE = 3.25 \text{ eV}$.

$$\underbrace{hf}_{6.5 \text{ eV}} = \underbrace{\phi}_{3.25 \text{ eV}} + \underbrace{KE}_{3.25 \text{ eV}}$$

Planck's constant from the photoelectric effect

$V_0 (V)$	$\lambda (nm)$
3.47	276
5.12	216

$$hf = \phi + eV_0$$

$$\frac{hc}{\lambda} = \phi + eV_0$$

$$\left. \begin{aligned} h \frac{c}{\lambda_1} &= \phi + eV_1 \\ h \frac{c}{\lambda_2} &= \phi + eV_2 \end{aligned} \right\} \text{Two equations, two unknowns: } h, \phi$$

$$h \frac{c}{\lambda_1} - eV_1 = h \frac{c}{\lambda_2} - eV_2$$

$$e(V_2 - V_1) = hc \left(\frac{1}{\lambda_2} - \frac{1}{\lambda_1} \right)$$

$$h = \frac{e(V_2 - V_1)}{c \left(\frac{1}{\lambda_2} - \frac{1}{\lambda_1} \right)}$$

$$h = \frac{1e \cdot (5.12V - 3.47V)}{3 \cdot 10^8 \frac{m}{s} \left(\frac{1}{2.16 \cdot 10^{-7}m} - \frac{1}{2.76 \cdot 10^{-7}m} \right)}$$

$$h = 5.46 \cdot 10^{-15} \text{ eVs}$$

$$(\text{Official value: } 4.14 \cdot 10^{-15} \text{ eVs})$$

Planck's constant 2.

$$\begin{aligned}\phi &= \frac{hc}{\lambda_1} - eV_1 = \\ &= \frac{5.46 \cdot 10^{-15} \text{ eVs} \cdot 3 \cdot 10^8 \frac{\text{m}}{\text{s}}}{2.76 \cdot 10^{-7} \text{ m}} - 3.47 \text{ eV} =\end{aligned}$$

$$= 5.93 \text{ eV} - 3.47 \text{ eV} = 2.46 \text{ eV}$$

$$\phi = \frac{hc}{\lambda_2} - eV_2 = \dots$$

X-ray production by bremsstrahlung

Voltage applied to the X-ray tube:

$$\Delta V = 25.4 \text{ kV}$$

Energy balance:

$$e \Delta V = hf_{\max}$$

$$e \Delta V = \frac{hc}{\lambda_{\min}}$$

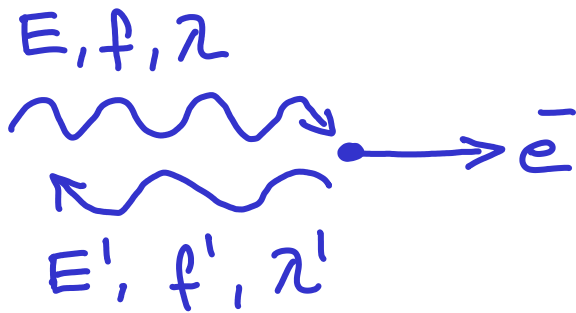
$$\lambda_{\min} = \frac{hc}{e \Delta V} = \frac{1240 \text{ eV nm}}{1e \cdot 25,400 \text{ V}}$$

$$\lambda_{\min} = 0.0488 \text{ nm} = 48.8 \text{ pm}$$

Maximum energy transfer in Compton scattering

$$\Delta\lambda = \lambda' - \lambda = \underbrace{\frac{h}{m_e c}}_{\lambda_c = 2.426 \text{ pm}} (1 - \cos\theta)$$

$\Delta\lambda$ is maximum when $\theta = 180^\circ$:



$$E = 42.2 \text{ keV}$$

$$\Delta\lambda = 2 \cdot \frac{h}{m_e c} = 2\lambda_c \quad \text{when } \theta = 180^\circ$$

$$\lambda' = \lambda + 2\lambda_c = \frac{hc}{E} + 2 \frac{h}{m_e c} = \frac{hc}{E} + 2 \frac{hc}{m_e c^2}$$

$$E' = \frac{hc}{\lambda'} = \frac{hc}{\frac{hc}{E} + 2 \frac{hc}{m_e c^2}} = \frac{1}{\frac{1}{E} + \frac{2}{m_e c^2}}$$

$$\Delta E = E - E' = E - \frac{1}{\frac{1}{E} + \frac{2}{m_e c^2}}$$

Maximum energy transfer 2.

$$\Delta E = E - \frac{E \cdot m_e c^2}{m_e c^2 + 2E} = \frac{2E^2}{m_e c^2 + 2E}$$

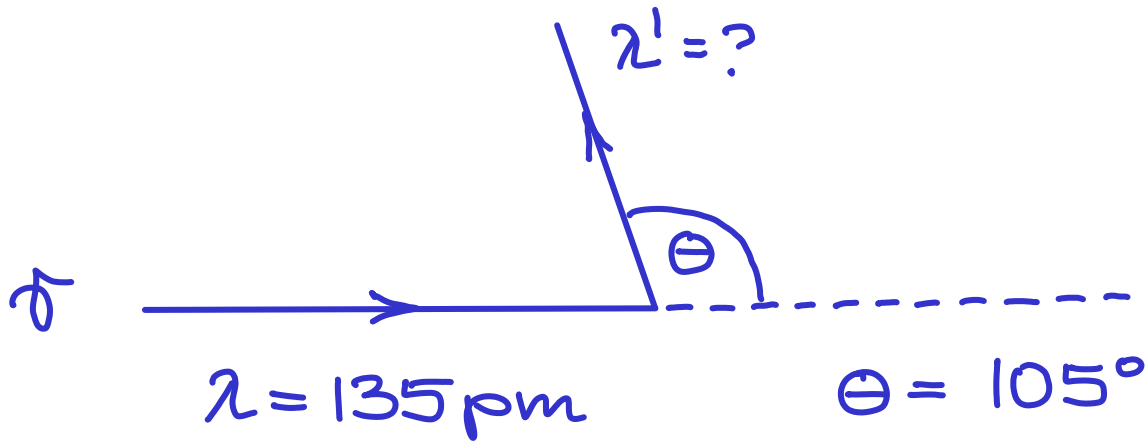
$$m_e c^2 = 511 \text{ keV} \quad E = 42.2 \text{ keV}$$

$$\Delta E = \frac{2 \cdot 42.2^2}{511 + 2 \cdot 42.2} = 5.98 \text{ keV}$$

$$E' = 36.2 \text{ keV}$$

This maximum energy transfer happens, when the photon bounces exactly backwards: $\Theta = 180^\circ$ and the electron is kicked exactly forward: $\phi = 0^\circ$.

Compton scattering



Compton:

$$\Delta\lambda = \frac{h}{m_e c} (1 - \cos\Theta)$$

$$\lambda' = \lambda + \Delta\lambda = \lambda + \frac{h}{m_e c} (1 - \cos\Theta) =$$

$$= \lambda + \frac{hc}{m_e c^2} (1 - \cos\Theta) =$$

$$= 0.135 \text{ nm} + \underbrace{\frac{1240 \text{ eV nm}}{511,000 \text{ eV}}}_{0.002426 \text{ nm}} (1 - \cos 105^\circ) =$$

$$= 135 \text{ pm} + 2.426 \text{ pm} \cdot (1 - \cos 105^\circ) =$$

$$= 138.05 \text{ pm}$$

$$\% = 100 \cdot \frac{\Delta\lambda}{\lambda} = 100 \cdot \frac{3.05}{135} = 2.26 \%$$