Thermo dy namics

Jan 2025

Laws of thermodynamics They state empirical facts found in the history of physics and chemistry.

The zeroth low of thermodynamics - Defines thermal equilibrium. If two systems are each in thermal equilibrium with a third system, then they are in thermal equilibrium with each other. - Forms a basis for the definition of Temperature. (1) Temperature reflects the degree of hotness or coldness of an object or a system. Temperature determines the director of heat transfer. Thigh Heat Teow Heat is a transfer of energy, always flows spontaneously from an object with higher T to an object with a lower T. Until the two objects reach the same T, which is called the thermal equilibrium is reached. If $T_A = T_c$ and $T_B = T_c$, The senth law then $T_A = T_B$. (3) Temperature is transitive:

Temperature - In the International system of Units (SI), the unit of temperature is the kelvin (Tk) other commonly used ones are Celsius (Tc) and • Fahrenheight to Celsius Fahrenheit (TF) $T_{\rm C} = \frac{5}{9} \left(T_{\rm F} - 32^{\circ} \mathrm{F} \right)$ Celsius to Fahrenheight $T_{\rm F} = \frac{9}{5}T_{\rm C} + 32 \,^{\circ}{\rm F}$ • Celsius to Kelvin $T_{\rm K} = T_{\rm C} + 273.15 \,^{\circ}{\rm C}$ Kelvin to Celsius $T_{\rm C} = T_{\rm K} - 273.15 \,^{\circ}{\rm C}$ - Absolute Zero is defined as TK=0. 1) In classical physics, it results in -zero kinetic energy. Z In Quantum physics, particles will have non-vanishing zero-point energy. due to the Heisenberg uncertaity principle. $\left(\Delta X^{\Delta} P \geq \frac{h}{4\pi}\right)$ h is the Planck Constant - Temperature Can be measured by a thermometer, such as thermometer. 98.6 100 2 4

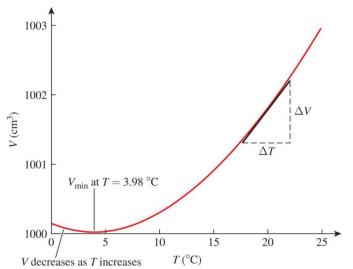
Thermal Expansion
- It's not always linear, but in the case of linear expansion:
() Define X as the coefficient of linear expansion for a solid object

$$\frac{\Delta L}{L_0} = \alpha \, \text{sT}$$
 for one-dimension (1-d), ST is small.
Where $\Delta T = T_F - T_i$
 $\Delta L = L_f - L_i$
The unit of $[\alpha] = \frac{1}{\text{temperature}} \propto \begin{pmatrix} \text{either} \\ \frac{1}{2C} \text{ or } \frac{1}{2F} \text{ or } \frac{1}{2K} \end{pmatrix}$
 $= \frac{1}{2C} \text{ or } \frac{1}{2F} \text{ or } \frac{1}{2K} \end{pmatrix}$
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 $= \frac{1}{2K} \text{ or } \frac{1}{2K} \text{$

Density of Water & Volume expansion of Water - Water is special ! 1.000 0.995 Highest 0.990 Density at 4° C 0.985

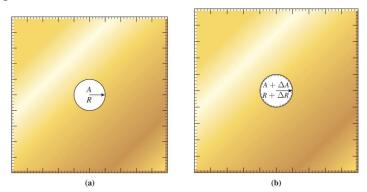
Density (gm/cm⁻³) 0.980 0.975 0.970 Ice Water 0.965 0.960 0.955 0 10 20 30 40 50 60 70 80 -50 -40 -30 -20 -10 90 100 110 {4°C} Temperature (°C) $\equiv \frac{M}{V_{(111)}} (Mass)$ (density) p **Volume Expansion of Water**

• The change in volume with temperature is never linear for water.



Hint to think about solving this problem

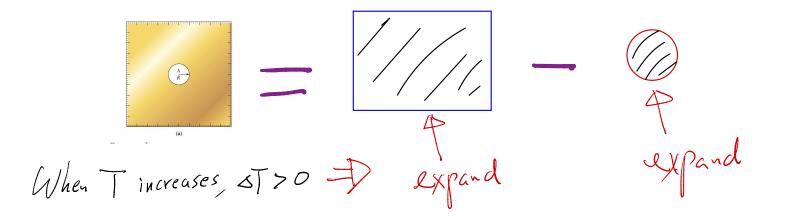
Expansion of a Plate with a Hole in It₃



The change in the area of the hole is $\frac{\Delta A}{\Lambda} = 2K \Delta T$

Research

• The area of the plate increases as the temperature increases and the area of the hole increases proportionally.



The First law of Thermodynamics

The first law of thermodynamics is a version of the law of conservation of energy, adapted for thermodynamic processes. In general, the conservation law states that the total energy of an isolated system is constant; energy can be transformed from one form to another, but can be neither created nor destroyed.

In a closed system (i.e. there is no transfer of matter into or out of the system), the first law states that the change in internal energy of the system ($\Delta U_{\rm system}$) is equal to the difference between the heat supplied to the system (Q) and the work (W) done by the system on its surroundings. (Note, an alternate sign convention, not used in this article, is to define W as the work done *on* the system by its surroundings): (all sys is denoted as the internal energy AEint.)

$$\Delta U_{\text{system}} = Q - W.$$

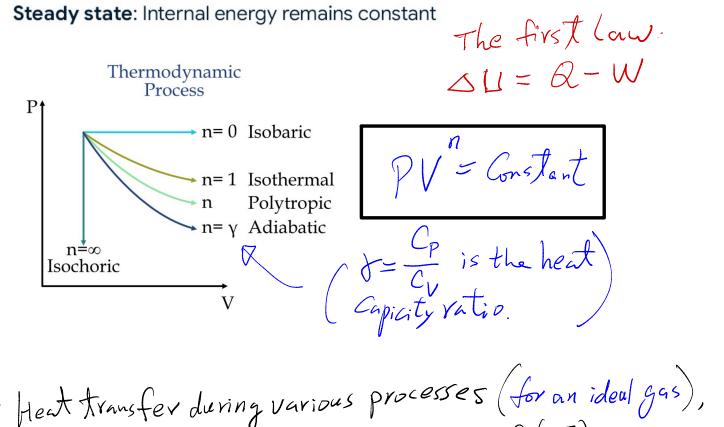
$$(d \mathcal{U} = d \mathcal{R} - p d \mathcal{I}$$

For processes that include the transfer of matter, a further statement is needed.

When two initially isolated systems are combined into a new system, then the total internal energy of the new system, $U_{\rm system}\!,$ will be equal to the sum of the internal energies of the two initial systems U_1 and U_2 . イト

Types of thermodynamic processes

- Adiabatic: No heat transfer into or out of the system
- Isobaric: Pressure remains constant
- Isochoric: Volume remains constant
- Isothermal: Temperature remains constant
- Isenthalpic: Enthalpy remains constant
- Isentropic: A reversible adiabatic process where entropy remains constant



Heat Arauster during various
$$f$$
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() Isobaric: $SP=D$, $Q = n C_p(T_T-T_i) = n C_p(ST)$
() Isobaric: $SP=D$, $Q = n C_p(T_T-T_i) = n C_p(ST)$
() heat capacity at
() for idea ges:
() $PV = nRT$
(

Heat transfer during various processes (for an ideal gas)
() Isobaric:
$$SP=0$$
, $R=n C(T_{T_{c}}) = n Cp(ST)$
 $PV=nRT$
 $W=\int pdV=p(Y-V_{c}) \implies SU=Q-W$ (V_{ros}
 $V=nRT$)
 $Charge of Extropy: $2S = \int \frac{dQ}{T} = \int_{T_{c}}^{T_{c}} nCp \frac{dT}{T} = MC ha(\frac{T_{c}}{T_{c}})$
(2) R_{c} thermal: $aT=0$, $aU=0$
 $aU=Q-W$ = $RT ln(\frac{V_{c}}{V_{c}})$
 $\Delta S = \frac{Q}{T} = nR ln(\frac{V_{c}}{V_{c}})$
 $S= \int \frac{dQ}{T} = nC_{c} h(\frac{T_{c}}{T_{c}})$ (a constant)
 $S= \int \frac{dQ}{T} = \int_{T_{c}}^{T_{c}} nC_{c} \frac{dT}{T} = nC_{c} h(\frac{T_{c}}{T_{c}})$
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 $M = \int pdV = \int \frac{C}{V^{c}} dV = \frac{PV^{2}-C^{2}}{V^{c}}$ ($q=C_{c}TR$)
 $W = \int pdV = \int \frac{C}{V^{c}} dV = \frac{PV^{2}-C^{2}}{V^{c}}$ ($q=C_{c}TR$)
 $T_{c} = \int \frac{dQ}{T} = -W$
 $\Delta S = \int \frac{dQ}{T} = 0$$

Mechanical Equivalent of Heat - In 1843, Joule's experiment Thermometer () Found W = G Q(mechanical) f f f (heat produced) Joule's constant G = 4.186 Joule Calorie = $\begin{pmatrix} W = Fd \\ = mad \\ (kg) \end{pmatrix} \begin{pmatrix} m \\ m \end{pmatrix} \begin{pmatrix} m \\ m \end{pmatrix} \end{pmatrix}$ Work: $I J = I (kg) \frac{vn^2}{S^2} = \begin{pmatrix} I newton of force \\ times one meter of \\ distance \end{pmatrix}$ Heat I Cal = the amount of heat needed to raise the temperature of 1 g water by 1°C. = 4.186 J 2 Unified Pifferent forms of energies, (mechanical energy versus heat) 3 Established the concept of Energy Conservation $\Delta E_{int} \equiv \Delta E_{int,f} - \Delta E_{int,i}$ = Q - W

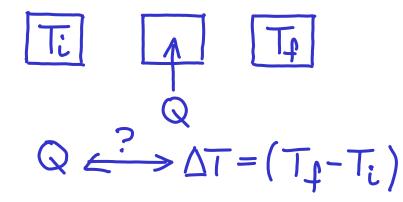
Calorimeter and Specific heat capacity To measure the amout of heat involved in a chemical or physical process. To measure the specific heat capacity (C) of substances. heat Q = mc at the tempeture change. heat the mass of the sample

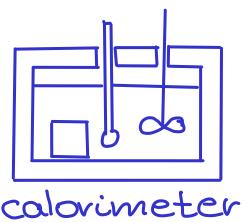
In a simple calorimetry process, (a) heat, q, is transferred from the hot metal, M, to the cool water, W, until (b) both are at the same temperature.

Energy Conservation $Q_{metal} + Q_{water} = 0$ = $\mathcal{V}_{\mathcal{M}} = \mathcal{O}_{\mathcal{M}} =$ where $(ST)_{M} \equiv T - T_{iM}$ $(ST)_{W} \equiv T - T_{iW}$ initial temperature $t = T - T_{iW}$ initial temperature $t = T - T_{iW}$ initial temperature $t = T - T_{iW}$ initial temperature Surroundings (a)

(b)

Heat and mechanical work



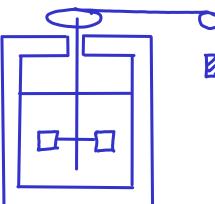


<u>Definition</u>: 1 calorie is the amount of heat which vaises the tem – perature of 1 g water by 1°C. (Warning: 1 Cal = 1000 cal = 1 k cal. Reminder: 1g is the mass of 1 cm³ water at 4°C temperature.)

Question: what is the relation between heat Q and mechanical work W=Fd?

james Prescott Joule (1818-1899):

1843 : 1 cal = 4.186 J



Heat capacity, specific heat

 $\begin{array}{c|c} \hline T_i & \hline T_f & Q = C \cdot \Delta T = C \cdot (T_f - T_i) \\ \hline Q & C = Q \\ \hline C & C \end{bmatrix} = J/K \end{array}$

C: heat capacity: amount of heat needed to raise the temperature of the object by 1K. (The object can be complex or composite.)

Homogeneous objects: $Q = cm \Delta T$ $C = \frac{Q}{m \Delta T} = \frac{C}{m}$ $[C] = \frac{J}{kgK}$

<u>C: Specific heat</u>: amount of heat needed to raise the temperature of a unit mass of material.

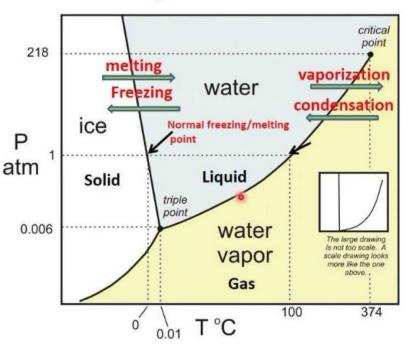
water : $1 \frac{cal}{g^2} = 4186.8 \frac{J}{kgK}$

 $AI : 0.215 \frac{cal}{gc} = 900 \frac{J}{kgK}$

 $: 0.0923 \frac{\text{Cal}}{92} = 386 \frac{\text{J}}{\text{kgK}}$ Cu

Molar heat = $C_m = \frac{C}{n}$

phase Diagram of Water



Phase Diagram for Water

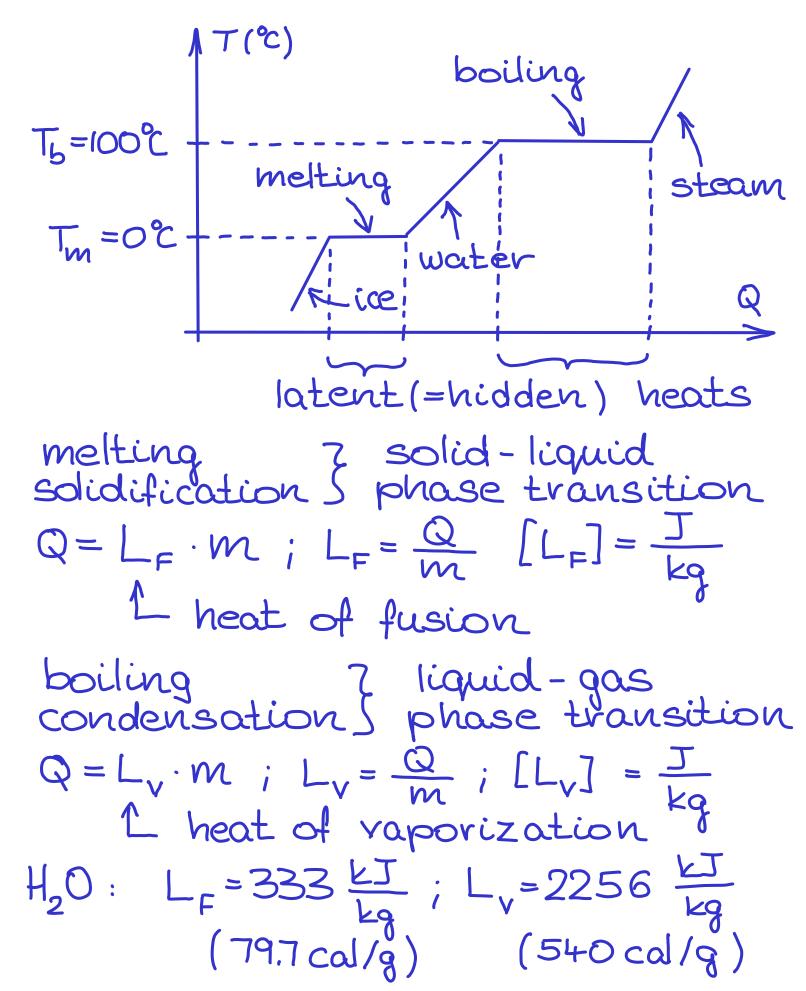
- Melting: When ice turns into water
- Boiling: When water turns into steam
- Condensation: When steam turns back into water
- Freezing: When water turns back into ice

The lines on the phase diagram are the lines of equilibrium. Along these lines, any two phases coexist in thermal equilibrium. However, a change in pressure or temperature will cause the phase to change abruptly. The AB line is the ice-vapor line, the BC line is the ice-water line, and the BD line is the water-vapor line. We can use the phase diagram to identify the state where water exists under specific temperature and pressure conditions. For example, water is ice at 50 kPa and – 10 $^{\circ}$ C.

Water has a triple point corresponding to the temperature and pressure at which ice, water, and vapor coexist in a stable equilibrium. It also has a critical point – the temperature at which water vapor cannot be forced back into water, no matter how much pressure is applied. The triple point and the critical point are listed in the table below.

Property	Temperature in Celsius (°C)	Temperature in Kelvin (K)	Pressure in atmosphere (atm)	Pressure in kilo Pascal (kPa)
Triple Point	0.01	273.16	0.006	0.608
Critical Point	373.99	647.14	217.7	22058.453

Heat(s) of transformation(s)



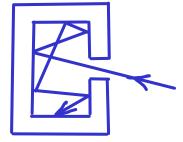
The three heat transfer mechanisms conduction, convection, radiation <u>Conduction</u>: $Q \propto \Delta T = (T_H - T_c)$ $Q \propto \Delta T = (T_H - T_c)$

k: thermal conductivity Examples for k in $\frac{W}{m \cdot K} = \frac{W \cdot m}{m^2 \cdot K}$ Silver: 428 Copper: 401 Steel: 14 Glass: 1-2 Fir: 0.026

<u>Convection</u>: fluid dynamics plus thermodynamics (way too complex) "Hot air rises."

Thermal radiation				
mostly infrared, can be visible or ultraviolet				
Stefan-Boltzmann law: P=5EAT ⁴				
T: absolute temperature in kelvin				
A: surface area of the object				
Eore: emissivity OSES1				
5: Stefan-Boltzmann constant:				
5.6703.108 $\frac{W}{m^2 K^4}$				

Black-body: a=1 absorbs everything e=1 emits the most



heated

cavity

(= oven)

The hole on the cavity behaves like a black-body: it absorbs all the radiation landing on it.

Blackbody radiation

$$T_{1} < T_{2} < T_{3}$$
 $I = I(\lambda, T)$:
intensity distribution:
power per area per
wavelength
 $[I] = \frac{W}{m^{2}} \cdot \frac{1}{m}$
blue-1 Lred
green
Wien's displacement law (1911 Nobel)
 $\lambda_{max} \cdot T = b$; $b = 2.898 \cdot 10^{3} \text{ m} \cdot \text{K}$
Sun: $\lambda_{max} = 550 \text{ m} \Rightarrow T = 5800 \text{ K}$
Stefan-Boltzmann law:
 $P = \overline{5} \text{EAT}^{4}$; $\overline{5} = 5.6705 \cdot 10^{3} \frac{W}{m^{2}} \text{ K}^{4}$

P: total area under the intensity distribution curve.

Max Planck (1918 Nobel): the exact shape of the intensity distribution curve. (You will learn about it in Quantum Mechanics.)